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Measurement of the Branching Fractions of the Radiative Charm Decays $D^0 \rightarrow \bar{K}^{*0}\gamma$ and $D^0 \rightarrow \phi\gamma$

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We present a measurement of the branching fractions for the Cabibbo-favored radiative decay, $D^0 \rightarrow \bar{K}^{*0}\gamma$, and the Cabibbo-suppressed radiative decay, $D^0 \rightarrow \phi\gamma$. These measurements are based on a data sample corresponding to an integrated luminosity of 387.1 fb^{-1} , recorded with

the *BABAR* detector at the PEP-II e^+e^- asymmetric-energy collider operating at center-of-mass energies 10.58 and 10.54 GeV. We measure the branching fractions relative to the well-studied decay $D^0 \rightarrow K^-\pi^+$ and find $\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma)/\mathcal{B}(D^0 \rightarrow K^-\pi^+) = (8.43 \pm 0.51 \pm 0.70) \times 10^{-3}$ and $\mathcal{B}(D^0 \rightarrow \phi\gamma)/\mathcal{B}(D^0 \rightarrow K^-\pi^+) = (7.15 \pm 0.78 \pm 0.69) \times 10^{-4}$, where the first error is statistical and the second is systematic. This is the first measurement of $\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma)$.

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In the b -quark sector, radiative decay processes have provided a rich field in which to study the Standard Model of particle physics. Decays such as $B \rightarrow \rho\gamma$ have yielded measurements of the Cabibbo-Kobayashi-Maskawa matrix element $|V_{tb}|$ [1, 2]. These decays are dominated by short-range electroweak processes, whereas long-range contributions are suppressed. The situation is reversed in the charm sector, where radiative decays are expected to be dominated largely by non-perturbative processes, examples of which are shown schematically in Fig. 1. Long-range contributions to radiative charm de-

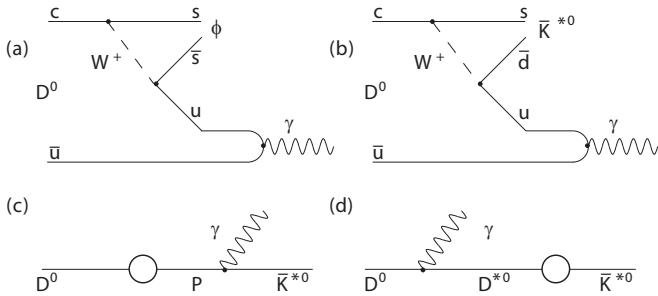


FIG. 1: Feynman diagrams for the long-range electromagnetic contributions to $D^0 \rightarrow V\gamma$, $V = \bar{K}^{*0}, \phi$. Figures (a) and (b) show sample vector dominance processes, while (c) and (d) are examples of pole diagrams, where the circles signify the weak transition and P represents a pseudoscalar meson.

cays are expected to increase the branching fractions for these modes to values of the order of 10^{-5} , whereas short-range interactions are predicted to yield rates at the 10^{-8} level. Given the expected dominance of long-range processes, radiative charm decays provide a laboratory in which to test these QCD-based calculations.

Numerous theoretical models have been developed to describe these radiative charm decays [3–9]. The two most comprehensive studies [5, 9] predict very similar amplitudes for the dominant diagrams shown in Fig. 1. The first paper bases predictions on Vector Meson Dominance (VMD) calculations, while the second paper uses Heavy-Quark Effective Theory in conjunction with Chiral-Lagrangians. Though each approach arrives at similar estimates for the magnitudes of the individual decay amplitudes, Ref. [5] predicts that the pole diagrams, shown in Figs. 1 (c) and (d), interfere destructively and cancel nearly completely. Ref. [9] makes no such predictions. Precise measurements of $\mathcal{B}(D^0 \rightarrow V\gamma, V = \bar{K}^{*0}, \phi)$ may provide insight into the

Mode	Experimental B.F. ($\times 10^{-5}$)	Theoretical[3–9] B.F. ($\times 10^{-5}$)
$D^0 \rightarrow \phi\gamma$	$(2.43^{+0.66}_{-0.57}(\text{stat.})^{+0.12}_{-0.14}(\text{sys.})) \times 10^{-5}$ [10]	$0.1 - 3.4$
$D^0 \rightarrow \bar{K}^{*0}\gamma$	< 76 (90% C.L.) [11]	$7 - 80$
$D^0 \rightarrow \rho^0\gamma$	< 24 (90% C.L.) [11]	$0.1 - 6.3$
$D^0 \rightarrow \omega\gamma$	< 24 (90% C.L.) [11]	$0.1 - 0.9$

TABLE I: The current experimental status and theoretical predictions for the branching fractions (B.F.) of radiative charm decays with vector mesons.

amount of interference between pole diagrams.

The first observation of a radiative, but color-suppressed, D^0 decay process was made by the Belle collaboration with a measurement of $\mathcal{B}(D^0 \rightarrow \phi\gamma) = (2.43^{+0.66}_{-0.57}(\text{stat.})^{+0.12}_{-0.14}(\text{sys.})) \times 10^{-5}$ [10]. CLEO II conducted searches for other radiative decays and established the current upper limit of $\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma) < 7.6 \times 10^{-4}$ at 90% confidence level (C.L.), as well as upper limits on $\mathcal{B}(D^0 \rightarrow \rho^0\gamma)$ and $\mathcal{B}(D^0 \rightarrow \omega\gamma)$ [11]. Table I summarizes theoretical predictions and current experimental results.

In this paper we present the first observation of the Cabibbo-favored radiative decay $D^0 \rightarrow \bar{K}^{*0}\gamma$, as well as an improved branching fraction measurement of the previously observed decay $D^0 \rightarrow \phi\gamma$. The analysis is based on 387.1 fb^{-1} of data recorded by the *BABAR* detector at the PEP-II e^+e^- asymmetric-energy collider operating at center-of-mass (CM) energies of $\sqrt{s} = 10.58 \text{ GeV}$ and 10.54 GeV , and uses approximately $5 \times 10^8 e^+e^- \rightarrow c\bar{c}$ events.

The *BABAR* detector is described in detail elsewhere [12]. Charged particle momenta are measured with a 5-layer double-sided silicon vertex tracker and a 40-layer drift chamber. Charged hadron identification is provided by measurements of the specific ionization energy loss, dE/dx , in the tracking system and of the Cherenkov angle obtained from a ring-imaging Cherenkov detector. An electromagnetic calorimeter consisting of 6580 CsI(Tl) crystals measures shower energy and position for electrons and photons. These detector elements are located inside, and coaxial with, the cryostat of a superconducting solenoidal magnet, which provides a 1.5 T magnetic field. The instrumented flux return of the magnet allows discrimination between muons and pions.

A detailed Monte Carlo (MC) simulation of the *BABAR* detector based on **GEANT 4** [13] is used to validate the analysis and determine the reconstruction efficiencies.

We optimize our selection criteria using simulated events by maximizing significance, defined as $N_S/\sqrt{N_S + N_B}$, where N_S and N_B denote the number of signal and background candidates in the MC simulation. We reconstruct radiative $D^0 \rightarrow V\gamma$, $V = \bar{K}^{*0}, \phi$ decays using the charged decay modes of the vector meson, $\bar{K}^{*0} \rightarrow K^-\pi^+$ ($\phi \rightarrow K^-K^+$) [14]. We form \bar{K}^{*0} (ϕ) candidates from pairs of oppositely charged tracks identified as $K^-\pi^+$ (K^-K^+) using the Cherenkov angle measurement of the DIRC and dE/dx measurements from the tracking system, and accept any $K^-\pi^+$ (K^-K^+) candidates with invariant mass in the range 0.848 to 0.951 GeV/ c^2 (1.01 to 1.03 GeV/ c^2). The charged track candidates are fit to a common vertex, and a fit probability greater than 0.1% is required.

A photon candidate is defined as an energy deposit in the EMC that is not associated with the trajectory of a charged track, and which exhibits the expected shower shape characteristics. Each such candidate is required to have CM energy greater than 0.54 GeV. The charged-particle vertex is assumed to be the production point of the photon. We suppress the significant background from $\pi^0 \rightarrow \gamma\gamma$ decays by rejecting a photon candidate which, when paired with another photon in the event, results in an invariant mass consistent with the π^0 mass, $(0.115 < M(\gamma\gamma) < 0.150)$ GeV/ c^2 .

Background from random $D^0 \rightarrow V\gamma$ candidates is reduced by requiring that the D^0 candidate be a product of the decay $D^{*+} \rightarrow D^0\pi^+$. A D^{*+} candidate is formed by combining a D^0 candidate with a low-momentum charged pion, denoted as π_s^+ . These pion candidates are required to have CM momentum less than 450 MeV/ c . We calculate the mass difference, $\Delta M = M(V\gamma\pi_s^+) - M(V\gamma)$ and require $(0.1435 < \Delta M < 0.1475)$ GeV/ c^2 . The ΔM distribution of candidates arising from signal decays is well-described by a Gaussian distribution function. Our selection corresponds to a six-standard deviation interval centered on the mean of the Gaussian, and hence retains almost all of the signal candidates. We reduce combinatoric background from $B\bar{B}$ events to a negligible level by requiring that the CM momentum of the D^{*+} candidate be greater than 2.62 GeV/ c .

The dominant background in our sample of $D^0 \rightarrow \bar{K}^{*0}\gamma$ candidates results from $D^0 \rightarrow K^-\pi^+\pi^0$ decays, where one of the photons from the π^0 decay is paired with the kaon and pion from the D^0 decay to closely mimic the signal mode. As described above, we use a π^0 veto to suppress such events but, given the large branching fraction of this mode, $\mathcal{B}(D^0 \rightarrow K^-\pi^+\pi^0) = (13.5 \pm 0.6)\%$ [15], a significant number of such candidates survives. We can separate this background from signal on a statistical basis because of differences in the $K^-\pi^+\gamma$ invariant mass distribution. The background distribution peaks slightly below the nominal D^0 mass, and has a different shape from that of signal events. An additional background arises from $D^0 \rightarrow \bar{K}^{*0}\eta$ events where the η decays to

two photons, one of which is combined with the $K^-\pi^+$ pair to form an invariant mass within our D^0 mass window. This contribution peaks well below the nominal D^0 mass, and it can be separated easily from correctly reconstructed $D^0 \rightarrow \bar{K}^{*0}\gamma$ decays.

The impact of both $D^0 \rightarrow \bar{K}^{*0}\pi^0$ and $D^0 \rightarrow \bar{K}^{*0}\eta$ is further reduced by using the \bar{K}^{*0} helicity angle θ_H . The helicity angle is defined as the angle between the momentum of the \bar{K}^{*0} meson parent particle (D^0) and the momentum of the \bar{K}^{*0} daughter kaon as measured in the \bar{K}^{*0} rest frame. Due to angular momentum conservation, $dN/d\cos\theta_H$ for signal candidates varies as $1 - \cos^2\theta_H$, whereas for $D^0 \rightarrow \bar{K}^{*0}\pi^0(\eta)$ events the cosine of the helicity angle is $\cos^2\theta_H$ distributed. The $\cos\theta_H$ distribution of $D^0 \rightarrow K^-\pi^+\pi^0$ candidates is complicated by the interference and overlap of resonant structure in the final state Dalitz plot. Based on a MC study an asymmetric selection of $-0.30 < \cos\theta_H < 0.65$ is chosen to maximize the signal significance.

Similarly, but to a lesser extent, the signal of the Cabibbo-suppressed radiative decay $D^0 \rightarrow \phi\gamma$ is obscured by backgrounds from $D^0 \rightarrow \phi\pi^0$ and $D^0 \rightarrow \phi\eta$ decays. Due to the small width of the ϕ meson, background from $D^0 \rightarrow K^-K^+\pi^0$ transitions with a K^+K^- invariant mass in the ϕ region yields a negligible contribution to $D^0 \rightarrow \phi\gamma$ [16]. Since angular momentum conservation dictates that the cosine of the helicity angle of the remaining $D^0 \rightarrow \phi\pi^0$ events follow a $\cos^2\theta_H$ distribution, we replace the tight $\cos\theta_H$ selection criterion used in the $D^0 \rightarrow \bar{K}^{*0}\gamma$ case with the looser requirement $|\cos\theta_H| < 0.9$ and include $\cos\theta_H$ as a variable in the fitting procedure. This retains a larger fraction of signal events, and so reduces statistical uncertainty.

We consider other radiative decays which might reflect into the $M(\bar{K}^{*0}\gamma)$ and $M(\phi\gamma)$ invariant mass distributions. Background to $D^0 \rightarrow \bar{K}^{*0}\gamma$ may arise from $D^0 \rightarrow \phi\gamma$ if a kaon from $\phi \rightarrow K^-K^+$ is mis-identified as a pion. Background from $D^0 \rightarrow \rho^0\gamma$ may arise if a pion from $\rho^0 \rightarrow \pi^-\pi^+$ is misidentified as a kaon. Real $D^0 \rightarrow \bar{K}^{*0}\gamma$ events can reflect into the $M(\phi\gamma)$ distributions if a π^+ is misidentified as a K^+ . Using MC simulations, all of these background contributions are found to be negligible.

We extract the $D^0 \rightarrow \bar{K}^{*0}\gamma$ yield using an unbinned extended maximum likelihood method (E-MLM) to fit the $M(\bar{K}^{*0}\gamma)$ invariant mass spectrum. The yield of $D^0 \rightarrow \phi\gamma$ events is extracted using an E-MLM to fit the two dimensional distribution of invariant mass, $M(\phi\gamma)$, and helicity, $\cos\theta_H$.

We use a Crystal Ball (CB) line shape [17] to model the invariant mass distributions for $D^0 \rightarrow \bar{K}^{*0}\gamma$ ($D^0 \rightarrow \phi\gamma$) signal events, and background reflections from $D^0 \rightarrow K^-\pi^+\pi^0$ ($D^0 \rightarrow \phi\pi^0$) decays. The invariant mass distributions of $D^0 \rightarrow \bar{K}^{*0}\eta$ and $D^0 \rightarrow \phi\eta$ background events are modeled with a Gaussian function and a first order Chebychev polynomial. The remaining combina-

toric background decays are modeled with a second order Chebychev polynomial. In the $\phi\gamma$, case the $\cos\theta_H$ distributions of $D^0 \rightarrow \phi\gamma$, $D^0 \rightarrow \phi\pi^0$, $D^0 \rightarrow \phi\eta$, and combinatoric background events are all modeled using second order Chebychev polynomials. The parameters of these probability distribution functions (PDFs) are obtained using simulated events and subsequently fixed when fitting the data.

We validate the invariant mass PDFs using data. To verify that the MC correctly simulates the backgrounds and the effects of the missing photon from the π^0 decay, we search our data sample for $D^0 \rightarrow K_S^0\gamma$ candidates. Since this decay is forbidden by angular momentum conservation, the candidates surviving our selection criteria are all combinatoric background or due to $D^0 \rightarrow K_S^0\pi^0$ or $D^0 \rightarrow K_S^0\eta$ decays.

We select K_S^0 candidates from pairs of oppositely charged tracks identified as pions. The pions are required to share a common production vertex and have an invariant mass in the range $(0.490 < M(\pi^+\pi^-) < 0.505) \text{ GeV}/c^2$. Selection criteria for the photon momentum, D^{*+} momentum, ΔM , and π^0 veto are identical to those used in the $D^0 \rightarrow V\gamma$ analyses. The resulting $K_S^0\gamma$ invariant mass spectrum is fit with a linear combination of three PDFs. The first PDF is used to model $D^0 \rightarrow K_S^0\pi^0$ candidates, and has the same functional form as the one used to model $D^0 \rightarrow K^-\pi^+\pi^0$ candidates. The second PDF is used to model $D^0 \rightarrow K_S^0\eta$ candidates, and has the same functional form as that used to model $D^0 \rightarrow \bar{K}^{*0}\eta$ candidates. The third PDF is a second order Chebychev polynomial used to model combinatoric background candidates. The shapes for both $D^0 \rightarrow K_S^0\eta$ and combinatoric background candidates are fixed using MC. The $D^0 \rightarrow K_S^0\pi^0$ signal shape is allowed to float in the final fit. Both MC and data are fit in this way and we find good agreement. The observed differences in the fit parameters are used to correct the CB line shape PDFs as described below.

A second test is performed using $D^0 \rightarrow \bar{K}^{*0}\gamma$ candidates taken from the sideband regions defined by $|\cos\theta_H| > 0.9$. Very few $D^0 \rightarrow \bar{K}^{*0}\gamma$ candidates are seen within this region, leading to a clean sample of $D^0 \rightarrow K^-\pi^+\pi^0$ decays. The resulting D^0 invariant mass spectrum is fit using a procedure similar to the one used for signal region $D^0 \rightarrow \bar{K}^{*0}\gamma$ candidates. The only differences are that the $D^0 \rightarrow \bar{K}^{*0}\gamma$ contribution is fixed to zero, and the $D^0 \rightarrow K^-\pi^+\pi^0$ signal shape is allowed to float freely. The $M(\bar{K}^{*0}\gamma)$ distribution of $D^0 \rightarrow K^-\pi^+\pi^0$ events is compared between data and MC and we find good agreement.

Potential differences between the $D^0 \rightarrow V\gamma$ invariant mass distributions for data and MC are evaluated by using $D^0 \rightarrow K_S^0\pi^0$ events. The selection criteria for K_S^0 mesons are identical to those applied in the $D^0 \rightarrow K_S^0\gamma$ analysis. The requirements on ΔM and D^{*+} CM momentum are as before. A π^0 candidate

consists of a photon pair with invariant mass satisfying $(0.110 < M(\gamma\gamma) < 0.150) \text{ GeV}/c^2$, and resultant laboratory momentum greater than $0.540 \text{ GeV}/c$. This resulting sample of $D^0 \rightarrow K_S^0\pi^0$ candidates is fit to a CB line shape and a linear background.

We used the average difference between the CB line shape parameters in MC and these data control samples to modify the PDF parameterizations used in the fit.

The fit results from data and expected signal and background contributions from MC are shown in Fig. 2(a-c). The event yields obtained from the E-MLM fit for both $D^0 \rightarrow \phi\gamma$ and $D^0 \rightarrow \bar{K}^{*0}\gamma$ are $N(D^0 \rightarrow \phi\gamma; \phi \rightarrow K^-K^+) = 242.6 \pm 24.8$ and $N(D^0 \rightarrow \bar{K}^{*0}\gamma; \bar{K}^{*0} \rightarrow K^-\pi^+) = 2285.8 \pm 113.2$. The reconstruction efficiencies, determined using MC, are found to be $\epsilon(D^0 \rightarrow \phi\gamma; \phi \rightarrow K^-K^+) = (10.8 \pm 0.1)\%$ and $\epsilon(D^0 \rightarrow \bar{K}^{*0}\gamma; \bar{K}^{*0} \rightarrow K^-\pi^+) = (6.4 \pm 0.1)\%$.

In order to avoid uncertainties in the overall normalization we measure the branching fraction of the radiative decays relative to $\mathcal{B}(D^0 \rightarrow K^-\pi^+)$. We prepare a $D^0 \rightarrow K^-\pi^+$ dataset following procedures similar to those described above, and find a yield $N(D^0 \rightarrow K^-\pi^+) = (335.1 \pm 4.0) \times 10^3$ with an efficiency of $\epsilon(D^0 \rightarrow K^-\pi^+) = (5.3 \pm 0.2)\%$.

We perform several consistency checks. Our result is compared to the $\cos\theta_H$ distribution expected for $D^0 \rightarrow \bar{K}^{*0}\gamma$ by refitting the data in intervals of $\cos\theta_H$ and measuring $N(D^0 \rightarrow \bar{K}^{*0}\gamma)$ in each interval. The normalized and efficiency-corrected result, shown in Fig. 2(d), compares well to the expected distribution. As an additional check, we divide the data into five distinct samples, one for each PEP-II run period, and perform the analysis on each subset independently. We see a $D^0 \rightarrow \bar{K}^{*0}\gamma$ signal for each run period, and find that the branching ratios are consistent within statistical uncertainties.

We evaluate the systematic uncertainties associated with our measurement in several different studies. Systematic effects due to the PDF parameterizations of signal and backgrounds are determined by generating an ensemble of 1,000 random numbers drawn from a normal distribution for each PDF parameter, including their correlations obtained from our fits. We refit the data using each of the 1,000 sets of random numbers. The resulting distribution of $N(D^0 \rightarrow V\gamma)$ is fit to a Gaussian function and the percent standard deviation is taken as the systematic error, 5.9% for $D^0 \rightarrow \phi\gamma$ and 4.4% for $D^0 \rightarrow \bar{K}^{*0}\gamma$.

Correcting the $D^0 \rightarrow V\gamma$ and $D^0 \rightarrow V\pi^0$ PDF parameters using the data control samples induces a second systematic uncertainty in the parameterization of the signal shapes. We estimate this effect by independently applying the corrections obtained using each of the three control samples. The largest percentage variation in $N(D^0 \rightarrow V\gamma)$ is taken as the systematic uncertainty associated with this correction; this leads to systematic uncertainties of 3.0% for $N(D^0 \rightarrow \phi\gamma)$ and 4.3% for

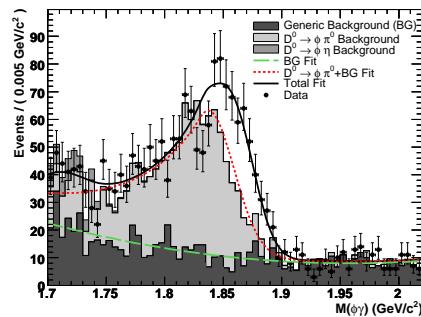
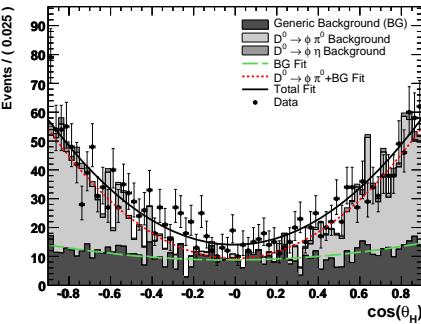
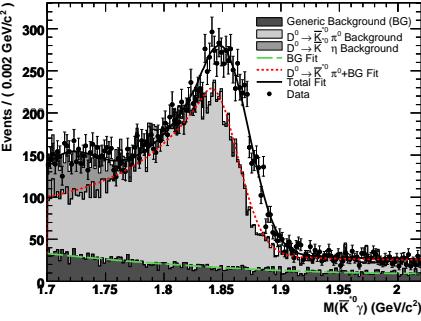
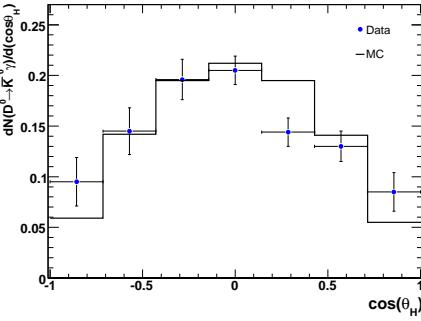
(a) The $\phi\gamma$ invariant mass distribution.(b) The $\phi\gamma$ helicity angle distribution.(c) The $\bar{K}^{*0}\gamma$ invariant mass distribution.(d) The $D^0 \rightarrow \bar{K}^{*0}\gamma$ helicity angle distribution.

FIG. 2: Invariant mass and $\cos\theta_H$ distributions for data (points) and simulated events (histograms). The curves show the fit results and the individual signal and background contributions. BG refers to the combinatoric background.

$$N(D^0 \rightarrow \bar{K}^{*0}\gamma).$$

We quantify the difference in particle identifica-

Systematic	$\sigma(D^0 \rightarrow \phi\gamma)$ (%)	$\sigma(D^0 \rightarrow K^{*0}\gamma)$ (%)
Tracking, vertexing	1.2	1.0
Particle ID	2.9	1.1
γ reconstruction	1.8	1.8
π^0 veto	1.8	1.8
PDF parameter	5.9	4.4
Correcting $\mathcal{P}_{D^0 \rightarrow V\gamma}$ and $\mathcal{P}_{D^0 \rightarrow V\pi^0}$	3.0	4.3
Ref. mode efficiency	1.5	1.5
Selection criteria	5.4	4.5
Total systematic effect	9.6	8.3

TABLE II: Summary of all systematic errors for each D^0 decay mode. The total systematic uncertainty is obtained by adding the individual systematic estimates in quadrature.

tion (PID) efficiency between data and simulation by means of a high-purity control sample of $D^{*+} \rightarrow D^0\pi^+$, $D^0 \rightarrow K^-\pi^+$ events, which we divide into intervals of polar angle and momentum. The change in yield when PID selection criteria are applied is computed separately for data and for simulated events and the difference is taken as a correction factor for that interval. We then weight the correction factors according to the expected momentum and polar-angle distributions of the $D^0 \rightarrow \bar{K}^{*0}\gamma$ signal. While a portion of the PID systematic uncertainty for our signal modes is canceled when measuring the branching fractions in ratio to $D^0 \rightarrow K^-\pi^+$, the residual uncertainty is found to be 2.88% for $D^0 \rightarrow \phi\gamma$ and 1.10% for $D^0 \rightarrow \bar{K}^{*0}\gamma$. By measuring $\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma)$ and $\mathcal{B}(D^0 \rightarrow \phi\gamma)$ with respect to $D^0 \rightarrow K^-\pi^+$, first-order effects from charged particle tracking also cancel, leaving only a second order systematic uncertainty of 1.00% for $D^0 \rightarrow \bar{K}^{*0}\gamma$ events and 1.20% for $D^0 \rightarrow \phi\gamma$. We summarize all systematic uncertainties in Table II.

In this paper, we report our observation of the Cabibbo-favored, but color-suppressed, radiative decay $D^0 \rightarrow \bar{K}^{*0}\gamma$. We also present confirmation of the previous measurement of the Cabibbo-suppressed radiative decay $\mathcal{B}(D^0 \rightarrow \phi\gamma)$, but with reduced statistical uncertainties. The measured branching ratios are

$$\begin{aligned} \frac{\mathcal{B}(D^0 \rightarrow \phi\gamma)}{\mathcal{B}(D^0 \rightarrow K^-\pi^+)} &= (7.15 \pm 0.78 \pm 0.69) \times 10^{-4} \\ \frac{\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma)}{\mathcal{B}(D^0 \rightarrow K^-\pi^+)} &= (8.43 \pm 0.51 \pm 0.70) \times 10^{-3} \end{aligned}$$

where the first uncertainty is statistical and the second is systematic. Using the current world average of $\mathcal{B}(D^0 \rightarrow K^-\pi^+) = (3.82 \pm 0.07)\%$ [15] we obtain the following absolute branching fractions:

$$\begin{aligned} \mathcal{B}(D^0 \rightarrow \phi\gamma) &= (2.73 \pm 0.30 \pm 0.26) \times 10^{-5} \\ \mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma) &= (3.22 \pm 0.20 \pm 0.27) \times 10^{-4}. \end{aligned}$$

These results are consistent with the theoretical expectations of Table I.

In the context of the vector dominance model the largest contribution to radiative D^0 decays is expected to come from a virtual ρ^0 coupling directly to a single photon, leading to the prediction that the branching ratios $\mathcal{B}(D^0 \rightarrow \phi\gamma)/\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma)$ and $\mathcal{B}(D^0 \rightarrow \phi\rho^0)/\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\rho^0)$ should be equal [5]. Comparing our measurements of the radiative D^0 decays with the current world averages [15] we find

$$\frac{\mathcal{B}(D^0 \rightarrow \phi\gamma)}{\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma)} = (6.27 \pm 0.71 \pm 0.79) \times 10^{-2}$$

$$\frac{\mathcal{B}(D^0 \rightarrow \phi\rho^0)}{\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\rho^0)} = (6.7 \pm 1.6) \times 10^{-2}$$

in agreement with this prediction.

If we assume all contributions are from VMD type processes and under the assumption that the ρ^0 meson is transversely polarized, as has been confirmed experimentally for $D^0 \rightarrow \bar{K}^{*0}\rho^0$ [15], we expect $\mathcal{B}(D^0 \rightarrow V\gamma) \approx \alpha_{EM}\mathcal{B}(D^0 \rightarrow V\rho^0)$ [5], where $\alpha_{EM} = 1/137$ is the fine structure constant. Using our results we find

$$\mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\gamma) = (0.021 \pm 0.005) \mathcal{B}(D^0 \rightarrow \bar{K}^{*0}\rho^0)$$

$$\mathcal{B}(D^0 \rightarrow \phi\gamma) = (0.020 \pm 0.003) \mathcal{B}(D^0 \rightarrow \phi\rho^0)$$

which in both cases is about a factor of three larger than the VMD prediction. This indicates that we are seeing enhancements from processes other than VMD, which might be explained by incomplete cancellation between pole diagrams.

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[1] A. Ali and A. Y. Parkhomenko, Eur. Phys. J. **C23**, 89 (2002).

[2] A. Ali, E. Lunghi, and A. Y. Parkhomenko, Phys. Lett. **B595**, 323 (2004).

[3] B. Bajc, S. Fajfer, and R. J. Oakes, Phys. Rev. **D51**, 2230 (1995).

[4] B. Bajc, S. Fajfer, and R. J. Oakes, Phys. Rev. **D54**, 5883 (1996).

[5] G. Burdman, E. Golowich, J. L. Hewett, and S. Pakvasa, Phys. Rev. **D52**, 6383 (1995).

[6] H.-Y. Cheng et al., Phys. Rev. **D51**, 1199 (1995).

[7] S. Fajfer, A. Prapotnik, S. Prelovsek, P. Singer, and J. Zupan, Nucl. Phys. Proc. Suppl. **115**, 93 (2003).

[8] S. Fajfer and P. Singer, Phys. Rev. **D56**, 4302 (1997).

[9] S. Fajfer, S. Prelovsek, and P. Singer, Eur. Phys. J. **C6**, 471 (1999).

[10] K. Abe et al., Phys. Rev. Lett. **92**, 101803 (2004), the published result has been rescaled using the latest values from [15].

[11] D. M. Asner et al., Phys. Rev. **D58**, 092001 (1998).

[12] B. Aubert et al., Nucl. Instrum. Meth. **A479**, 1 (2002).

[13] S. Agostinelli et al., Nucl. Instrum. Meth. **A506**, 250 (2003).

[14] Unless explicitly stated otherwise, charge conjugate reactions are included throughout this paper.

[15] W.-M. Yao et al. (Particle Data Group), J. Phys. **G33**, 1 (2006), and 2007 partial update for the 2008 edition.

[16] B. Aubert et al., Phys. Rev. **D76**, 011102 (2007).

[17] M. J. Oreglia, Ph.D Thesis, SLAC-236 (1980), J. E. Gaiser, Ph.D. Thesis, SLAC-255 (1982), T. Skwarnicki, Ph.D Thesis, DESY F31-86-02 (1986).