

## Search for Radiative Penguin Decays $B^+ \rightarrow \rho^+\gamma$ , $B^0 \rightarrow \rho^0\gamma$ , and $B^0 \rightarrow \omega\gamma$

- B. Aubert,<sup>1</sup> R. Barate,<sup>1</sup> D. Boutigny,<sup>1</sup> F. Couderc,<sup>1</sup> J.-M. Gaillard,<sup>1</sup> A. Hicheur,<sup>1</sup> Y. Karyotakis,<sup>1</sup> J. P. Lees,<sup>1</sup> V. Tisserand,<sup>1</sup> A. Zghiche,<sup>1</sup> A. Palano,<sup>2</sup> A. Pompili,<sup>2</sup> J. C. Chen,<sup>3</sup> N. D. Qi,<sup>3</sup> G. Rong,<sup>3</sup> P. Wang,<sup>3</sup> Y. S. Zhu,<sup>3</sup> G. Eigen,<sup>4</sup> I. Oftet,<sup>4</sup> B. Stugu,<sup>4</sup> G. S. Abrams,<sup>5</sup> A. W. Borgland,<sup>5</sup> A. B. Breon,<sup>5</sup> D. N. Brown,<sup>5</sup> J. Button-Shafer,<sup>5</sup> R. N. Cahn,<sup>5</sup> E. Charles,<sup>5</sup> C. T. Day,<sup>5</sup> M. S. Gill,<sup>5</sup> A. V. Gritsan,<sup>5</sup> Y. Groysman,<sup>5</sup> R. G. Jacobsen,<sup>5</sup> R. W. Kadel,<sup>5</sup> J. Kadyk,<sup>5</sup> L. T. Kerth,<sup>5</sup> Yu. G. Kolomensky,<sup>5</sup> G. Kukartsev,<sup>5</sup> G. Lynch,<sup>5</sup> L. M. Mir,<sup>5</sup> P. J. Oddone,<sup>5</sup> T. J. Orimoto,<sup>5</sup> M. Pripstein,<sup>5</sup> N. A. Roe,<sup>5</sup> M. T. Ronan,<sup>5</sup> V. G. Shelkov,<sup>5</sup> W. A. Wenzel,<sup>5</sup> M. Barrett,<sup>6</sup> K. E. Ford,<sup>6</sup> T. J. Harrison,<sup>6</sup> A. J. Hart,<sup>6</sup> C. M. Hawkes,<sup>6</sup> S. E. Morgan,<sup>6</sup> A. T. Watson,<sup>6</sup> M. Fritsch,<sup>7</sup> K. Goetzen,<sup>7</sup> T. Held,<sup>7</sup> H. Koch,<sup>7</sup> B. Lewandowski,<sup>7</sup> M. Pelizaeus,<sup>7</sup> M. Steinke,<sup>7</sup> J. T. Boyd,<sup>8</sup> N. Chevalier,<sup>8</sup> W. N. Cottingham,<sup>8</sup> M. P. Kelly,<sup>8</sup> T. E. Latham,<sup>8</sup> F. F. Wilson,<sup>8</sup> T. Cuhadar-Donszelmann,<sup>9</sup> C. Hearty,<sup>9</sup> N. S. Knecht,<sup>9</sup> T. S. Mattison,<sup>9</sup> J. A. McKenna,<sup>9</sup> D. Thiessen,<sup>9</sup> A. Khan,<sup>10</sup> P. Kyberd,<sup>10</sup> L. Teodorescu,<sup>10</sup> A. E. Blinov,<sup>11</sup> V. E. Blinov,<sup>11</sup> V. P. Druzhinin,<sup>11</sup> V. B. Golubev,<sup>11</sup> V. N. Ivanchenko,<sup>11</sup> E. A. Kravchenko,<sup>11</sup> A. P. Onuchin,<sup>11</sup> S. I. Serednyakov,<sup>11</sup> Yu. I. Skoppen,<sup>11</sup> E. P. Solodov,<sup>11</sup> A. N. Yushkov,<sup>11</sup> D. Best,<sup>12</sup> M. Bruinsma,<sup>12</sup> M. Chao,<sup>12</sup> I. Eschrich,<sup>12</sup> D. Kirkby,<sup>12</sup> A. J. Lankford,<sup>12</sup> M. Mandelkern,<sup>12</sup> R. K. Mommsen,<sup>12</sup> W. Roethel,<sup>12</sup> D. P. Stoker,<sup>12</sup> C. Buchanan,<sup>13</sup> B. L. Hartfiel,<sup>13</sup> S. D. Foulkes,<sup>14</sup> J. W. Gary,<sup>14</sup> B. C. Shen,<sup>14</sup> K. Wang,<sup>14</sup> D. del Re,<sup>15</sup> H. K. Hadavand,<sup>15</sup> E. J. Hill,<sup>15</sup> D. B. MacFarlane,<sup>15</sup> H. P. Paar,<sup>15</sup> Sh. Rahatlou,<sup>15</sup> V. Sharma,<sup>15</sup> J. W. Berryhill,<sup>16</sup> C. Campagnari,<sup>16</sup> B. Dahmes,<sup>16</sup> O. Long,<sup>16</sup> A. Lu,<sup>16</sup> M. A. Mazur,<sup>16</sup> J. D. Richman,<sup>16</sup> W. Verkerke,<sup>16</sup> T. W. Beck,<sup>17</sup> A. M. Eisner,<sup>17</sup> C. A. Heusch,<sup>17</sup> J. Kroseberg,<sup>17</sup> W. S. Lockman,<sup>17</sup> G. Nesom,<sup>17</sup> T. Schalk,<sup>17</sup> B. A. Schumm,<sup>17</sup> A. Seiden,<sup>17</sup> P. Spradlin,<sup>17</sup> D. C. Williams,<sup>17</sup> M. G. Wilson,<sup>17</sup> J. Albert,<sup>18</sup> E. Chen,<sup>18</sup> G. P. Dubois-Felsmann,<sup>18</sup> A. Dvoretskii,<sup>18</sup> D. G. Hitlin,<sup>18</sup> I. Narsky,<sup>18</sup> T. Piatenko,<sup>18</sup> F. C. Porter,<sup>18</sup> A. Ryd,<sup>18</sup> A. Samuel,<sup>18</sup> S. Yang,<sup>18</sup> S. Jayatilleke,<sup>19</sup> G. Mancinelli,<sup>19</sup> B. T. Meadows,<sup>19</sup> M. D. Sokoloff,<sup>19</sup> T. Abe,<sup>20</sup> F. Blanc,<sup>20</sup> P. Bloom,<sup>20</sup> S. Chen,<sup>20</sup> W. T. Ford,<sup>20</sup> U. Nauenberg,<sup>20</sup> A. Olivas,<sup>20</sup> P. Rankin,<sup>20</sup> J. G. Smith,<sup>20</sup> J. Zhang,<sup>20</sup> L. Zhang,<sup>20</sup> A. Chen,<sup>21</sup> J. L. Harton,<sup>21</sup> A. Soffer,<sup>21</sup> W. H. Toki,<sup>21</sup> R. J. Wilson,<sup>21</sup> Q. L. Zeng,<sup>21</sup> D. Altenburg,<sup>22</sup> T. Brandt,<sup>22</sup> J. Brose,<sup>22</sup> M. Dickopp,<sup>22</sup> E. Feltresi,<sup>22</sup> A. Hauke,<sup>22</sup> H. M. Lacker,<sup>22</sup> R. Müller-Pfefferkorn,<sup>22</sup> R. Nogowski,<sup>22</sup> S. Otto,<sup>22</sup> A. Petzold,<sup>22</sup> J. Schubert,<sup>22</sup> K. R. Schubert,<sup>22</sup> R. Schwierz,<sup>22</sup> B. Spaan,<sup>22</sup> J. E. Sundermann,<sup>22</sup> D. Bernard,<sup>23</sup> G. R. Bonneau,<sup>23</sup> F. Brochard,<sup>23</sup> P. Grenier,<sup>23</sup> S. Schrenk,<sup>23</sup> Ch. Thiebaux,<sup>23</sup> G. Vasileiadis,<sup>23</sup> M. Verderi,<sup>23</sup> D. J. Bard,<sup>24</sup> P. J. Clark,<sup>24</sup> D. Lavin,<sup>24</sup> F. Muheim,<sup>24</sup> S. Playfer,<sup>24</sup> Y. Xie,<sup>24</sup> M. Andreotti,<sup>25</sup> V. Azzolini,<sup>25</sup> D. Bettoni,<sup>25</sup> C. Bozzi,<sup>25</sup> R. Calabrese,<sup>25</sup> G. Cibinetto,<sup>25</sup> E. Luppi,<sup>25</sup> M. Negrini,<sup>25</sup> L. Piemontese,<sup>25</sup> A. Sarti,<sup>25</sup> E. Treadwell,<sup>26</sup> F. Anulli,<sup>27</sup> R. Baldini-Ferroli,<sup>27</sup> A. Calcaterra,<sup>27</sup> R. de Sangro,<sup>27</sup> G. Finocchiaro,<sup>27</sup> P. Patteri,<sup>27</sup> I. M. Peruzzi,<sup>27</sup> M. Piccolo,<sup>27</sup> A. Zallo,<sup>27</sup> A. Buzzo,<sup>28</sup> R. Capra,<sup>28</sup> R. Contri,<sup>28</sup> G. Crosetti,<sup>28</sup> M. Lo Vetere,<sup>28</sup> M. Macri,<sup>28</sup> M. R. Monge,<sup>28</sup> S. Passaggio,<sup>28</sup> C. Patrignani,<sup>28</sup> E. Robutti,<sup>28</sup> A. Santroni,<sup>28</sup> S. Tosi,<sup>28</sup> S. Bailey,<sup>29</sup> G. Brandenburg,<sup>29</sup> K. S. Chaisanguanthum,<sup>29</sup> M. Morii,<sup>29</sup> E. Won,<sup>29</sup> R. S. Dubitzky,<sup>30</sup> U. Langenegger,<sup>30</sup> W. Bhimji,<sup>31</sup> D. A. Bowerman,<sup>31</sup> P. D. Dauncey,<sup>31</sup> U. Egede,<sup>31</sup> J. R. Gaillard,<sup>31</sup> G. W. Morton,<sup>31</sup> J. A. Nash,<sup>31</sup> M. B. Nikolich,<sup>31</sup> G. P. Taylor,<sup>31</sup> M. J. Charles,<sup>32</sup> G. J. Grenier,<sup>32</sup> U. Mallik,<sup>32</sup> J. Cochran,<sup>33</sup> H. B. Crawley,<sup>33</sup> J. Lamsa,<sup>33</sup> W. T. Meyer,<sup>33</sup> S. Prell,<sup>33</sup> E. I. Rosenberg,<sup>33</sup> A. E. Rubin,<sup>33</sup> J. Yi,<sup>33</sup> M. Biasini,<sup>34</sup> R. Covarelli,<sup>34</sup> M. Pioppi,<sup>34</sup> M. Davier,<sup>35</sup> X. Giroux,<sup>35</sup> G. Grosdidier,<sup>35</sup> A. Höcker,<sup>35</sup> S. Laplace,<sup>35</sup> F. Le Diberder,<sup>35</sup> V. Lepeltier,<sup>35</sup> A. M. Lutz,<sup>35</sup> T. C. Petersen,<sup>35</sup> S. Plaszczynski,<sup>35</sup> M. H. Schune,<sup>35</sup> L. Tantot,<sup>35</sup> G. Wormser,<sup>35</sup> C. H. Cheng,<sup>36</sup> D. J. Lange,<sup>36</sup> M. C. Simani,<sup>36</sup> D. M. Wright,<sup>36</sup> A. J. Bevan,<sup>37</sup> C. A. Chavez,<sup>37</sup> J. P. Coleman,<sup>37</sup> I. J. Forster,<sup>37</sup> J. R. Fry,<sup>37</sup> E. Gabathuler,<sup>37</sup> R. Gamet,<sup>37</sup> D. E. Hutchcroft,<sup>37</sup> R. J. Parry,<sup>37</sup> D. J. Payne,<sup>37</sup> R. J. Sloane,<sup>37</sup> C. Touramanis,<sup>37</sup> J. J. Back,<sup>38</sup>, \* C. M. Cormack,<sup>38</sup> P. F. Harrison,<sup>38</sup>, \* F. Di Lodovico,<sup>38</sup> G. B. Mohanty,<sup>38</sup>, \* C. L. Brown,<sup>39</sup> G. Cowan,<sup>39</sup> R. L. Flack,<sup>39</sup> H. U. Flaecher,<sup>39</sup> M. G. Green,<sup>39</sup> P. S. Jackson,<sup>39</sup> T. R. McMahon,<sup>39</sup> S. Ricciardi,<sup>39</sup> F. Salvatore,<sup>39</sup> M. A. Winter,<sup>39</sup> D. Brown,<sup>40</sup> C. L. Davis,<sup>40</sup> J. Allison,<sup>41</sup> N. R. Barlow,<sup>41</sup> R. J. Barlow,<sup>41</sup> P. A. Hart,<sup>41</sup> M. C. Hodgkinson,<sup>41</sup> G. D. Lafferty,<sup>41</sup> A. J. Lyon,<sup>41</sup> J. C. Williams,<sup>41</sup> A. Farbin,<sup>42</sup> W. D. Hulsbergen,<sup>42</sup> A. Jawahery,<sup>42</sup> D. Kovalskyi,<sup>42</sup> C. K. Lae,<sup>42</sup> V. Lillard,<sup>42</sup> D. A. Roberts,<sup>42</sup> G. Blaylock,<sup>43</sup> C. Dallapiccola,<sup>43</sup> K. T. Flood,<sup>43</sup> S. S. Hertzbach,<sup>43</sup> R. Kofler,<sup>43</sup> V. B. Koptchev,<sup>43</sup> T. B. Moore,<sup>43</sup> S. Saremi,<sup>43</sup> H. Staengle,<sup>43</sup> S. Willocq,<sup>43</sup> R. Cowan,<sup>44</sup> G. Sciolla,<sup>44</sup> S. J. Sekula,<sup>44</sup> F. Taylor,<sup>44</sup> R. K. Yamamoto,<sup>44</sup> D. J. J. Mangeol,<sup>45</sup> P. M. Patel,<sup>45</sup> S. H. Robertson,<sup>45</sup> A. Lazzaro,<sup>46</sup> V. Lombardo,<sup>46</sup> F. Palombo,<sup>46</sup>

Work supported in part by the Department of Energy Contract DE-AC02-76SF00515

Stanford Linear Accelerator Center, Stanford University, Stanford, CA 94309

Contributed to the 32nd International Conference on High-Energy Physics (ICHEP 04),

August 16-22, 2004, Beijing, China

- J. M. Bauer,<sup>47</sup> L. Cremaldi,<sup>47</sup> V. Eschenburg,<sup>47</sup> R. Godang,<sup>47</sup> R. Kroeger,<sup>47</sup> J. Reidy,<sup>47</sup> D. A. Sanders,<sup>47</sup>  
D. J. Summers,<sup>47</sup> H. W. Zhao,<sup>47</sup> S. Brunet,<sup>48</sup> D. Côté,<sup>48</sup> P. Taras,<sup>48</sup> H. Nicholson,<sup>49</sup> N. Cavallo,<sup>50</sup>, † F. Fabozzi,<sup>50</sup>, †  
C. Gatto,<sup>50</sup> L. Lista,<sup>50</sup> D. Monorchio,<sup>50</sup> P. Paolucci,<sup>50</sup> D. Piccolo,<sup>50</sup> C. Sciacca,<sup>50</sup> M. Baak,<sup>51</sup> H. Bulten,<sup>51</sup>  
G. Raven,<sup>51</sup> H. L. Snoek,<sup>51</sup> L. Wilden,<sup>51</sup> C. P. Jessop,<sup>52</sup> J. M. LoSecco,<sup>52</sup> T. Allmendinger,<sup>53</sup> K. K. Gan,<sup>53</sup>  
K. Honscheid,<sup>53</sup> D. Hufnagel,<sup>53</sup> H. Kagan,<sup>53</sup> R. Kass,<sup>53</sup> T. Pulliam,<sup>53</sup> A. M. Rahimi,<sup>53</sup> R. Ter-Antonyan,<sup>53</sup>  
Q. K. Wong,<sup>53</sup> J. Brau,<sup>54</sup> R. Frey,<sup>54</sup> O. Ignokina,<sup>54</sup> C. T. Potter,<sup>54</sup> N. B. Sinev,<sup>54</sup> D. Strom,<sup>54</sup> E. Torrence,<sup>54</sup>  
F. Colecchia,<sup>55</sup> A. Dorigo,<sup>55</sup> F. Galeazzi,<sup>55</sup> M. Margoni,<sup>55</sup> M. Morandin,<sup>55</sup> M. Posocco,<sup>55</sup> M. Rotondo,<sup>55</sup>  
F. Simonetto,<sup>55</sup> R. Stroili,<sup>55</sup> G. Tiozzo,<sup>55</sup> C. Voci,<sup>55</sup> M. Benayoun,<sup>56</sup> H. Briand,<sup>56</sup> J. Chauveau,<sup>56</sup> P. David,<sup>56</sup>  
Ch. de la Vaissière,<sup>56</sup> L. Del Buono,<sup>56</sup> O. Hamon,<sup>56</sup> M. J. J. John,<sup>56</sup> Ph. Leruste,<sup>56</sup> J. Malcles,<sup>56</sup> J. Ocariz,<sup>56</sup>  
M. Pivk,<sup>56</sup> L. Roos,<sup>56</sup> S. T'Jampens,<sup>56</sup> G. Therin,<sup>56</sup> P. F. Manfredi,<sup>57</sup> V. Re,<sup>57</sup> P. K. Behera,<sup>58</sup> L. Gladney,<sup>58</sup>  
Q. H. Guo,<sup>58</sup> J. Panetta,<sup>58</sup> C. Angelini,<sup>59</sup> G. Batignani,<sup>59</sup> S. Bettarini,<sup>59</sup> M. Bondioli,<sup>59</sup> F. Bucci,<sup>59</sup> G. Calderini,<sup>59</sup>  
M. Carpinelli,<sup>59</sup> F. Forti,<sup>59</sup> M. A. Giorgi,<sup>59</sup> A. Lusiani,<sup>59</sup> G. Marchiori,<sup>59</sup> F. Martinez-Vidal,<sup>59</sup>, † M. Morganti,<sup>59</sup>  
N. Neri,<sup>59</sup> E. Paoloni,<sup>59</sup> M. Rama,<sup>59</sup> G. Rizzo,<sup>59</sup> F. Sandrelli,<sup>59</sup> J. Walsh,<sup>59</sup> M. Haire,<sup>60</sup> D. Judd,<sup>60</sup>  
K. Paick,<sup>60</sup> D. E. Wagoner,<sup>60</sup> N. Danielson,<sup>61</sup> P. Elmer,<sup>61</sup> Y. P. Lau,<sup>61</sup> C. Lu,<sup>61</sup> V. Miftakov,<sup>61</sup> J. Olsen,<sup>61</sup>  
A. J. S. Smith,<sup>61</sup> A. V. Telnov,<sup>61</sup> F. Bellini,<sup>62</sup> G. Cavoto,<sup>61</sup>,<sup>62</sup> R. Faccini,<sup>62</sup> F. Ferrarotto,<sup>62</sup> F. Ferroni,<sup>62</sup>  
M. Gaspero,<sup>62</sup> L. Li Gioi,<sup>62</sup> M. A. Mazzoni,<sup>62</sup> S. Morganti,<sup>62</sup> M. Pierini,<sup>62</sup> G. Piredda,<sup>62</sup> F. Safai Tehrani,<sup>62</sup>  
C. Voena,<sup>62</sup> S. Christ,<sup>63</sup> G. Wagner,<sup>63</sup> R. Waldi,<sup>63</sup> T. Adye,<sup>64</sup> N. De Groot,<sup>64</sup> B. Franek,<sup>64</sup> N. I. Geddes,<sup>64</sup>  
G. P. Gopal,<sup>64</sup> E. O. Olaiya,<sup>64</sup> R. Aleksan,<sup>65</sup> S. Emery,<sup>65</sup> A. Gaidot,<sup>65</sup> S. F. Ganzhur,<sup>65</sup> P.-F. Giraud,<sup>65</sup>  
G. Hamel de Monchenault,<sup>65</sup> W. Kozanecki,<sup>65</sup> M. Legendre,<sup>65</sup> G. W. London,<sup>65</sup> B. Mayer,<sup>65</sup> G. Schott,<sup>65</sup>  
G. Vasseur,<sup>65</sup> Ch. Yèche,<sup>65</sup> M. Zito,<sup>65</sup> M. V. Purohit,<sup>66</sup> A. W. Weidemann,<sup>66</sup> J. R. Wilson,<sup>66</sup> F. X. Yumiceva,<sup>66</sup>  
D. Aston,<sup>67</sup> R. Bartoldus,<sup>67</sup> N. Berger,<sup>67</sup> A. M. Boyarski,<sup>67</sup> O. L. Buchmueller,<sup>67</sup> R. Claus,<sup>67</sup> M. R. Convery,<sup>67</sup>  
M. Cristinziani,<sup>67</sup> G. De Nardo,<sup>67</sup> D. Dong,<sup>67</sup> J. Dorfan,<sup>67</sup> D. Dujmic,<sup>67</sup> W. Dunwoodie,<sup>67</sup> E. E. Elsen,<sup>67</sup> S. Fan,<sup>67</sup>  
R. C. Field,<sup>67</sup> T. Glanzman,<sup>67</sup> S. J. Gowdy,<sup>67</sup> T. Hadig,<sup>67</sup> V. Halyo,<sup>67</sup> C. Hast,<sup>67</sup> T. Hryna'ova,<sup>67</sup> W. R. Innes,<sup>67</sup>  
M. H. Kelsey,<sup>67</sup> P. Kim,<sup>67</sup> M. L. Kocian,<sup>67</sup> D. W. G. S. Leith,<sup>67</sup> J. Libby,<sup>67</sup> S. Luitz,<sup>67</sup> V. Luth,<sup>67</sup> H. L. Lynch,<sup>67</sup>  
H. Marsiske,<sup>67</sup> R. Messner,<sup>67</sup> D. R. Muller,<sup>67</sup> C. P. O'Grady,<sup>67</sup> V. E. Ozcan,<sup>67</sup> A. Perazzo,<sup>67</sup> M. Perl,<sup>67</sup>  
S. Petrak,<sup>67</sup> B. N. Ratcliff,<sup>67</sup> A. Roodman,<sup>67</sup> A. A. Salnikov,<sup>67</sup> R. H. Schindler,<sup>67</sup> J. Schwiening,<sup>67</sup> G. Simi,<sup>67</sup>  
A. Snyder,<sup>67</sup> A. Soha,<sup>67</sup> J. Stelzer,<sup>67</sup> D. Su,<sup>67</sup> M. K. Sullivan,<sup>67</sup> J. Va'vra,<sup>67</sup> S. R. Wagner,<sup>67</sup> M. Weaver,<sup>67</sup>  
A. J. R. Weinstein,<sup>67</sup> W. J. Wisniewski,<sup>67</sup> M. Wittgen,<sup>67</sup> D. H. Wright,<sup>67</sup> A. K. Yarritu,<sup>67</sup> C. C. Young,<sup>67</sup>  
P. R. Burchat,<sup>68</sup> A. J. Edwards,<sup>68</sup> T. I. Meyer,<sup>68</sup> B. A. Petersen,<sup>68</sup> C. Roat,<sup>68</sup> S. Ahmed,<sup>69</sup> M. S. Alam,<sup>69</sup>  
J. A. Ernst,<sup>69</sup> M. A. Saeed,<sup>69</sup> M. Saleem,<sup>69</sup> F. R. Wappler,<sup>69</sup> W. Bugg,<sup>70</sup> M. Krishnamurthy,<sup>70</sup> S. M. Spanier,<sup>70</sup>  
R. Eckmann,<sup>71</sup> H. Kim,<sup>71</sup> J. L. Ritchie,<sup>71</sup> A. Satpathy,<sup>71</sup> R. F. Schwitters,<sup>71</sup> J. M. Izen,<sup>72</sup> I. Kitayama,<sup>72</sup>  
X. C. Lou,<sup>72</sup> S. Ye,<sup>72</sup> F. Bianchi,<sup>73</sup> M. Bona,<sup>73</sup> F. Gallo,<sup>73</sup> D. Gamba,<sup>73</sup> L. Bosisio,<sup>74</sup> C. Cartaro,<sup>74</sup> F. Cossutti,<sup>74</sup>  
G. Della Ricca,<sup>74</sup> S. Dittongo,<sup>74</sup> S. Grancagnolo,<sup>74</sup> L. Lanceri,<sup>74</sup> P. Poropat,<sup>74</sup>,<sup>8</sup> L. Vitale,<sup>74</sup> G. Vuagnin,<sup>74</sup>  
R. S. Panvini,<sup>75</sup> Sw. Banerjee,<sup>76</sup> C. M. Brown,<sup>76</sup> D. Fortin,<sup>76</sup> P. D. Jackson,<sup>76</sup> R. Kowalewski,<sup>76</sup> J. M. Roney,<sup>76</sup>  
R. J. Sobie,<sup>76</sup> H. R. Band,<sup>77</sup> B. Cheng,<sup>77</sup> S. Dasu,<sup>77</sup> M. Datta,<sup>77</sup> A. M. Eichenbaum,<sup>77</sup> M. Graham,<sup>77</sup> J. J. Hollar,<sup>77</sup>  
J. R. Johnson,<sup>77</sup> P. E. Kutter,<sup>77</sup> H. Li,<sup>77</sup> R. Liu,<sup>77</sup> A. Mihalyi,<sup>77</sup> A. K. Mohapatra,<sup>77</sup> Y. Pan,<sup>77</sup> R. Prepost,<sup>77</sup>  
P. Tan,<sup>77</sup> J. H. von Wimmersperg-Toeller,<sup>77</sup> J. Wu,<sup>77</sup> S. L. Wu,<sup>77</sup> Z. Yu,<sup>77</sup> M. G. Greene,<sup>78</sup> and H. Neal<sup>78</sup>

(The BABAR Collaboration)

<sup>1</sup>Laboratoire de Physique des Particules, F-74941 Annecy-le-Vieux, France

<sup>2</sup>Università di Bari, Dipartimento di Fisica and INFN, I-70126 Bari, Italy

<sup>3</sup>Institute of High Energy Physics, Beijing 100039, China

<sup>4</sup>University of Bergen, Inst. of Physics, N-5007 Bergen, Norway

<sup>5</sup>Lawrence Berkeley National Laboratory and University of California, Berkeley, CA 94720, USA

<sup>6</sup>University of Birmingham, Birmingham, B15 2TT, United Kingdom

<sup>7</sup>Ruhr Universität Bochum, Institut für Experimentalphysik 1, D-44780 Bochum, Germany

<sup>8</sup>University of Bristol, Bristol BS8 1TL, United Kingdom

<sup>9</sup>University of British Columbia, Vancouver, BC, Canada V6T 1Z1

<sup>10</sup>Brunel University, Uxbridge, Middlesex UB8 3PH, United Kingdom

<sup>11</sup>Budker Institute of Nuclear Physics, Novosibirsk 630090, Russia

<sup>12</sup>University of California at Irvine, Irvine, CA 92697, USA

<sup>13</sup>University of California at Los Angeles, Los Angeles, CA 90024, USA

<sup>14</sup>University of California at Riverside, Riverside, CA 92521, USA

<sup>15</sup>University of California at San Diego, La Jolla, CA 92093, USA

<sup>16</sup>University of California at Santa Barbara, Santa Barbara, CA 93106, USA

- <sup>17</sup>University of California at Santa Cruz, Institute for Particle Physics, Santa Cruz, CA 95064, USA  
<sup>18</sup>California Institute of Technology, Pasadena, CA 91125, USA  
<sup>19</sup>University of Cincinnati, Cincinnati, OH 45221, USA  
<sup>20</sup>University of Colorado, Boulder, CO 80309, USA  
<sup>21</sup>Colorado State University, Fort Collins, CO 80523, USA  
<sup>22</sup>Technische Universität Dresden, Institut für Kern- und Teilchenphysik, D-01062 Dresden, Germany  
<sup>23</sup>Ecole Polytechnique, LLR, F-91128 Palaiseau, France  
<sup>24</sup>University of Edinburgh, Edinburgh EH9 3JZ, United Kingdom  
<sup>25</sup>Università di Ferrara, Dipartimento di Fisica and INFN, I-44100 Ferrara, Italy  
<sup>26</sup>Florida A&M University, Tallahassee, FL 32307, USA  
<sup>27</sup>Laboratori Nazionali di Frascati dell'INFN, I-00044 Frascati, Italy  
<sup>28</sup>Università di Genova, Dipartimento di Fisica and INFN, I-16146 Genova, Italy  
<sup>29</sup>Harvard University, Cambridge, MA 02138, USA  
<sup>30</sup>Universität Heidelberg, Physikalisches Institut, Philosophenweg 12, D-69120 Heidelberg, Germany  
<sup>31</sup>Imperial College London, London, SW7 2AZ, United Kingdom  
<sup>32</sup>University of Iowa, Iowa City, IA 52242, USA  
<sup>33</sup>Iowa State University, Ames, IA 50011-3160, USA  
<sup>34</sup>Università di Perugia, Dipartimento di Fisica and INFN, I-06100 Perugia, Italy  
<sup>35</sup>Laboratoire de l'Accélérateur Linéaire, F-91898 Orsay, France  
<sup>36</sup>Lawrence Livermore National Laboratory, Livermore, CA 94550, USA  
<sup>37</sup>University of Liverpool, Liverpool L69 7ZE, United Kingdom  
<sup>38</sup>Queen Mary, University of London, E1 4NS, United Kingdom  
<sup>39</sup>University of London, Royal Holloway and Bedford New College, Egham, Surrey TW20 0EX, United Kingdom  
<sup>40</sup>University of Louisville, Louisville, KY 40292, USA  
<sup>41</sup>University of Manchester, Manchester M13 9PL, United Kingdom  
<sup>42</sup>University of Maryland, College Park, MD 20742, USA  
<sup>43</sup>University of Massachusetts, Amherst, MA 01003, USA  
<sup>44</sup>Massachusetts Institute of Technology, Laboratory for Nuclear Science, Cambridge, MA 02139, USA  
<sup>45</sup>McGill University, Montréal, QC, Canada H3A 2T8  
<sup>46</sup>Università di Milano, Dipartimento di Fisica and INFN, I-20133 Milano, Italy  
<sup>47</sup>University of Mississippi, University, MS 38677, USA  
<sup>48</sup>Université de Montréal, Laboratoire René J. A. Lévesque, Montréal, QC, Canada H3C 3J7  
<sup>49</sup>Mount Holyoke College, South Hadley, MA 01075, USA  
<sup>50</sup>Università di Napoli Federico II, Dipartimento di Scienze Fisiche and INFN, I-80126, Napoli, Italy  
<sup>51</sup>NIKHEF, National Institute for Nuclear Physics and High Energy Physics, NL-1009 DB Amsterdam, The Netherlands  
<sup>52</sup>University of Notre Dame, Notre Dame, IN 46556, USA  
<sup>53</sup>Ohio State University, Columbus, OH 43210, USA  
<sup>54</sup>University of Oregon, Eugene, OR 97403, USA  
<sup>55</sup>Università di Padova, Dipartimento di Fisica and INFN, I-35131 Padova, Italy  
<sup>56</sup>Universités Paris VI et VII, Laboratoire de Physique Nucléaire et de Hautes Energies, F-75252 Paris, France  
<sup>57</sup>Università di Pavia, Dipartimento di Elettronica and INFN, I-27100 Pavia, Italy  
<sup>58</sup>University of Pennsylvania, Philadelphia, PA 19104, USA  
<sup>59</sup>Università di Pisa, Dipartimento di Fisica, Scuola Normale Superiore and INFN, I-56127 Pisa, Italy  
<sup>60</sup>Prairie View A&M University, Prairie View, TX 77446, USA  
<sup>61</sup>Princeton University, Princeton, NJ 08544, USA  
<sup>62</sup>Università di Roma La Sapienza, Dipartimento di Fisica and INFN, I-00185 Roma, Italy  
<sup>63</sup>Universität Rostock, D-18051 Rostock, Germany  
<sup>64</sup>Rutherford Appleton Laboratory, Chilton, Didcot, Oxon, OX11 0QX, United Kingdom  
<sup>65</sup>DSM/Dapnia, CEA/Saclay, F-91191 Gif-sur-Yvette, France  
<sup>66</sup>University of South Carolina, Columbia, SC 29208, USA  
<sup>67</sup>Stanford Linear Accelerator Center, Stanford, CA 94309, USA  
<sup>68</sup>Stanford University, Stanford, CA 94305-4060, USA  
<sup>69</sup>State University of New York, Albany, NY 12222, USA  
<sup>70</sup>University of Tennessee, Knoxville, TN 37996, USA  
<sup>71</sup>University of Texas at Austin, Austin, TX 78712, USA  
<sup>72</sup>University of Texas at Dallas, Richardson, TX 75083, USA  
<sup>73</sup>Università di Torino, Dipartimento di Fisica Sperimentale and INFN, I-10125 Torino, Italy  
<sup>74</sup>Università di Trieste, Dipartimento di Fisica and INFN, I-34127 Trieste, Italy  
<sup>75</sup>Vanderbilt University, Nashville, TN 37235, USA  
<sup>76</sup>University of Victoria, Victoria, BC, Canada V8W 3P6  
<sup>77</sup>University of Wisconsin, Madison, WI 53706, USA  
<sup>78</sup>Yale University, New Haven, CT 06511, USA

(Dated: January 3, 2005)

A search for the decays  $B \rightarrow \rho(770)\gamma$  and  $B^0 \rightarrow \omega(782)\gamma$  is performed on a sample of 211 million  $\Upsilon(4S) \rightarrow B\bar{B}$  events collected by the *BABAR* detector at the PEP-II asymmetric-energy  $e^+e^-$  storage ring. No evidence for the decays is seen. We set the following limits on the individual branching fractions  $\mathcal{B}(B^+ \rightarrow \rho^+\gamma) < 1.8 \times 10^{-6}$ ,  $\mathcal{B}(B^0 \rightarrow \rho^0\gamma) < 0.4 \times 10^{-6}$ , and  $\mathcal{B}(B^0 \rightarrow \omega\gamma) < 1.0 \times 10^{-6}$  at the 90% confidence level (C.L.). We use the quark model to limit the combined branching fraction  $\overline{\mathcal{B}}[B \rightarrow (\rho/\omega)\gamma] < 1.2 \times 10^{-6}$ , from which we determine a constraint on the ratio of CKM matrix elements  $|V_{td}|/|V_{ts}|$ .

PACS numbers: 12.15.Hh, 13.25.Hw

Within the Standard Model (SM), the decays  $B \rightarrow \rho\gamma$  and  $B^0 \rightarrow \omega\gamma$  proceed primarily through a  $b \rightarrow d\gamma$  electromagnetic penguin process that contains a top quark within the loop [1]. The rates for  $B^+ \rightarrow \rho^+\gamma$ ,  $B^0 \rightarrow \rho^0\gamma$ , and  $B^0 \rightarrow \omega\gamma$  [2] are related by the spectator-quark model, and we define the average branching fraction [3],  $\overline{\mathcal{B}}[B \rightarrow (\rho/\omega)\gamma] = \frac{1}{2} \left\{ \mathcal{B}(B^+ \rightarrow \rho^+\gamma) + \frac{\tau_{B^+}}{\tau_{B^0}} [\mathcal{B}(B^0 \rightarrow \rho^0\gamma) + \mathcal{B}(B^0 \rightarrow \omega\gamma)] \right\}$ , where  $\frac{\tau_{B^+}}{\tau_{B^0}}$  is the ratio of  $B$ -meson lifetimes. Recent calculations of  $\overline{\mathcal{B}}[B \rightarrow (\rho/\omega)\gamma]$  in the SM indicate a range of  $(0.9 - 1.8) \times 10^{-6}$  [3, 4]. There may also be contributions resulting from physics beyond the SM [5]. The ratio between the branching fractions for  $B \rightarrow (\rho/\omega)\gamma$  and  $B \rightarrow K^*\gamma$  is related in the SM to the ratio of Cabibbo-Kobayashi-Maskawa (CKM) matrix elements  $|V_{td}|/|V_{ts}|$  [3, 6]. Previous searches by *BABAR* [7] and CLEO [8] have found no evidence for  $B \rightarrow (\rho/\omega)\gamma$  decays.

We search for  $B \rightarrow \rho\gamma$  and  $B^0 \rightarrow \omega\gamma$  decays in a data sample containing  $211 \pm 2$  million  $\Upsilon(4S) \rightarrow B\bar{B}$  decays, collected by the *BABAR* detector [9] at the PEP-II asymmetric-energy  $e^+e^-$  storage ring. The data correspond to an integrated luminosity of  $191 \text{ fb}^{-1}$ .

The decay  $B \rightarrow \rho\gamma$  is reconstructed with  $\rho^0 \rightarrow \pi^+\pi^-$  and  $\rho^+ \rightarrow \pi^+\pi^0$ , while  $B^0 \rightarrow \omega\gamma$  is reconstructed with  $\omega \rightarrow \pi^+\pi^-\pi^0$ . Background comes primarily from  $e^+e^- \rightarrow q\bar{q}$  continuum events, where  $q = u, d, s, c$ , in which a high-energy photon is produced through  $\pi^0/\eta \rightarrow \gamma\gamma$  decays or via initial-state radiation (ISR). There are also significant  $B\bar{B}$  backgrounds:  $B \rightarrow K^*\gamma$ ,  $K^* \rightarrow K\pi$ , where a  $K^\pm$  is misidentified as a  $\pi^\pm$ ;  $B \rightarrow (\rho/\omega)\pi^0$  and  $B \rightarrow (\rho/\omega)\eta$ , where a high-energy photon comes from the  $\pi^0$  or  $\eta$  decay; and combinatorial background, mostly from high multiplicity  $b \rightarrow s\gamma$  decays.

We select  $\pi^\pm$  candidates from tracks with a momentum transverse to the beam direction greater than  $100 \text{ MeV}/c$ . The  $\pi^\pm$  selection algorithm combines measurements of energy loss in the tracking system with any associated Cherenkov photons measured by the ring imaging Cherenkov detector. The algorithm is optimized to reduce backgrounds from  $K^\pm$  produced in  $b \rightarrow s\gamma$  processes [7].

Neutral pion candidates are identified as pairs of neutral energy-deposits reconstructed in the CsI crystal calorimeter, each with an energy greater than  $50 \text{ MeV}$

in the laboratory frame. For  $B^0 \rightarrow \omega\gamma$  ( $B^+ \rightarrow \rho^+\gamma$ ) decays, the invariant mass of the pair is required to satisfy  $110 < m_{\gamma\gamma} < 150 \text{ MeV}/c^2$  ( $117 < m_{\gamma\gamma} < 145 \text{ MeV}/c^2$ ). To reduce combinatorial background, we require the cosine of the opening angle between the daughter photons in the laboratory frame be greater than 0.6; this selection retains 98% of  $\pi^0$  from signal decays.

A  $\rho^0$  candidate is reconstructed by selecting two identified pions that have opposite charge and a common vertex. We obtain  $\rho^+$  candidates by pairing  $\pi^0$  candidates with an identified  $\pi^+$ . The  $\omega$  candidates are reconstructed by combining a  $\pi^0$  candidate with pairs of oppositely charged pion candidates that originate from a common vertex; the charged pion pair must be consistent with originating from the interaction region to suppress  $K_s^0$  decays. We select  $\rho$  ( $\omega$ ) candidates with an invariant mass satisfying  $630 < m_{\pi\pi} < 940 \text{ MeV}/c^2$  ( $764 < m_{\pi^+\pi^-\pi^0} < 795 \text{ MeV}/c^2$ ).

The high-energy photon from the signal  $B$  decay is identified as a neutral energy-deposit in the calorimeter. We require that the deposit meets a number of criteria designed to eliminate background from charged particles and hadronic showers [10]. We veto photons from  $\pi^0(\eta)$  decay by requiring that the invariant mass of the candidate combined with any other photon of laboratory energy greater than  $30$  ( $250$ ) MeV not to be within the range  $105$  to  $155 \text{ MeV}/c^2$  ( $500$  to  $590 \text{ MeV}/c^2$ ).

The photon and  $\rho/\omega$  candidates are combined to form the  $B$ -meson candidates. We define  $\Delta E^* \equiv E_B^* - E_{\text{beam}}^*$ , where  $E_B^*$  is the center-of-mass (CM) energy of the  $B$ -meson candidate and  $E_{\text{beam}}^*$  is the CM beam energy. The  $\Delta E^*$  distribution of Monte Carlo (MC) simulated signal events is centered at zero, with a resolution of about  $50 \text{ MeV}$ . We also define the beam-energy-substituted mass  $m_{\text{ES}} \equiv \sqrt{E_{\text{beam}}^{*2} - p_B^{*2}}$ , where  $p_B^*$  is the CM momentum of the  $B$  candidate. Signal MC events peak in  $m_{\text{ES}}$  at the mass of the  $B$ -meson,  $m_B$ , with a resolution of  $3 \text{ MeV}/c^2$ . The distribution of continuum and combinatorial  $B\bar{B}$  background peaks in neither  $m_{\text{ES}}$  nor  $\Delta E^*$ ; the background distributions of  $B \rightarrow K^*\gamma$ ,  $B \rightarrow (\rho/\omega)\pi^0$  and  $B \rightarrow (\rho/\omega)\eta$  peak at  $m_B$  in  $m_{\text{ES}}$  and between  $-190 \text{ MeV}$  and  $-60 \text{ MeV}$  in  $\Delta E^*$ . We consider candidates in the ranges  $-0.3 < \Delta E^* < 0.3 \text{ GeV}$  and  $5.20 < m_{\text{ES}} < 5.29 \text{ GeV}/c^2$  to incorporate sidebands that allow the combinatorial background yields to be extracted from a fit to the data.

Several variables that distinguish between signal and continuum events are combined in a neural network [11]. The input variables depend mainly on the rest of the event (ROE), defined to be all charged tracks and neutral energy deposits in the calorimeter not used to reconstruct the  $B$  candidate. To reject ISR events, we compute the ratio of second-to-zeroth order Fox-Wolfram moments [12] for the ROE and the  $\rho/\omega$  candidate, in the frame recoiling against the photon momentum. To discriminate between the jet-like continuum background and the more spherically-symmetric signal events, we compute the angle between the photon and the thrust axis of the ROE in the CM frame and the moments  $L_i \equiv \sum_j p_j^* \cdot |\cos \theta_j^*|^i / \sum_j p_j^*$ , where  $p_j^*$  and  $\theta_j^*$  are the momentum and angle with respect to an axis, respectively, for each particle  $j$  in the ROE. We use  $L_1$ ,  $L_2$ , and  $L_3$  with respect to the thrust axis of the ROE, as well as  $L_1$  with respect to the photon direction. Differences in lepton and kaon production between background and  $B$  decays are exploited by including *BABAR* flavor tagging variables [13] as well as the maximum CM momentum and number of  $K^\pm$  and  $K_s^0$  in the ROE. For the  $B^0 \rightarrow (\rho^0/\omega)\gamma$  modes, we also use the separation along the beam axis of the  $B$ -meson candidate and ROE vertices; to remove poorly reconstructed events we require the separation be less than 4 mm. A separate neural network is trained for each mode. We make a loose selection on the output of the neural network,  $\mathcal{N}$ , that retains around 80% of the signal events.

To suppress background, we combine a number of signal-decay variables in a Fisher discriminant [14],  $\mathcal{F}$ , separately for each mode. We calculate the  $B$ -meson production angle  $\theta_B^*$ , the  $\rho/\omega$  helicity angle  $\theta_H$ , which is defined with respect to the normal of the decay plane for  $\omega$  candidates, and the  $\omega$  Dalitz angle  $\theta_D$  [7]. To reject  $B \rightarrow \rho(\pi^0/\eta)$  and  $B \rightarrow \omega(\pi^0/\eta)$  events in the  $B^+ \rightarrow \rho^+\gamma$  and  $B^0 \rightarrow \omega\gamma$  ( $B^0 \rightarrow \rho^0\gamma$ ) selection, we require  $|\cos \theta_H| < 0.70$  (0.75).

After applying the  $\mathcal{N}$  and  $|\cos \theta_H|$  criteria the expected average candidate multiplicity in signal events is 1.15, 1.03 and 1.14 for  $B^+ \rightarrow \rho^+\gamma$ ,  $B^0 \rightarrow \rho^0\gamma$  and  $B^0 \rightarrow \omega\gamma$ , respectively; in events with multiple candidates the one with the smallest value of  $|\Delta E^*|$  is retained.

The signal yield is determined from an extended maximum likelihood fit to the selected data. We fit the four-dimensional distribution of  $m_{\text{ES}}$ ,  $\Delta E^*$ ,  $\mathcal{F}$  and  $\mathcal{N}$ . For the  $B \rightarrow \rho\gamma$  fits, five event hypotheses are considered: signal, continuum background, combinatorial  $B$ -background, peaking  $B \rightarrow \rho(\pi^0/\eta)$  background and peaking  $B \rightarrow K^*\gamma$  background. For the  $B^0 \rightarrow \omega\gamma$  fit we consider only signal, continuum background, and peaking  $B \rightarrow \omega(\pi^0/\eta)$  background. The correlations among the observables are small; therefore, we assume that the probability density function (PDF)  $\mathcal{P}(\vec{x}_j; \vec{\alpha}_i)$  for each hypothesis is the product of individual PDFs for the variables  $\vec{x}_j = \{m_{\text{ES}}, \Delta E^*, \mathcal{F}, \mathcal{N}\}$  given the set of parameters

$\vec{\alpha}_i$ . The likelihood function is a product over all  $N_k$  candidate events of the sum of the PDFs,

$$\mathcal{L}_k = \exp \left( - \sum_{i=1}^{N_{\text{hyp}}} n_i \right) \cdot \left[ \prod_{j=1}^{N_k} \left( \sum_{i=1}^{N_{\text{hyp}}} n_i \mathcal{P}_i(\vec{x}_j; \vec{\alpha}_i) \right) \right],$$

where  $n_i$  is the yield of each hypothesis,  $k$  is the  $B \rightarrow (\rho/\omega)\gamma$  mode, and  $N_{\text{hyp}} = 5(3)$  for  $B \rightarrow \rho\gamma$  ( $B \rightarrow \omega\gamma$ ).

The  $m_{\text{ES}}$  and  $\Delta E^*$  PDFs are parameterized by a Crystal Ball function [15] for both the signal and peaking background. The parameterization is determined from signal MC samples, except the mean of the  $\Delta E^*$  distribution, which is offset by the observed difference between data and MC samples of  $B \rightarrow K^*\gamma$  decays. The continuum background  $m_{\text{ES}}$  and  $\Delta E^*$  distributions are parameterized by an ARGUS threshold function [16] and a second-order polynomial, respectively. The combinatorial  $B$  background is described by a smoothed distribution [17] determined from MC events in both  $m_{\text{ES}}$  and  $\Delta E^*$ . The distribution of  $\mathcal{N}$  for signal and  $B\bar{B}$  background is parameterized by a Crystal Ball function. The  $\mathcal{N}$  distribution for continuum is determined from sideband data, and a histogram is used as the PDF. The distribution of  $\mathcal{F}$  is parameterized by smoothed histograms of sideband data for the continuum background and MC events for all other hypotheses.

The fit to the data determines the shape parameters of the continuum background  $m_{\text{ES}}$  and  $\Delta E^*$  PDFs, as well as the signal, continuum background and combinatorial  $B\bar{B}$  background yields. All other parameters are fixed, including the peaking  $B\bar{B}$  background yields. A combined fit is also performed relating the modes using the definition of  $\bar{\mathcal{B}}[B \rightarrow (\rho/\omega)\gamma]$  to determine an effective yield ( $n_{\text{eff}}$ ) assuming  $n(B^+ \rightarrow \rho^+\gamma) = n_{\text{eff}} \cdot \epsilon(B^+ \rightarrow \rho^+\gamma)$  and  $n(B^0 \rightarrow (\rho^0/\omega)\gamma) = \frac{1}{2} \frac{\tau_{B^0}}{\tau_{B^+}} n_{\text{eff}} \cdot \epsilon(B^0 \rightarrow (\rho^0/\omega)\gamma)$  where  $n$  and  $\epsilon$  are the yields and reconstruction efficiencies of each mode; the efficiencies include the daughter branching fractions. We take  $\frac{\tau_{B^+}}{\tau_{B^0}} = 1.086 \pm 0.017$  [18]. Fig. 1 shows the projections of the combined fit results compared to the data. The results for the individual mode signal yields and  $n_{\text{eff}}$  are given in Table I. The significance is computed as  $\sqrt{2\Delta \log \mathcal{L}}$  where  $\Delta \log \mathcal{L}$  is the log likelihood difference between the best fit and the null-signal hypothesis. No significant signal is observed.

The most important systematic uncertainties are associated with the modeling of  $B\bar{B}$  backgrounds, the fixed parameters of the PDFs used in the fit, and the signal reconstruction efficiency. The first two contribute to the uncertainties on the signal yields.

The uncertainty on the peaking  $B \rightarrow K^*\gamma$  background is dominated by the  $K^\pm$  misidentification rate; the rate is corrected by the difference in  $K^\pm$  misidentification between data and MC samples of  $D^*$  decays, with the whole correction taken as the uncertainty. For the  $B^+ \rightarrow \rho^+(\pi^0/\eta)$ ,  $B^0 \rightarrow \rho^0\eta$  and  $B^0 \rightarrow \omega(\pi^0/\eta)$

TABLE I: The signal yield ( $n_{\text{sig}}$ ), continuum background yield ( $n_{\text{cont}}$ ), peaking background ( $n_{\text{peak}}$ ), significance in standard deviations  $\sigma$ , efficiency ( $\epsilon$ ), and branching fraction ( $\mathcal{B}$ ) central value and upper limit at the 90% C.L. for each mode. The results of the combined fit are shown in the bottom row where  $n_{\text{sig}}$  is equal to  $n_{\text{eff}}$ , which is described in the text. When two errors are quoted, the first is statistical and the second is systematic.

Mode	$n_{\text{sig}}$	$n_{\text{cont}}$	$n_{\text{peak}}$	Significance ( $\sigma$ )	$\epsilon(\%)$	$\mathcal{B}(10^{-6})$	$\mathcal{B}(10^{-6})$	90% C.L.
$B^+ \rightarrow \rho^+ \gamma$	$26^{+15+2}_{-14-2}$	$6850 \pm 90$	$18 \pm 4$	1.9	$13.2 \pm 1.4$	$0.9^{+0.6}_{-0.5} \pm 0.1$	—	< 1.8
$B^0 \rightarrow \rho^0 \gamma$	$0.3^{+7.2+1.7}_{-5.4-1.6}$	$4269 \pm 73$	$18 \pm 7$	0.0	$15.8 \pm 1.9$	$0.0 \pm 0.2 \pm 0.1$	—	< 0.4
$B^0 \rightarrow \omega \gamma$	$8.3^{+5.7+1.3}_{-4.5-1.9}$	$1378 \pm 37$	$2.6^{+0.8}_{-1.2}$	1.5	$8.6 \pm 0.9$	$0.5 \pm 0.3 \pm 0.1$	—	< 1.0
Combined	$269^{+126+40}_{-120-45}$	—	—	2.1	—	$0.6 \pm 0.3 \pm 0.1$	—	< 1.2

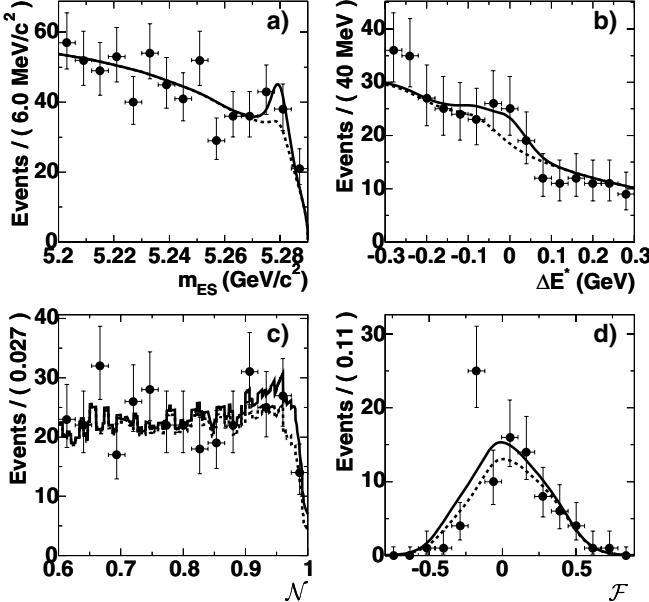


FIG. 1: Projections of the combined fit to  $B \rightarrow \rho \gamma$  and  $B^0 \rightarrow \omega \gamma$  in the four discriminating variables: (a)  $m_{\text{ES}}$ , (b)  $\Delta E^*$ , (c)  $N$ , and (d)  $F$ . The points are data, the solid line is the total PDF and the dashed line is the background only PDF. The selections applied, unless the variable is projected, are:  $5.272 < m_{\text{ES}} < 5.286 \text{ GeV}/c^2$ ,  $-0.10 < \Delta E^* < 0.05 \text{ GeV}$  and  $N > 0.9$ ; the selection efficiencies for signal events are 45%, 57%, 70%, and 44% for the  $m_{\text{ES}}$ ,  $\Delta E^*$ ,  $N$  and  $F$  projections, respectively.

peaking background decays, we vary the branching fractions by either one standard deviation from the measured values or between zero and the measured upper limit if the decay has not been observed [19, 20]; the value of the  $B^0 \rightarrow \rho^0 \pi^0$  branching fraction is varied between zero and  $5.1 \times 10^{-6}$  [20, 21]. The uncertainty on the peaking background of each mode is shown in Table I. We find the bias from neglecting the  $B \rightarrow K^* \gamma$  background and combinatorial  $B\bar{B}$  background in the fit to  $B^0 \rightarrow \omega \gamma$  candidates is  $1.1^{+1.9}_{-1.1}$  events; the corrected yield is given in Table I. To estimate the uncertainty related to the extraction of the signal  $m_{\text{ES}}$  and  $\Delta E^*$  PDFs from MC distributions, we vary the parameters within their errors. The variation in the fitted signal yield is taken as a systematic uncertainty. The uncertainty related to the statistics of the

histogram PDF that describes the continuum  $\mathcal{N}$  distribution is evaluated by varying the binning and by using a fifth-order polynomial as an alternative PDF. Several different control samples of data and MC events were used to determine alternative PDFs for the different hypotheses; none of these resulted in a significant change to the fitted signal yield.

The signal efficiency systematic error contains uncertainties from tracking, particle identification, photon/ $\pi^0$  reconstruction, photon selection and the neural network selection that are determined as in Ref. [22]. We determine the effect of correlations among the fit variables by using an ensemble of MC experiments of parameterized continuum background simulations embedded in samples of fully simulated signal and  $B\bar{B}$  background events. No bias is observed within the statistical error on the mean yields from this ensemble, which is taken as a multiplicative systematic uncertainty. The total multiplicative systematic error values are 11%, 13% and 10% for  $B^+ \rightarrow \rho^+ \gamma$ ,  $B^0 \rightarrow \rho^0 \gamma$  and  $B^0 \rightarrow \omega \gamma$ , respectively. The corrected signal efficiencies and their uncertainties are shown in Table I.

In calculating branching fractions, we assume  $\mathcal{B}(Y(4S) \rightarrow B^0 \bar{B}^0) = \mathcal{B}(Y(4S) \rightarrow B^+ B^-) = 0.5$ . The 90% C.L. is taken as the largest value of the efficiency-corrected signal yield at which  $2\Delta \log \mathcal{L} = 1.28^2$ . We include systematic uncertainties by increasing the efficiency corrected signal yield by 1.28 times its systematic uncertainty. Table I shows the resulting upper limits on the branching fractions.

Using the measured value of  $\mathcal{B}(B \rightarrow K^* \gamma)$  [22], we calculate a limit of  $\overline{\mathcal{B}}[B \rightarrow (\rho/\omega)\gamma]/\mathcal{B}(B \rightarrow K^* \gamma) < 0.029$  at 90% C.L. This limit is used to constrain the ratio of CKM elements  $|V_{td}/V_{ts}|$  by means of the equation [3, 6]:

$$\frac{\overline{\mathcal{B}}[B \rightarrow (\rho/\omega)\gamma]}{\mathcal{B}(B \rightarrow K^* \gamma)} = \left| \frac{V_{td}}{V_{ts}} \right|^2 \left( \frac{1 - m_\rho^2/M_B^2}{1 - m_{K^*}^2/M_B^2} \right)^3 \zeta^2 [1 + \Delta R],$$

where  $\zeta$  describes the flavor-SU(3) breaking between  $\rho/\omega$  and  $K^*$ , and  $\Delta R$  accounts for annihilation diagrams. Both  $\zeta$  and  $\Delta R$  must be taken from theory [3, 6, 23]. Following [3], we choose the values  $\zeta = 0.85 \pm 0.10$ , and  $\Delta R = 0.10 \pm 0.10$ , which is the average over the values given for the three modes. We find the limit

$|V_{td}|/|V_{ts}| < 0.19$  at 90% C.L. ignoring the theoretical uncertainties. Our upper limit on  $|V_{td}|/|V_{ts}|$  constrains  $|V_{td}| < 0.008$  at 90% C.L. assuming  $|V_{ts}| = |V_{cb}|$  [18]; this lies within the current 90% confidence interval  $0.005 < |V_{td}| < 0.014$ , which is obtained from a fit to experimental results on the CKM matrix elements [18]. Varying the values of  $\zeta$  and  $\Delta R$  within their uncertainties leads to changes in the limits by  $\pm 0.03$  and  $\pm 0.001$  for  $|V_{td}|/|V_{ts}|$  and  $|V_{td}|$ , respectively.

In conclusion, we have found no evidence for the exclusive  $b \rightarrow d\gamma$  transitions  $B \rightarrow \rho\gamma$  and  $B^0 \rightarrow \omega\gamma$  in 211 million  $\Upsilon(4S) \rightarrow B\bar{B}$  decays studied with the *BABAR* detector. The 90% C.L. upper limits on the branching fractions and  $|V_{td}|/|V_{ts}|$  are significantly lower than our previous values [7] and restrict the range indicated by SM predictions [3, 4].

We are grateful for the excellent luminosity and machine conditions provided by our PEP-II colleagues, and for the substantial dedicated effort from the computing organizations that support *BABAR*. The collaborating institutions wish to thank SLAC for its support and kind hospitality. This work is supported by DOE and NSF (USA), NSERC (Canada), IHEP (China), CEA and CNRS-IN2P3 (France), BMBF and DFG (Germany), INFN (Italy), FOM (The Netherlands), NFR (Norway), MIST (Russia), and PPARC (United Kingdom). Individuals have received support from CONACyT (Mexico), A. P. Sloan Foundation, Research Corporation, and Alexander von Humboldt Foundation.

- [2] The named member of a charge-conjugate pair of particles stands for either.
- [3] A. Ali, E. Lunghi, and A. Parkhomenko, Phys. Lett. B **595**, 323 (2004).
- [4] S. W. Bosch and G. Buchalla, Nucl. Phys. B **621**, 459 (2002).
- [5] See, for example, S. Bertolini, F. Borzumati, and A. Masiero, Nucl. Phys. B **294**, 321 (1987); H. Baer and M. Brhlik, Phys. Rev. D 55, 3201 (1997); J. Hewett and J. Wells, Phys. Rev. D 55, 5549 (1997); M. Carena *et al.*, Phys. Lett. B **499**, 141 (2001).
- [6] A. Ali and A. Y. Parkhomenko, Eur. Phys. J. C **23**, 89 (2002).
- [7] B. Aubert *et al.* [*BABAR* Collaboration], Phys. Rev. Lett. **92**, 111801 (2004).
- [8] T.E. Coan *et al.* [*CLEO* Collaboration], Phys. Rev. Lett. **84**, 5283 (2000).
- [9] B. Aubert *et al.* [*BABAR* Collaboration], Nucl. Instrum. Meth. A **479**, 1 (2002).
- [10] B. Aubert *et al.* [*BABAR* Collaboration], Phys. Rev. Lett. **88**, 101805 (2002).
- [11] We use the Stuttgart Neural Network Simulator (<http://www-ra.informatik.uni-tuebingen.de/SNNS>) to train a neural net with one hidden layer of 5 or 6 nodes.
- [12] G.C. Fox and S. Wolfram, Nucl. Phys. B **149**, 413 (1979).
- [13] B. Aubert *et al.* [*BABAR* Collaboration], Phys. Rev. Lett. **89**, 201802 (2002).
- [14] R.A. Fisher, Annals Eugen. **7**, 179 (1936).
- [15] E.D. Bloom and C.W. Peck, Ann. Rev. Nucl. Part. Sci., **33** 143 (1983).
- [16] H. Albrecht *et al.* [*ARGUS* Collaboration], Z. Phys. **C48**, 543 (1990).
- [17] K. Cranmer, Comput. Phys. Commun. **136**, 198 (2001).
- [18] S. Eidelman *et al.* [Particle Data Group Collaboration], Phys. Lett. B **592**, 1 (2004).
- [19] B. Aubert *et al.* [*BABAR* Collaboration], hep-ex/0403046; B. Aubert *et al.* [*BABAR* Collaboration], hep-ex/0403025.
- [20] B. Aubert *et al.* [*BABAR* Collaboration], Phys. Rev. Lett. **93**, 051802 (2004).
- [21] J. Dragic *et al.* [*Belle* Collaboration], hep-ex/0405068.
- [22] B. Aubert *et al.* [*BABAR* Collaboration], hep-ex/0407003.
- [23] B. Grinstein and D. Pirjol, Phys. Rev. D **62**, 093002 (2000).

\* Now at Department of Physics, University of Warwick, Coventry, United Kingdom

† Also with Università della Basilicata, Potenza, Italy

‡ Also with IFIC, Instituto de Física Corpuscular, CSIC-Universidad de Valencia, Valencia, Spain

§ Deceased

[1] For a review, see K. Lingel, T. Skwarnicki, and J.G. Smith, Ann. Rev. Nucl. Part. Sci. **48**, 253 (1998).