# BEAM TRANSPORT SYSTEMS OR CHARGED PARTICLE SPECTROMETERS * 

Karl L. Brown<br>Stanford Linear Accelerator Center Stanford, California

## Summary

By extrapolating the systematics of the general first- and second-order theory of beam transport optics ( $1,2,3$ ) to include higher order multipole terms, it has been possible to evolve a simple, step by step, procedure for the design of high resolving power static-magnetic beam transport systems. The choice of the appropriate dipole and quadrupole elements for a given system may be determined once the resolving power, solid angle, momentum range and detector system of the instrument have been specified. The partial derivative of any nth-order aberration coefficient with respect to an nth-order multipole component located anywhere in the system has been derived. From this "coupling coefficient", the strength and the optimum location of multipole element(s) to correct or modify a given aberration or group of aberrations is uniquely determined.

```
\({ }^{*}\) Work supported by the U.S. Atomic Energy Commission.
    (Submitted to the Third International Magnet Symposium, Hamburg, Germany,
    May 1970.)
```

coefficients used in the expansion of the magnetic scalar potential $\varphi$.
a the radial distance from the central trajectory to the pole of a multipole element.
$B_{0} \quad$ the magnetic field at the pole of a multipole element.
$B_{x}, B_{y}, B_{t} \quad$ the magnetic field components corresponding to the curvilinear coondinates $x, y$ and $t$.
$B \rho=\frac{B}{h}=\frac{P_{0}}{e}$ the magnetic rigidity of the central trajectory.
$\left.\begin{array}{l}c_{x}=\left(x \mid x_{0}\right) \\ c_{y}=\left(y \mid y_{0}\right)\end{array}\right\}$ First-order Taylor coefficients in the trajectory equations.
$\begin{array}{ll}d_{x}=(x \mid \delta) & \text { the spatial dispersion, a first-order Taylor coefficient in the } \\ \text { trajectory equation. }\end{array}$
e : the charge of the particle
$G_{i}(t, \tau) \quad$ the Green's functions used to evaluated the Taylor coefficients.
$h(\tau)=\frac{I}{\rho_{0}}$ the curvature of the central trajectory.
$K_{n}(\tau) \quad$ the multipole strength per unit length .
$\boldsymbol{l}_{0} \quad$ the path length of the central trajectory.
$\boldsymbol{\ell}$ the path length of an arbitrary trajectory.
I. the length of a multipole element.
$P_{0} \quad$ the momentum of the central trajectory.
$P \quad$ the momentum of an arbitrary trajectory.
$R \quad$ the radius of curvature of an entrance or exit boundary of a magnet.
$R_{1} \quad$ the first-order momentum resolving power of a system.
$S_{n}=\int_{0}^{L} K_{n}(\tau) d \tau \quad$ the total strength of a multipole element.
$\left.\begin{array}{l}s_{x}=\left(x \mid x_{0}^{\prime}\right) \\ s_{y}=\left(y \mid y_{0}^{\prime}\right)\end{array}\right\}$ First-order Taylor coefficients in the trajectory equations.
$x$ the transverse coordinates of the curvilinear coordinate system $(x, y, t) . \quad x^{\prime}=\frac{d x}{d t} ; \quad y^{\prime}=\frac{d y}{d t}$.
$Z(x) \quad$ the coordinate perpendicular to $x$ and lying in the midplane which describes the contour of an entrance or exit field boundary of a magnet.
$\alpha \quad$ the angle of bend of the central trajectory in a dipole magnet.
B the angle of rotation of an entrance or exit field boundary.
the sextupole strength per unit length of a non-uniform field (not to be confused with the previous $\beta$ ).
$\gamma h^{4}$
the octupole strength per unit length of a non-uniform field expansion.
$\delta=\frac{\Delta P}{P_{0}} \quad$ the fractional momentum deviation of an arbitrary trajectory from that of the central trajectory.
$\boldsymbol{\kappa}, \boldsymbol{\lambda}, \mu, \boldsymbol{\nu}, \boldsymbol{X}$ symbols used as exponents in the Taylor coefficients where $\mathrm{n}=(\boldsymbol{\kappa}+\lambda+\mu+\nu+X)$ is the aberration or multipole order in the expansion.
prime (') means $\frac{d}{d t}$, the derivative with respect to the curvilinear coordinate $t$.
$\tau$
the location of a multipole component ; a variable of integration.
$\varphi$ the magnetic scalar potential.

## I. Introduction

Within the last two decades, significant advances have been made in the understanding of charged particle optics. Perhaps the first major contribution was the development of the theory of the Alternating Gradient Synchrotron (A.G.S.) by Courant, Livingston, and Snyder(4) which led to the first-order matrix algebra formulation of beam-transport optics. Subsequent to this a second-order matrix algebra was developed by Brown, Belbeoch, and Bounin(5); followed by the development at SLAC of the digital computer program called TRANSPORT(6) that is widely used today in many laboratories for solving first- and second-order staticmagnetic beam transport problems. In principle, the second-order matrix formalism may be extended to any order, but in practice this approach has proved to be too cumbersome. Thus beyond second-order it has been more efficient to use computer ray-tracing programs which integrate the basic differential equation of motion of the charged particles through the known or assumed magnetic fields. The fundamental difficulty with ray-tracing has been the required computational time to complete a design involving the minimization of many higher-order aberrations.

In this report, we will outline a systematic procedure for the design of high-resolution systems based upon the extrapolation of the first-and secondorder theory ( $1,2,3$ ) to include higher-order multipole components. A general equation has been derived for the coupling coefficient of an nth-order multipole to any given nth-order aberration coefficient. As will be shown later, these coupling coefficients are a function only of the characteristic first-order trajectories (matrix elements) introduced and defined in References 1 and 2.

Given this information, a systematic procedure for designing high resolution beam transport systems is as follows:

1) Find a satisfactory first-order solution to the problem using TRANSPORT or its equivalent.
2) Calculate and make the necessary corrections to the second-order aberrations by introducing sextupole components into the system. The "best" locations and strengths of the sextupole components required may be selected via the coupling coefficients for the aberrations to be minimized.
3) Calculate and make the necessary corrections (via ray-tracing) to the third-order aberrations by introducing octupole components into the system. (Note that an nth-order multipole couples with terms of order $n$ or higher but not wilh terms of order lower than $n$. Thus an octupole component will not disturb the first- and second-order solutions already found from steps 1 and 2.)
4) Repeat the above procedure up to the multipole order desired or needed to achieve the design objectives.

If the design requires a solution to nth-order and multipoles at each order are necessary to minimize the aberrations, the number of computer runs previously needed to complete a design was at least $(n+m)^{2}$. Having a knowledge of the coupling coefficients, after the first-order design has been selected, now (in principle) reduces the number of computer runs required to n. Since ray-tracing is very time consuming, this is indeed a significant saving.

## II. Theory*

The following results are applicable to static-magnetic charged particle optical systems possessing median plane symmetry. As in Ref. 1, we shall use a right-handed curvilinear coordinate system ( $x, y, t$ ) where $x$ and $y$ are the transverse coordinates. $x$ is the outward normal distance in the median plane away from the central trajectory, $y$ is the perpendicular distance from the median plane, $t$ is the distance along the central trajectory, and $h=h(t)$ is the curvature of the central trajectory.

The existence of the median plane requires that the scalar potential $\varphi$ be an odd function of $y$, i.e., $\varphi(x, y, t)=-\varphi(x,-y, t)$. The most general form of $\varphi$ may therefore, be expressed as follows:

$$
\begin{equation*}
\varphi(x, y, t)=\sum_{m=0}^{\infty} \sum_{n=0}^{\infty} A_{2 m+1, n} \frac{x^{n}}{n!} \frac{y^{2 m+1}}{(2 m+1)!} \tag{I}
\end{equation*}
$$

where the coefficients $A_{2 m+1, n}$ are functions of $t$.
In this coordinate system, the differential line element $d T$ is given by

$$
d T^{2}=d x^{2}+d y^{2}+(1+h x)^{2} d t^{2}
$$

The Laplace equation has the form

$$
\begin{equation*}
\nabla^{2} \varphi=\frac{1}{(1+h x)} \frac{\partial}{\partial x}\left[(1+h x) \frac{\partial \varphi}{\partial x}\right]+\frac{\partial^{2} \varphi}{\partial y^{2}}+\frac{1}{(1+h x)} \frac{\partial}{\partial t}\left[\frac{1}{(1+h x)} \frac{\partial \varphi}{\partial t}\right]=0 \tag{2}
\end{equation*}
$$

Substitution of (1) into (2) gives the following recursion formula for the coefficients:

$$
\begin{align*}
-A_{2 m+3, n}= & A_{2 m+1, n}^{\prime \prime}+n h A_{2 m+1, n-1}^{\prime \prime}-n h^{1} A_{2 m+1, n-1}^{1}+A_{2 m+1, n+2}  \tag{3}\\
& +(3 n+1) h A_{2 m+1, n+1}+n(3 n-1) h^{2} A_{2 m+1, n}+n(n-1)^{2} h^{3} A_{2 m+1, n-1} \\
& +3 n h A_{2 m+3, n-1}+3 n(n-1) h^{2} A_{2 m+3, n-2}+n(n-1)(n-2) h^{3} A_{2 m+3, n-3}
\end{align*}
$$

where prime means $\frac{d}{d t}$, and where it is understood that all coefficients $A$ with one or more negative subscripts are zero. This recursion formula expresses all

* The notation used in this report follows that used in Ref. 1 unless otherwise indicated.
the coefficients in terms of the midplane field $B_{y}(x, 0, t)$ :
where

$$
\begin{equation*}
A_{1, n}=\left(\frac{\partial^{n} B_{y}}{\partial x^{n}}\right)_{\substack{x=0 \\ y=0}}=\text { functions of } t . \tag{4}
\end{equation*}
$$

Since $\varphi$ is an odd function of $y$, on the median plane we have $B_{x}=B_{t}=0$. The normal (in $x$ direction) derivatives of $B_{y}$ on the reference curve defines $B_{y}$ over the entire median plane, hence the magnetic field $\vec{B}$ over the whole space. The components of the field are expressed in terms of $\varphi$ explicitly by $\vec{B}=\vec{\nabla} \varphi$ or

$$
\begin{align*}
& B_{x}=\frac{\partial \varphi}{\partial x}=\sum_{m=0}^{\infty} \sum_{n=0}^{\infty} A_{2 m+1, n+1} \frac{x^{n}}{n!} \frac{y^{2 m+1}}{(2 m+1)!} \\
& B_{y}=\frac{\partial \varphi}{\partial y}=\sum_{m=0}^{\infty} \sum_{n=0}^{\infty} A_{2 m+1, n} \frac{x^{n}}{n!} \frac{y^{2 m}}{(2 m)!} \\
& B_{t}=\frac{1}{(1+h x)} \frac{\partial \varphi}{\partial t}=\frac{1}{(1+h x)} \sum_{m=0}^{\infty} \sum_{n=0}^{\infty} A_{2 m+1, n} \frac{x^{n}}{n!} \frac{y^{2 m+1}}{(2 m+1)!} \tag{5}
\end{align*}
$$

The expression for the magnetic field on the midplane is

$$
\begin{equation*}
B_{y}(x, 0, t)=\sum_{n=0}^{\infty} A_{l, n} \frac{x^{n}}{n!} \tag{6}
\end{equation*}
$$

At this point we deviate from the notation and formalism of Ref. I and introduce $K_{n}(t)$, the multipole strength per unit length; and $S_{n}$, the total multipole strength of a static-magnetic field.

We rewrite equation (6) as

$$
\begin{equation*}
B_{y}(x, o, t)=B p \quad \sum_{n=0}^{\infty} K_{n}(t) x^{n} \tag{7}
\end{equation*}
$$

Where $B \rho=\frac{B}{h}=\frac{P_{0}}{e}$ is the magnetic rigidity of a particle of monentum $P_{0}$ and charge e along the central trajectory; from which

$$
\begin{equation*}
K_{n}(t)=\left(\frac{1}{B \rho}\right)\left(\frac{1}{n!}\right)\left(A_{I}, n^{(t)}\right)=\left(\frac{1}{B \rho}\right)\left(\frac{1}{n!}\right)\left(\frac{\partial^{n} B_{y}}{\partial x^{n}}\right)_{x=y=0} \tag{8}
\end{equation*}
$$

We define $S_{n}$ as

$$
\begin{equation*}
s_{n}=\int_{0}^{L} K_{n}(t) d t \tag{9}
\end{equation*}
$$

$S_{n}$ so defined is the strength of the nth-orier multipole component of a field over the interval of integration.

Multipole Strengths for Pure Multipole Fields
Consider the scalar potential of an nth-order $[2(n+1)$ pole $]$ pure multipole element:

$$
\begin{equation*}
\varphi=\frac{B_{0} r^{n+1}}{(n+1) a^{n}}[\sin (n+1) \theta] \tag{10}
\end{equation*}
$$

where

$$
x=r \cos \theta \text { and } y=r \sin \theta
$$

$B_{0}$ is the field at the pole and a is the radial distance to the pole from the central trajectory.

Expanding $\varphi$ as a function of $x$ and $y$ and differentiating, we have

$$
B_{y}=\frac{\partial \varphi}{\partial y}=\frac{B_{o}}{a^{n}}\left[x^{n}+\cdots \cdots \cdot \cdot\right]
$$

From which

$$
K_{n}=\left(\frac{B_{0}}{a^{n}}\right)\left(\frac{1}{B \rho}\right)
$$

and

$$
\begin{equation*}
S_{n}=\left(\frac{B_{o}}{a^{n}}\right)\left(\frac{I_{1}}{B P}\right) \tag{II}
\end{equation*}
$$

Where $I$ is the Iength of the multipole element.

For a dipole $n=0$ and the dipole strength is

$$
S_{0}=\frac{L}{\rho}=\alpha \text { (The angle of bend of the central trajectory) }
$$

For a quadrupole $n=1$ and

$$
S_{1}=\left(\frac{B_{0}}{a}\right)\left(\frac{L}{B \rho}\right)
$$

For a sextupole $\mathrm{n}=2$ and

$$
S_{2}=\left(\frac{B_{0}}{a^{2}}\right)\left(\frac{L}{B p}\right)
$$

etc. for higher-order multipoles.

Multipole Strengths for a Non-Uniform Field Expansion
From the midplane field expansion of a non-uniform magnetic field

$$
\begin{equation*}
B_{y}(x, 0, t)=B_{y}(0,0, t)\left[1-n h x+\beta(h x)^{2}+\gamma(h x)^{3}+\cdots \cdot \cdot \cdot\right] \tag{12}
\end{equation*}
$$

the multipole strength factors are:

$$
K_{0}=h, \quad K_{1}=-n h^{2}, \quad K_{2}=\beta h^{3}, \quad \text { etc. }
$$

and $S_{n}$ evaluated over the length $L$ of the central trajectory is:

$$
\begin{aligned}
& S_{0}=h L=\alpha \text { as before, } \\
& S_{1}=-n h^{2} L, \text { and } S_{2}=\beta h^{3} L_{1}, \text { etc. }
\end{aligned}
$$

Multipole Strengths for a Contoured Entrance or Exit Boundary of a Magnet
A third method of introducing multipole components is via a curved entrance or exit boundary of a magnet. To calculate the multipole strengths in this case, we integrate equation (7), holding $x$ constant, as follows:

$$
\begin{equation*}
\int_{0}^{L} B_{y}(x, 0, t) d t=B \rho \sum x^{n} \int_{0}^{L} K_{n}(t) d t=B \rho \sum S_{n} x^{n} \tag{13}
\end{equation*}
$$

To relate this to the field boundary, we assume $B_{y}$ to be a constant inside the effective field boundary and zero outside (i.e., we ignore the finite extent of the fringing field). In this sharp-cutoff approximation, the field boundary $Z=Z(x)$ is:

$$
\begin{equation*}
z=\frac{1}{B_{y}} \int_{C}^{I} B_{y}(x, 0, t) d t=\frac{1}{h} \sum S_{n} x^{n}=-x \tan \beta+\frac{S_{2}}{h} x^{2}+\cdots \tag{14}
\end{equation*}
$$

where $h=\frac{1}{\rho}$ and $\beta$ is the angle of rotation of the entrance or exit face of the magnet at $x=0$. A positive $\beta$ implies radial ( $x$ ) defocusing and transverse (y) focusing. We note that:

$$
S_{1}=-h \tan \beta=\text { The "quadrupole strength" }
$$

The radius of curvature of the boundary is related to the sextupole strength as follows:

$$
\frac{1}{R}=\frac{Z^{\prime \prime}}{\left(1+Z^{\prime 2}\right)^{3 / 2}}=\frac{2 s_{2}}{h \sec ^{3} \beta}
$$

or

$$
S_{2}=\frac{h \sec ^{3} \beta}{2 R}=\text { The "sextupole strength" }
$$

From equation (13), we note that a positive multipole component of the field increases the $\int$ Bdt for a positive $x$; thus a positive sextupole is represented by a concave surface of the entrance or exit boundary.

The Description of the Trajectories as a Taylor's Expansion
The deviation of an arbitrary trajectory from the central trajectory is described by expressing $x$ and $y$ as functions of $t$. The expressions will also contain $x_{0}, y_{0}, x_{0}^{\prime}, y_{0}^{\prime}$ and $\delta$, where the subscript 0 indicates that the quantity is evaluated at $t=0$. The prime (') denotes the derivative with respect to $t$, and $\delta=\frac{\Delta P}{P_{0}}$ is the fractional momentum deviation of the ray from that of the central trajectory. These five initial boundary values will have the value zero for the central trajectory itself. $x$ and $y$ are expressed as a five-fold Taylor expansion using these initial boundary values. The expansions are written:

$$
\begin{align*}
& x(t)=\sum\left(x \mid x_{0}^{\kappa} y_{0}^{\lambda} x_{0}^{\prime \mu} y_{0}^{\prime v} \delta_{\delta} x_{0} x_{0}^{\kappa} y y_{0}^{\lambda} x_{0}^{\prime \mu} y_{0}^{\prime \nu} \delta_{\delta}^{X}\right. \\
& y(t)=\sum\left(y \mid x_{0}^{\kappa} y_{0}^{\lambda} x_{0}^{\prime \mu} y_{0}^{\prime v} \delta x_{0} x_{0}^{\kappa} y_{0}^{\lambda} x_{0}^{\prime \mu} y_{0}^{\prime \nu} \delta_{\delta} x\right. \tag{15}
\end{align*}
$$

Here, the parentheses are symbols for the Taylor coefficients; the first part of the symin identifies the coordinate represented by the expansion, and the second indicates the term in question. These coefficients are functions of $t$ to be determined. The symbol $\sum$ indicates summation over zero and all positive integer values of the exponents $\kappa, \lambda, \mu, \nu, X ;$. The constant term is zero, and the terms that would indicate a coupling between the coordinates $x$ and $y$ are also zero; this results from the midplane symmetry. Thus we have

$$
\begin{align*}
& (x \mid 1)=(y \mid 1)=0 \\
& \left(x \mid y_{0}\right)=\left(y \mid x_{0}\right)=0 \\
& \left(x \mid y_{0}^{\prime}\right)=\left(y \mid x_{0}^{\prime}\right)=0 \tag{16}
\end{align*}
$$

Here, the first line is a consequence of choosing the central trajectory as the reference axis, while the second and third lines follow directly from considerations of median plane symmetry.

Since they will appear often in the formalism, it is convenient to introduce the following abbreviations for the first-order Taylor coefficients:

$$
\begin{array}{lll}
\left(x \mid x_{0}\right)=c_{x}(t) & \left(x \mid x_{0}^{1}\right)=s_{x}(t) & (x \mid \delta)=d_{x}(t) \\
\left(y \mid y_{0}\right)=c_{y}(t) & \left(y \mid y_{0}^{1}\right)=s_{y}(t) \tag{17}
\end{array}
$$

When the transverse position of an arbitrary trajectory at position $t$ is written as a first-order Taylor's expansion as a function of the initial boundary conditions, the above five quantities are just the coefficients appearing in the expansion for the transverse coordinates $x$ and $y$ as follows:

$$
x(t)=c_{x}(t) x_{0}+s_{x}(t) x_{0}^{\prime}+d_{x}(t) \delta+\text { higher-order terms }
$$

and

$$
y(t)=c_{y}(t) y_{0}+s_{y}(t) y_{0}^{\prime}+\text { higher-order terms } .
$$

III. Solution of the Equations of Motion

The general differential equation of motion of a charged particle in a static-magnetic field valid to all orders in $x$ and $y$ and their derivatives as derived in Ref. 1, equation (5) is:

$$
\begin{align*}
& \hat{x}\left\{\left[x^{\prime \prime}-h(1+h x)\right]-\frac{x^{\prime}}{\left(T^{\prime}\right)^{2}}\left[x^{\prime} x^{\prime \prime}+y^{\prime} y^{\prime \prime}+(1+h x)\left(h x^{\prime}+h^{\prime} x\right)\right]\right\} \\
& +\hat{y}\left\{y^{\prime \prime}-\frac{y^{\prime}}{\left(T^{\prime}\right)^{2}}\left[x^{\prime} x^{\prime \prime}+y^{\prime} y^{\prime \prime}+(1+h x)\left(h x^{\prime}+h^{\prime} x\right)\right]\right\} \\
& +\hat{t}\left\{\left(2 h x^{\prime}+h^{\prime} x\right)-\frac{(1+h x)}{\left(T^{\prime}\right)^{2}}\left[x^{\prime} x^{\prime \prime}+y^{\prime} y^{\prime \prime}+(1+h x)\left(h x^{\prime}+h^{\prime} x\right)\right]\right\} \\
& =\frac{e}{P} T^{\prime}\left(\overrightarrow{T^{\prime}} \times \vec{B}\right)=\frac{e}{P} T^{\prime}\left\{\hat{x^{\prime}\left[y^{\prime} B_{t}-(1+h x) B_{y}\right]} \begin{array}{r}
+\hat{y}\left[(1+h x) B_{x}-x^{\prime} B_{t}\right] \\
\\
\end{array} \begin{array}{l}
\left.+\hat{t}\left[x^{\prime} B_{y}-y^{\prime} B_{x}\right]\right\}
\end{array}\right.
\end{align*}
$$

If this equation is solved to nth-order for the Taylor's coefficients of equation (15), it will be observed that the result has the remarkably simple form:

$$
\begin{align*}
& + \text { Terms containing } K_{o}, \cdots \ldots . . K_{n-1} \tag{19}
\end{align*}
$$

where the variable of integration is $\tau$ and $n=(\kappa+\lambda+\mu+\nu+X)$.
The $x_{i}$ have the following meaning:

$$
x_{1}=x(t) \quad x_{2}=x^{\prime}(t) \quad x_{3}=y(t) \quad x_{4}=y^{\prime}(t)
$$

$c_{x}, c_{y}, s_{x}, s_{y}$, and $d_{x}$ are defined by equation (17) and in general are functions of the variable of integration $\tau$ over the interval of integration. $K_{n}$ is defined by equation (8) and in general is also a function of $T$.

The $G_{i}$ 's are Green's functions where:

$$
\begin{align*}
& G_{1}(t, \tau)=\left(x(t) \mid x^{\prime}(\tau)\right)=s_{x}(t) c_{x}(\tau)-c_{x}(t) s_{x}(\tau) \\
& G_{2}(t, \tau)=\left(x^{\prime}(t) \mid x^{\prime}(\tau)\right)=s_{x}^{\prime}(t) c_{x}(\tau)-c_{x}^{\prime}(t) s_{x}(\tau) \\
& G_{3}(t, \tau)=\left(y(t) \mid y^{\prime}(\tau)\right)=s_{y}(t) c_{y}(\tau)-c_{y}(t) s_{y}(\tau) \\
& G_{4}(t, \tau)=\left(y^{\prime}(t) \mid y^{\prime}(\tau)\right)=s_{y}^{\prime}(t) c_{y}(\tau)-c_{y}^{\prime}(t) s_{y}(\tau) \tag{20}
\end{align*}
$$

Note that the $G_{i}$ 's are just first-order Taylor's coefficients measured from the location ( $\tau$ ) of the multipole component to the end of the system ( $t$ ).

Thus we see that the coupling coefficient to an nth-order multipole is a function only of the first-order matrix elements $c_{x}, c_{y}, s_{x}, s_{y}, d_{x}$ and their derivatives with respect to $t$.

From median-plane symmetry considerations, the allowed aberrations are those with $y$ and/or $y^{\prime}$ appearing an even number of times in the Taylor coefficient. For example $\left(x \mid x_{0}^{2}\right),\left(x \mid y_{0} y_{0}^{1}\right)$ and ( $\left.y \mid y_{0}^{2} y_{0}^{1}\right)$ are allowed aberrations; whereas $\left(x \mid y_{0}\right),\left(x \mid x_{0}^{2} y_{0}^{1}\right)$ or $\left(y \mid y_{0}^{2}\right)$ are not allowed and are therefore equal to zero.

The minus sign is used when $y$ and/or $y^{\prime}$ appear $0,4,8,12 \cdots$ times and the plus sign is used when $y$ and/or $y^{\prime}$ appear 2, 6, $10 \ldots .$. times. For example for the coefficients $\left(x \mid x_{0}^{2}\right)$ and $\left(y \mid y_{0}^{3}\right)$, the minus sign is applicable; whereas for the coefficients $\left(x \mid y_{0}^{2}\right)$ and $\left(y^{\prime} \mid y_{0}^{3} y_{0}^{\prime 2}\right)$ the plus sign is applicable.

Equation (19) is derived by observing in the pattern of the solution of the differential equation that an nth-order aberration term containing the nth-order multipole strength factor $K_{n}$ cannot include multipole strength factors of lower order than $n$; or stated physically, an nth-order multipole cannot couple to aberrations (terms) of order lower than $n$. This fact allows the recursion formula equation (3) to be reduced to the simple form

$$
\begin{equation*}
A_{2 m+3, n}=-A_{2 m+1, n+2} \tag{21}
\end{equation*}
$$

in so far as it applies to the derivation of nth-order terms containing only $K_{n}$. As a consequence, the scalar potential for deriving these terms assumes the simplied form

$$
\begin{equation*}
\varphi(x, y, t)=\sum_{m=0}^{\infty} \sum_{n=0}^{\infty}(-1)^{m} A_{1}, 2 m+n \frac{x^{n}}{n!} \frac{y^{2 m+1}}{(2 m+1)!} \tag{22}
\end{equation*}
$$

From which, it follows that

$$
B_{x}(x, y, t)=\left(\frac{P_{0}}{e}\right) \sum_{m=0}^{\infty} \sum_{n=0}^{\infty}(-1)^{m} K_{2 m+n+1} \frac{(2 m+n+1)!}{n!(2 m+1)!} x^{n} y^{2 m+1}
$$

and

$$
\begin{equation*}
B_{y}(x, y, t)=\left(\frac{P_{0}}{e}\right) \sum_{m=0}^{\infty} \sum_{n=0}^{\infty}(-1)^{m} n_{2 m+n} \frac{(2 m+n)!}{n!2 m!} x^{n} y^{2 m} \tag{23}
\end{equation*}
$$

For terms containing only $K_{n}$, the basic differential equations assume the form:

$$
\begin{align*}
& x^{\prime \prime}+\cdots \cdots \cdot=-\frac{e}{P} B_{y} \\
& y^{\prime \prime}+\cdots \cdots \cdots=\frac{e}{P} B_{x} \tag{24}
\end{align*}
$$

Substituting the Taylor's expansion of equation (15) and solving for the nth-order terms using a conventional Green's function solution (see Ref. 1) yields equation (19) above.

## IV. Interpretation and Use of Equation (19)

For most practical cases of interest, $K_{n}$ will be a constant over the interval of integration. In this event we may define the coupling coefficient of an nth-order multipole to an nth-order aberration as the partial derivative of equation (19) with respect to the $K_{n}$ in question as follows:

where now the interval of integration is over the multipole length $L$ represented by $K_{n}$. For a distributed multipole component (such as in a non-uniform field bending magnet), equation (25) is used.

In many cases where a curved entrance or exit pole contour is used or a short multipole magnet is used such that the characteristic first-order functions $c_{x}, c_{y}, s_{x}, s_{y}$ and $d_{x}$ are essentially constants over the interval of integration (the length of one multipole), then the coupling coefficient is best defined as the partial derivative of equation (19) with respect to $S_{n}$ as follows:

Examples
Assume a situation where the end of the system is a point-to-point image or the origin (i.e., $s_{x}(t)=0$ ), then using equation 26 , the coupling coefficients of a sextupule of strength $S_{2}$ to various second order aberration coefficients are:

$$
\begin{align*}
& \frac{\partial\left(x \mid x_{0}^{\prime} \delta\right)}{\partial s_{2}}=c_{x}(t) s_{x}^{2} d_{x} \\
& \frac{\partial\left(x \mid x_{0}^{\prime}\right)}{\partial s_{2}}=c_{x}(t) s_{x}^{3} \\
& \frac{\partial\left(x \mid y y_{0} y_{0}^{\prime}\right)}{\partial s_{2}}=-2 c_{x}(t) c_{y} s_{x} s_{y} \tag{27}
\end{align*}
$$

etc. Where the Green's function used in these examples is

$$
G_{1}=s_{x}(t) c_{x}-c_{x}(t) s_{x}=-c_{x}(t) s_{x} \quad\left(\text { since } s_{x}(t)=0\right. \text { for point-to-point }
$$

The aberration and $c_{x}(t)$ are evaluated at the end of the system. $c_{x}(t)$ is equal to the magnification $M_{x}$ in the examples given. The remaining coefficients $c_{y}, s_{x}, s_{y}$ and $d_{x}$ are evaluated at the location of the sextupole $S_{2}$. The above results are in agreement with Table VII of Ref. 1.

To illustrate a more complex example, consider the fourth-order aberration coefficient ( $y \mid y_{0}^{2} y_{0}^{\prime \delta}$ ) and assume parallel-to-point imaging in the $y$ coordinate (i.e., $c_{y}(t)=0$ ). The appropriate Green's function is:

$$
G_{3}=s_{y}(t) c_{y}-c_{y}(t) s_{y}=s_{y}(t) c_{y}
$$

and the coupling coefficient to a fourth-order multipole of strength $S_{4}$ is:

$$
\begin{equation*}
\frac{\partial\left(y \mid y_{0}^{2} y_{0}^{\prime} \delta\right)}{\partial s_{4}}=-\left(\frac{4!}{2!}\right) \quad s_{y}(t) c_{y}^{3} s_{y} d x \tag{28}
\end{equation*}
$$

where again the aberration coefficient $\left(y \mid y_{0}^{2} y_{o}^{\prime} \delta\right)$ and $s_{y}(t)$ are evaluated at the end of the optical system and $c y, s_{y}$, and $d_{x}$ are evaluated at the location of the fourth-order multipole $S_{4}$.

## V. A Systematic Procedure for Designing High Resolution Systems <br> First-Order Considerations

In meny respects; the determination of a satisfactory first-order magneticoptical design is more difficult to achieve than is the subsequent higher-order design. This is true not only because the basic equipment configuration is dominated by first-order optical considerations but also because the choice of the first-order optics affects the magnitude of all higher-order aberrations and the ease with which these aberrations may be minimized by introducing multipole components into the design.

The dominating design parameters that must be clearly specified in order to evolve a first-order design are the momentum resolving power; the spatial resolution of the particle detector system to be used (this determines the momentum dispersion required); the required phase space acceptance (the solid angle, the source size, and the momentum range) of the instrument, and the first-order imaging requirements in both the $x$ and $y$ coordinates.

Given the above specifications (assuming they are self-consistent), the optical mode and physical configuration of the instrument may be determined. Often, more than one theoretical solution exists; in which case the choice is usually resolved by practical or economic considerations. In other cases, no solution is evident and the basic specifications must be modified accordingly. In any event, the following equations and discussion are applicable to the solution of the problem.

## 1) First-Order Resolving Power

A general equation for the first-order resolving power has been derived in References ( 1,2 , and 3). For point-to-point imaging the first-order momentum resolving power $R_{1}$ is defined as the ratio of the momentum dispersion at the image plane to the total image size. Thus if $2 x_{0}$ is the total source size then from Reference 1 we have:

$$
\begin{equation*}
R_{1}=\frac{P}{\Delta P}=\left|\frac{d_{x}(t)}{2 x_{0} c_{x}(t)}\right|=\frac{1}{2 x_{0}}\left|\int_{0}^{t} s_{x}(\tau) h(\tau) d \tau\right| \tag{29}
\end{equation*}
$$

Note that $h(\tau) d \tau=d \alpha$ is the differential angle of bend of the central trajectory of the optical system.

Equation (29) may be expressed in a number of useful forms. If we consider a particle originating at the source with $x_{0}=0$ and $\delta=\frac{\Delta P}{P_{0}}=0$ and lying in the midplane (i.e., a monoenergetic point source), the first-order equation of its trajectory is

$$
\begin{equation*}
x(\tau)=s_{x}(\tau) x_{0}^{;} \tag{30}
\end{equation*}
$$

We may then rewrite equation (29) as follows:

$$
\begin{equation*}
\left.R_{1}=\left.\frac{1}{2 x_{0} x_{0}^{\prime}}\right|_{0} ^{t} x(\tau) h(\tau) d \tau \right\rvert\,=\frac{\left(\ell-\ell_{0}\right)}{2 x_{0} x_{0}^{\prime}} \tag{3I}
\end{equation*}
$$

where $\left(\mathbb{L}-l_{0}\right)$ is the path length difference between the trajectory described by equation (30) and the central trajectory. Or we may also write equation (31) in the form

$$
\begin{equation*}
R_{1}=\frac{1}{2 x_{0} x_{0}^{1}}\left|\int_{0}^{t} \frac{B x(\tau) d \tau}{B p}\right|=\left(\frac{1}{2 x_{0} x_{0}^{\tau}}\right)\left(\frac{1}{B \rho}\right)\left|\int_{0}^{t} B d A\right| \tag{32}
\end{equation*}
$$

where $\int B d A$ is the magnetic flux enclosed between the central trajectory and the trajectory described by equation (30), and $B \rho$ is the magnetic rigidity of the central trajectory. Please note, however, that if the trajectory of equation (30) crosses the central trajectory or the sign of $B$ changes, this changes the sign of the integration. From equation (32) we may define resolving power as the magnetic flux enclosed per unit phase space area ( $2 x_{0} x_{0}^{\prime}$ ), per unit momentum ( $B p$ ) of the central ray.

In any given design, one or more of the above equations may be used as a guide toward achieving the required resolving power. One of the design decisions that must be made is the appropriate choice of the dipole magnet parameters (width and length) to achieve the required $\int B d A$. From first-order considerations, this choice is dominated primarily by practical and economic factors. However, a study of the nature of the origin of aberrations (see for example Ref. 1) suggests that it is advisable to keep the amplitude of $s_{x}$ small. In order to simultaneously satisfy this requirement and meet the required resolving power $R_{1}$, we see from equation (29) that the total angle of bend $\alpha$ of the central trajectory should be chosen as large as is practical. Also, in general, the focal plane angle tends to be more normal to the optic axis for larger $\alpha$ - a property usually desired in most designs.
2) Dispersion

From Reference 1, 2, or 3; for point-to-poini imaging ( $s_{x}(t)=0$ ) the dispersion at the image plane is

$$
\begin{equation*}
d_{x}(t)=-c_{x}(t) \int_{0}^{t} s_{x}(\tau) h(\tau) d \tau \tag{33}
\end{equation*}
$$

where $c_{X}(t)$ is the magnification at the image plane.
The dispersion and hence the magnification in the design of a spectrometer is dominated almost entirely by a compromise between the spatial resolution of the particle detectors used at the image plane and the momentum range to be covered by the instrument; or in the case of a momentum defining (analyzing) system, by the acceptable momentum-defining slit spacings.
3) The Selection of the Optical Mode

By optical mode, we mean the type of imaging (e.g., point-to-point or parallel-to-point, etc.) required at the image plane in both the $x$ and $y$ coordinates, and the number of intermediate images imposed between the source and image planes. The imaging requirements at the image plane are usually dominated by the physics to be performed by the instrument and the nature of the particle detectors used. However often (especially at low energies) the imaging in the $y$ plane may be unimportant as far as the physics requirements are concerned which in turn provides some additional flexibility in the optics design.

A study of the coupling coefficients to the aberration coefficients (equation 19) shows the not surprising result, that multipoles located at intermediate images in a system do not couple to aberrations in the plane in which the intermediate image occurs. Hence it often proves beneficial to intentionally create an intermediate image in the y plane of an optical system so as to achieve some degree of "orthogonality" in the minimizing of $x$ and $y$ aberrations.

The considerations of 1), 2), and 3) above are the determining factors in the selection of the first-order solution of a system design.

The optical mode and dispersion of the system are determined to a great extent by the choice of the quadrupole components choscn to achieve the firstorder imaging although it is clear that the dipole elements also influence the first-order imaging to a greater or lesser extent depending upon the total angle of bend of the system.
4) Aberrations and their Correction

A study of the source of second- and higher-order aberrations (see for example Ref. l) suggests that it is advisable to maintain the characteristic
first-order functions $c_{x}, s_{x}, d_{x}$ and $c_{y}, s_{y}$ and their derivatives as small as is feasible through the magnetic elements of a system when choosing the firstorder design. This procedure will tend to reduce the initial size of the aberrations and hence simplify the problem of minimizing them by the addition of. multipole components to the system design.

The procedure for minimizing aberrations has already been outlined in the Introduction and as such will not be repeated here. The "key" to the minimization procedure is the coupling coefficient. given by the integral expression in equation (19). The "best" location for the correcting multipole is where the coupling coefficient has its maximum value.

The preferred method of introducing the multipole components, i.e., via pure multipoles, contoured entrance or exit boundaries, or non-uniform fields is a combination of practical and economical considerations and, of course, personal taste and experience. All three methods have been used with pure multipoles dominating the situation for higher energy physics and the other two methods dominating medium and low-energy physics applications. All three techniques should be considered in any given design situation to be certain that an important economic or practical advantage has not been ignored.

## REFERENCES

(1) "A First- and Second-Order Matrix Theory for the Design of Beam Transport Systems and Charged Particle Spectrometers" SLAC Report No. 75 by K. L. Brown or Advances in Particle Physics Interscience 1, 71-134 (1967).
(2) "A General First- and Second-Order Theory of Beam Transport Optics", Proc. of the International Symposium on Magnet Technology (1965); p. 141.
(3) "Some First- and Second-Order Magnetic Optics Theorems Applicable to the Design of Beam Transport Systems and Charged Particlc Spectrometcre", Proc. of the International Symposium on Magnet Technology (1967); p. 40.
(4) E.D. Courant and H. S. Snyder, "Theory of the Alternating Gradient Synchrotron", Annals of Physics 3, pp. 1-48 (1958).
(5) K. L. Brown (with R. Belbeoch and P. Bounin), "First- and Second-Order Magnetic Optics Matrix Equations for the Midplane of Uniform-Field Wedge Magnets, Rev. Sci. Instr. 35, 481 (1964).
(6) K. L. Brown, B. K. Kear, S. K. Howry, "TRANSPORT/360 A Computer Program for Designing Beam Transport Systems", SLAC Report No. 9i, Stanford Linear Accelerator Center, Stanford University, Stanford, California (1970).

