

DEUTERON MODEL CALCULATION
 OF PHOTONUCLEON YIELDS

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Levinger's¹ deuteron model for the high energy nuclear photo-effect has provided a successful explanation of experiments where photon energies between 90 Mev and 340 Mev are employed². The model states that the nuclear cross-section is a multiple of the deuteron photo cross-section. We shall assume that this model is valid at the high photon energies expected in the proposed Stanford accelerator and shall calculate the photonucleon yields produced by the soft-shower photons.

The deuteron cross-section will be taken to be isotropic at high photon energies³, although this is not strictly true. Experimental values of the total cross-section are shown in Fig. 1. These results were all obtained at Cal Tech.^{3,4} In Wilson's⁵ shielding calculations for the Cambridge Electron Accelerator, the total deuteron cross-section is represented by

$$\sigma_0 (E_\gamma) = \begin{cases} 70 \text{ microbarns;} & 50 \text{ Mev} < E_\gamma < 300 \text{ Mev} \\ 6/E_\gamma^2 \text{ microbarns;} & E_\gamma > 300 \text{ Mev} \\ (E_\gamma \text{ in Bev}) & \end{cases} \quad (1)$$

¹J. S. Levinger, Phys. Rev. 84, 43 (1951)

²Oadian et al, Phys. Rev. 102, 837 (1956)

³See for example: J. C. Keck and A. V. Tollestrup, Phys. Rev. 101, 360 (1956)

⁴R. Gomez (private communication).

⁵R. Wilson, "A Revision of Shielding Calculations," CEA-73, May 12, 1959.

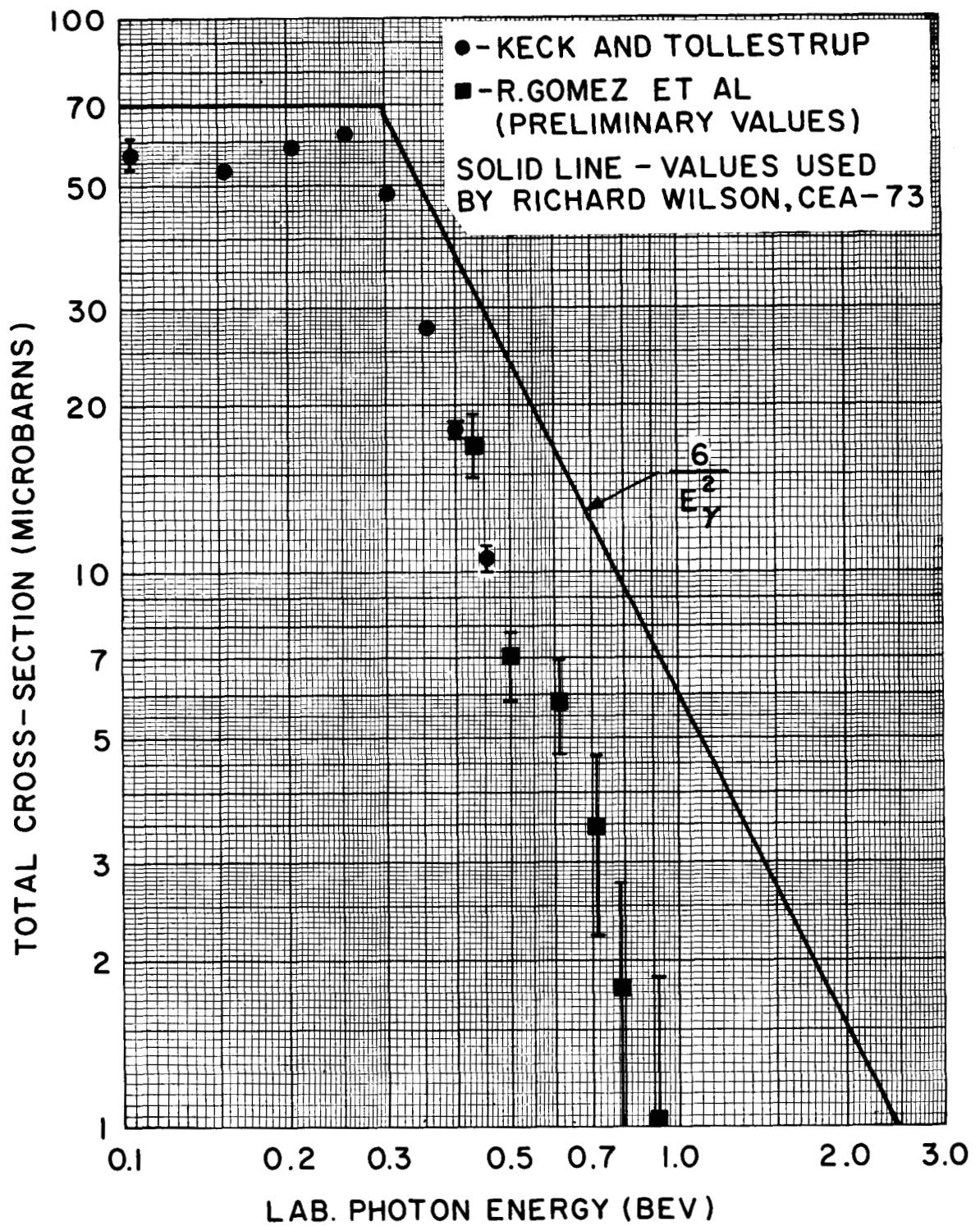


FIG. 1 Deuteron photoeffect cross section vs laboratory photon energy.

This cross-section is also shown in Fig. 1 and is clearly a conservative representation over the range of photon energies shown. We use Eq. (1) in the calculations to follow.

A further assumption is that a photon of energy E_γ yields a unique value of laboratory photonucleon energy $E_n(E_\gamma, \theta)$ at a given laboratory angle θ . This means that the momentum distribution of nucleons within the nucleus is ignored. Hence E_n is calculated using relativistic kinematics appropriate to the situation where a photon of energy E_γ is incident on a "quasi-deuteron", and we finally have two photonucleons in the laboratory. In Fig. 2 the photonucleon energies E_n are plotted vs laboratory angle θ for several values of E_γ . Figure 3 shows the angle transformation between the center-of-mass and laboratory systems.

The distribution of photonucleons in the laboratory is given by⁶

$$W(E_n, \theta) = W'(E_\gamma, \theta') \left[\left(\frac{\partial E_n}{\partial E_\gamma} \right)_\theta \left(\frac{\partial (\cos \theta)}{\partial (\cos \theta')} \right)_{E_\gamma} \right]^{-1} \quad (2)$$

where $W(E_n, \theta)$ is the photonucleon spectrum in the laboratory (photonucleons per second per unit laboratory energy E_n per unit laboratory solid angle), and $W'(E_\gamma, \theta')$ is the rate of production of these photonucleons which occur in the corresponding unit solid angle in the center of mass and are due to photons with laboratory energies between E_γ and $E_\gamma + dE_\gamma$. The derivatives in Eq. (2) are calculated from the Lorentz transformations giving $E_n(E_\gamma, \theta)$ and $\theta(E_\gamma, \theta')$.

All that remains, then, is the problem of calculating $W'(E_\gamma, \theta')$, the distribution in the center of mass. As mentioned before, the photons arise from a soft shower. The spectrum is given by the familiar

⁶You'll have to figure out how to get this expression all by yourself (hint: see M-226).

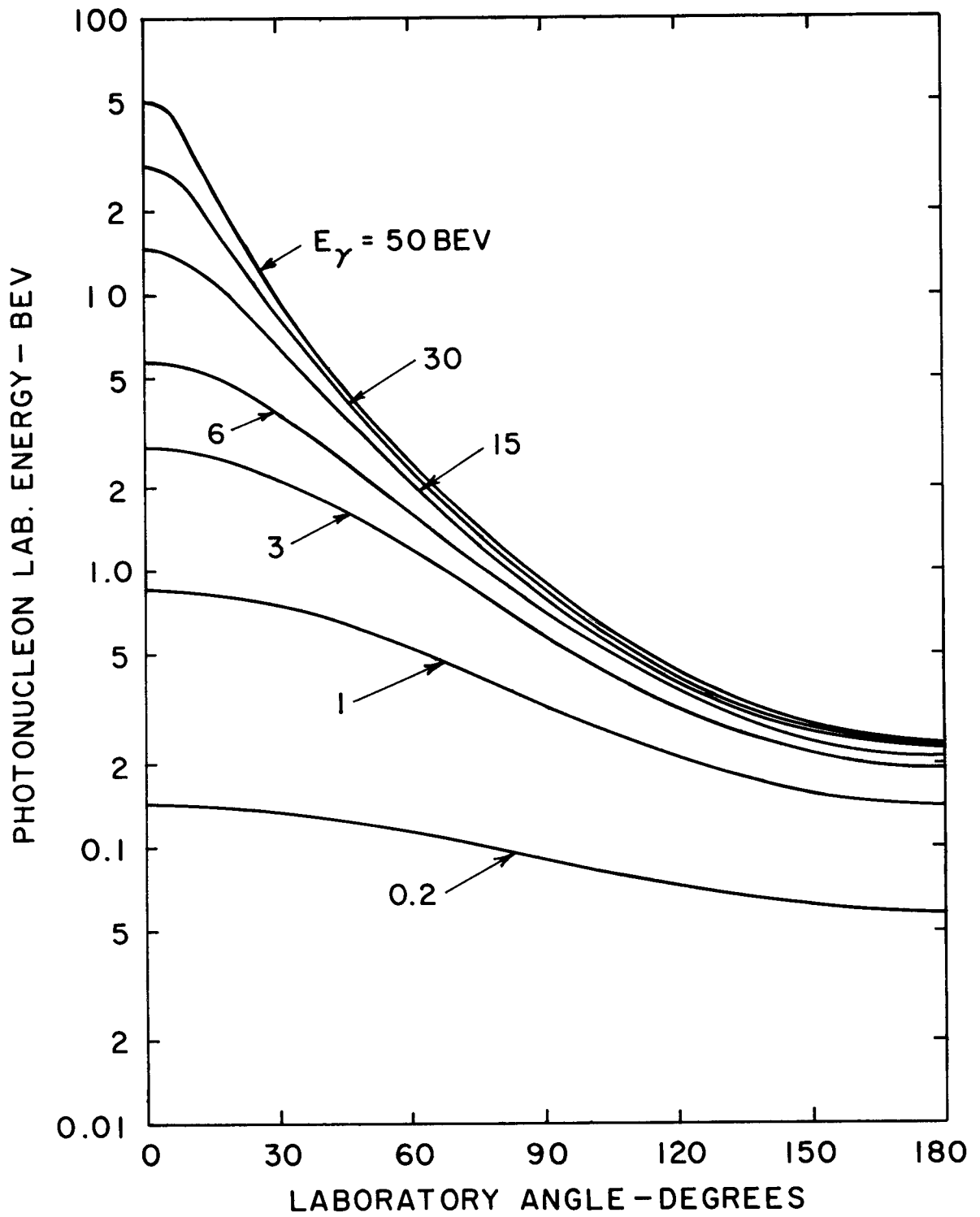


FIG. 2 Deuteron model photoeffect kinematics.

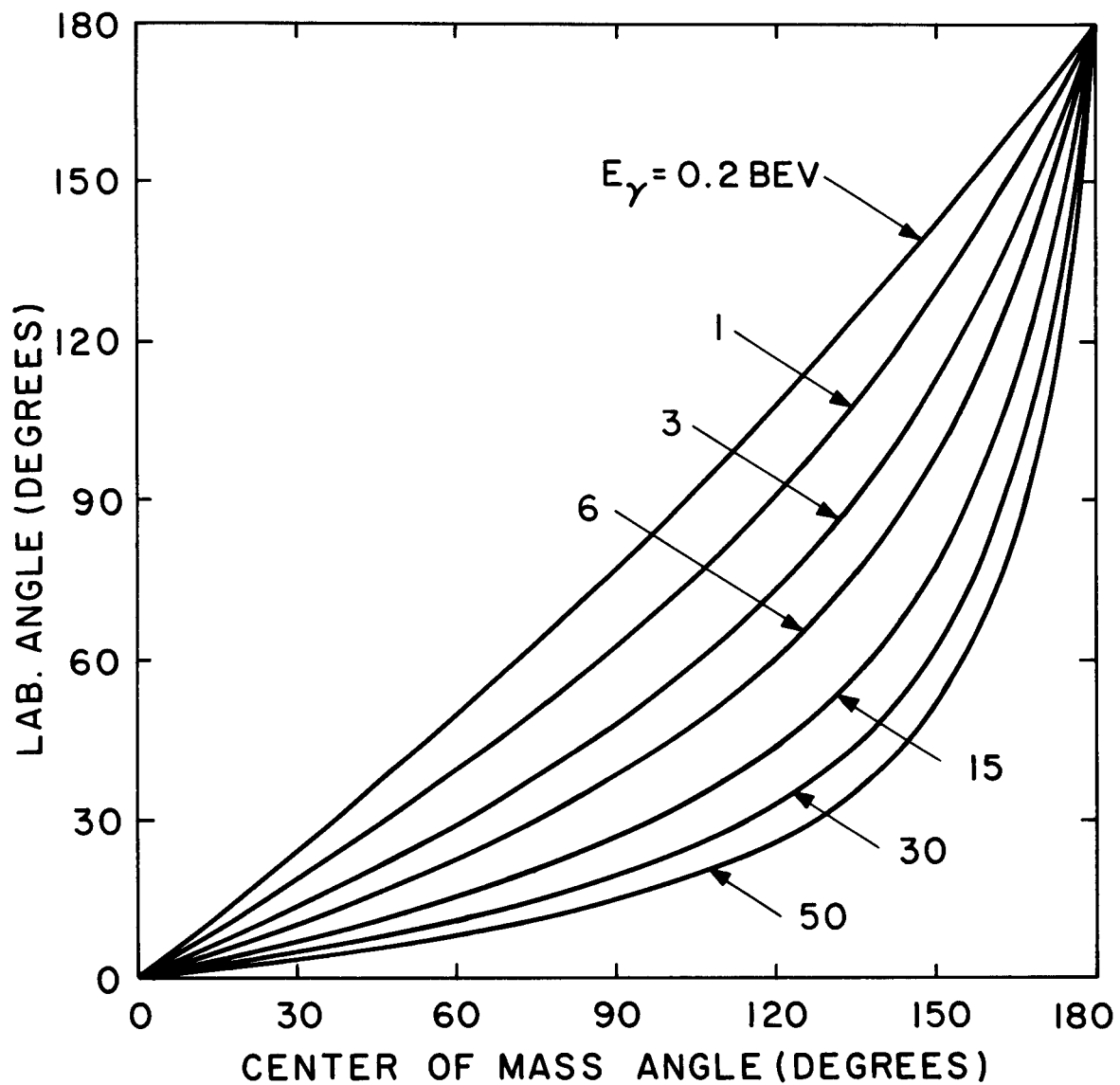


FIG. 3 Angle transformation for quasi-deuteron photoeffect kinematics.

track length expression⁷

$$(0.572) \frac{E_0}{E_\gamma^2} dE_\gamma \quad (3)$$

where E_0 is the kinetic energy of the incident electron. This expression gives the track length in radiation lengths of the photons with energies between E_γ and $E_\gamma + dE_\gamma$. Multiplication of Eq. (3) by (X_0/ρ) gives the track length in centimeters, where X_0 is the value of the radiation length in grams/cm², and ρ is the target density in grams/cm³. The yield $W'(E_\gamma, \theta')$ is then given by

$$W'(E_\gamma, \theta') = N_1 (0.572) \frac{X_0}{\rho E_\gamma^2} \sigma_a(E_\gamma, \theta') E_0 \quad (4)$$

where N_1 is the number of nuclei per cm³ in the target, and $\sigma_a(E_\gamma, \theta')$ is the photoeffect differential cross-section of these nuclei. N_1 is given by $(N_0 \rho/A)$, where N_0 is Avogadro's number, and A is the atomic weight.

According to the deuteron model,

$$\sigma_a(E_\gamma, \theta') = \alpha \frac{NZ}{A} \sigma_d(E_\gamma, \theta') \quad (5)$$

where N is the number of neutrons in the nucleus, Z is the atomic number, A is the mass number, and α is a constant which is different for all nuclei⁸. We shall take α to be 1.5 for all nuclei.

Equations (2), (4) and (5) are now combined to give

⁷See for example: B. Rossi, High Energy Particles, Prentice-Hall, New York, 1952, p. 244.

⁸See references 2 and 5 for further discussion.

$$W(E_n, \theta) = (0.572)(1.5) \frac{N \times E_0}{AE_\gamma^2} \left(\frac{NZ}{A} \right) \frac{\sigma_0(E_\gamma)}{4\pi} \left[\left(\frac{\partial E_n}{\partial E_\gamma} \right)_\theta \left(\frac{\partial (\cos \theta)}{\partial (\cos \theta')} \right)_{E_\gamma} \right]^{-1} \quad (6)$$

where Eq. (1) for $\sigma_0(E_\gamma)$ will be used in the calculations based on Eq. (6). This expression gives the number of neutrons in unit solid angle in the laboratory per unit nucleon kinetic energy E_n when one electron is incident on the target.

We are mainly interested in the yields from a copper target, since many electrons are expected to strike the accelerator walls on their long journey.

Figures 4, 5 and 6 show the results obtained for $E_0 = 15, 30$ and 45 Bev in the case of a copper target.

Example:

What is the flux of photoneutrons at laboratory angle $\theta = 40^\circ$ with energies between $E_n = 2$ Bev and 2.10 Bev when a 30 Bev, one micro-ampere electron beam strikes a copper target?

Use Fig. 5, find 10^{-9} neutrons/Mev-steradian for one incident electron. The flux is then

$$\frac{100}{1.6 \times 10^{-19}} \times 10^{-9} = 6.25 \times 10^{11} \text{ neutrons/steradian}$$

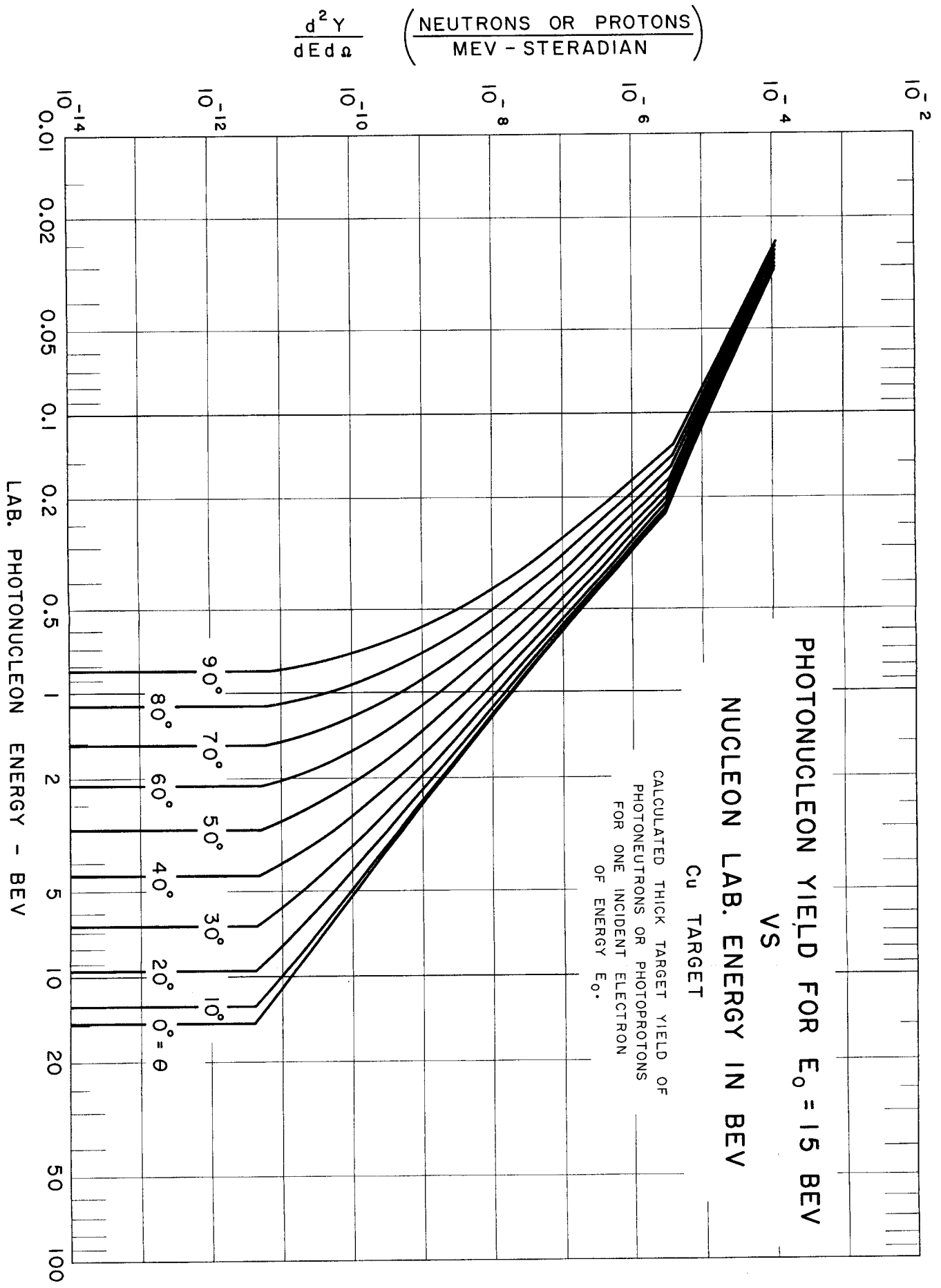


FIG. 4

LAB. PHOTONUCLEON ENERGY - BEV

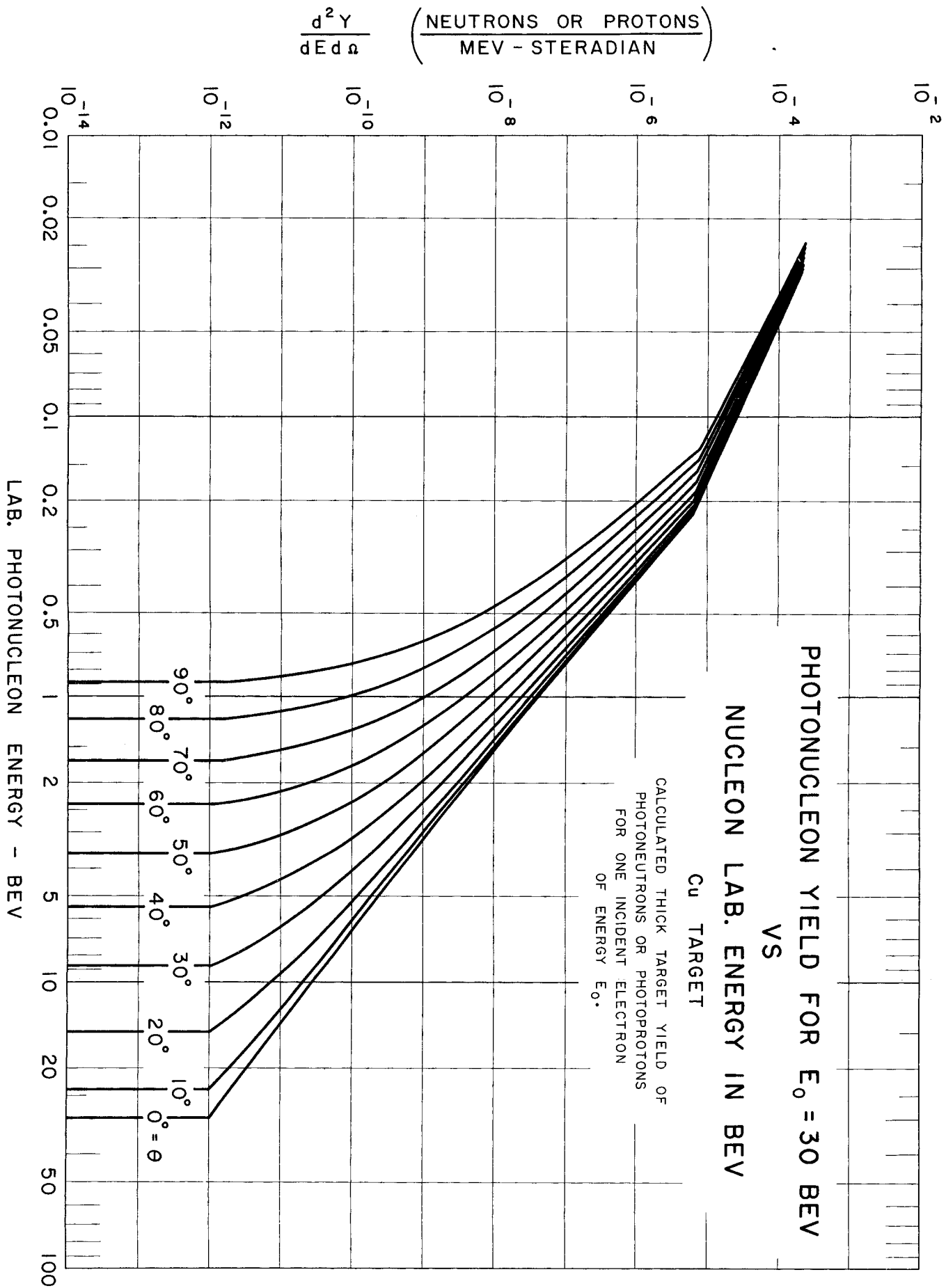


FIG. 5

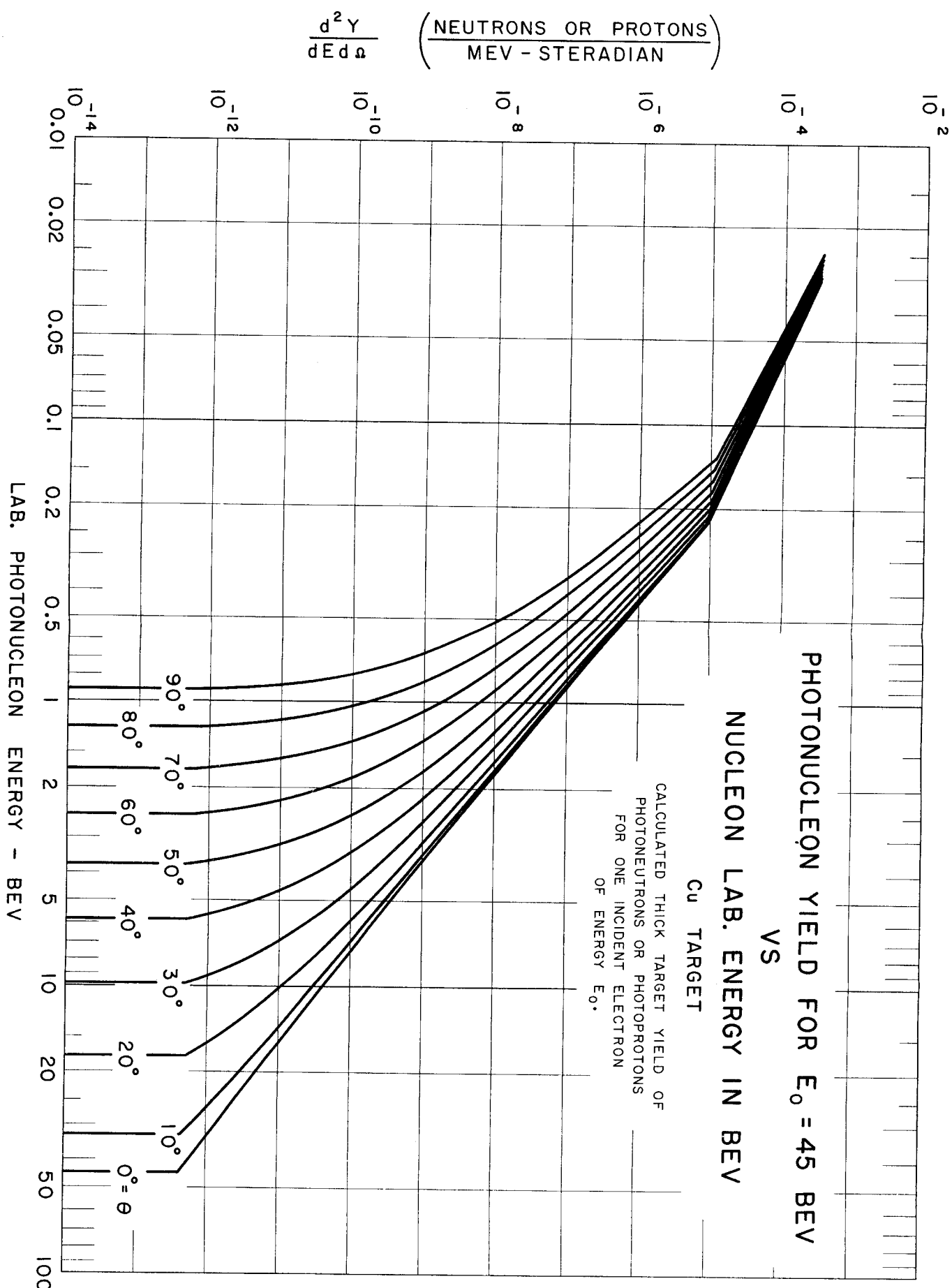


FIG. 6