

Electroweak Symmetry Breaking

in the

Early
Universe

~||~

COSMOLOGY

The

Next

Three

Minutes

COSMOLOGY

THE UNIVERSE

is

EXPANDING

!



$$d(t) = d_0 R(t)$$

\therefore All distances on sufficiently large scales

Cosmological Principle

\Rightarrow Uniform Scaling

$$\frac{\dot{R}}{R} \equiv H$$

$$H_0 \sim 70 \frac{\text{km}}{\text{s}} \frac{1}{\text{Mpc}}$$

GRAVITY

$$H^2 = \frac{8\pi G_N}{3} \rho \leftarrow \begin{array}{l} \text{Energy} \\ \text{Density} \end{array}$$

ρ {	radiation	T^4
	dust	T^3
	curvature	$1/R^2$

THERMAL

ISENTROPIC

$$S \sim (TR)^3 = \text{constant}$$

$$T \propto 1/R$$

\therefore Not always true:

- $e^+ + e^- \leftrightarrow \gamma$

- phase transition

EXPANSION

radiation

$$H^2 \equiv \left(\frac{\dot{R}}{R}\right)^2 = \frac{8\pi}{3} G_N \sigma T^4 \rightarrow \frac{1}{R^4}$$

$$R(t) = R_0 \left[\frac{t}{t_0}\right]^{1/2}$$

$$H(t) = \frac{1}{2t} \sim \frac{T^2}{M_p}$$

$\therefore H$ measures age of the Universe

ELECTROWEAK

$$M \sim 100 \text{ GeV}$$

$$t_{ew} \sim \frac{M_p}{M^2} \sim 10^{-11} \text{ s}$$

Causal Domain

cm

$$\tilde{\text{today}}: T_0 \sim 3 \text{ K} \sim 10^{-4} \text{ eV}$$

$$D_{\text{today}} \sim 10^{15} \text{ cm}$$

8 REMNANTS

Domain Size Small

∴ Not Observable

BUT :

* BARYON VIOLATION

$$B=0$$

↓ t evolves

$$B \neq 0$$

BARYOGENESIS

Uniform Effect on Average
Large Scale

BARYON VIOLATION
CP VIOLATION
Non EQUILIBRIUM

Sakharov 1968!

★ C, CP VIOLATION

Quarks



anti-Quarks



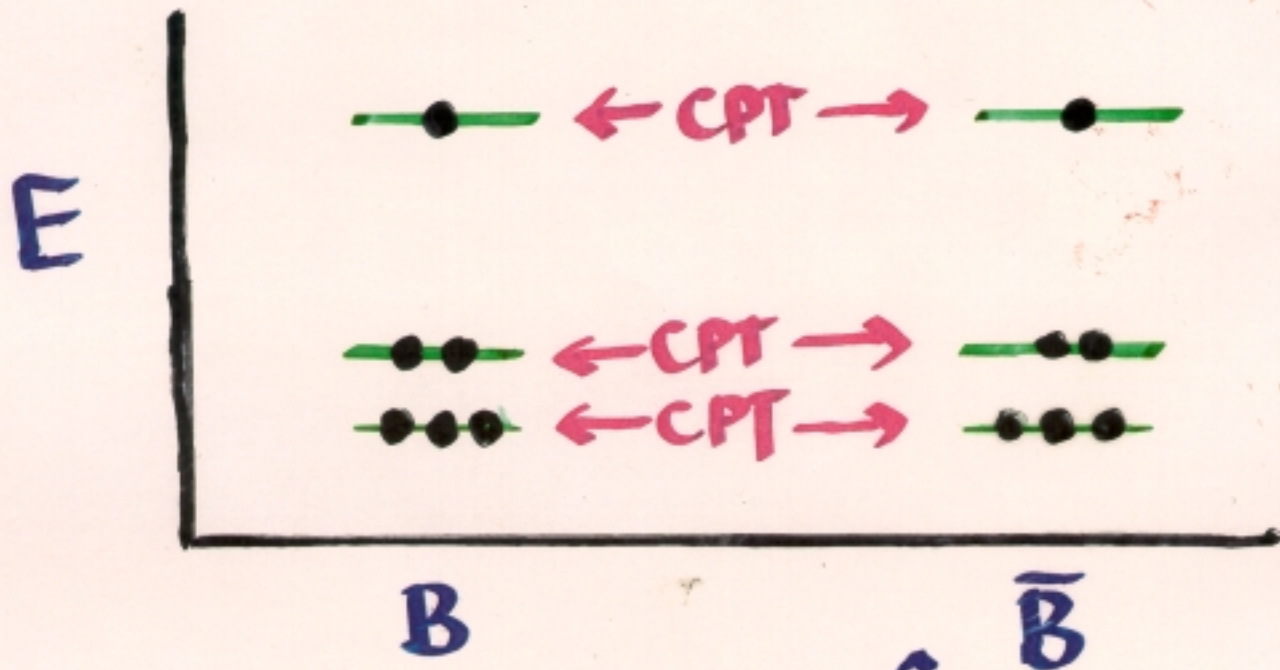
\hat{B} odd under C, CP

$$\frac{\text{Rate}[K_L \rightarrow e^+ + \pi^- + \nu]}{\text{Rate}[K_L \rightarrow e^- + \pi^+ + \bar{\nu}]} \approx 1.006$$

$\sim 10^{-3}$ effect

3

* DEPARTURE FROM THERMAL EQUILIBRIUM



$$\langle B \rangle = \text{Tr} \hat{B} e^{-\beta \hat{H}} = 0$$

Odd under CPT

$$\hat{B} \rightarrow -\hat{B}$$

Even under CPT

$$\hat{H} \rightarrow \hat{H}$$

ELECTROWEAK BARYOGENESIS

All of Sakharov's criteria
may be met in the

Standard Model

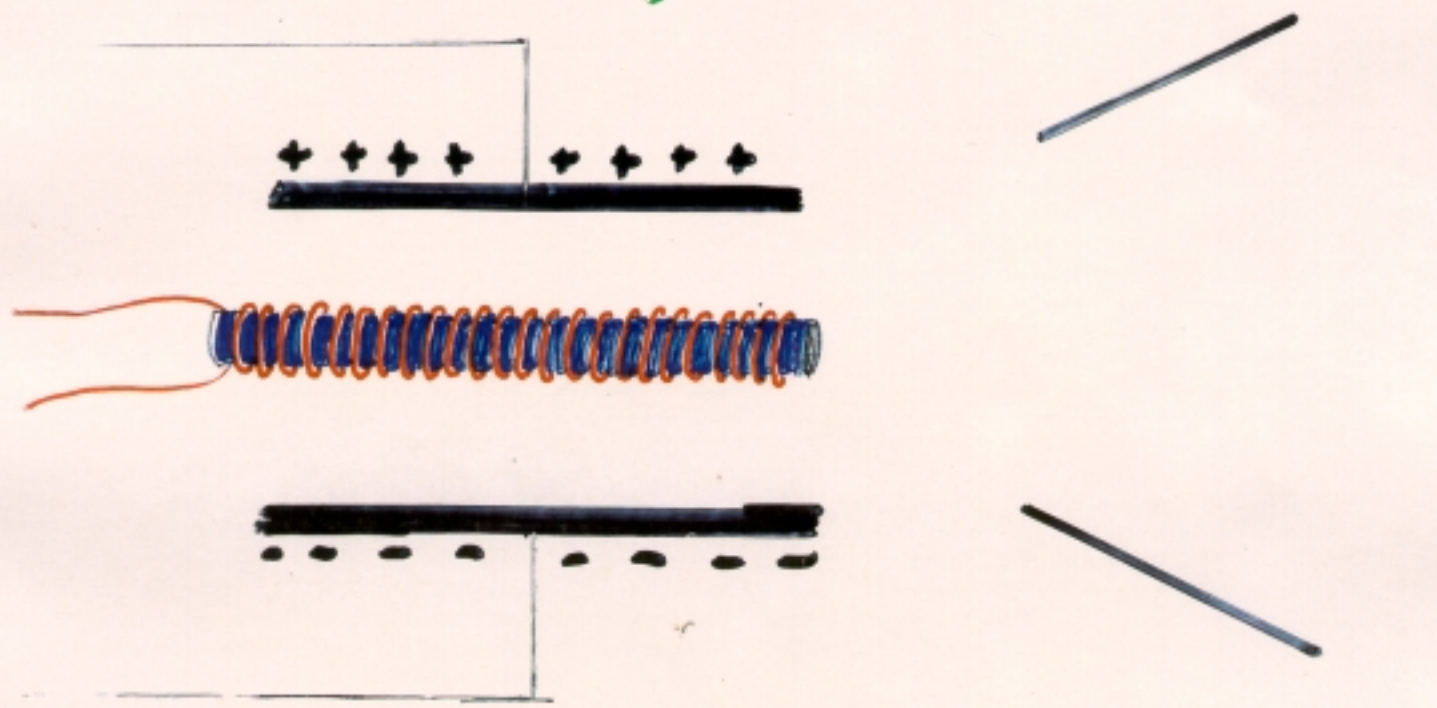
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BARYON VIOLATION

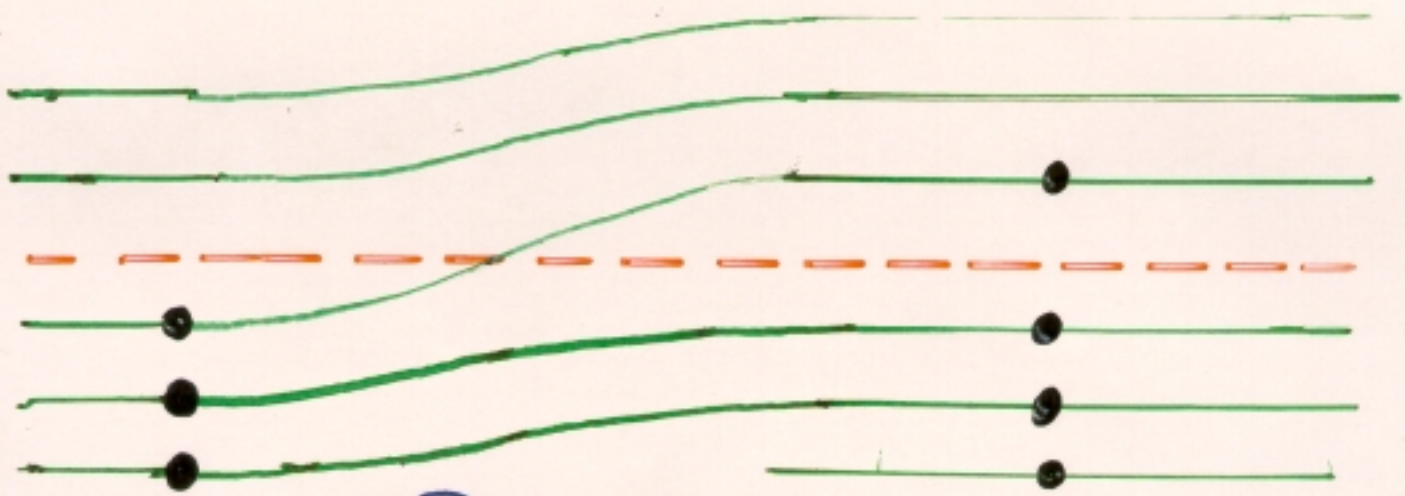
in

WEAK INTERACTION

? ?



WIKAL DEHA



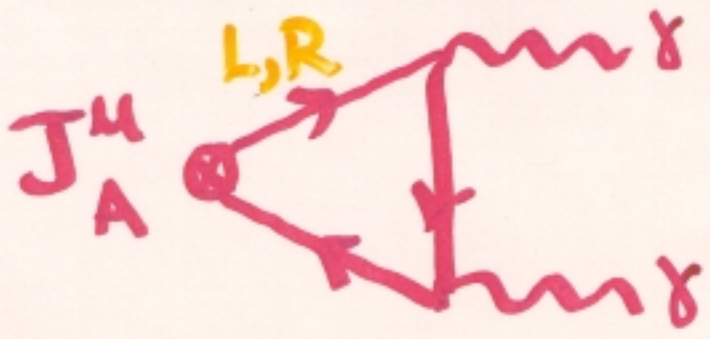
$B=0$



'KNOT' OF W & Z

$B \neq 0$

QED



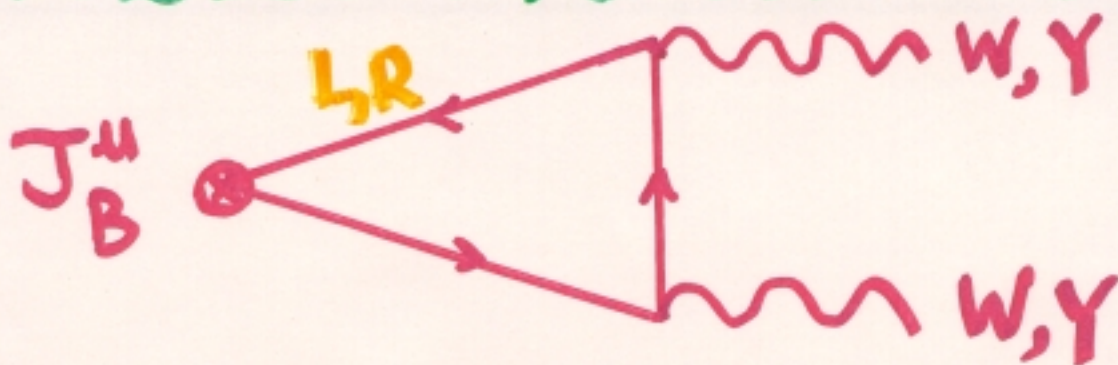
$$\partial_\mu J_A^\mu = \frac{e^2}{32\pi^2} F_{\mu\nu} \tilde{F}^{\mu\nu} \quad (\vec{E} \cdot \vec{B})$$



chiral

'Anomaly'

ELECTROWEAK



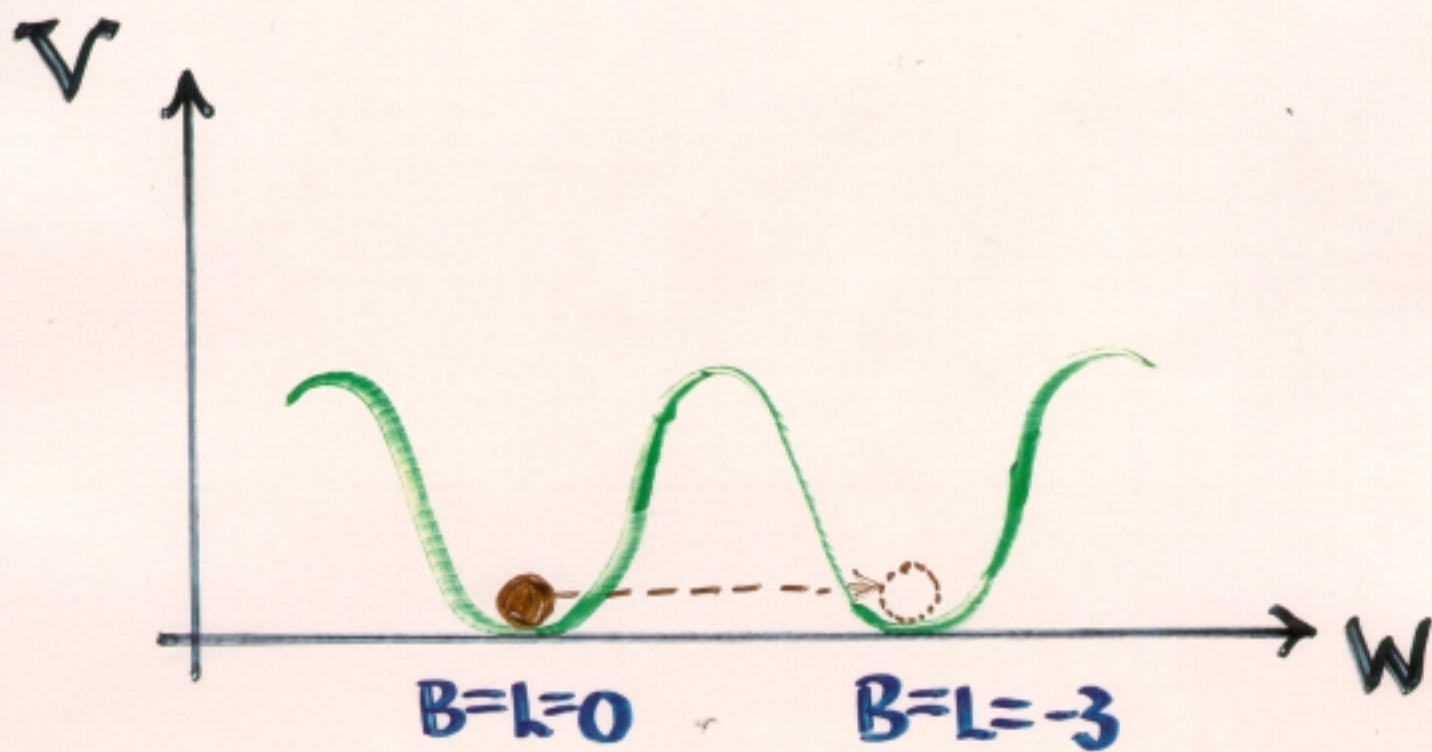
$$\partial_\mu J_B^\mu = 3 \left[\frac{g^2}{32\pi^2} W_{\mu\nu}^a \tilde{W}^{a\mu\nu} + \frac{g'^2}{32\pi^2} F_{\mu\nu} \tilde{F}^{\mu\nu} \right]$$

↑
chiral

$$= \partial_\mu J_L^\mu$$

B-L conserved!

16 WHY IS THE PROTON STABLE ?



Quantum Tunneling

$$\Gamma \propto e^{-4\pi/d_{WK}} \sim 10^{-140} !$$

Weak Fine Structure Constant

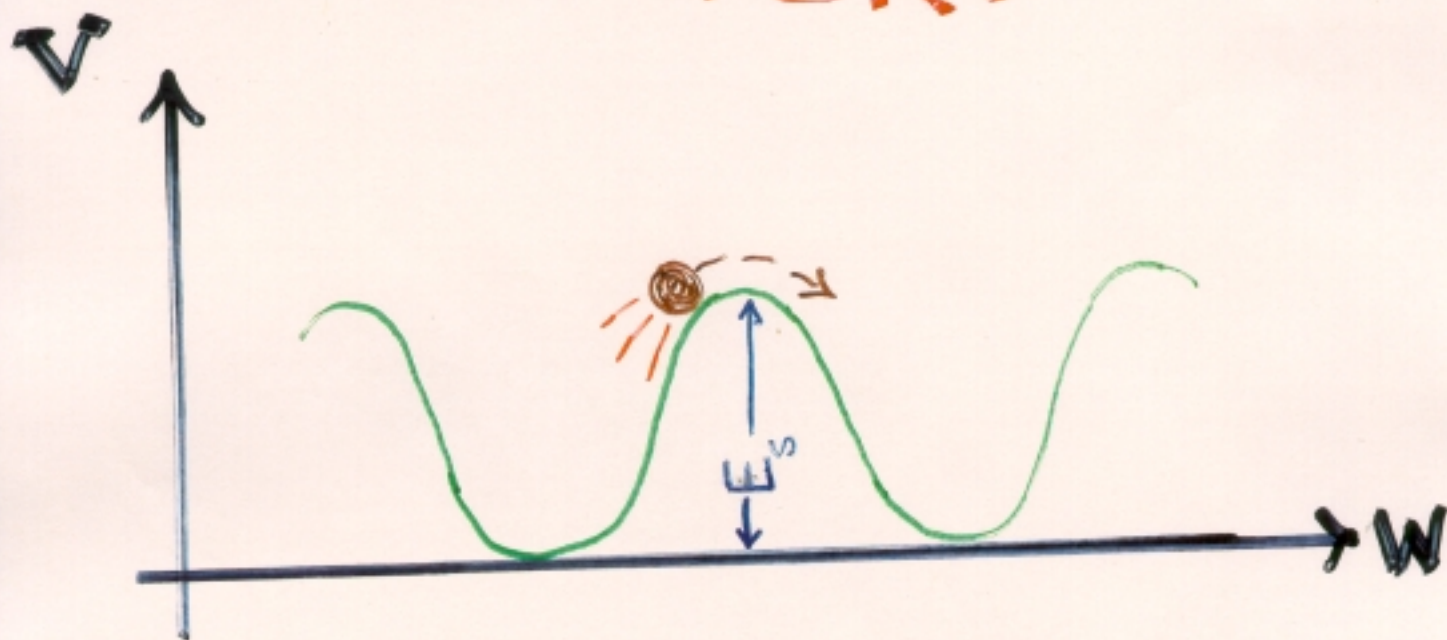
Really Small number

5

ACTUALLY...

- Proton absolutely stable
 $\Delta B = 3$

HOWEVER:



THERMAL ACTIVATION

$$\Gamma \propto e^{-\beta E_s}$$

← 'sphaleron'

$$E_s \sim \frac{M_Z}{\alpha_{WK}} \leftarrow 95$$

Barrier rather than
Mass suppresses low Energy
B-violation

$$E_S \sim \text{TeV}$$



$$T \gtrsim \text{TeV}$$

B violation unsuppressed



GUT Wipeout!

B-L ←

DEPARTURE FROM EQUILIBRIUM

Expansion Rate $H \sim -\frac{\dot{T}}{T} \sim \frac{T^2}{M_{Pl}^2}$

Very small at $T \lesssim \text{TeV}$.

($\Gamma \sim \alpha_{WK} T \sim 10^{13} \text{ H}$)

Need something more

Violent

PHASE TRANSITION

UNBROKEN PHASE

$M_W = M_Z = 0 \rightarrow$ Long Range Forces

$$v = \langle \phi \rangle = 0$$

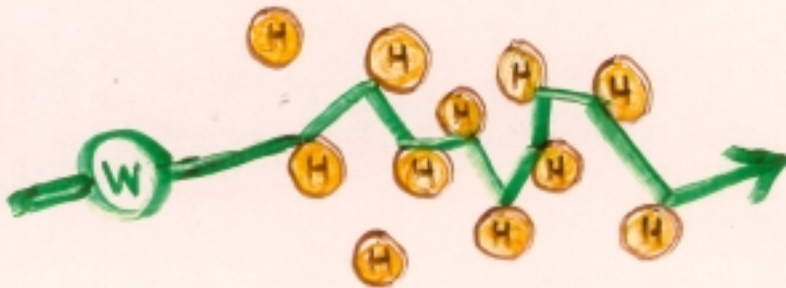


$$\text{Rate} \propto e^{-\beta v/g} \sim 1$$

$T \sim 100 \text{ GeV}$ — ~~WEAK PHASE TRANSITION~~

$M_W, M_Z \neq 0 \rightarrow$ Short Range Forces

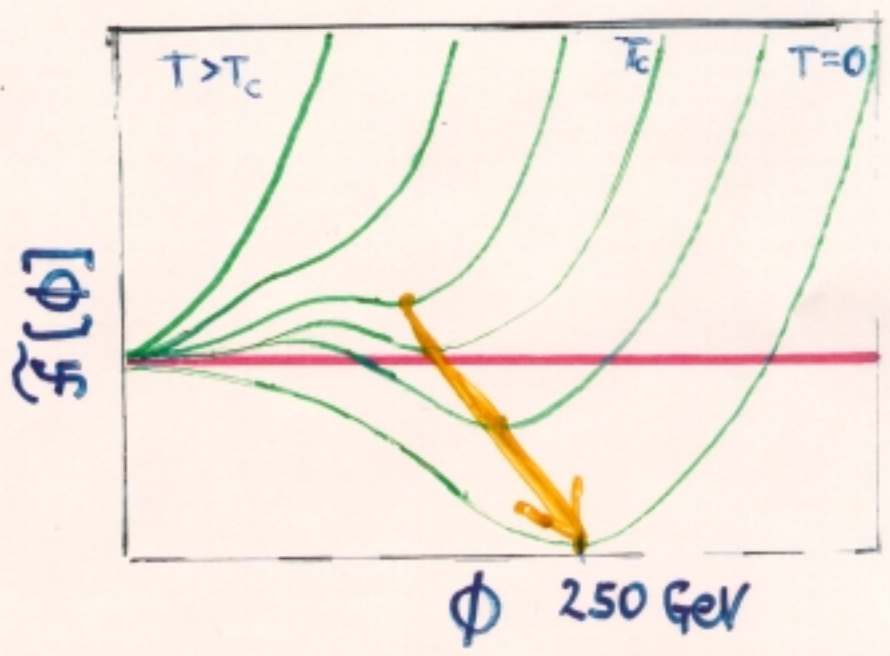
$$v = \langle \phi \rangle \sim 250 \text{ GeV}$$



$$\text{Rate} \propto e^{-\beta v/g} \ll 1$$

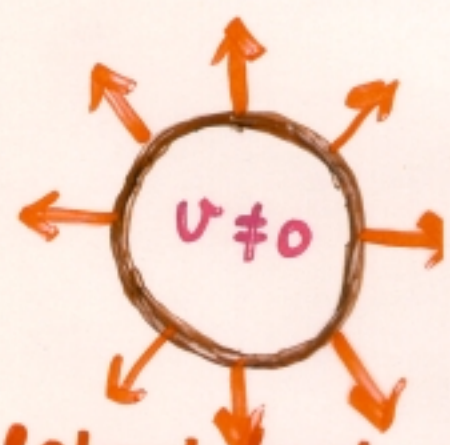
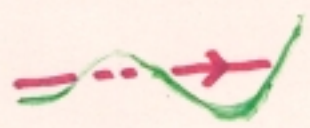
BROKEN PHASE

1st ORDER PHASE TRANSITION



• Supercool below T_c
($\phi=0$)

• Nucleate Bubbles



• Bubbles Expand, release latent heat

OUT-OF-EQUILIBRIUM B-VIOLATION

Outside Bubble

$$\Gamma_{\Delta B} \sim \alpha_{wk}^5 T \rightarrow \text{rapid}$$

Inside Bubble

$$\Gamma_{\Delta B} \propto e^{-\frac{4\pi v}{g} \frac{1}{T}} \ll \frac{T^2}{M_{pl}}$$

$$v \neq 250 \text{ GeV.}$$

Transition must be strongly

1st Order: $\Delta\phi$ large

PHASE TRANSITION

?

What breaks the Weak Symmetry

??

SM → probably not 1st order

MSSM → small region of parameter space

NMSSM → Easy!

NEW MODELS

'Deconstructed'
Models

Phase Transitions

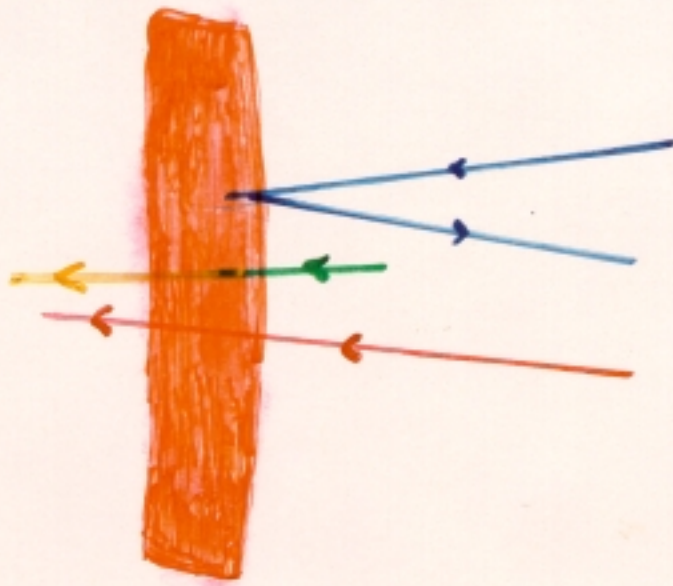
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Strongly 1st Order

CP Violation

!

CP VIOLATION



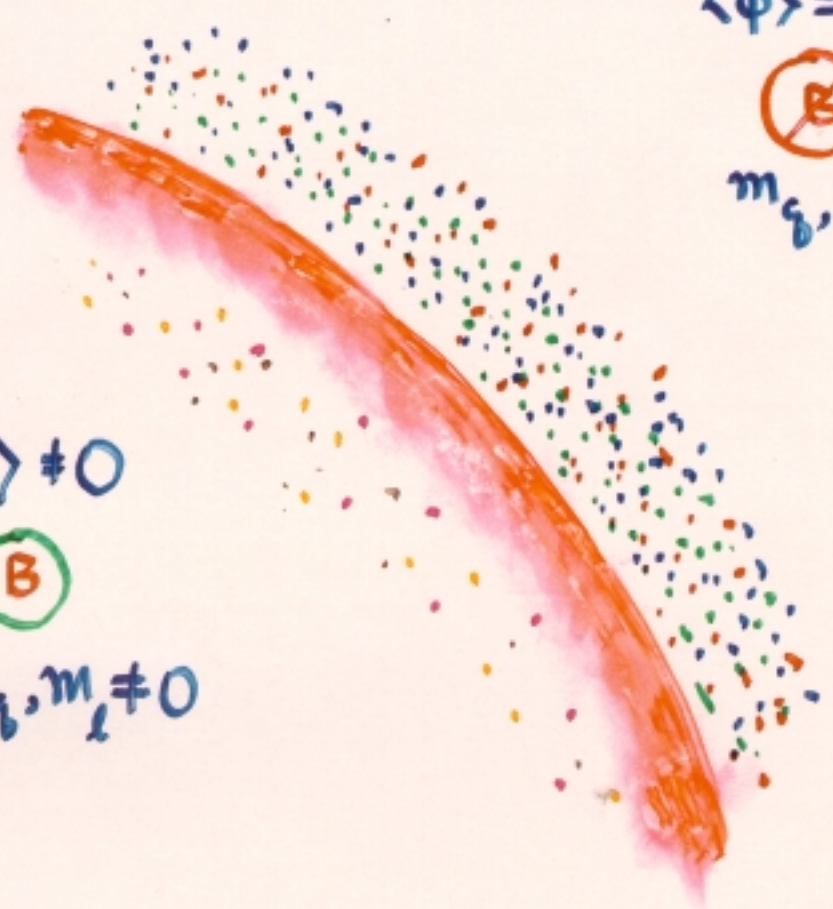
- 1) Interaction of quarks, leptons, ...
with **CP** violating phases
- 2) Plasma transport properties
move charges around
- 3) Thick Wall - $\textcircled{\cancel{B}}$
Thin Wall - $\textcircled{\cancel{CP}}$ $\textcircled{\cancel{B}}$
- 4) Expanding Wall

HOW DOES THIS MAKE AN ASYMMETRY?

$$\langle \phi \rangle = 0$$

~~B~~

$$m_\nu, m_\ell = 0$$



$$\langle \phi \rangle \neq 0$$

B

$$m_\nu, m_\ell \neq 0$$

Non Equilibrium - Spatial Separation
of B violation

CP VIOLATION

CKM

$$\delta_{CP} \sim \text{Im} [M_u^3 M_d^2 M_u M_d]$$

$$M_d \equiv \frac{g_W^2}{2M_W^2} K M_d^2 K^\dagger$$

$$M_u \equiv \frac{g_W^2}{2M_W^2} M_u^2$$

*
$$\delta_{CP} \sim \left[\frac{g_W^2}{2M_W^2} \right]^7 S_1^2 S_2^2 S_3 S_\delta m_t^6 m_b^4 m_c^2 m_s^2$$

$\sim 10^{-22}$

NEW CP VIOLATION!

CKM CP VIOLATION

$$\delta_{CP} \sim \text{Im}[VV, V^*V^*]$$

$$V \sim Y_u Y_d^\dagger$$

8 Yukawa Couplings

4 Weak Interactions

$$d^2 \lambda_t^4 \lambda_b^2 \lambda_s \lambda_d s_1^2 s_2 s_3 s_\delta$$

$$\sim 10^{-16}$$

CKM in K System?

$$E \ll v$$

\therefore Effective Field Theory is different!

$$\frac{1}{m_b^2}, \frac{1}{m_t^2}$$



\bullet m_i dependence may be different.

NEW CP VIOLATION

Multi-Higgs ✓

MSSM ✓

NMSSM ✓

Deconstruction ✓

THE GOOD NEWS

New CP violation is }
observable!

Electroweak Phase } If we
Transition } know
Computable } EW Symmetry
breaking

We're going to find
out

.

THE BAD NEWS

No generic predictions

EDM

B,D mixing

top physics

Higgs physics

THE REALLY BAD NEWS



Non-Electroweak models also work, but are untestable

Ex: CP violation in super heavy right handed neutrinos

Yukawa Couplings \rightarrow Late Decays violate lepton number

\therefore Electroweak B, L violation
Equilibrates with

$$B \neq 0$$

EW Symmetry Breaking &

The Origin of Matter

Baryogenesis may be connected with the physics of EW symmetry breaking.