

Neutrino Mass Models

- Why BSM?
- Neutrino mass models roadmap
- Survey of approaches
- TBM, A_4 , CSD
- Family symmetry and GUTs
- Sum rules and predictions



Great interest in neutrino theory, e.g. Melbourne Participants:

- Kev Abazajian (Maryland)
Carl Albright (Fermilab)
Evgeny Akhmedov (Max Planck, Heidelberg)
Matthew Baring (Rice)
Pasquale Di Bari (Padova)
Nicole Bell (Melbourne)
Mu-Chun Chen (UC Irvine)
Vincenzo Cirigliano (LANL)
Roland Crocker (Monash)
Basudeb Dasgupta (Tata Institute)
Amol Dighe (Tata Institute)
Andreu Esteban-Pretel (Valencia)
Ferruccio Feruglio (Padua/INFN)
Robert Foot (Melbourne)
George Fuller (UC San Diego)
Alex Friedland (LANL)
Julia Garayoa Roca (Valencia)
Vladimir N. Gavrin (Moscow, INR)
Damien George (Melbourne)
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Tom Griffin (Melbourne)
Gary Hill (Madison)
Martin Hirsch (Valencia)
Thomas Jacques (Melbourne)
Girish Joshi (Melbourne)
Sin Kyu Kang (Seoul National University of Technology)
Boris Kayser (Fermilab)
- Steve King (Southampton)
Archil Kobakhidze (Melbourne)
Sandy Law (Melbourne)
Manfred Lindner (Max Planck, Heidelberg)
Ernest Ma (UC Riverside)
Kristian McDonald (TRIUMF)
Bruce McKellar (Melbourne)
Hitoshi Murayama (UC Berkeley)
Sandip Pakvasa (Hawaii)
Sergio Palomares-Ruiz (Durham)
Stephen Parke (Fermilab)
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Michael Pluemacher (Max Planck, Munich)
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Ricard Tomas (Hamburg)
Timur Rashba (Max Planck, Munich)
Ray Sawyer (UC Santa Barbara)
Alexei Smirnov (ICTP, Trieste)
Gerard Stephenson (UNM)
Alexander Studenikin (Moscow State University)
Jayne Thompson (Melbourne)
Shoichi Uchinami (Tokyo Metropolitan U.)
Raoul Viollier (Cape Town)
Ray Volkas (Melbourne)
Renata Zukanovich-Funchal (São Paulo)

■ Why Beyond Standard Model?

1. There are no right-handed neutrinos ν_R
2. There are only Higgs doublets of $SU(2)_L$
3. There are only renormalizable terms

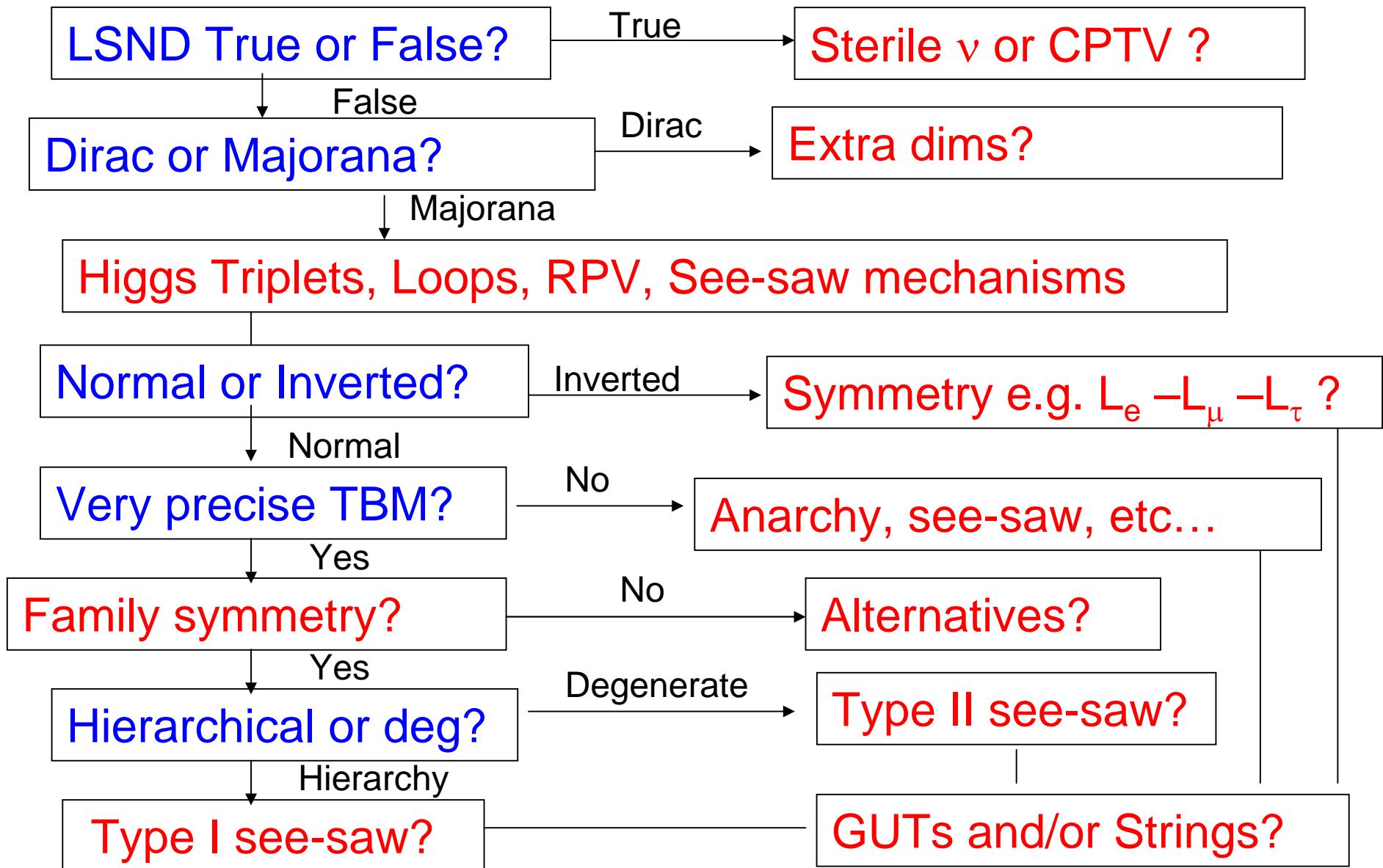
In the **Standard Model** these conditions all apply so neutrinos are **massless**, with ν_e , ν_μ , ν_τ distinguished by separate lepton numbers L_e , L_μ , L_τ

Neutrinos and anti-neutrinos are distinguished by the total conserved lepton number $L=L_e+L_\mu+L_\tau$

To generate neutrino mass we must relax 1 and/or 2 and/or 3

Staying within the SM is not an option – but what direction?

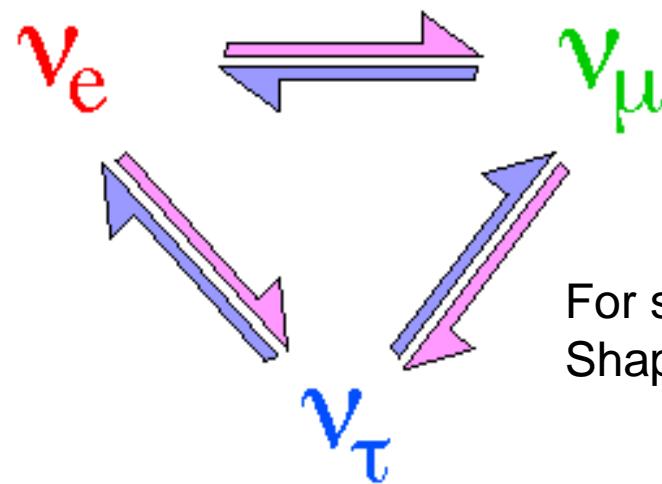
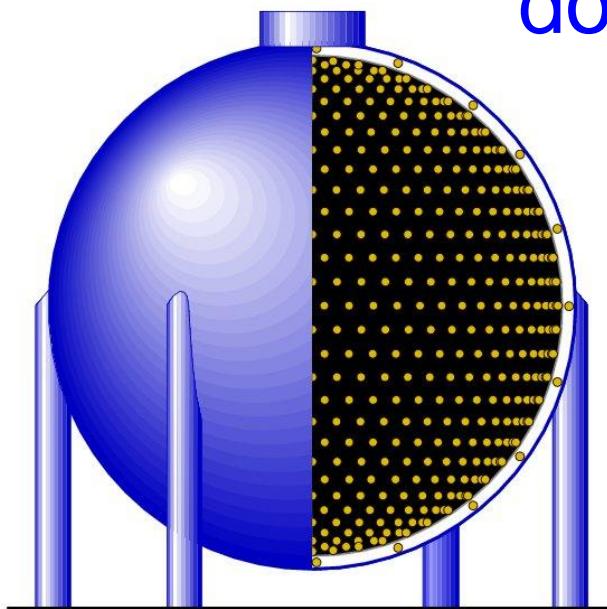
■ Neutrino mass models roadmap



LSND True or False?

MiniBoone does not support LSND result

does support three neutrinos

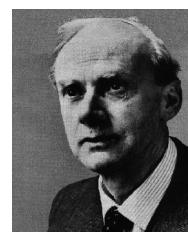


For steriles see
Shaposhnikov talk

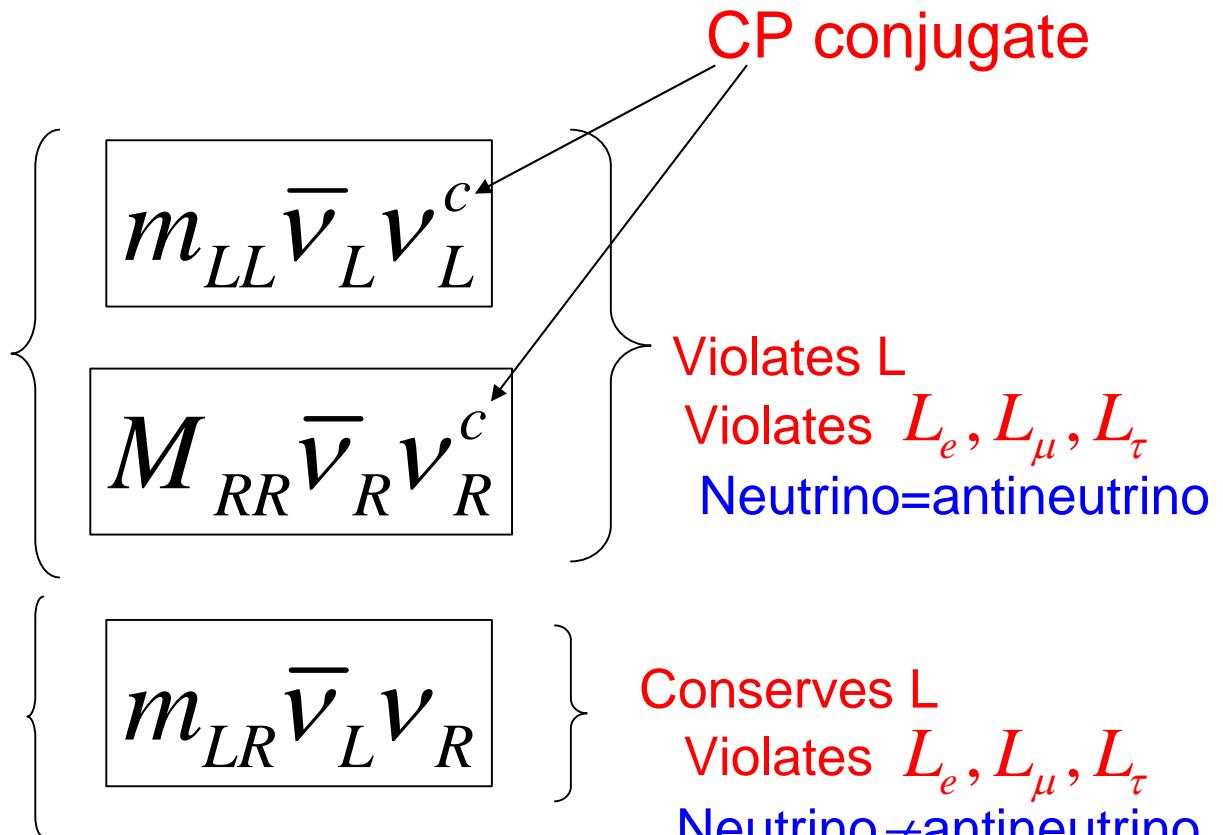
In this talk we assume that LSND is **false**

Dirac or Majorana?

Majorana masses



Dirac mass



1st Possibility: Dirac

Recall origin of electron mass in SM with $L = \begin{pmatrix} \nu_e \\ e^- \end{pmatrix}_L, \quad e_R^-, \quad H = \begin{pmatrix} H^+ \\ H^0 \end{pmatrix}$

$$\lambda_e \bar{L} H e_R^- = \lambda_e \langle H^0 \rangle \bar{e}_L^- e_R^-$$

Yukawa coupling λ_e must be small since $\langle H^0 \rangle = 175 \text{ GeV}$

$$m_e = \lambda_e \langle H^0 \rangle \approx 0.5 \text{ MeV} \Leftrightarrow \lambda_e \approx 3.10^{-6}$$

Introduce right-handed neutrino ν_{eR} with zero Majorana mass

$$\lambda_\nu \bar{L} H^c \nu_{eR} = \lambda_\nu \langle H^0 \rangle \bar{\nu}_{eL} \nu_{eR}$$

then Yukawa coupling generates a Dirac neutrino mass

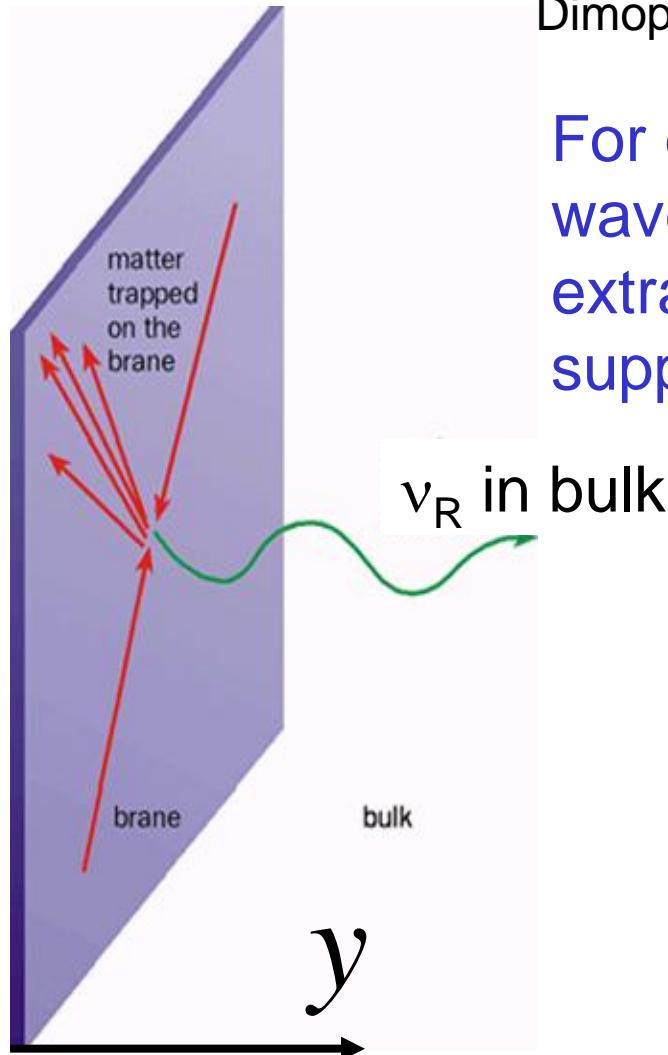
$$m_{LR}^\nu = \lambda_\nu \langle H^0 \rangle \approx 0.2 \text{ eV} \Leftrightarrow \lambda_\nu \approx 10^{-12}$$

Why so small?
– extra dimensions

■ Flat extra dimensions with RH neutrinos in the bulk

Dienes, Dudas, Gherghetta; Arkhani-Hamed, Dimopoulos, Dvali, March-Russell

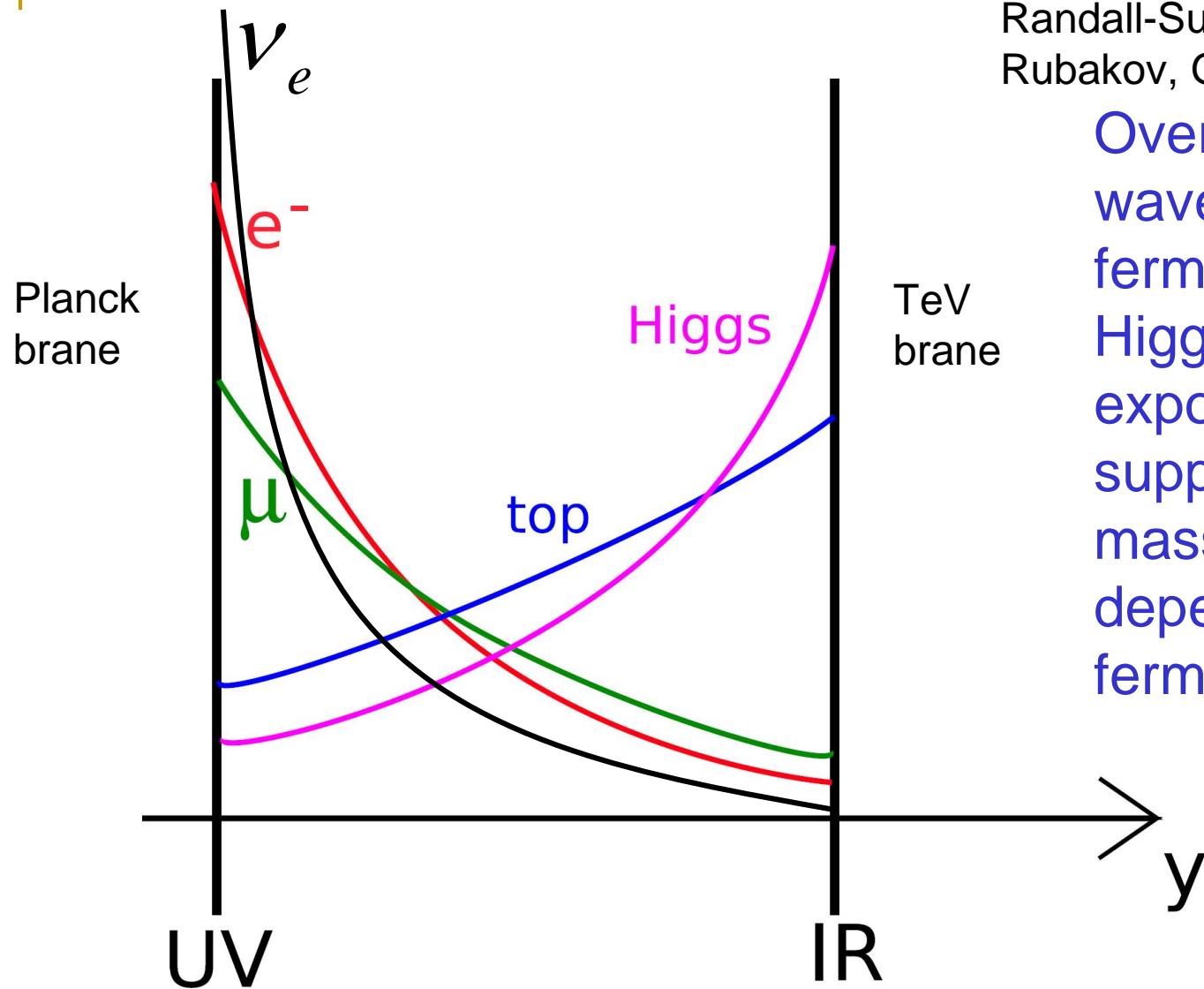
For one extra dimension y the ν_R wavefunction spreads out over the extra dimension, leading to a volume suppressed Yukawa coupling at $y=0$



$$\rightarrow m_{LR}^\nu = \frac{\lambda \langle H^0 \rangle}{\sqrt{V}} = \lambda \langle H^0 \rangle \frac{M_{string}}{M_{Planck}}$$

$$e.g. \quad \frac{M_{string}}{M_{Planck}} = \frac{10^7}{10^{19}} = 10^{-12}$$

■ Warped extra dimensions with SM in the bulk



Randall-Sundrum;
Rubakov, Gherghetta,...

Overlap
wavefunction of
fermions with
Higgs gives
exponentially
suppressed Dirac
masses,
depending on the
fermion profiles

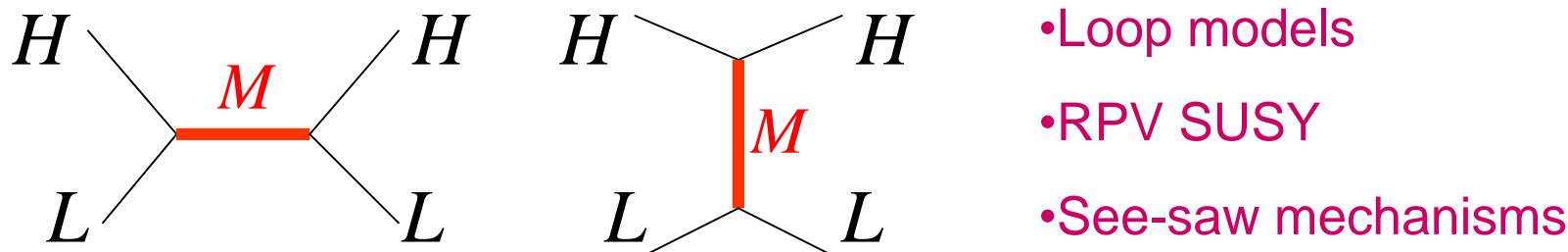
2nd Possibility: Majorana

Renormalisable $\Delta L = 2$ operator $\lambda_\nu \bar{L} L \Delta$ where Δ is light Higgs triplet with VEV $< 8\text{GeV}$ from ρ parameter

Non-renormalisable $\Delta L = 2$ operator $\frac{\lambda_\nu}{M} \bar{L} L H H = \frac{\lambda_\nu}{M} \langle H^0 \rangle^2 \bar{\nu}_{eL} \nu_{eL}^c$ Weinberg

This is nice because it gives naturally small Majorana neutrino masses $m_{LL} \sim \langle H^0 \rangle^2 / M$ where M is some high energy scale

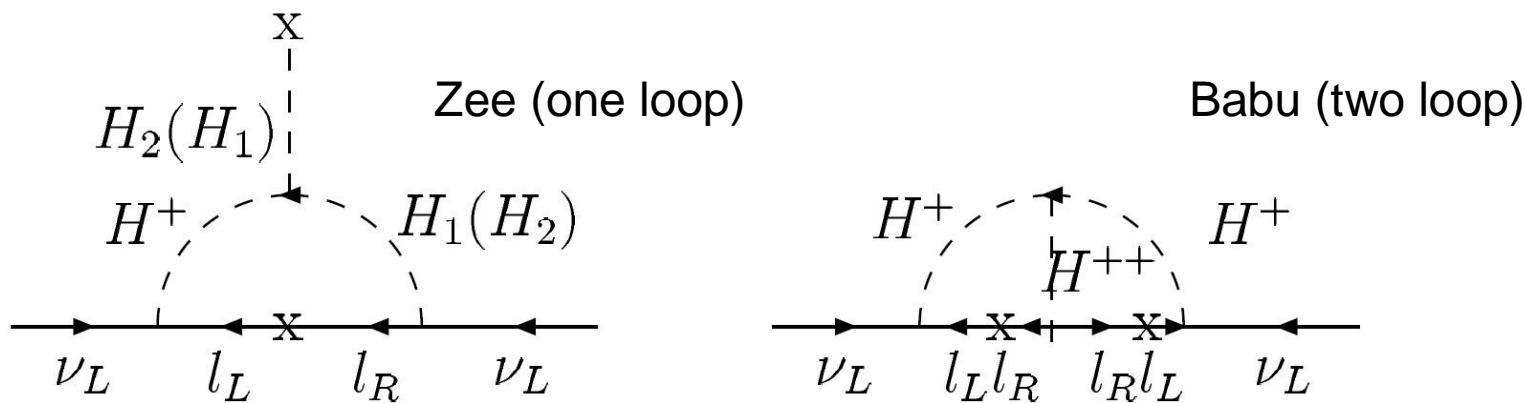
The high mass scale can be associated with some heavy particle of mass M being exchanged (can be singlet or triplet)



- Loop models

Introduce Higgs singlets and triplets with couplings to leptons

$$-\mathcal{L}^{yuk} = f_{ij} H^{++} l_i l_j + g_{ij} H^+ l_i \nu_j + h_{ij} H^0 \nu_i \nu_j$$

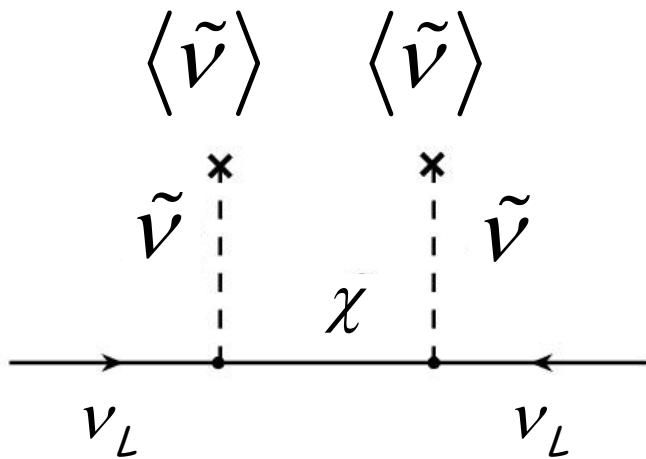


• RPV SUSY

Another way to generate Majorana masses is via SUSY

Scalar partners of lepton doublets (slepton doublets)
have same quantum numbers as Higgs doublets

If R-parity is violated then sneutrinos may get (small)
VEVs inducing a mixing between neutrinos and
neutralinos χ



$$m_{LL}^\nu \approx \frac{\langle \tilde{\nu} \rangle^2}{M_\chi} \approx \frac{MeV^2}{TeV} \approx eV$$

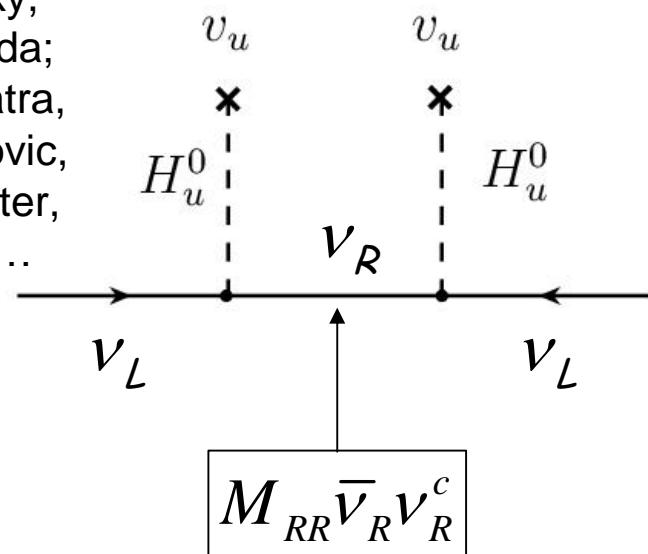
Also need loops

Drees, Dreiner, Diaz, Hirsch, Porod,
Romao, Valle, ...

See Senjanovic talk for type III

• Type I and II see-saw mechanism

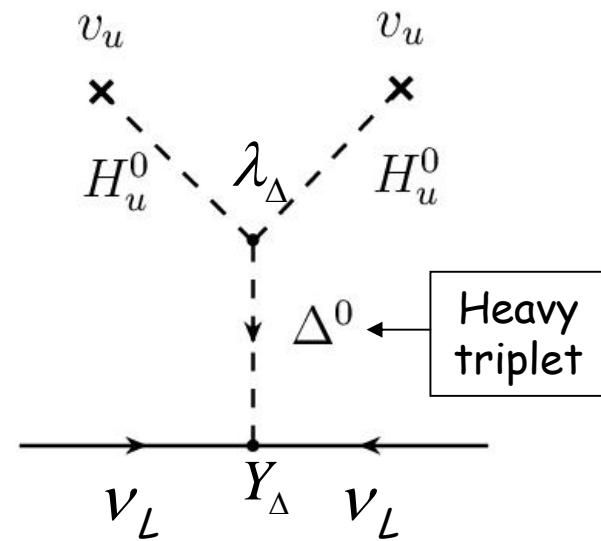
P.Minkowski,
Gell-Mann,
Ramond,
Slansky,
Yanagida;
Mohapatra,
Senjanovic,
Schechter,
Valle,...



$$m_{LL}^I \approx -m_{LR} M_{RR}^{-1} m_{LR}^T$$

Type I see-saw mechanism Type II see-saw mechanism

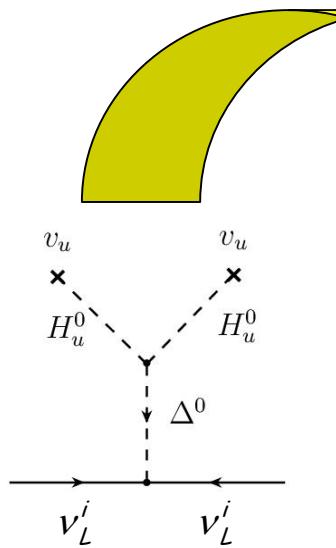
Lazarides,
Magg,
Mohapatra,
Senjanovic,
Shafi,
Wetterich
(1981)



$$m_{LL}^{II} \approx \lambda_\Delta Y_\Delta \frac{v_u^2}{M_\Delta}$$

• Type II upgrade of type I models

Antusch, SFK

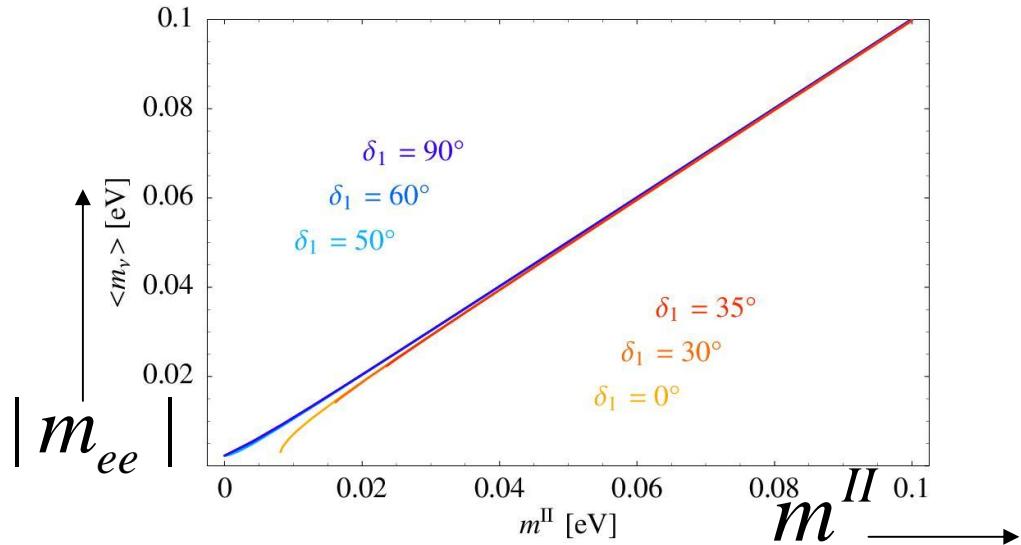


$$m_{LL}^\nu = m^{II} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} - m_{LR} M_{RR}^{-1} m_{LR}^T$$

Unit matrix type II contribution from an SO(3) family symmetry

Hierarchical type I contribution controls the neutrino mixings and mass splittings

Type II contribution governs the neutrino mass scale and renders neutrinoless double beta decay observable



Very precise Tri-bimaximal mixing (TBM) ?

Harrison, Perkins, Scott

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\theta_{12} = 35^\circ, \quad \theta_{23} = 45^\circ, \quad \theta_{13} = 0^\circ.$$

c.f. data

$$\theta_{12} = 33.8^\circ \pm 1.4^\circ, \theta_{23} = 45^\circ \pm 3^\circ, \theta_{13} < 12^\circ$$

- Current data is consistent with TBM
- But no convincing reason for exact TBM – expect deviations

It is useful to consider the following parametrization of the PMNS mixing matrix in terms of deviations from TBM

$$s_{13} = \frac{r}{\sqrt{2}}, \quad s_{12} = \frac{1}{\sqrt{3}}(1 + s), \quad s_{23} = \frac{1}{\sqrt{2}}(1 + a)$$

$$0 < r < 0.22, \quad -0.11 < s < 0.04, \quad -0.12 < a < 0.13.$$

r = reactor

s = solar

a = atmospheric

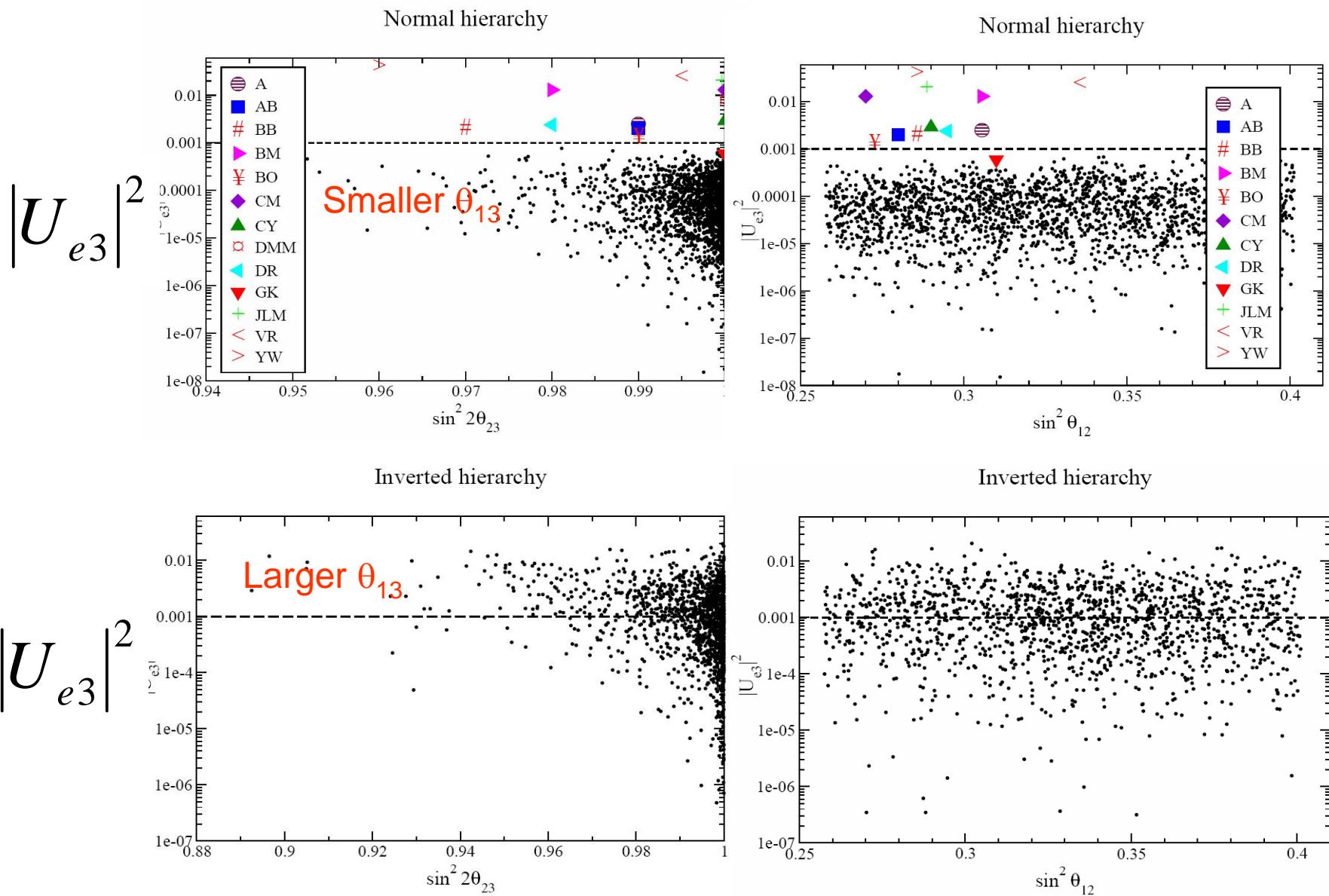
SFK;
see also
Pakvasa,
Rodejohann,
Wyler;
Bjorken,
Harrison,
Scott,
Parke,...

$$U \approx \begin{pmatrix} \sqrt{\frac{2}{3}}(1 - \frac{1}{2}s) & \frac{1}{\sqrt{3}}(1 + s) & \frac{1}{\sqrt{2}}re^{-i\delta} \\ -\frac{1}{\sqrt{6}}(1 + s - a + re^{i\delta}) & \frac{1}{\sqrt{3}}(1 - \frac{1}{2}s - a - \frac{1}{2}re^{i\delta}) & \frac{1}{\sqrt{2}}(1 + a) \\ \frac{1}{\sqrt{6}}(1 + s + a - re^{i\delta}) & -\frac{1}{\sqrt{3}}(1 - \frac{1}{2}s + a + \frac{1}{2}re^{i\delta}) & \frac{1}{\sqrt{2}}(1 - a) \end{pmatrix}$$

For a list of oscillation formulae in terms of r,s,a see SFK [arXiv:0710.0530](https://arxiv.org/abs/0710.0530)

Perturbing the TBM neutrino mass matrix

Albright, Rodejohann



TBM mass matrices in three different bases

1. Diagonal charged lepton basis $U_{MNS} = V^{\nu_L \dagger}$

$$m_{LR}^E = \begin{pmatrix} m_e & 0 & 0 \\ 0 & m_\mu & 0 \\ 0 & 0 & m_\tau \end{pmatrix} \quad m_{LL} = \frac{m_3}{2} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 1 \\ 0 & 1 & 1 \end{pmatrix} + \frac{m_2}{3} \begin{pmatrix} 1 & 1 & -1 \\ 1 & 1 & -1 \\ -1 & -1 & 1 \end{pmatrix} + \frac{m_1}{6} \begin{pmatrix} 4 & -2 & 2 \\ -2 & 1 & -1 \\ 2 & -1 & 1 \end{pmatrix}$$

2. Cabibbo-Wolfenstein basis $U_{MNS} = V^{E_L} V^{\nu_L \dagger}$ $\omega = e^{2\pi i/3}$

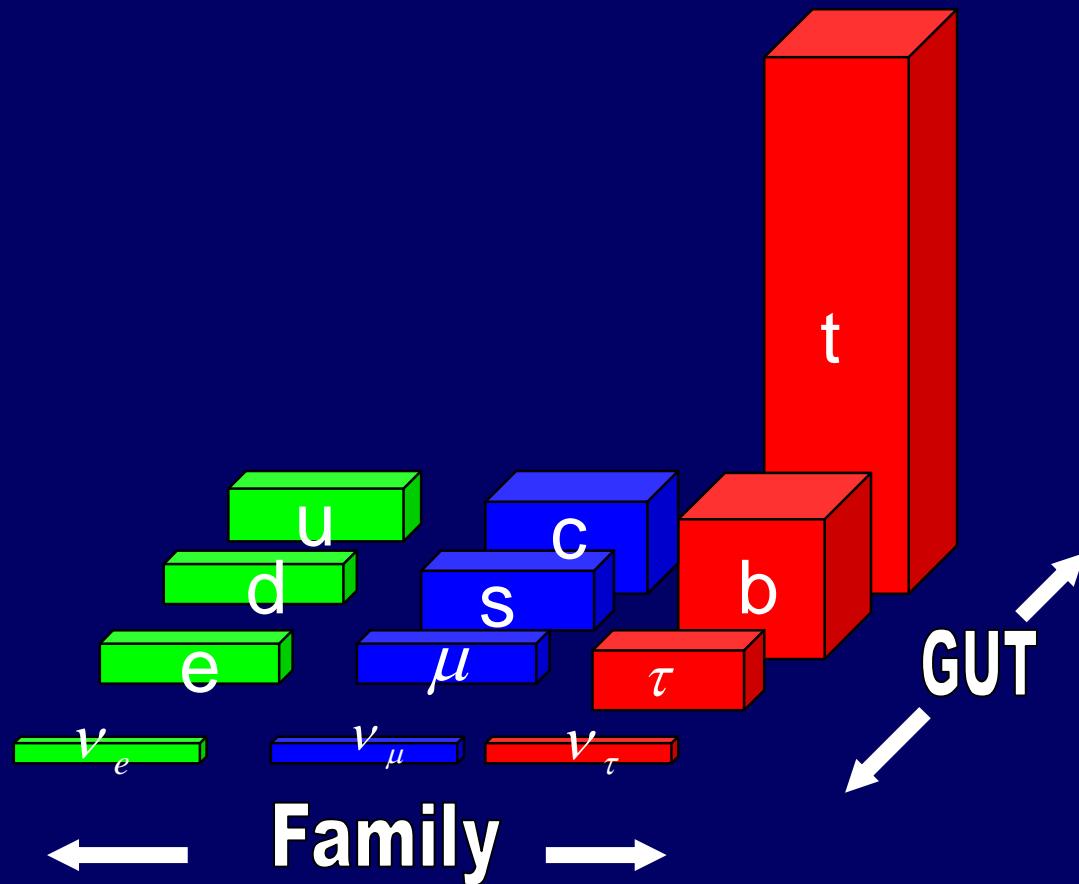
$$m_{LR}^E = \begin{pmatrix} m_e & m_\mu & m_\tau \\ m_e & \omega^2 m_\mu & \omega m_\tau \\ m_e & \omega m_\mu & \omega^2 m_\tau \end{pmatrix}, \quad m_{LL} = \begin{pmatrix} m_2 & 0 & 0 \\ 0 & m_2 & \Delta \\ 0 & \Delta & m_2 \end{pmatrix} \rightarrow V^{E_L} = \frac{1}{\sqrt{3}} \begin{pmatrix} 1 & 1 & 1 \\ 1 & \omega^2 & \omega \\ 1 & \omega & \omega^2 \end{pmatrix}, \quad V^{\nu_L \dagger} = \begin{pmatrix} 0 & 1 & 0 \\ \frac{1}{\sqrt{2}} & 0 & \frac{-i}{\sqrt{2}} \\ \frac{1}{\sqrt{2}} & 0 & \frac{i}{\sqrt{2}} \end{pmatrix}$$

3. Diagonal neutrino basis $U_{MNS} = V^{E_L}$

$$m_{LR}^E \approx \begin{pmatrix} \sqrt{\frac{2}{3}} m_e & -\sqrt{\frac{1}{6}} m_\mu & \sqrt{\frac{1}{6}} m_\tau \\ \sqrt{\frac{1}{3}} m_e & \sqrt{\frac{1}{3}} m_\mu & -\sqrt{\frac{1}{3}} m_\tau \\ 0 & \sqrt{\frac{1}{2}} m_\mu & \sqrt{\frac{1}{2}} m_\tau \end{pmatrix} \quad m_{LL} = \begin{pmatrix} m_1 & 0 & 0 \\ 0 & m_2 & 0 \\ 0 & 0 & m_3 \end{pmatrix}$$

Low energy physics doesn't care about the choice of basis, but the high scale theory does

■ Family symmetry



The basic idea of family symmetry is to assign each family a new type of ~~colour~~ charge

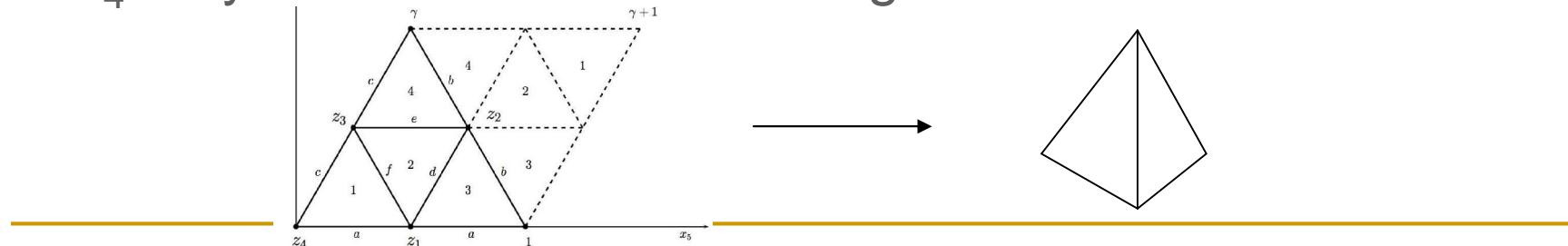
The magic symmetry A_4

- A_4 is symmetry group of the tetrahedron (Plato's ``fire'')
reproduces the TBM form of the charged lepton mass
matrix in the *Cabibbo-Wolfenstein* basis 2 Ma, Rajasakaran

- TBM form of the neutrino mass matrix then requires a
delicate Higgs vacuum alignment Ma; Altarelli,Feruglio

- A_4 may also be used to give the TBM neutrino mass
matrix in the *Flavour* basis 1 Altarelli,Feruglio

- A_4 may arise from 6D orbifolding Altarelli,Feruglio,Lin



Deriving TBM from see-saw mechanism

SFK

Diagonal RH nu basis

$$M_{RR} = \begin{pmatrix} X & 0 & 0 \\ 0 & Y & 0 \\ 0 & 0 & Z \end{pmatrix}$$

columns

$$Y_{LR}^\nu = (A \quad B \quad C)$$

See-saw I $\rightarrow m_{LL}^\nu = \frac{AA^T}{X} + \frac{BB^T}{Y} + \frac{CC^T}{Z}$

Sequential dominance \rightarrow Dominant Subdominant Decoupled

$$\left. \begin{array}{l} m_3 \\ m_2 \\ m_1 \end{array} \right\} \rightarrow m_{LL}^\nu = \frac{m_3}{2} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 1 \\ 0 & 1 & 1 \end{pmatrix} + \frac{m_2}{3} \begin{pmatrix} 1 & 1 & -1 \\ 1 & 1 & -1 \\ -1 & -1 & 1 \end{pmatrix}$$

$|A_1| = 0,$
 $|A_2| = |A_3|,$
 $|B_1| = |B_2| = |B_3|,$
 $A^\dagger B = 0$

Constrained SD

TBM mass matrix (~ 2 RHN)

This requires a non-Abelian family symmetry

Need

$$Y_{LR}^\nu = \begin{pmatrix} 0 & B_1 & - \\ A_2 & B_2 & - \\ A_3 & B_3 & - \end{pmatrix}$$

with

$$\begin{aligned} |A_1| &= 0, \\ |A_2| &= |A_3|, \\ |B_1| &= |B_2| = |B_3|, \\ A^\dagger B &= 0 \end{aligned}$$

$2 \leftrightarrow 3$ symmetry (from maximal atmospheric mixing)

$1 \leftrightarrow 2 \leftrightarrow 3$ symmetry (from tri-maximal solar mixing)

Several examples of suitable non-Abelian Family Symmetries:

SFK, Ross; Velasco-Sevilla; Varzelas

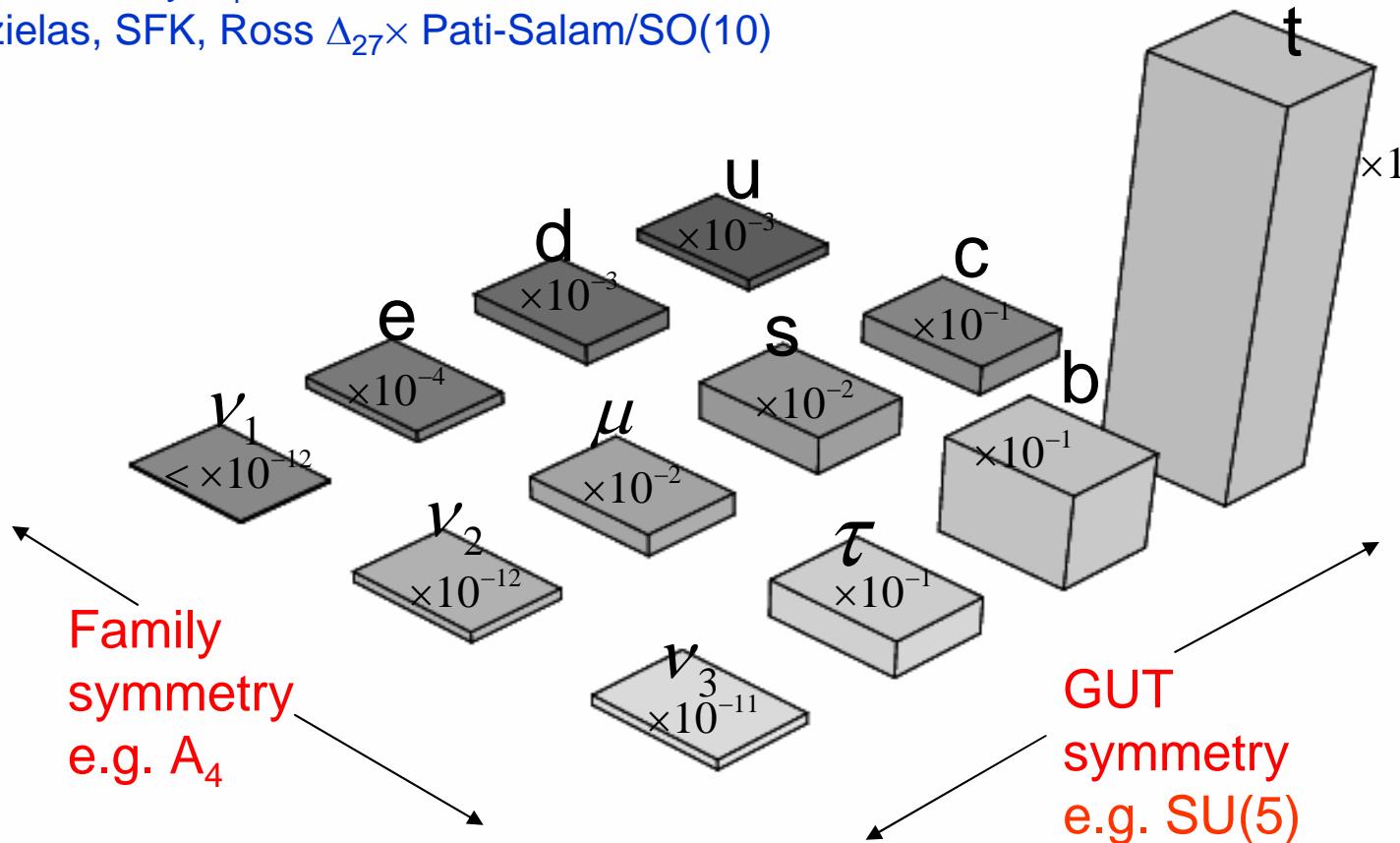
$$\left. \begin{array}{ll} SU(3) & \Delta_{27} \\ SO(3) & A_4 \end{array} \right\}$$

Discrete subgroups
preferred by vacuum
alignment

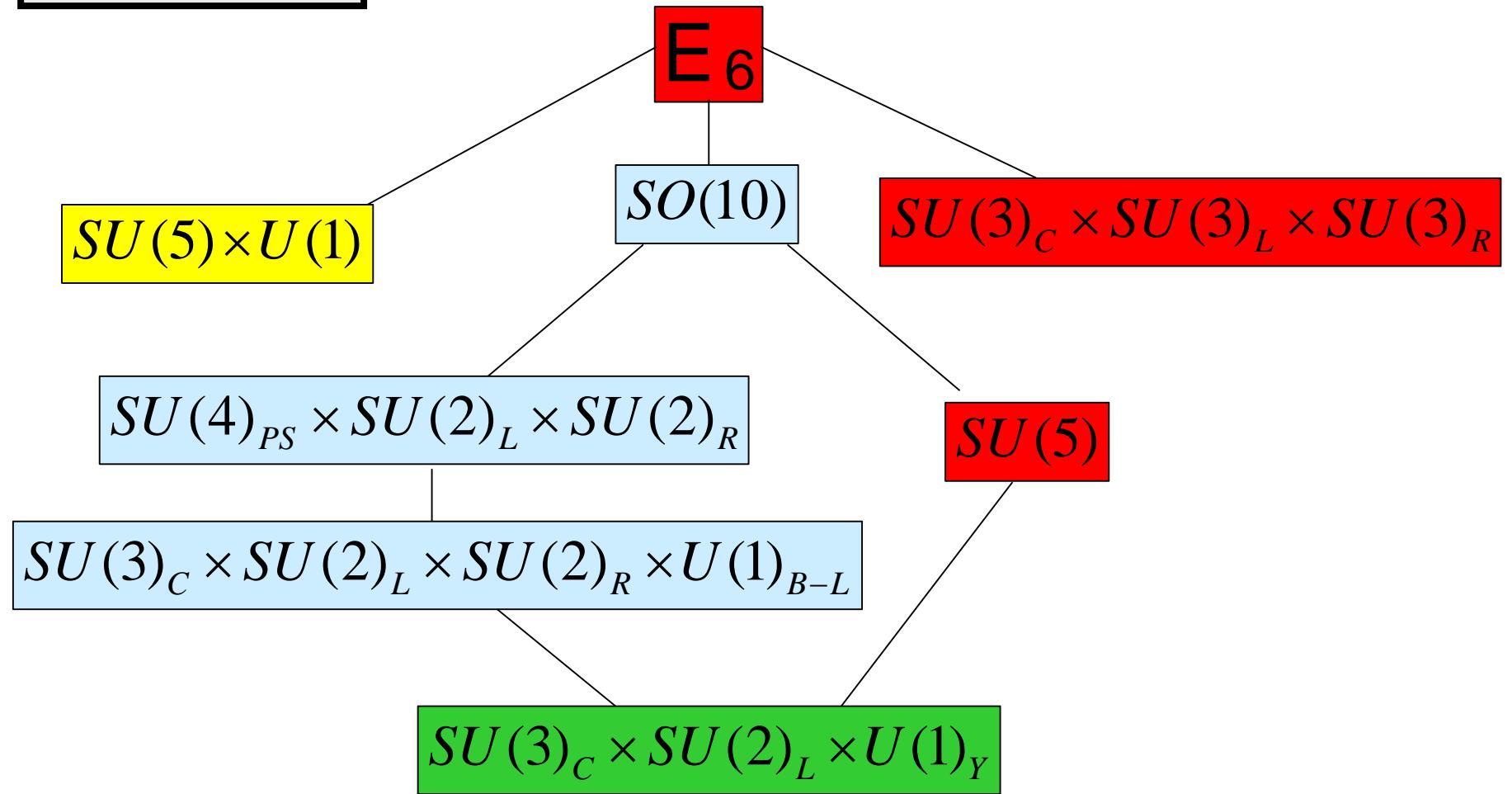
SFK, Malinsky

■ Family \times GUT symmetry

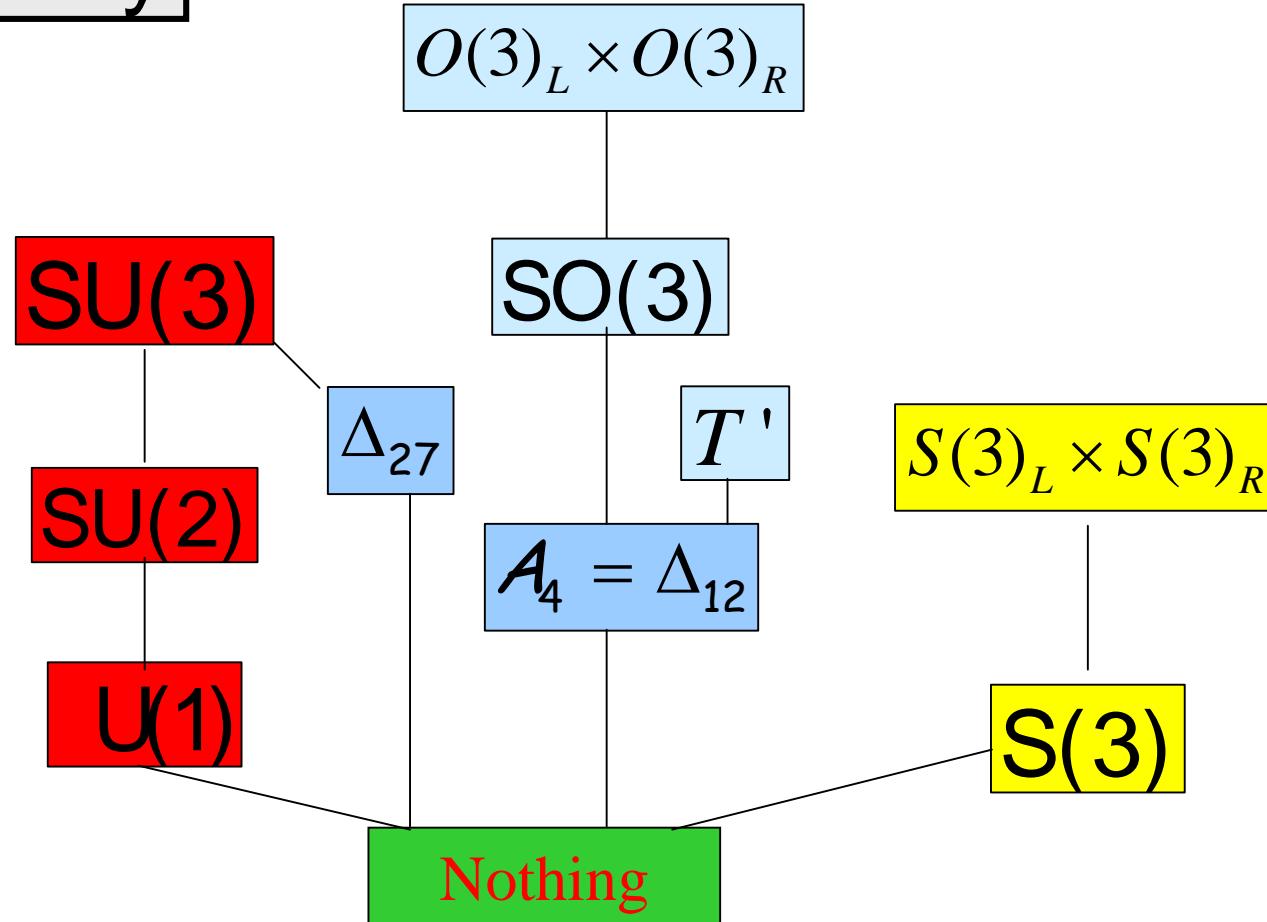
e.g. Chen and Mahanthappa $T' \times \text{SU}(5)$
Altarelli, Feruglio, Hagedorn $A_4 \times \text{SU}(5)$ (in 5d)
SFK, Malinsky $A_4 \times \text{Pati-Salam}$
Varzielas, SFK, Ross $\Delta_{27} \times \text{Pati-Salam/SO}(10)$



G_{GUT}



GFamily



General Strategy

Choose a GUT and family symmetry and write down the reps

Assign quarks, leptons, Higgs to reps $L_i = \begin{pmatrix} \nu_i \\ E_i^- \end{pmatrix}$, E_j^c , $H = \begin{pmatrix} H^+ \\ H^0 \end{pmatrix}$

Renormalizable Yukawas requires extended Higgs $H \rightarrow H^{ij}$

$$\lambda_{ij} H L_i E_j^c \rightarrow \lambda H^{ij} L_i E_j^c \quad \text{Machado,Pleitez}$$

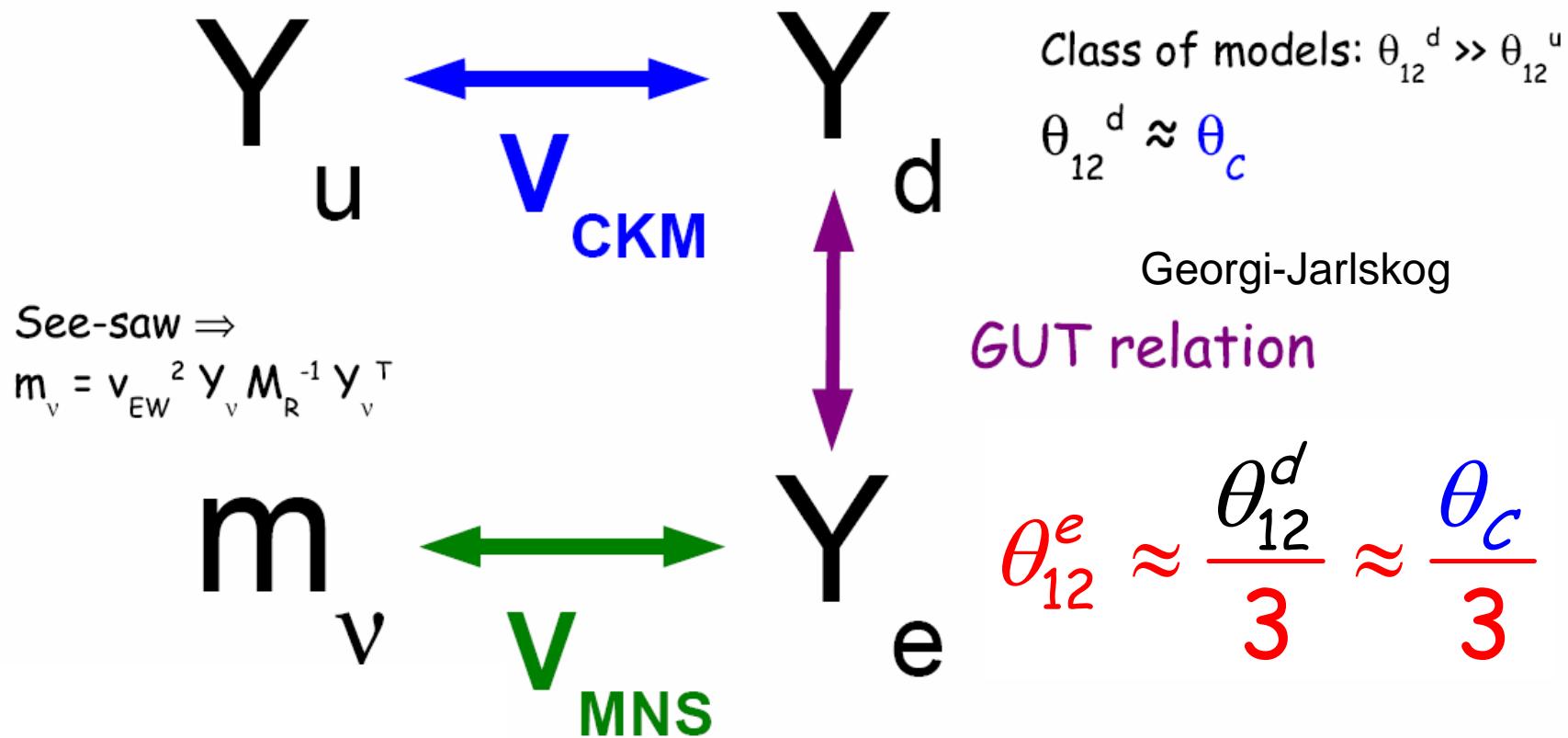
Alternatively promote Yukawas to non-renormalizable terms involving the usual Higgs H plus SM singlet *flavon fields* ϕ

$$\lambda_{ij} H L_i E_j^c \rightarrow \lambda \frac{\phi^{ij}}{M} H L_i E_j^c \quad \text{or} \quad \lambda_{ij} H L_i E_j^c \rightarrow \lambda \frac{\phi^i \phi^j}{M^2} H L_i E_j^c$$

Koide; Stech

SFK,Ross

GUT relations



Can this lead to Quark-Lepton Complementarity (QLC)?

$$\theta_{12} + \theta_c = 45^\circ$$

Petcov, Smirnov; Raidal; Ohlsson, Seidl



Sum Rules

$$U_{MNS} = V^{E_L} V^{\nu_L \dagger}$$

Cabibbo-like
Bimaximal or
Tri-bimaximal

Bjorken; Ferrandis, Pakvasa; SFK

$\theta_{13} \approx \frac{\theta_{12}^e}{\sqrt{2}} \approx \frac{\theta_C}{3\sqrt{2}} \approx 3^\circ,$

$\theta_{12} = 45^\circ (35)^\circ + \frac{\theta_C}{3\sqrt{2}} \cos \delta$

Antusch,SFK

Oscillation phase

$\theta_{12}^o = 45^\circ (35)^\circ + \theta_{13}^o \cos \delta$

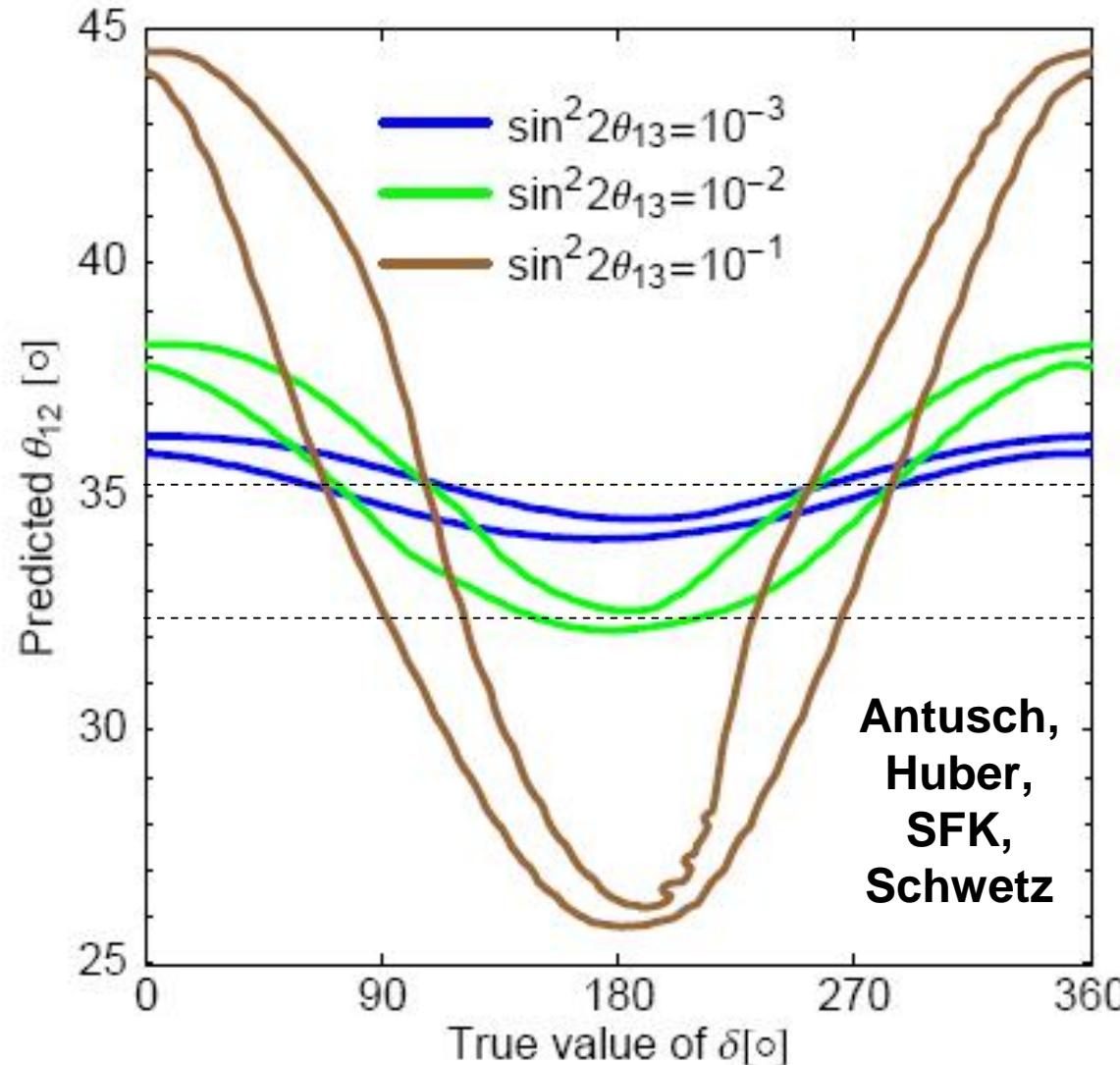
Sum Rule
SFK; Antusch,SFK; Masina

Bimaximal sum rule with 45° requires $\theta_{13} \approx \theta_C$ and $\delta \approx \pi$
 \rightarrow QLC is only achieved for a special phase and large θ_{13}

Antusch,SFK,Mohapatra

What about tri-bimaximal sum rule with 35° ?

Tri-bimaximal sum rule $\theta_{12} \approx 35.3^\circ + \theta_{13} \cos \delta$



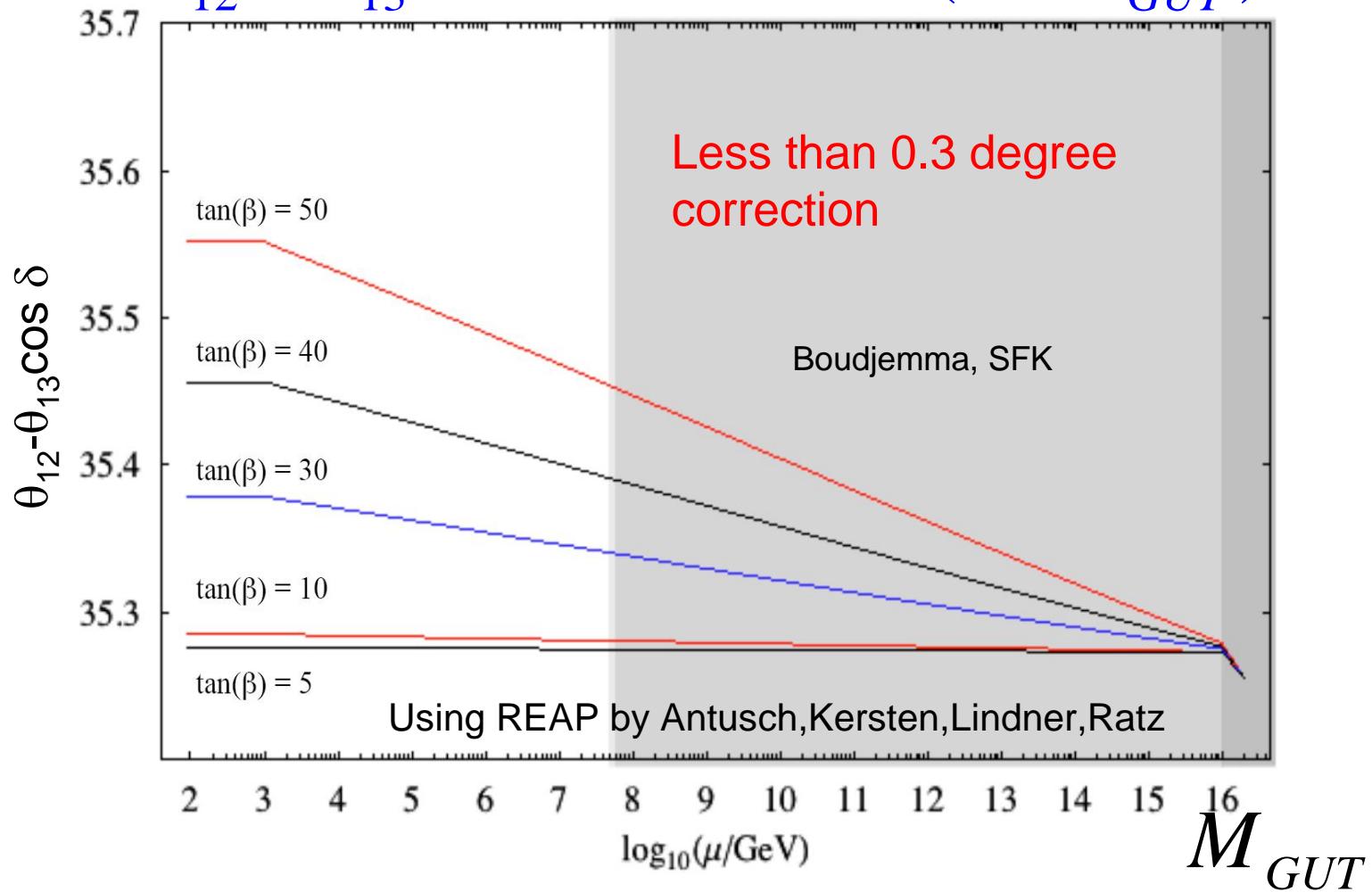
Bands show
3 σ error for an
optimized
neutrino factory
determination
of $\theta_{13} \cos \delta$

$\theta_{12} = 33.8^\circ \pm 1.4^\circ$
(current value)

Tri-bimaximal sum
rule works
incredibly well !!

RGE corrections to TBM sum rule

$$\theta_{12}^o - \theta_{13}^o \cos \delta \approx 35.3^\circ \quad (\text{at } M_{GUT})$$



■ Alternative Ideas

- Accidental family symmetry from messenger dominance
Ferretti, SFK, Romanino; Barr
- SU(8) GUTs Barr
- Mass matrices from shift symmetry: $\nu_R \rightarrow \nu_R + \eta \theta$ Friedberg, Lee; Jarlskog
- Extremization of mass matrix Jarlskog invariants Harrison, Scott
- Theories of the Koide mass formula Koide, ...
- Dirac screening in the double see-saw Lindner, Smirnov, Schmidt
- Low energy see-saw models with gauged B-L SFK, Yanagida
- Anarchy/Landscape (large θ_{13} only) Hall, Murayama, Weiner
- RH Neutrino masses in string theory Antusch, Ibanez; Nilles, Langacker
- Invariant classification of see-saw models SFK

Conclusion

- **Neutrino mass and mixing requires new physics BSM**
- Many roads for model building, but answers to key experimental questions will provide the signposts
- **One key question is how accurately is TBM realised?**
- Goal of next generation of oscillation experiments is to show that the deviations from TBM r,s,a are non-zero and measure them and δ
- If TBM is accurately realised this may imply a new symmetry of nature: family symmetry
- GUTs \times family symmetry with see-saw + CSD is very attractive framework for TBM \rightarrow sum rule prediction
- Few realistic models, complicated vacuum alignment
- **Status quo is not an option – neutrino physics demands a theory of flavour, and may provide further clues**