



Double Beta Decay: Scintillators

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Queen's University

Neutrino 2008
Christchurch, New Zealand

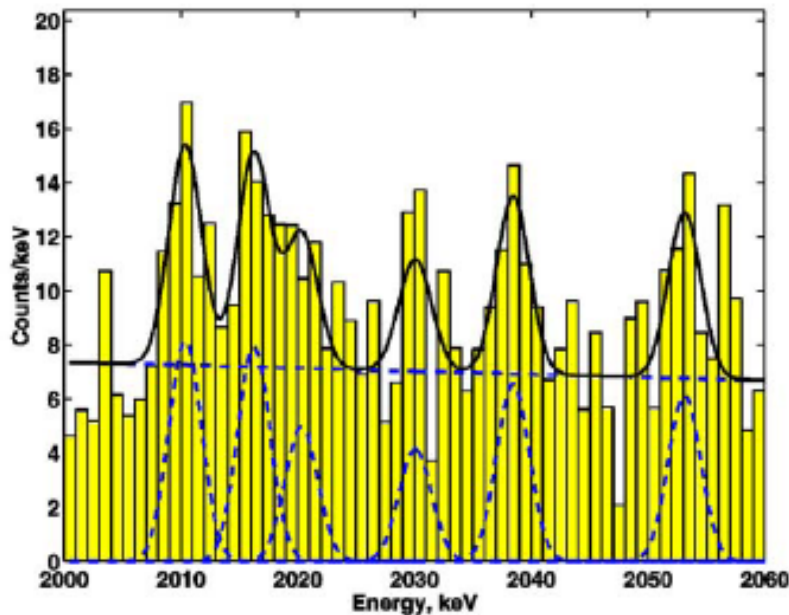


Talk Outline

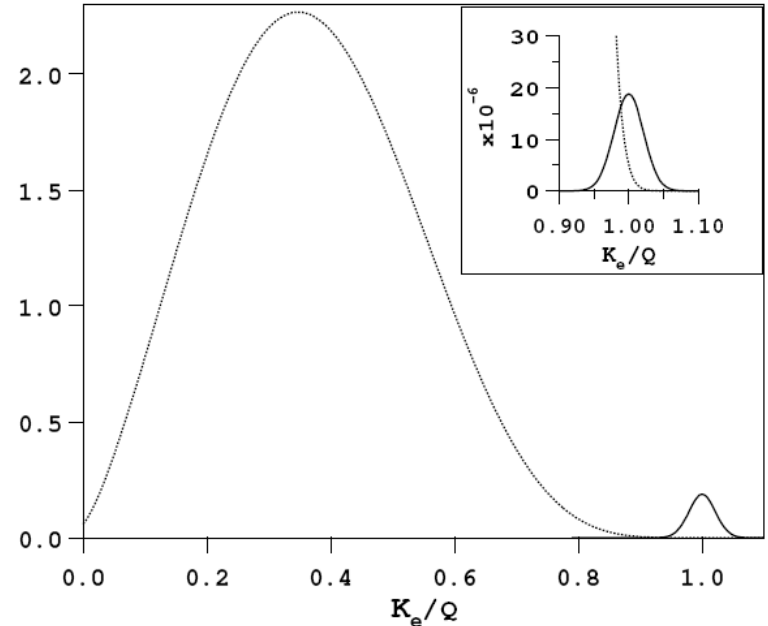
- scintillators for double beta decay:
 - what they can offer
- survey of experimental programs
 - XMASS double beta decay [liquid xenon]
 - ELEGANT / CANDLES [inorganic crystals]
 - Kiev group [inorganic crystals]
 - SNO+ double beta decay [loaded liquid]
- summary

Why Good Energy Resolution is Needed?

- to separate $0\nu\beta\beta$ from $2\nu\beta\beta$
- to separate $0\nu\beta\beta$ signal from other gamma lines



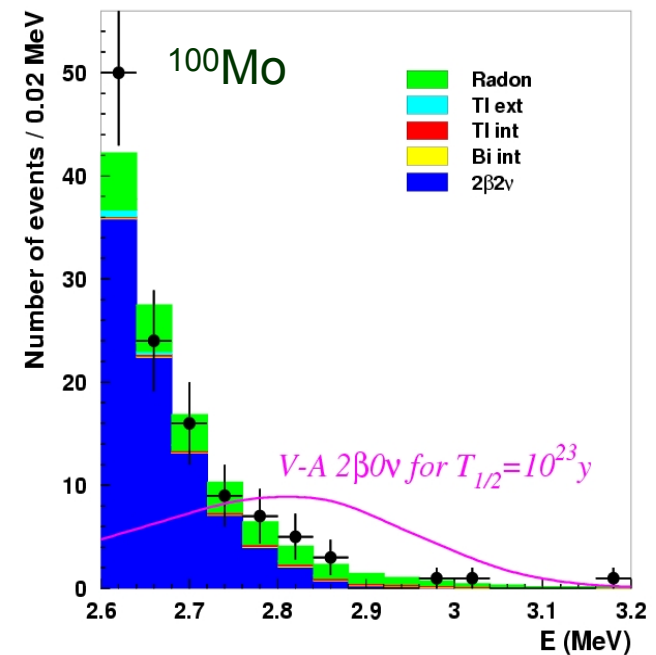
from H.V. Klapdor-Kleingrothaus et al.



from S. Elliott and P. Vogel

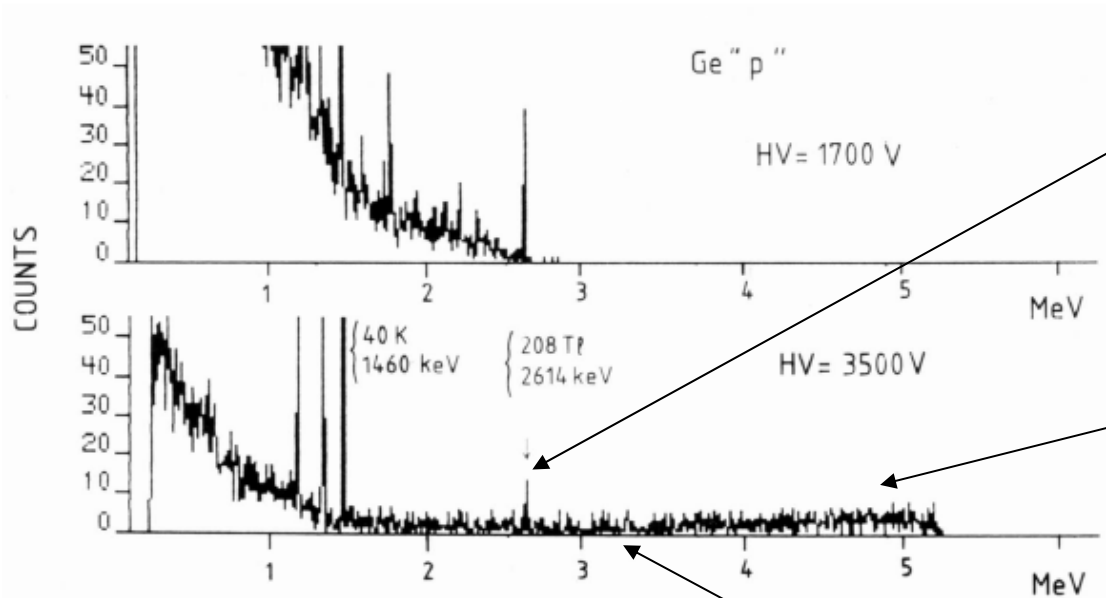
Can You Live With Worse Resolution?

- to separate $0\nu\beta\beta$ from $2\nu\beta\beta$
 - **YES!** by fitting the endpoint shape...resolution is far less important when fitting spectral shapes than simply counting signal and background events in an energy bin
 - this is already done (e.g. NEMO-3)
- to separate $0\nu\beta\beta$ signal from other gam
 - **YES!** if there are no backgrounds!
- how to achieve zero (low) γ background
 - use B-field tracking detector: tags $\beta^-\beta^-$ fr or
 - choose a high Q-value isotope
 - with an ultra-low background detector



from F. Piquemal

Above the ^{208}Tl Line at 2.614 MeV



highest energy line from natural radioactivity

continuum background from internal Th chain contamination

typical Ge spectrum from Ph. Hubert

continuum background from internal U chain (radon) contamination ends at 3.2 MeV

if you are searching for a peak, you can live with a low-level continuum background

$\beta\beta$ Isotopes with High Q-values

isotope	Q-value [MeV]	natural abundance
^{48}Ca	4.27	0.19%
^{150}Nd	3.37	5.6%
^{96}Zr	3.35	2.8%
^{100}Mo	3.03	9.6%
^{82}Se	3.00	9.2%
^{116}Cd	2.80	7.5%

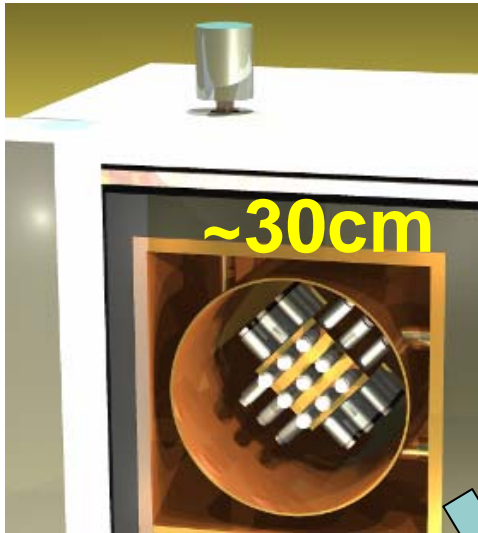
What Do Scintillators Offer?

- “economical” way to build a detector with a **large** amount of isotope
- several isotopes can be made into (or put in) a scintillator
- ultra-low background environment can be achieved (e.g. phototubes stand off from the scintillator, self-shielding of fiducial volume)
- with a liquid scintillator, possibility to purify *in-situ* to further reduce backgrounds

Experimental Programs – I

- XMASS double beta decay
 - liquid xenon scintillation
 - ^{136}Xe , Q-value = 2.48 MeV
 - slides from S. Moriyama

Strategy of the XMASS project



Prototype detector
(FV 3kg)

Confirmation of
feasibilities of
the ~1ton detector



~1 ton detector
(FV 100kg)

Dark matter search
Under construction

BG reduction with
self shield is
effective $< 500\text{keV}$

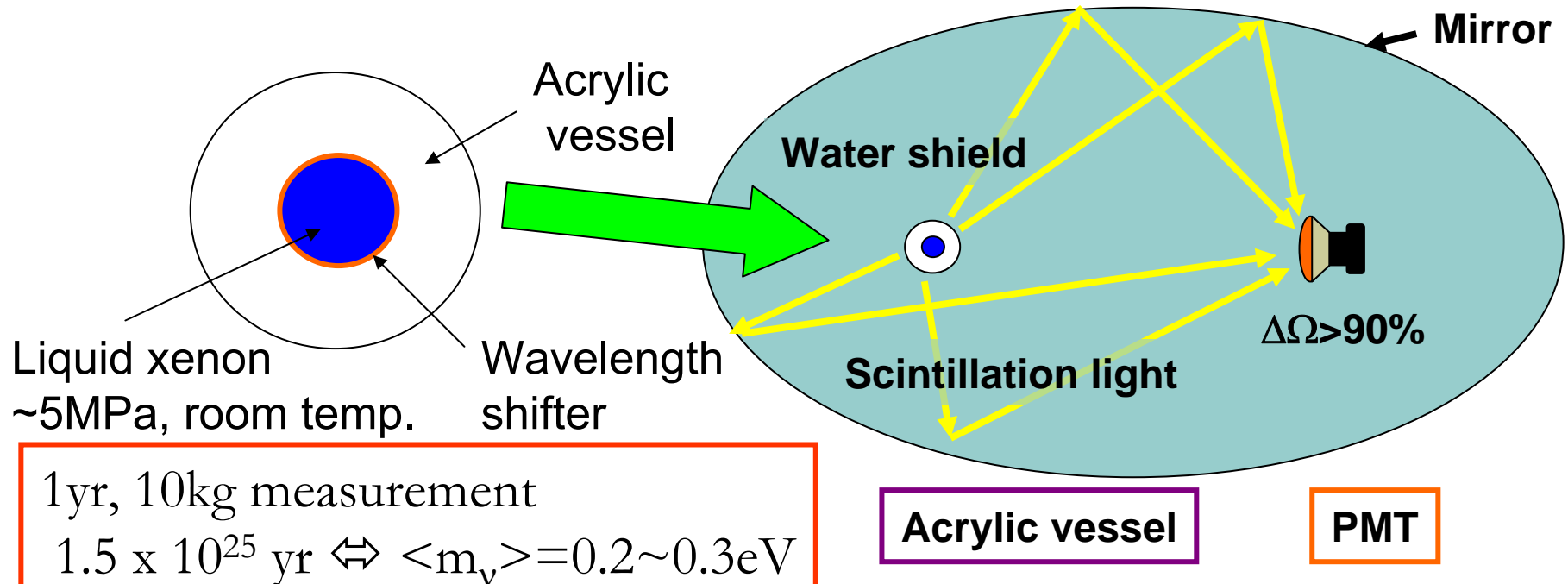
~20 ton detector
(FV 10ton)

Solar neutrinos
Dark matter search

Double beta decay option
w/ different design to realize
low background at ~MeV.

XMASS: ^{136}Xe double beta decay

- One possible method utilizes a high pressure liquid-xenon detector under room temperature. R&D ongoing.



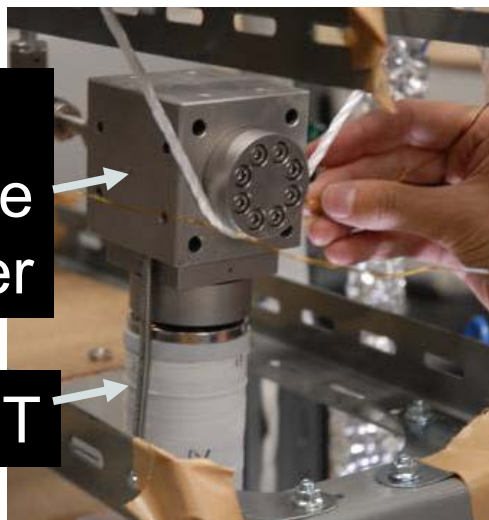
RI contamination in the acrylic vessel ($\sim 10^{-12}\text{g/gU}$) limits its sensitivity.

- Photon yield at room temp. ~ 29 photons/keV
(K. Ueshima *et al.*, arXiv0803.2888)
- Energy resolution needs to be evaluated.

Photon yield measurement at room temp./ wavelength shifter/elliptic water tank

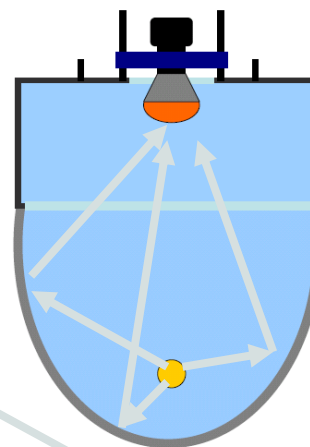
High
pressure
chamber

PMT



~29photons/keV
arXiv0803.2888

D.N.Mckinsey et.al.
NIMB 132 (1997) 351



Elliptic tank



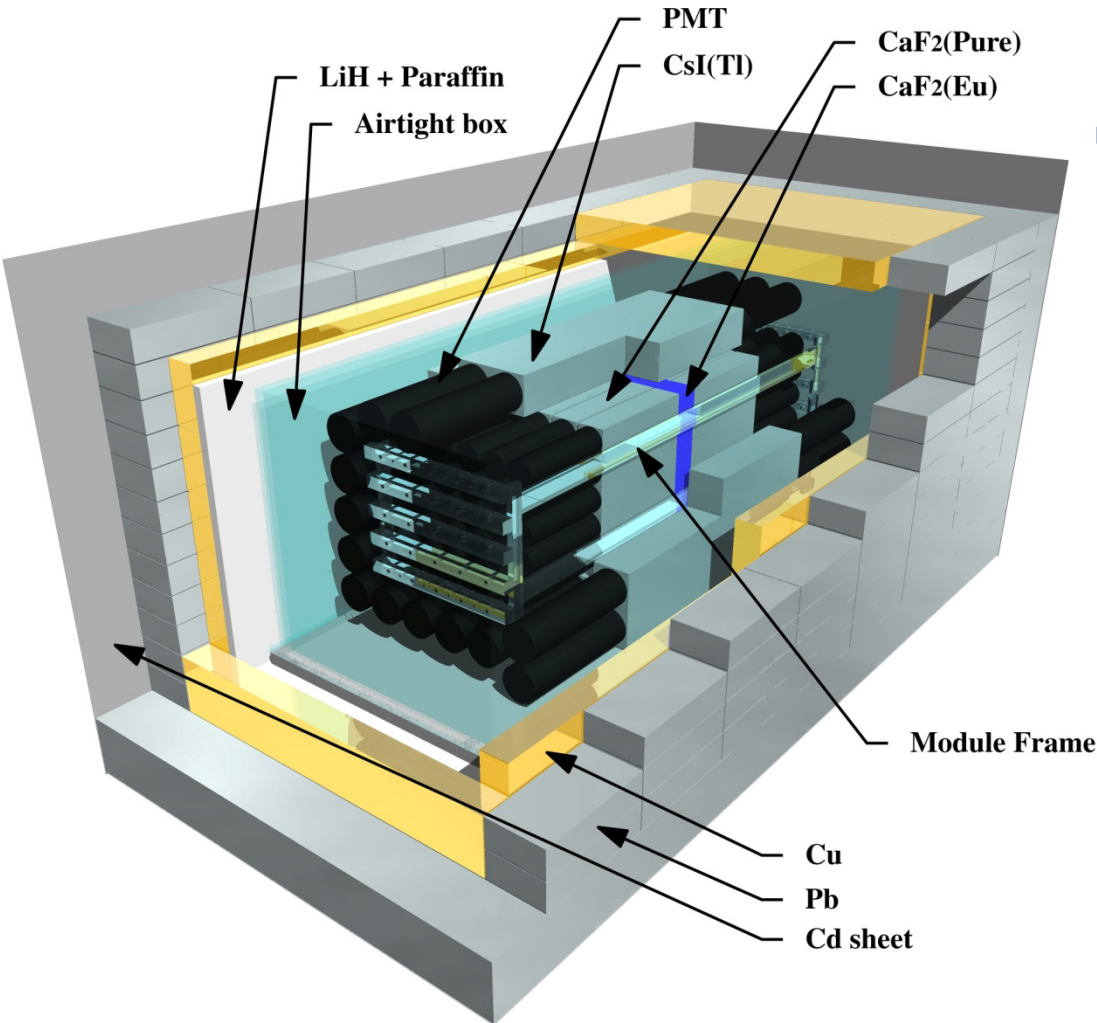
0.5% TPB doped PS, 100μm

- Three components were developed. Under examination.

Experimental Programs – II

- Osaka group: slides from T. Kishimoto and S. Umehara
- past: ELEGANT VI
 - $\text{CaF}_2(\text{Eu})$ scintillating crystals
 - 7.7 g of ^{48}Ca , Q-value = 4.27 MeV
 - ran at Oto Cosmo Observatory
- future: CANDLES III
 - pure CaF_2 scintillating crystals
 - U chain: $\sim 36 \mu\text{Bq/kg}$ (30 times better than ELEGANT VI)
 - Th chain: $\sim 29 \mu\text{Bq/kg}$ (3 times better than ELEGANT VI)
 - ~ 300 kg of crystals (that's $\sim 400\text{-}450$ g of ^{48}Ca)
 - expected resolution: $\sim 3.5\%$ FWHM at endpoint
 - will run in Kamioka

ELEGANT VI



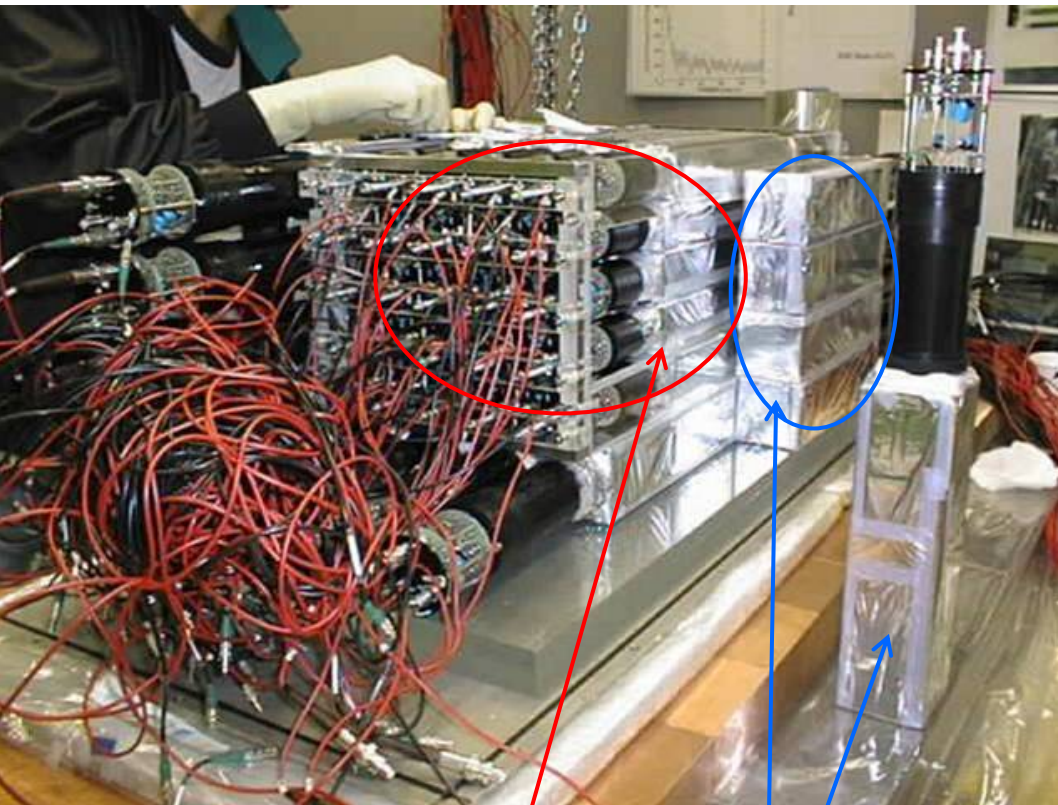
Surrounded by H_3BO_3 loaded-water tank

- 23 CaF_2 modules
 - $\sim 3.5 \text{ kg } ^{19}\text{F}$
 - $\sim 7.7 \text{ g } ^{48}\text{Ca}$
- Background reduction
 - least material : non hygroscopic
 - **4π active shield**
 - ♦ $\text{CaF}_2(\text{Eu}) + \text{CaF}_2(\text{pure})$
 - roll-off ratio
 - ♦ segmentation
 - ♦ CsI(Tl) veto detector
 - passive shield
 - ♦ OFHC Cu(t:5 cm), Pb(t:10 cm)
 - ♦ air-tight box + N_2 gas purge
 - Rn in the air
 - ♦ LiH + paraffin(t:15 mm), Cd sheet(t:0.6 mm), and $\text{H}_3\text{BO}_3 + \text{H}_2\text{O}$ tank
 - neutron

at **Oto Cosmo Observatory**

ELEGANT VI

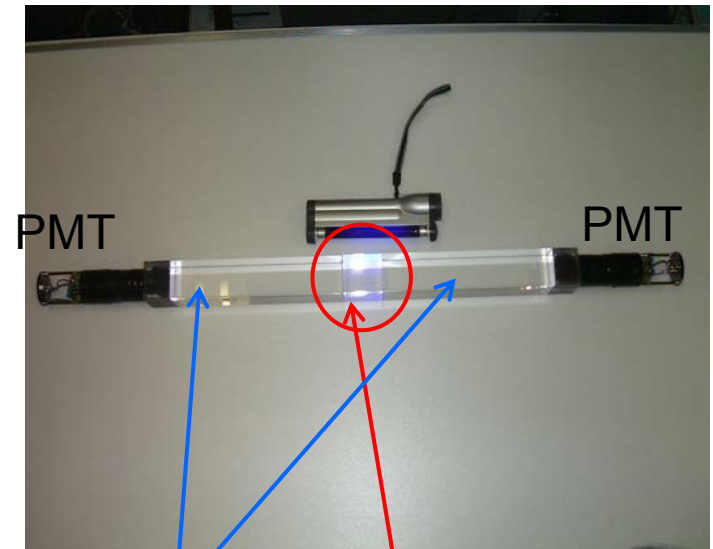
Scintillators in ELEGANT VI System



CaF₂ Module

CsI(Tl)

CaF₂ Module



CaF₂(pure)

CaF₂(Eu)

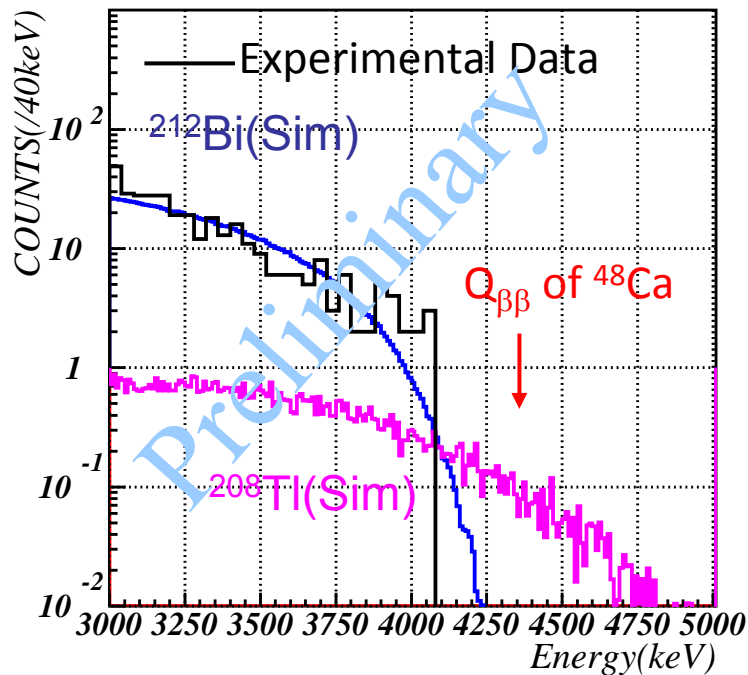
... UV Light

... Visible Light

Active Shielding Technique
in ELEGANT VI system

Double beta decay of ^{48}Ca

Double beta decay of ^{48}Ca



Date	Analysis	Number of Event	Expected BG(^{212}Bi , ^{214}Bi , ^{208}Tl)	Detection Efficiency	Live Time kg· day
Jun 1998 -	without FADC	0	1.30	0.55	1553
Jan 2003 -	with FADC	0	0.27	0.53	1114
Jan 2004 -	with FADC	0	0.43	0.53	2280

$0\nu\beta\beta$ Half-Life of ^{48}Ca :
 $> 6 \times 10^{22}$ year (90% C.L.)
 Preliminary

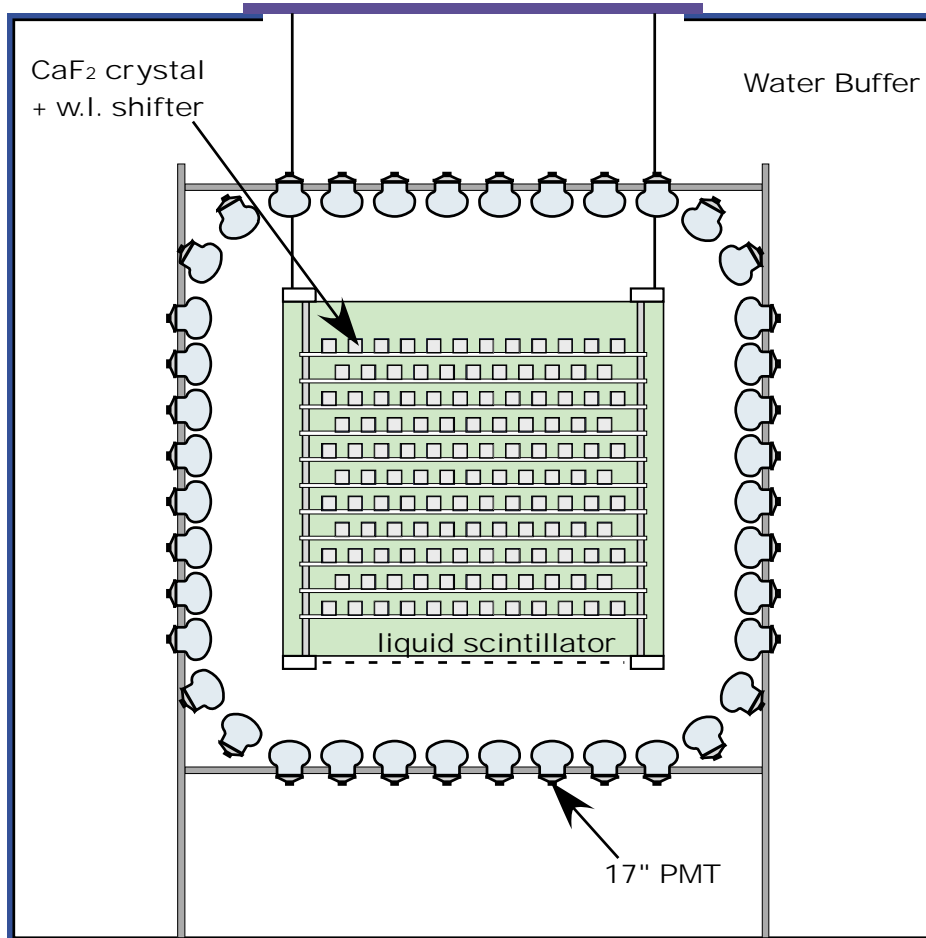


Next step: **CANDLES**

CANDLES



Calcium fluoride for studies of Neutrino and Dark matters
by Low Energy Spectrometer



- undoped CaF_2 ($\text{CaF}_2(\text{pure})$)
 - ^{48}Ca (0.187%)
 - 305 kg (III-chika)
 - 3.4 t (IV) $\sim 0.1 \text{ eV}$
 - 30 t, 2% enriched (V) $\sim 30 \text{ meV}$ (best NME)
- Liquid Scintillator (LS)
 - 4π active shield
 - also wavelength shifts light
- Photomultiplier
 - large photo-coverage
- Water buffer
 - Passive shield

Two-Layer Wavelength Shifter

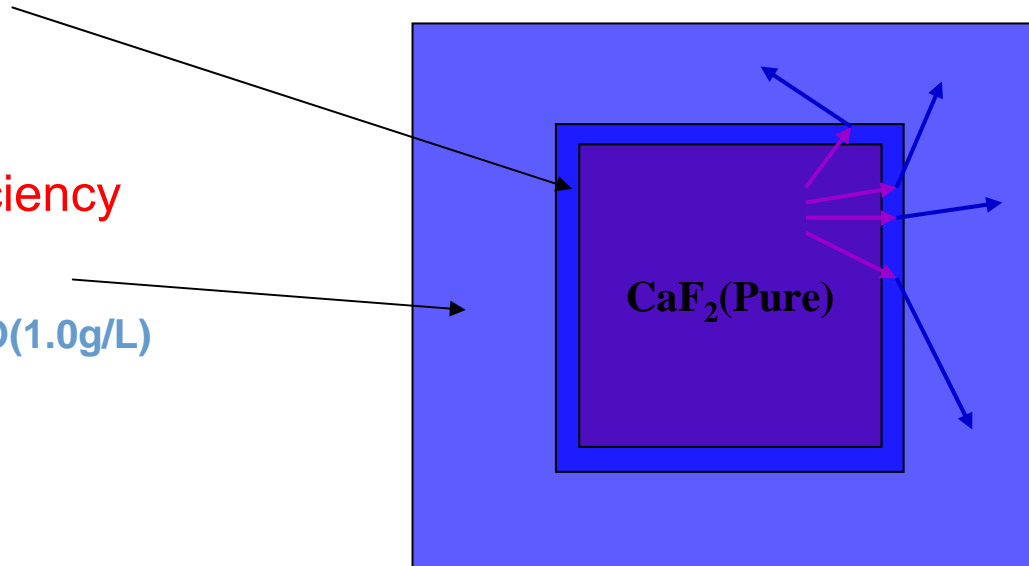
Concept of Method

□ WLS Phase

- M.O.(100%)+PPO(0.3g/L)
- 5-10 mm in thickness
- large conversion efficiency

□ Veto Phase

- M.O.(80%)+PC(20%)+PPO(1.0g/L)
- large light output

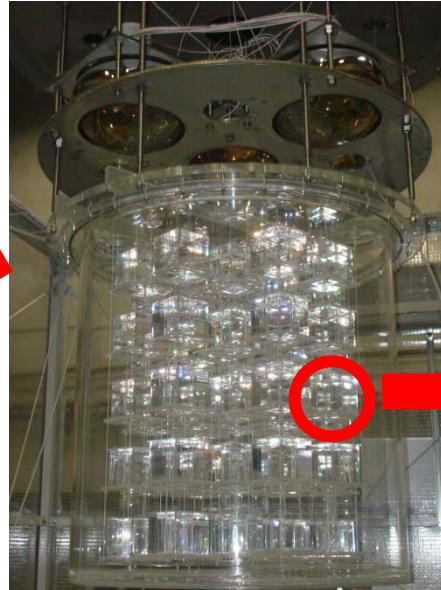
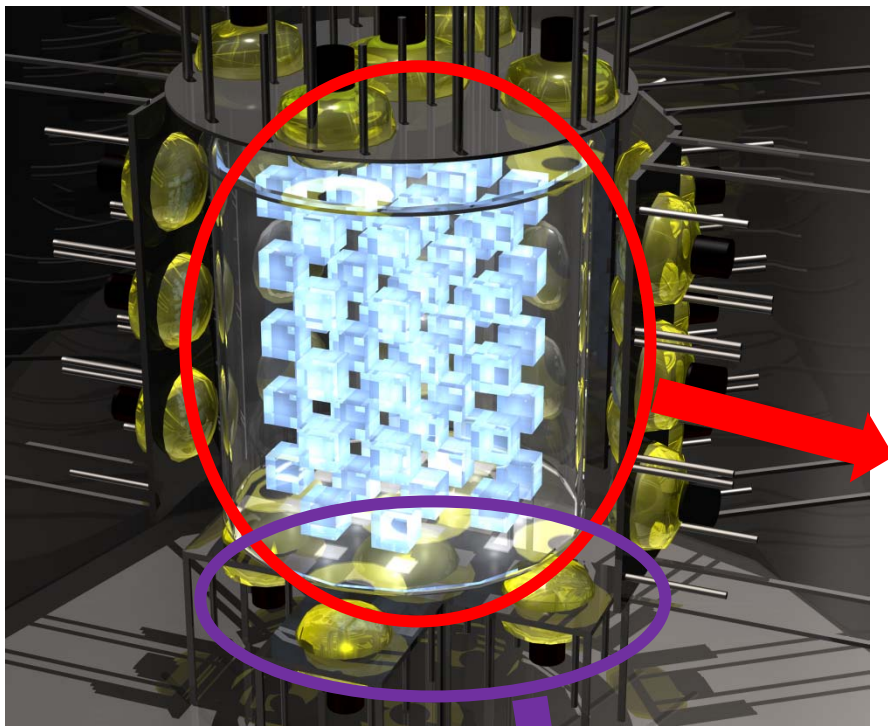


CANDLES III



Candles

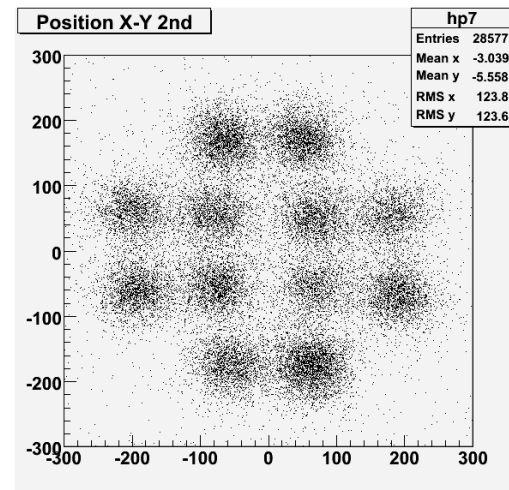
CaF_2 : 191 kg
 $10^3 \text{ cm}^3 \times 60$



Tank: $\phi 2.8 \times h 2.6 \text{ m}$



40 PMTs



Position
reconstruction
in X-Y plane

CANDLES III Status

- ICRR completed the excavation of two new chambers in Kamioka (one for XMASS and one for CANDLES)
 - new space ready for occupancy at the end of this year
- ~300 kg of crystal will be installed in detector
 - $\langle m_\nu \rangle < 0.5 \text{ eV}$
- data taking starts in 2009
- also in the process of requesting funding to enlarge for the future

Experimental Programs – III

- Kiev group: slides from F. Danevich
 - experiments developed and/or considered in the past with different scintillating crystals with different isotopes
 - e.g. CAMEO, CARVEL, etc.
 - possible deployment of crystals in large, existing detectors (e.g. Borexino, SNO)
 - currently the following scintillating crystals (and experiments) are being developed
 - $^{116}\text{CdWO}_4$ with ^{116}Cd , Q-value = 2.80 MeV
 - also ^{106}Cd $\beta^+\beta^+$ decay, Q-value = 2.77 MeV
 - CaMO_4 with ^{100}Mo , Q-value = 3.03 MeV
 - ZnWO_4 with ^{64}Zn , Q-value = 1.10 MeV

Kyiv Institute for Nuclear Research, Solotvina Underground Laboratory (Ukraine)

Main results

2β decay of ^{116}Cd

$$T_{1/2}^{2\nu} = 2.9 \times 10^{19} \text{ yr}$$

$$T_{1/2}^{0\nu} \geq 1.7 \times 10^{23} \text{ yr @ 90\% CL}$$

$$\langle m_\nu \rangle \leq 1.7 \text{ eV}$$

[PRC 68 (2003) 035501]

Search for ^{160}Gd 2β decay

$$T_{1/2}^{0\nu} \geq 1.3 \times 10^{21} \text{ yr}$$

[NPA 694 (2001) 375]

First experiment to search for 2β of ^{64}Zn by using ZnWO_4

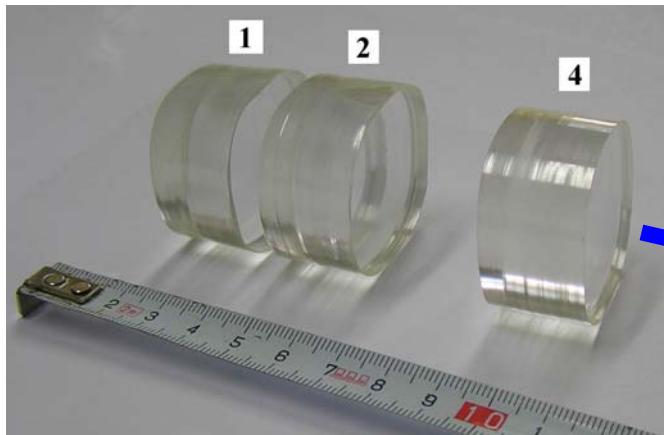
[NIMA 544 (2005) 553]

Present projects

- R&D of advanced ^{116}Cd 2β experiment (producing of $\sim 1.2\text{-}1.8$ kg $^{116}\text{CdWO}_4$ crystal scintillators) in collaboration with ITEP (Moscow, Russia), KIMS (Korea), NIIC (Novosibirsk, Russia)
- Search for 2β of Zinc and Tungsten by ZnWO_4 in collaboration with DAMA (experiments are running in the LNGS, Italy)
- CaMoO_4 crystal scintillators for 2β decay of ^{100}Mo [large collaboration, see NIMA 584 (2008) 334]
- R&D of ≈ 0.15 kg $^{106}\text{CdWO}_4$ in collaboration with DAMA (Italy), JINR (Dubna, Russia) crystal producers in Ukraine and Russia

CaMoO₄ crystal scintillators

2 β decay of ¹⁰⁰Mo

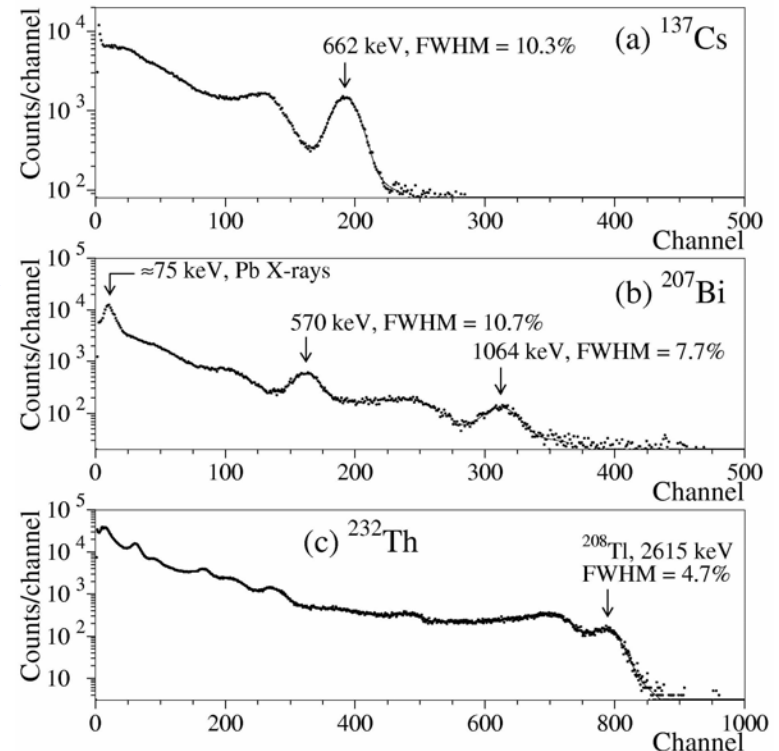


CARAT, Lviv, Ukraine



ICMSAI, Moscow, Russia

NIMA 584 (2008) 334

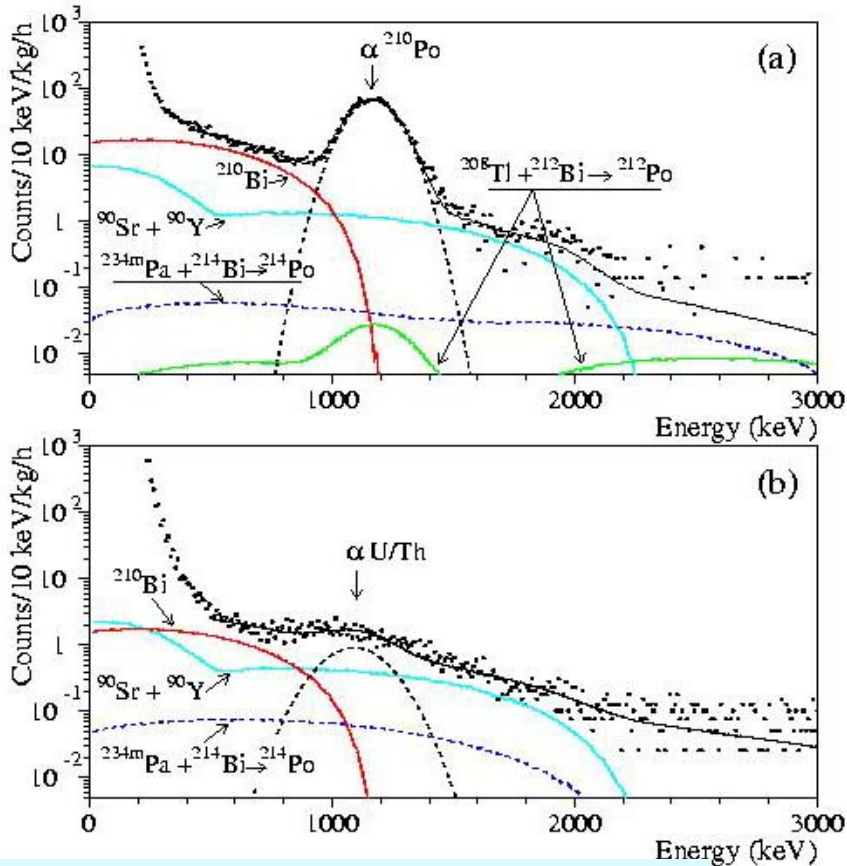


Energy resolution

FWHM=10.3% for 662 keV γ line of ¹³⁷Cs was obtained with CaMoO₄ crystal scintillators produced by CARAT

CaMoO₄ radiopurity

CARAT, Lviv, Ukraine



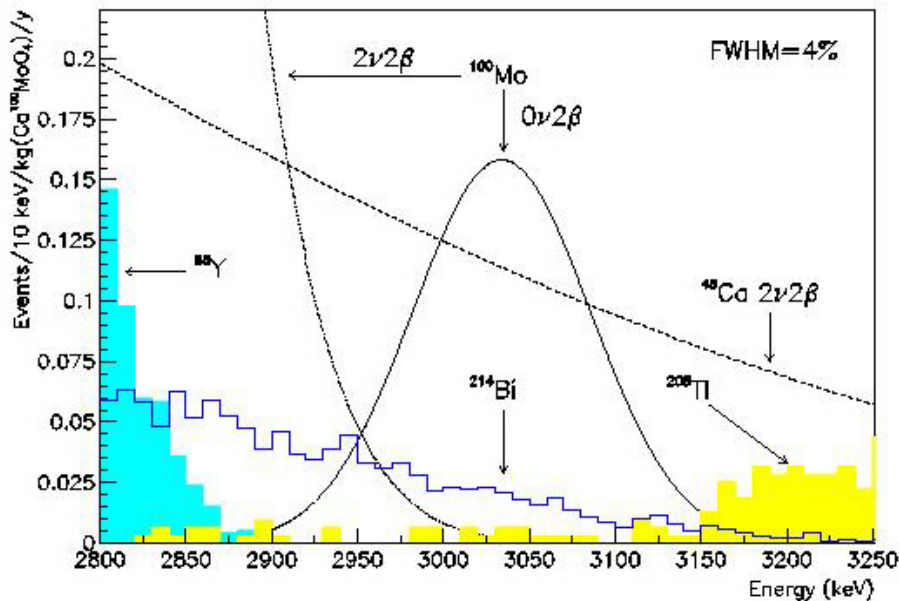
ICMSAI, Moscow, Russia

NIMA 584 (2008) 334

Source	Activity (mBq/kg)	
	CARAT	ICMSAI
²³²Th	< 0.7	< 1.5
²²⁸Th	0.2-0.4	0.04
²³⁸U	< 0.5	< 1.5
²²⁶Ra	2.1-2.5	0.13
²¹⁰Pb	< 400	< 17
²¹⁰Po	400-500	< 8
⁴⁰K	< 1 - <3	< 3
⁹⁰Sr	<60 - <180	< 23

measured in the Solotvina
Underground Lab

Sensitivity of a pilot experiment with ≈ 1 kg of $\text{Ca}^{100}\text{MoO}_4$



Main sources of background

- $2\nu 2\beta$ decay of ^{48}Ca ($T_{1/2}^{2\nu} = 4 \times 10^{19}$ yr)
- ^{208}Tl and ^{214}Bi (both with 0.1 mBq/kg)

A suppression of factor 10 for ^{214}Bi by pulse-shape analysis is supposed.

- ^{88}Y cosmogenic negligible
- Distributions for ^{100}Mo are shown for:
 - $T_{1/2}^{2\nu} = 7 \times 10^{18}$ yr
 - $T_{1/2}^{0\nu} = 10^{24}$ yr.

$T_{1/2}^{0\nu} > 4 \times 10^{23}$ yr at 90% CL over 1 years with 1 kg $\text{Ca}^{100}\text{MoO}_4$

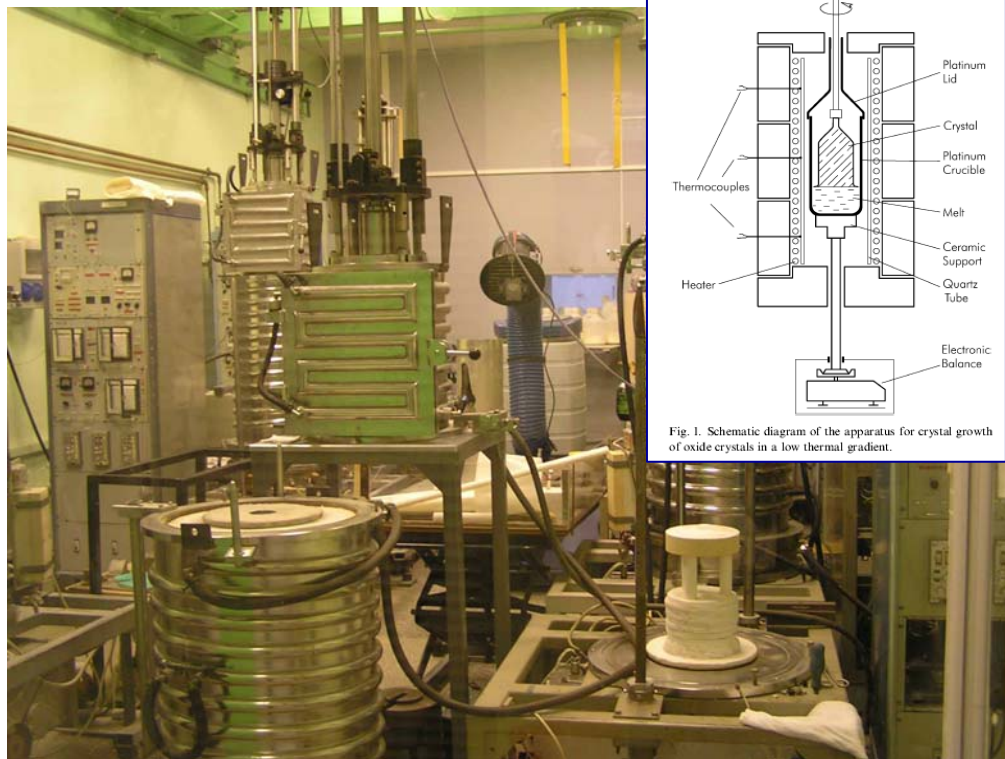
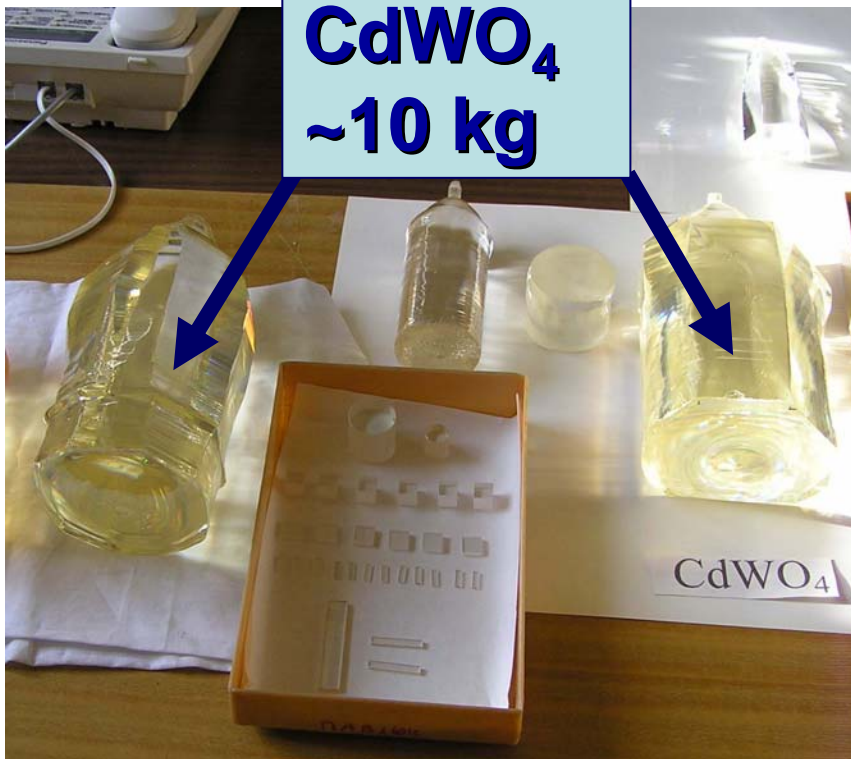
$T_{1/2}^{0\nu} > 2 \times 10^{24}$ yr 10 kg×yr $\text{Ca}^{100}\text{MoO}_4$

$T_{1/2}^{0\nu} \sim 10^{25}$ yr 200 kg×yr $\text{Ca}^{100}\text{MoO}_4$

$T_{1/2}^{0\nu} \sim 10^{26}$ yr 1000 kg×yr $\text{Ca}^{100}\text{MoO}_4$ as low temperature bolometer

Unique low-thermal-gradient Czochralski technology in the Nikolaev Institute of Inorganic Chemistry (Novosibirsk, Russia)

**CdWO_4
~10 kg**



R&D to produce ~1.2-1.8 kg of enriched $^{116}\text{CdWO}_4$ crystals is in progress in collaboration with ITEP (Moscow, Russia) and KIMS (Korea)

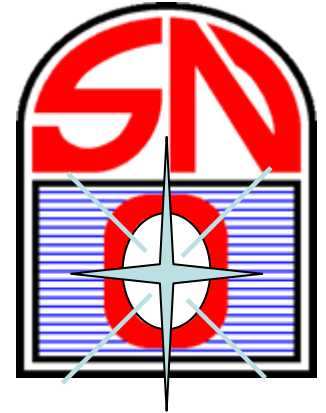
R&D of ≈ 0.15 kg $^{106}\text{CdWO}_4$ to search for 2β processes in ^{106}Cd

- Technology to purify Cd at the level of 0.1 ppm was developed
- $^{\text{nat}}\text{Cd}$ and ^{106}Cd were purified
- Technology to produce raw material for crystal growing was developed and CdWO_4 and $^{106}\text{CdWO}_4$ powders were produced (NeoChem company, Moscow, Russia)
- $^{\text{nat}}\text{CdWO}_4$ scintillator ~ 0.12 kg with 8% energy resolution for 662 keV ^{137}Cs was produced (ISMA, Kharkov, Ukraine)
- *The next step: $^{106}\text{CdWO}_4$ (NIIC, Novosibirsk, Russia)*

Minimization and careful control of ^{106}Cd losses at all the steps

collaboration with DAMA (Italy), JINR (Dubna, Russia) and crystal producers in Ukraine and Russia. Experiment in the LNGS (Italy).

Experimental Programs – IV



- SNO+ with Nd-loaded liquid scintillator
 - ...also called SNO++
- 0.1% Nd in 1000 tons of scintillator
 - with natural Nd corresponds to 56 kg of ^{150}Nd isotope
- sensitivity below 100 meV with natural Nd
- meters of ultra-low background self-shielding against gammas and neutrons
 - leads to well-defined background model
- liquid detector allows for additional *in-situ* purification
- possibility to enrich neodymium at French AVLIS facility

Why ^{150}Nd ?

- 3.37 MeV endpoint (2nd highest of all $\beta\beta$ isotopes)
 - above most backgrounds from natural radioactivity
- largest phase space factor of all $\beta\beta$ isotopes
 - factor of 33 greater compared with ^{76}Ge
 - for the same effective Majorana neutrino mass, the $0\nu\beta\beta$ rate in ^{150}Nd is the fastest
- cost of NdCl_3 is \$86,000 for 1 ton (not expensive)
- upcoming experiments use Ge, Xe, Te; we can deploy a large and comparable amount of Nd

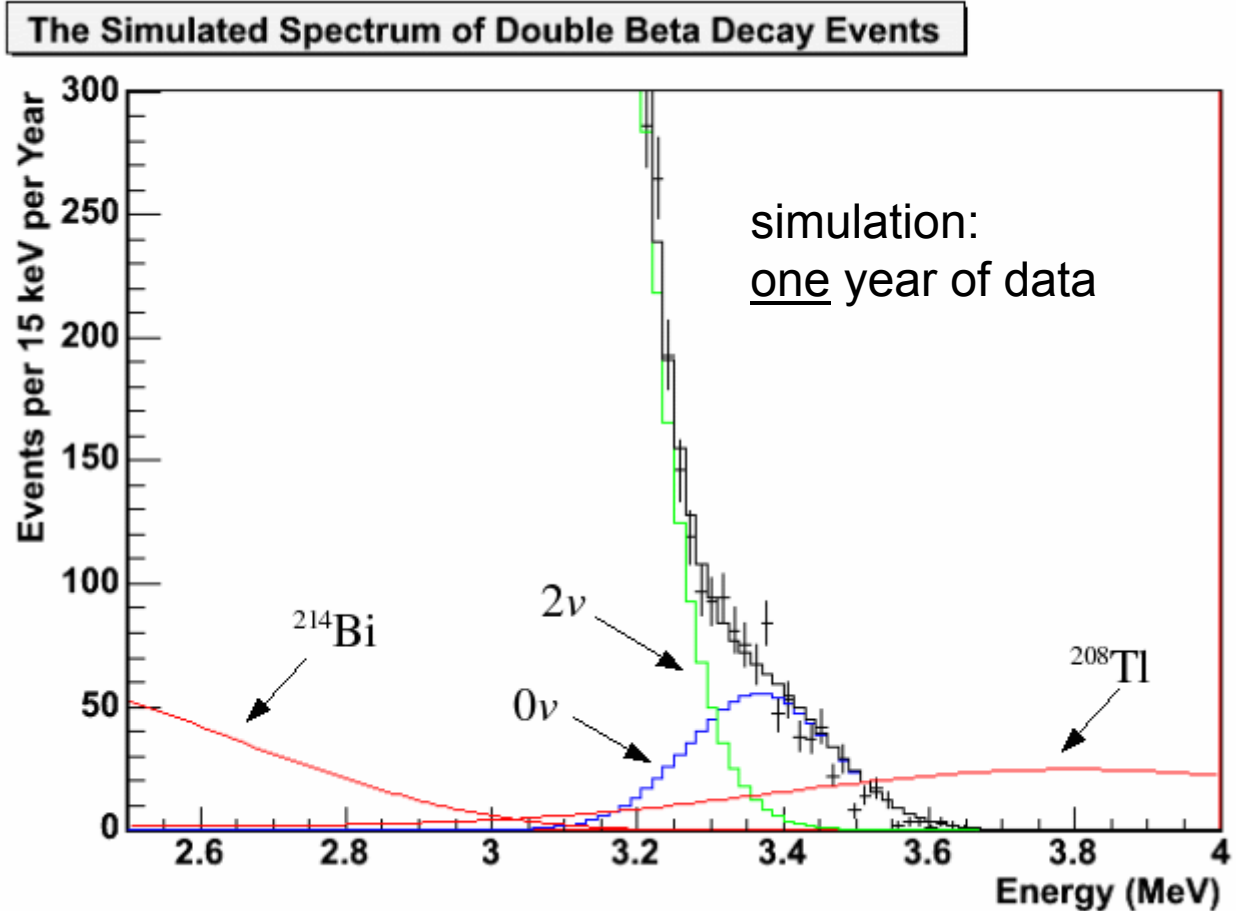
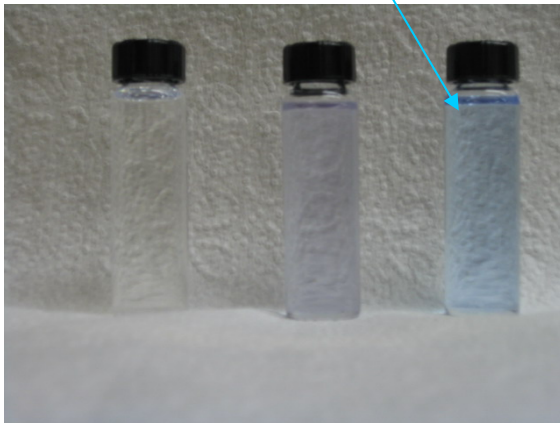
How Does ^{150}Nd Compare?

- 56 kg of ^{150}Nd is equivalent to:
- considering only the phase space factor
 - ~220 kg of ^{136}Xe
 - ~230 kg of ^{130}Te
 - ~950 kg of ^{76}Ge
- including QRPA matrix element calculations
 - ~1500 kg of ^{136}Xe
 - ~400 kg of ^{130}Te
 - ~570 kg of ^{76}Ge

thanks L. Simard and F. Piquemal

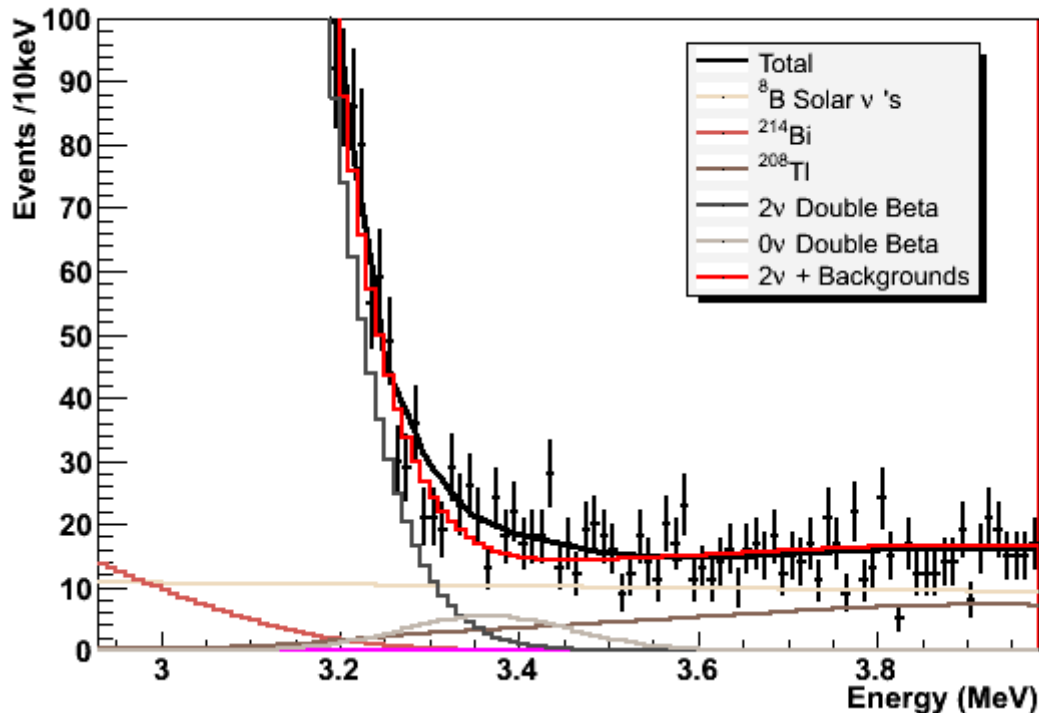
$0\nu\beta\beta$ Signal for $\langle m_\nu \rangle = 0.150$ eV

0ν : 1000 events per year with 1% natural Nd-loaded liquid scintillator in SNO++



56 kg of ^{150}Nd and $\langle m_\nu \rangle = 100 \text{ meV}$

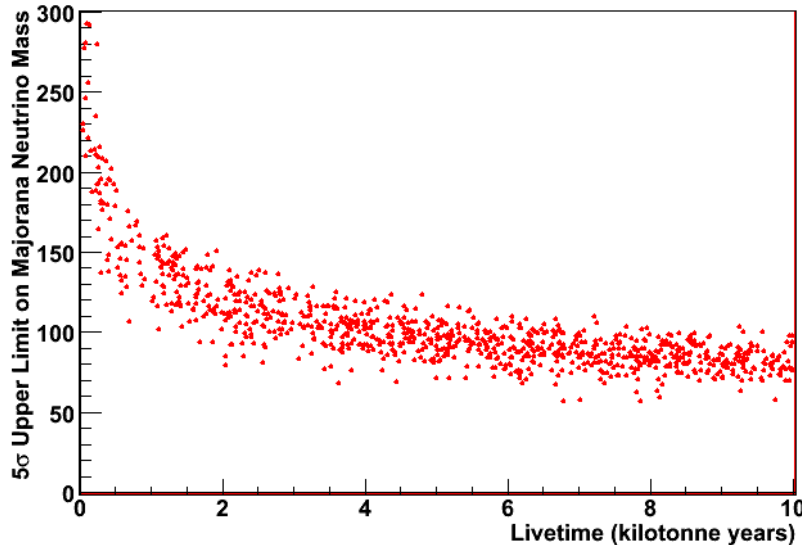
Simulated SNO+ Energy Spectrum



- 6.4% FWHM at Q-value
- 3 years livetime
- U, Th at Borexino levels
- 5σ sensitivity
- note: the dominant background is ^8B solar neutrinos!
- ^{214}Bi (from radon) is almost negligible
- ^{212}Po - ^{208}Tl tag (3 min) might be used to veto ^{208}Tl backgrounds; ^{212}Bi - ^{212}Po (300 ns) events constrain the amount of ^{208}Tl

Neutrino Mass Sensitivity

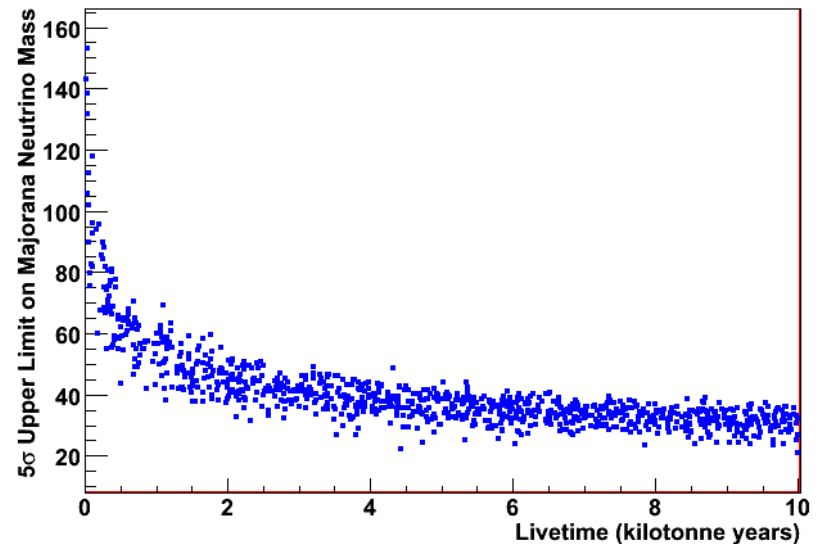
The D.B.D. Limit as a Function of Livetime



With natural Nd SNO+ is sensitive to effective neutrino masses as low as 100 meV.

With 10X enriched Nd our sensitivity extends to 40 meV.

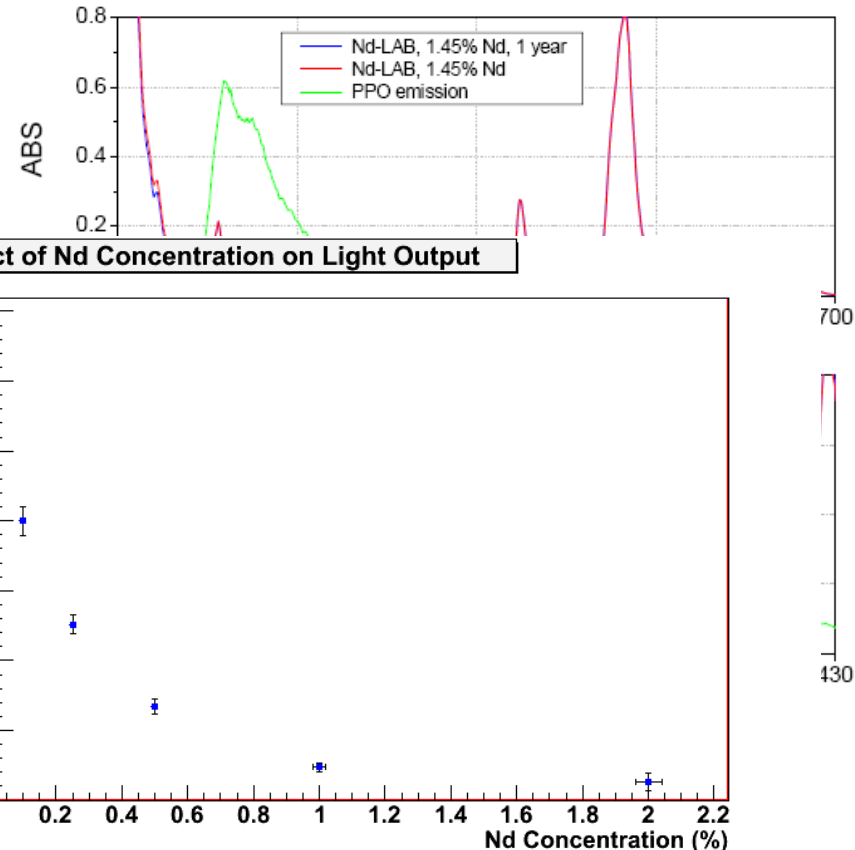
The D.B.D. Limit as a Function of Livetime



^{150}Nd Scintillator Properties

- stable Nd-loaded liquid scintillator optic
- scintillation optical properties studied
- target background levels achievable w

- NdCl_3 scintill
- purified BaSO_4
- using s demon
- remen are the organi



Status of SNO+

- funded by NSERC for final design/engineering and initial construction 2008-2010
- submission of full capital proposal to CFI in Q4 2008 with decision in Q2 2009
- construction of hold-down net begins in 2009
- construction of scintillator process and purification begins in 2010
- end of 2010 → ready for scintillator filling
- *new collaborators welcome!*

Summary

- future double beta decay experiments with scintillators utilize many different double beta decay candidate isotopes
 - it's very useful to search for neutrinoless double beta decay in many isotopes
- particularly interesting (IMHO) are experiments that will have a large mass of the high Q-value isotopes like ^{48}Ca and ^{150}Nd