

Tehran, 15-17 May '06

# Models of Neutrino Masses & Mixings

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Some recent work by our group

G.A., F. Feruglio, I. Masina, hep-ph/0402155,

G.A., F. Feruglio, hep-ph/0504165, hep-ph/0512103

G.A., R. Franceschini, hep-ph/0512202.

Reviews:

G.A., F. Feruglio, New J.Phys.6:106,2004 [hep-ph/0405048];

G.A., hep-ph/0410101; F. Feruglio, hep-ph/0410131

The current experimental situation is still unclear

- LSND: true or false? -> MiniBooNE soon will tell
- what is the absolute scale of  $\nu$  masses?
- no detection of  $0\nu\beta\beta$  (proof that  $\nu$ 's are Majorana)

.....

Different classes of models are still possible:

If LSND true

sterile  $\nu$ (s)??

CPT violat'n??

• "3-1" or "3-n"

$\nu_{\text{sterile}}$



$m^2 \sim 1-2 \text{ eV}^2$

If LSND false



3 light  $\nu$ 's are OK

We assume this case here

- Degenerate ( $m^2 \gg \Delta m^2$ )   $m^2 < o(1) \text{ eV}^2$

- Inverse hierarchy



- Normal hierarchy

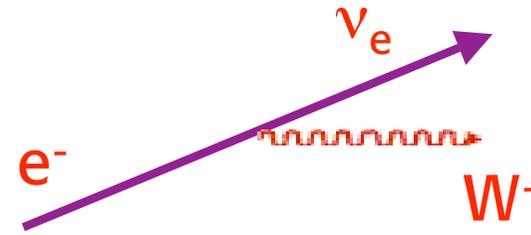


# 3-ν Models

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

flavour

mass



$$U = U_{\text{P-MNS}}$$

Pontecorvo

Maki, Nakagawa, Sakata

In basis where  $e^-, \mu^-, \tau^-$  are diagonal:  $\delta$ : CP violation

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \sim$$

$s$  = solar: large

$$\sim \begin{pmatrix} c_{13} & c_{12} & c_{13} & s_{12} & s_{13}e^{-i\delta} \\ \dots & \dots & \dots & \dots & c_{13} & s_{23} \\ \dots & \dots & \dots & \dots & c_{13} & c_{23} \end{pmatrix}$$

CHOOZ:  $|s_{13}| < \sim 0.2$

atm.:  $\sim \text{max}$



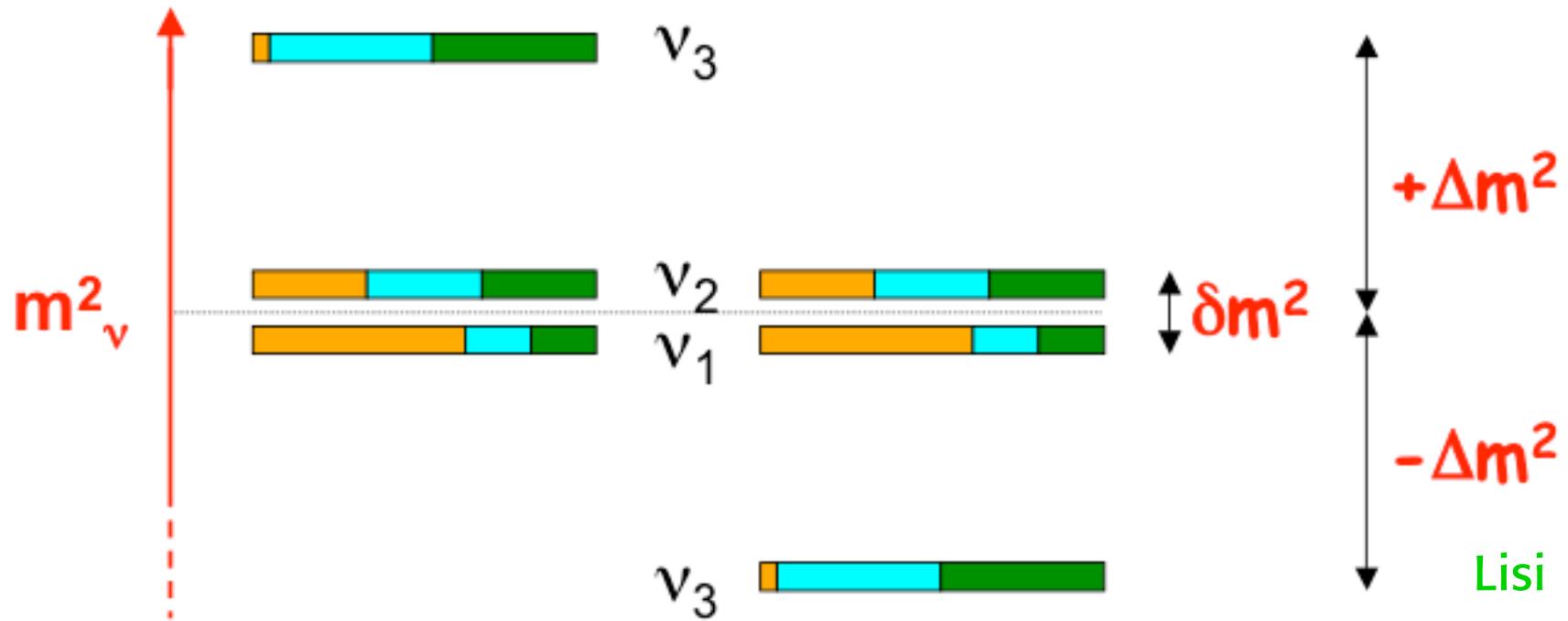
$$U \equiv \begin{pmatrix} c & -s & 0 \\ s & c & -1 \\ \frac{s}{\sqrt{2}} & \frac{c}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ \frac{s}{\sqrt{2}} & \frac{c}{\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

(some signs are conventional)

In general:  $U = U_e^+ U_\nu$



Abs. scale Normal hierarchy... OR... Inverted hierarchy mass<sup>2</sup> splittings



$\delta m^2 \sim 8 \times 10^{-5} \text{ eV}^2$   
 $\Delta m^2 \sim 3 \times 10^{-3} \text{ eV}^2$

$\sin^2 \theta_{12} \sim 0.3$   
 $\sin^2 \theta_{23} \sim 0.5$

$m_\nu < O(1) \text{ eV}$

$\sin^2 \theta_{13} < \text{few}\%$

sign( $\pm \Delta m^2$ ) unknown

$\delta$  (CP) unknown



$0\nu\beta\beta$  would prove that L is not conserved and  $\nu$ 's are Majorana  
 Also can tell degenerate, inverted or normal hierarchy

$$|m_{ee}| = c_{13}^2 [m_1 c_{12}^2 + e^{i\alpha} m_2 s_{12}^2] + m_3 e^{i\beta} s_{13}^2$$

Degenerate:  $\sim |m| |c_{12}^2 + e^{i\alpha} s_{12}^2| \sim |m| (0.3-1)$

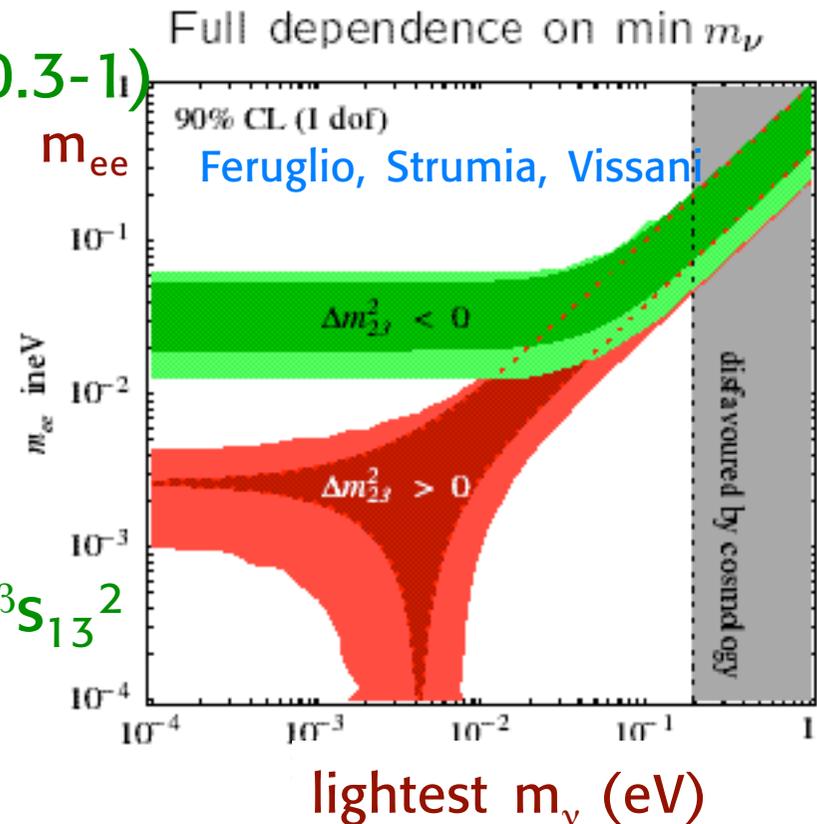
$$|m_{ee}| \sim |m| (0.3-1) < 0.23-1 \text{ eV}$$

IH:  $\sim (\Delta m_{\text{atm}}^2)^{1/2} |c_{12}^2 + e^{i\alpha} s_{12}^2|$

$$|m_{ee}| \sim (1.6-5) 10^{-2} \text{ eV}$$

NH:  $\sim (\Delta m_{\text{sol}}^2)^{1/2} s_{12}^2 + (\Delta m_{\text{atm}}^2)^{1/2} e^{i\beta} s_{13}^2$

$$|m_{ee}| \sim (\text{few}) 10^{-3} \text{ eV}$$

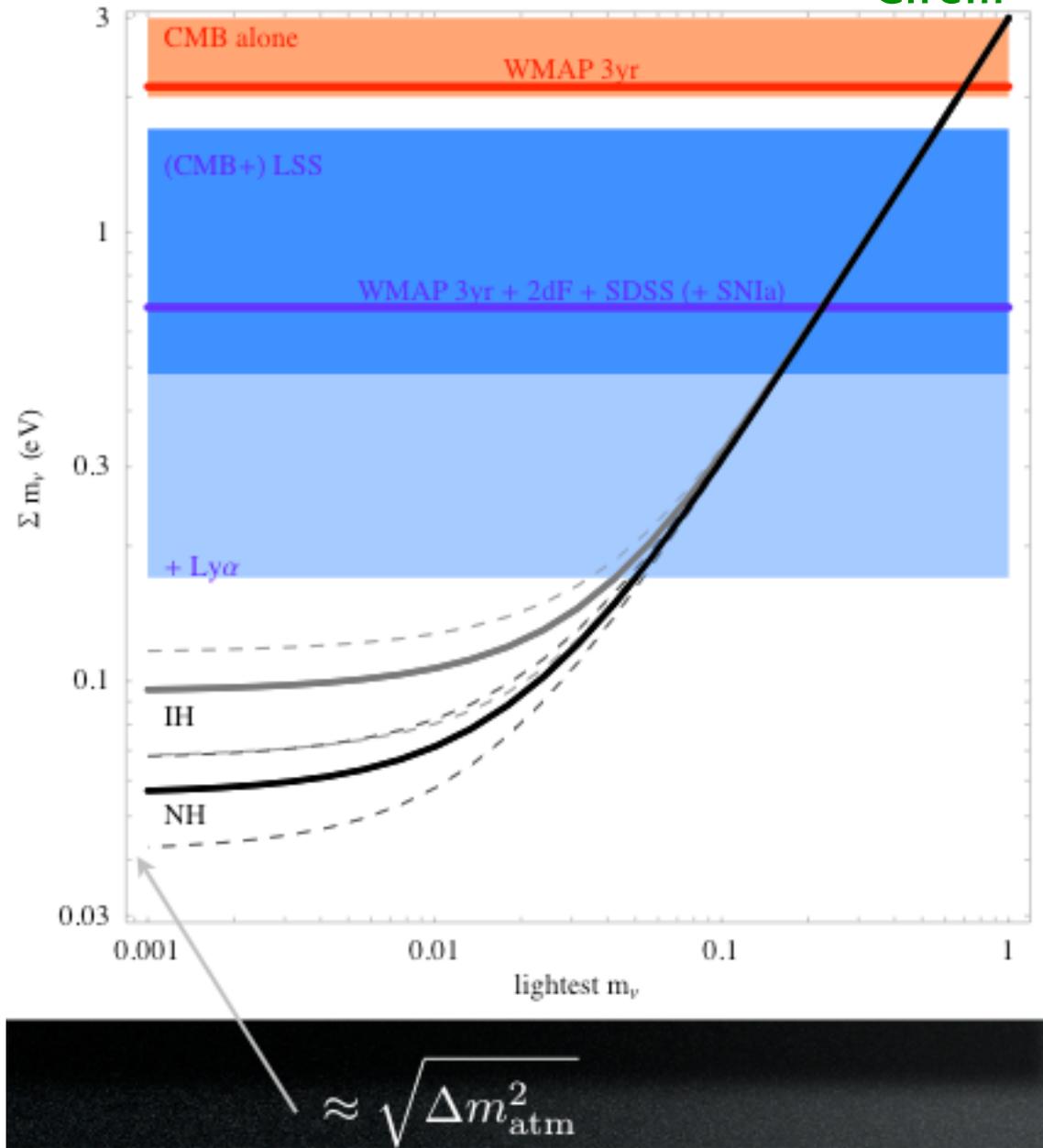


Present exp. limit:  $m_{ee} < 0.3-0.5 \text{ eV}$   
 (and a hint of signal????? Klapdor  
 Kleingrothaus)



# Summary

Cirelli



## Model building

## Quality factors for models:

- Based on the most general lagrangian compatible with some simple symmetry or dynamical principle
  - Should be complete: address at least charged leptons and neutrinos ( $U_{P-NMS} = U_e^+ U_\nu$ , and the gauge symmetry connects ch. leptons and LH neutrinos)
  - As many as possible small parameters (masses and mixings) should be naturally explained as a consequence.
  - The necessary vev configuration should be a minimum of the most general potential for a region of parameter space
  - Checked stability under radiative corrections and higher dim operators
    - Elegance, simplicity, economy of fields and parameters,
- ⊕ predictivity

# General remarks

- After KamLAND, SNO and WMAP.... not too much hierarchy is needed for  $\nu$  masses:

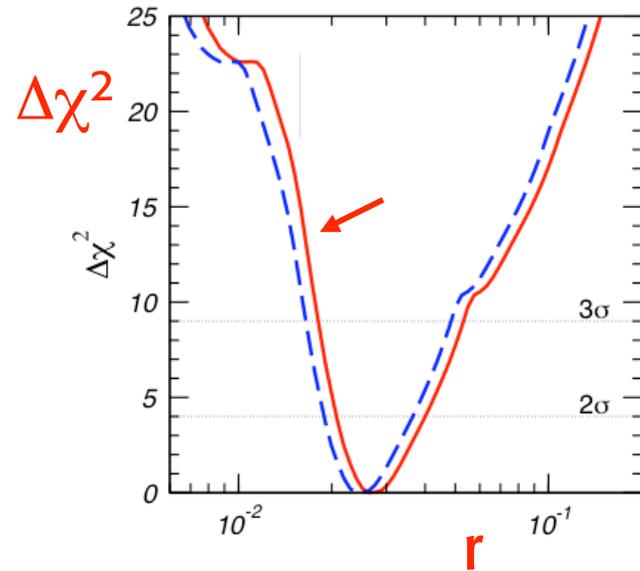
$$r \sim \Delta m^2_{\text{sol}} / \Delta m^2_{\text{atm}} \sim 1/35$$

Precisely at  $3\sigma$ :  $0.018 < r < 0.053$

or

$$m_{\text{heaviest}} < 0.2 - 0.7 \text{ eV}$$

$$m_{\text{next}} > \sim 8 \cdot 10^{-3} \text{ eV}$$



For a hierarchical spectrum:  $\frac{m_2}{m_3} \approx \sqrt{r} \approx 0.2$

Comparable to:  $\lambda_C \approx 0.22$  or  $\sqrt{\frac{m_\mu}{m_\tau}} \approx 0.24$

Suggests the same "hierarchy" parameters for  $q, l, \nu$



e.g.  $\theta_{13}$  not too small!

- Still large space for non maximal 23 mixing

$$3\text{-}\sigma \text{ interval } 0.31 < \sin^2\theta_{23} < 0.72$$

Maximal  $\theta_{23}$  theoretically hard

- $\theta_{13}$  not necessarily too small  
probably accessible to exp.

Very small  $\theta_{13}$  theoretically hard

Normal models:  $\theta_{23}$  large but not maximal,  
 $\theta_{13}$  not too small ( $\theta_{13}$  of order  $\lambda_C$  or  $\lambda_C^2$ )

Exceptional models:  $\theta_{23}$  maximal and/or  $\theta_{13}$  very small  
or: a special value for  $\theta_{12}$ ....



Natural models of the “normal” type are not too difficult to build up

It is reasonable to attribute hierarchies in masses and mixings to differences in some flavour quantum number(s).

A simplest flavour (or horizontal) symmetry is  $U(1)_F$

For example, some models based on see-saw and  $U(1)_F$  work for all quark and lepton masses and mixings, are natural and compatible with (SUSY) GUT's, e.g  $SU(5) \times U(1)_F$ .



# Hierarchy for masses and mixings via horizontal $U(1)_F$ charges.

Froggatt, Nielsen '79

**Principle:**

A generic mass term

$$\bar{R}_1 m_{12} L_2 H$$

is forbidden by  $U(1)$

if  $q_1 + q_2 + q_H$  not 0

$q_1, q_2, q_H$ :  
 $U(1)$  charges of  
 $\bar{R}_1, L_2, H$

$U(1)$  broken by vev of "flavon" field  $\theta$  with  $U(1)$  charge  $q_\theta = -1$ .  
If vev  $\theta = w$ , and  $w/M = \lambda$  we get for a generic interaction:

$$\bar{R}_1 m_{12} L_2 H \left(\frac{\theta}{M}\right)^{q_1 + q_2 + q_H} \quad m_{12} \rightarrow m_{12} \lambda^{q_1 + q_2 + q_H}$$

$\Delta_{\text{charge}}$  (indicated by a green arrow pointing to the exponent)

Hierarchy: More  $\Delta_{\text{charge}} \rightarrow$  more suppression ( $\lambda$  small)

One can have more flavons ( $\lambda, \lambda', \dots$ )  
with different charges ( $>0$  or  $<0$ ) etc  $\rightarrow$  many versions



## Degenerate $\nu$ 's

$$m^2 \gg \Delta m^2$$

- Limits on  $m_{ee}$  from  $0\nu\beta\beta$

→  $m_{ee} < 0.3-0.7 \text{ eV}$  (Exp)

$$m_{ee} = c_{13}^2 (m_1 c_{12}^2 + m_2 s_{12}^2) + s_{13}^2 m_3 \sim m_1 c_{12}^2 + m_2 s_{12}^2$$

are not very demanding: for  $\sin^2\theta \sim 0.3$   $\cos^2\theta - \sin^2\theta \sim 0.4$

and  $|m_1| \sim |m_2| \sim |m_3| \sim 1-2 \text{ eV}$  (with  $m_1 = -m_2$ ) would be perfectly fine

However, WMAP&LSS:  $|m| < 0.23 \text{ eV}$ , is very constraining

Only a moderate degeneracy is still allowed:

$$m / (\Delta m_{\text{atm}}^2)^{1/2} < 5, \quad m / (\Delta m_{\text{sol}}^2)^{1/2} < 30.$$

If so, constraints from  $0\nu\beta\beta$  are satisfied  
(both  $m_1 = \pm m_2$  allowed)



It is difficult to marry degenerate models with see-saw

$$m_\nu \sim m_D^T M^{-1} m_D$$

(needs a sort of conspiracy between  $M$  and  $m_D$ )

So most degenerate models deny all relation to  $m_D$  and directly work in the LL Majorana sector

For  $|m| < 0\nu\beta\beta$  bound we can have  $m_1 = m_2 = -m_3$  and

$$m_\nu \sim m \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{bmatrix}$$

equality of  $m_{11} = m_{23} = m_{32}$   
requires a symmetry  
eg  $O(3)$  Barbieri, Hall, Kane, Ross

$\theta_{23}$  maximal,  $\theta_{12}$  large,  $\theta_{13} = 0$   
in symmetric limit

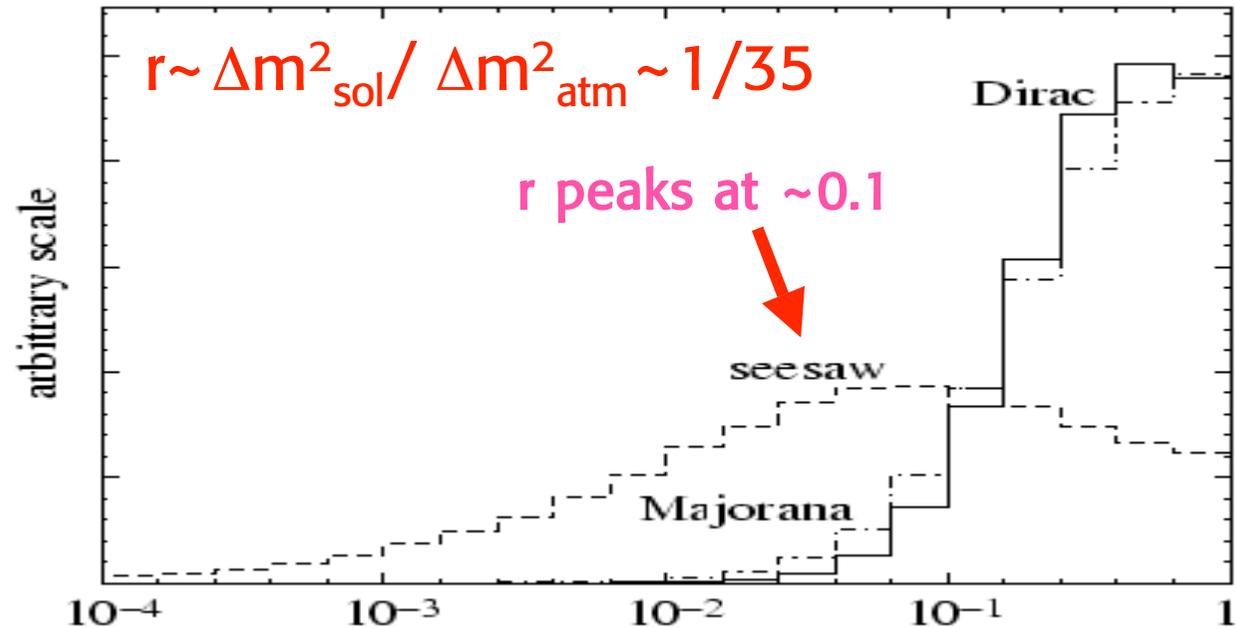


Anarchy (or accidental hierarchy):  
No structure in the leptonic sector

Hall, Murayama, Weiner

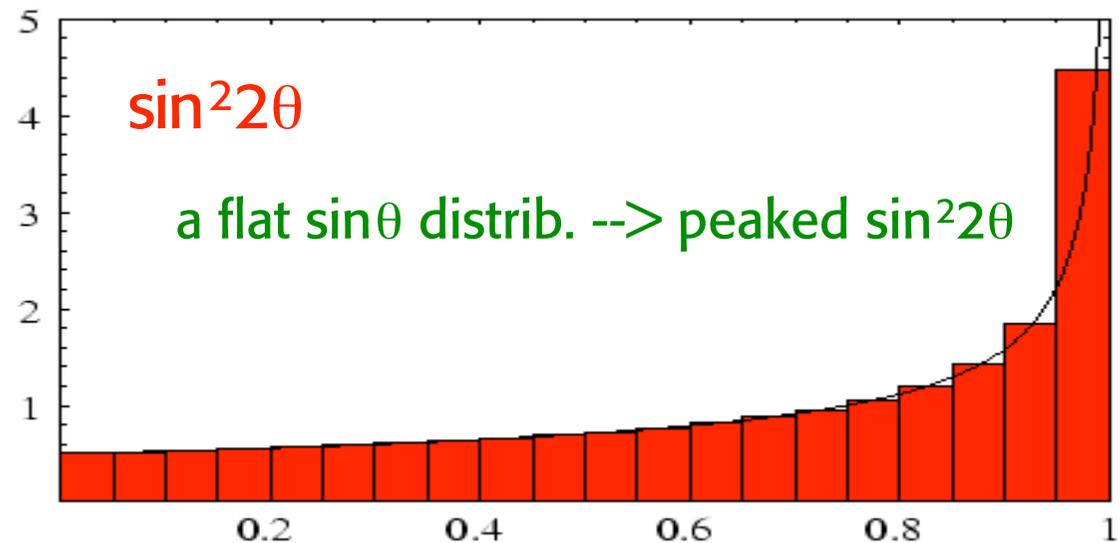
See-Saw:  
 $m_\nu \sim m^2/M$   
produces hierarchy  
from random  $m, M$

could fit the data



But: all mixing angles  
should be not too large,  
not too small →

Marginal: predicts  
 $\theta_{13}$  near bound



## Semianarchy: no structure in 23

Consider a matrix like  $m_\nu \sim \begin{pmatrix} \varepsilon^2 & \varepsilon & \varepsilon \\ \varepsilon & 1 & 1 \\ \varepsilon & 1 & 1 \end{pmatrix}$  **Note:**  $\theta_{13} \sim \varepsilon$   
 $\theta_{23} \sim 1$

with coeff.s of  $o(1)$  and  $\det 23 \sim o(1)$   
[ $\varepsilon \sim 1$  corresponds to anarchy]

After 23 and 13 rotations  $m_\nu \sim \begin{pmatrix} \varepsilon^2 & \varepsilon & 0 \\ \varepsilon & \eta & 0 \\ 0 & 0 & 1 \end{pmatrix}$

Normally two masses are of  $o(1)$  or  $r \sim 1$  and  $\theta_{12} \sim \varepsilon$   
But if, accidentally,  $\eta \sim \varepsilon$ , then  $r$  is small and  $\theta_{12}$  is large.

The advantage over anarchy is that  $\theta_{13}$  is small, but  $\theta_{12}$  large  
and the hierarchy  $m^2_3 \gg m^2_2$  are accidental

Ramond et al, Buchmuller et al



$q(\bar{5}) \sim (2, 0, 0)$  with no see-saw  $\rightarrow$  no structure in 23

Consider a matrix like  $m_\nu \sim L^T L \sim \begin{pmatrix} \lambda^4 & \lambda^2 & \lambda^2 \\ \lambda^2 & 1 & 1 \\ \lambda^2 & 1 & 1 \end{pmatrix}$  Note:  $\theta_{13} \sim \lambda^2$   
 $\theta_{23} \sim 1$

with coeff.s of  $o(1)$  and  $\det 23 \sim o(1)$

[semianarchy, while  $\lambda \sim 1$  corresponds to anarchy]

After 23 and 13 rotations  $m_\nu \sim \begin{pmatrix} \lambda^4 & \lambda^2 & 0 \\ \lambda^2 & \eta & 0 \\ 0 & 0 & 1 \end{pmatrix}$

Normally two masses are of  $o(1)$  or  $r \sim 1$  and  $\theta_{12} \sim \varepsilon$

But if, accidentally,  $\eta \sim \varepsilon$ , then  $r$  is small and  $\theta_{12}$  is large.

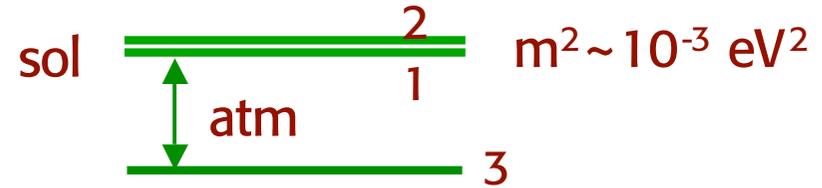
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Ramond et al, Buchmuller et al

⊕ With see-saw, one can do much better (see later)

# Inverted Hierarchy

Zee, Joshipura et al;  
 Mohapatra et al; Jarlskog et al;  
 Frampton, Glashow; Barbieri et al  
 Xing; Giunti, Tanimoto.....



An interesting model:

An exact  $U(1)_{L_e-L_\mu-L_\tau}$  symmetry for  $m_\nu$  predicts:  
 (a good 1<sup>st</sup> approximation)

$$m_\nu = U m_{\nu\text{diag}} U^T = m \begin{pmatrix} 0 & a & -b \\ a & 0 & 0 \\ -b & 0 & 0 \end{pmatrix} \quad \text{with} \quad m_{\nu\text{diag}} = \begin{pmatrix} m & 0 & 0 \\ 0 & -m & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

- $\theta_{13} = 0$
  - $\theta_{12} = \pi/4$
  - $\sin^2 \theta_{23} = b^2$
- $\theta_{\text{sun}}$  maximal!       $\theta_{\text{atm}}$  generic

Can arise from see-saw or dim-5  $L^T H H^T L$   
 • 1-2 degeneracy stable under rad. corr.'s



1<sup>st</sup> approximation

$$m_{\nu\text{diag}} = \begin{pmatrix} m & 0 & 0 \\ 0 & -m & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad m_{\nu} = U m_{\nu\text{diag}} U^T = m \begin{pmatrix} 0 & a & -b \\ a & 0 & 0 \\ -b & 0 & 0 \end{pmatrix}$$

- Data? This texture prefers  $\theta_{\text{sol}}$  closer to maximal than  $\theta_{\text{atm}}$

In fact: 12  $\rightarrow$   $\begin{pmatrix} 0 & a \\ a & 0 \end{pmatrix} \rightarrow$  Pseudodirac  $\theta_{12}$  maximal  $23 \rightarrow$   $\begin{pmatrix} 0 & 0 \\ 0 & 0 \end{pmatrix} \rightarrow \theta_{23} \sim o(1)$

With perturbations:  $\begin{pmatrix} 0 & a & -b \\ a & 0 & 0 \\ -b & 0 & 0 \end{pmatrix} \rightarrow \begin{pmatrix} \delta & 1 & 1 \\ 1 & \eta & \eta \\ 1 & \eta & \eta \end{pmatrix}$  (modulo  $o(1)$  coeff.s)

one gets  $1 - \tan^2 \theta_{12} \sim o(\delta + \eta) \sim (\Delta m_{\text{sol}}^2 / \Delta m_{\text{atm}}^2)$

Exp. ( $3\sigma$ ): 0.39-0.70 0.018-0.053

- In principle one can use the charged lepton mixing to go away from  $\theta_{12}$  maximal.  
In practice constraints from  $\theta_{13}$  small ( $\delta\theta_{12} \sim \theta_{13}$ )

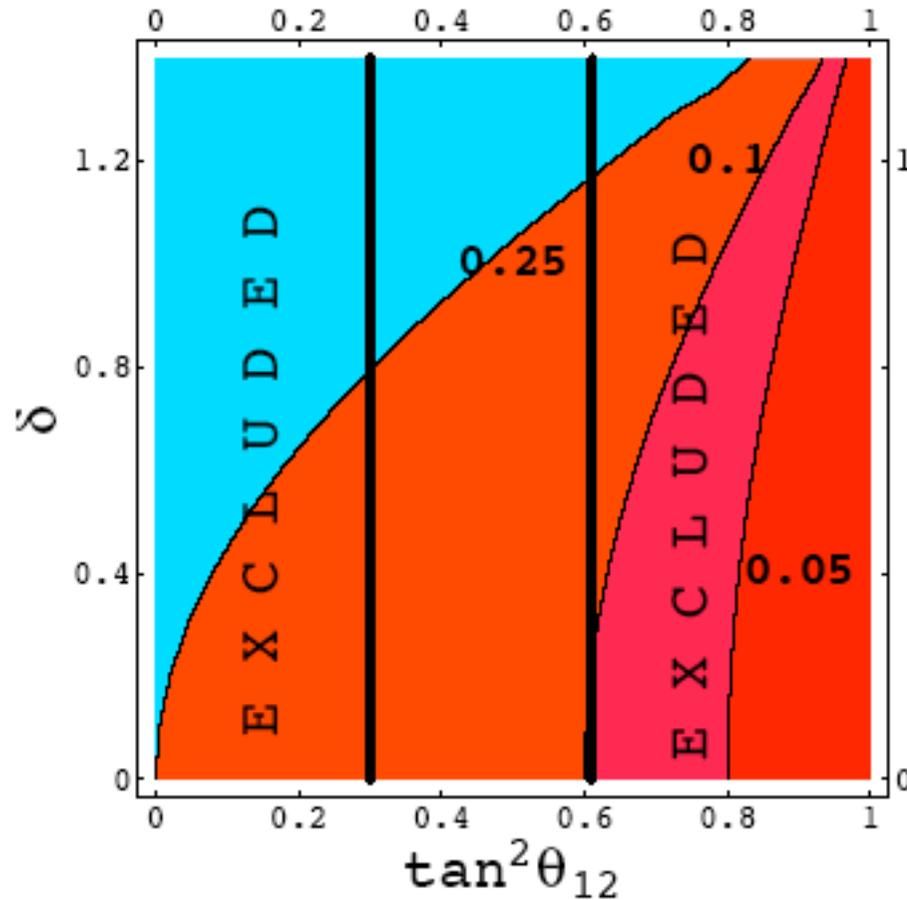


Frampton et al; GA, Feruglio, Masina '04



For the corrections from the charged lepton sector,  
 typically  $|\sin\theta_{13}| \sim (1 - \tan^2\theta_{12})/4\cos\delta \sim 0.15$

GA, Feruglio, Masina '04



$$\bar{U}_{12} = -\frac{e^{-i(\alpha_1+\alpha_2)}}{\sqrt{2}} + \frac{s_{12}^e e^{-i\alpha_2} + s_{13}^e e^{i\delta_e}}{2}$$

$$\bar{U}_{13} = \frac{s_{12}^e e^{-i\alpha_2} - s_{13}^e e^{i\delta_e}}{\sqrt{2}}$$

$$\bar{U}_{23} = -e^{-i\alpha_2} \frac{1 + s_{23}^e e^{i\alpha_2}}{\sqrt{2}}$$

Corr.'s from  $s_{12}^e, s_{13}^e$  to  $U_{12}$  and  $U_{13}$  are of first order  
 (2nd order to  $U_{23}$ )



**But:** a realistic model (eg  $\tan^2 2\theta_{12}$  large ) with IH,  $\theta_{13}$  small  
can be obtained from see-saw, if  $L_e-L_\mu-L_\tau$  is badly broken in  $M_{RR}$   
Grimus, Lavoura; G.A., Franceschini

As  $\nu_R$  are gauge singlets the large soft breaking in  $M_{RR}$  does  
not invade all other sectors when we do rad. corr's

By adding a small flavon breaking of  $U(1)_F$  symmetry with  
parameter  $\lambda \sim m_\mu/m_\tau$  the lepton spectrum is made natural  
and leads to  $\theta_{13} \sim m_\mu/m_\tau \sim 0.05$  or even smaller.



## $U(1)_F$ charges

$$l_i \sim L_e - L_\mu - L_\tau \sim (1, -1, -1)$$

$$l_{Ri} \sim (Q_e, Q_\mu, -1)$$

$$\nu_{Ri} \sim (-Q_R, Q_R, 0)$$

$\lambda$	$Q_e$	$Q_\mu$	$Q_\tau$
$0.25 \sim \xi^{\frac{1}{2}}$	7	-3	-1
$0.15 \sim \xi^{\frac{2}{3}}$	$\frac{11}{2}$	$-\frac{5}{2}$	-1
$0.06 \sim \xi$	4	-2	-1
$4 \times 10^{-3} \sim \xi^2$	$\frac{5}{2}$	$-\frac{3}{2}$	-1
$2 \times 10^{-4} \sim \xi^3$	2	$-\frac{4}{3}$	-1

## Charged lepton sector

$$m^l \sim \bar{l}_R \sim m_\tau \begin{pmatrix} \lambda^{-1+Q_e} & \lambda^{-1+Q_\mu} & \lambda^2 \\ \lambda^{1+Q_e} & \lambda^{1+Q_\mu} & 1 \\ \lambda^{1+Q_e} & \lambda^{1+Q_\mu} & 1 \end{pmatrix} \sim m_\tau \begin{pmatrix} \xi' & \xi\epsilon & \epsilon \\ \xi'\epsilon & \xi & 1 \\ \xi'\epsilon & \xi & 1 \end{pmatrix}$$

$$\epsilon \sim \lambda^2, \quad \xi \sim \lambda^{1+Q_\mu} \sim \frac{m_\mu}{m_\tau} \sim 6 \cdot 10^{-2}, \quad \xi' \sim \lambda^{-1+Q_e} \sim \frac{m_e}{m_\tau} \sim 3 \cdot 10^{-4}$$

## Diagonalisation

$$\begin{pmatrix} \xi' & \xi\epsilon & \epsilon \\ \xi'\epsilon & \xi & 1 \\ \xi'\epsilon & \xi & 1 \end{pmatrix} = U_l \begin{pmatrix} \xi' & 0 & 0 \\ 0 & \xi & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad U_l = R_{23}(\theta_l)R_{13}(\epsilon)R_{12}(\epsilon)$$



$\theta_l \rightarrow$  large shift to  $\theta_{23}$ ,  $O(\lambda^2)$  contrib'ns to  $\theta_{13}, \theta_{12}$

# Neutrino sector

$$\nu_{Ri} \sim (-Q_R, Q_R, 0) \quad Q_R=1$$

Dirac:  $m_\nu^D \sim \bar{\nu}_R l \sim m \begin{pmatrix} y_{11}\lambda^2 & a & b \\ d & y_{22}\lambda^2 & y_{23}\lambda^2 \\ y_{31}\lambda & y_{32}\lambda & y_{33}\lambda \end{pmatrix}$

a,b,d  
W,Z do not  
break U(1)

Majorana:

$$m_{RR}^{-1} \sim \frac{1}{M} \begin{pmatrix} x_{11}\lambda^2 & W & x_{13}\lambda \\ W & x_{22}\lambda^2 & x_{23}\lambda \\ x_{13}\lambda & x_{23}\lambda & Z \end{pmatrix}$$

no soft breaking

$$m_{RR}^{-1} \sim \frac{1}{M} \begin{pmatrix} A & W & B \\ W & C & D \\ B & D & Z \end{pmatrix}$$

with soft breaking

$$m_\nu = m_\nu^{DT} m_{RR}^{-1} m_\nu^D \sim m_{\nu 0} + \lambda m_{\nu 1} + \dots \sim$$

$$\sim \frac{m^2}{M} \begin{pmatrix} d^2 C & adW & bdW \\ adW & a^2 A & abA \\ bdW & abA & b^2 A \end{pmatrix} +$$

after see-saw

$$+ \lambda \frac{m^2}{M} \begin{pmatrix} 2y_{31}dD & y_{31}aB + y_{32}dD & y_{31}bB + y_{33}dD \\ y_{31}aB + y_{32}dD & 2y_{32}aB & y_{32}bB + y_{33}aB \\ y_{31}bB + y_{33}dD & y_{32}bB + y_{33}aB & 2y_{33}bB \end{pmatrix} + o(\lambda^2)$$



## Various stages:

exact U(1)  $r=0$

$$m_1 = -m_2; \quad m_3 = 0; \quad \theta_{13} = 0; \quad \theta_{12} = \frac{\pi}{4}; \quad \tan \theta_{23} = \frac{b}{a}$$



pure  $L_e-L_\mu-L_\tau$

only soft breaking ( $\lambda=0$ )

$$m_3 = 0; \quad m_1 + m_2 = \bar{C} + \bar{A}; \quad m_1 - m_2 = \sqrt{(\bar{A} - \bar{C})^2 + \bar{W}^2};$$

$$\theta_{13} = 0; \quad \tan \theta_{23} = \left| \frac{b}{a} \right|; \quad \tan^2 2\theta_{12} \sim \frac{\bar{W}^2}{(\bar{A} - \bar{C})^2}$$

with

$$\bar{A} = \frac{m^2}{M} A (a^2 + b^2); \quad \bar{C} = \frac{m^2}{M} C d^2; \quad \bar{W}^2 = \frac{m^2}{M} 4W^2 (a^2 + b^2) d^2$$

$$\theta_{12} \rightarrow \frac{|\bar{A} - \bar{C}|}{|\bar{W}|} \sim 0.40.$$



requires large soft breaking

$$r \sim 1/30 \rightarrow \left| \frac{\bar{A} + \bar{C}}{\bar{A} - \bar{C}} \right| \cos \delta \sim 0.02$$

requires some fine tuning

## Summarising this model with IH

In the limit of exact  $U(1)_F$   $\theta_{12} = \pi/4$  and  $r$ ,  $\theta_{13}$ , as well as  $m_e/m_\tau$  and  $m_\mu/m_\tau$  (for our choice of charges) are all zero.

In general a small symmetry breaking will make them different from zero but small. And  $\theta_{12}$  will only be slightly displaced from  $\pi/4$  (bad)

A large soft explicit mixing in the  $M_{RR}$  sector can decouple  $\theta_{12}$ , which gets a large shift, from  $\theta_{13}$ ,  $m_e/m_\tau$  and  $m_\mu/m_\tau$  which remain small.

The only remaining imperfection is that a moderate fine tuning is needed for  $r$ .



## Normal hierarchy

- A crucial point: in the 2-3 sector we need both large  $m_3 - m_2$  splitting and large mixing.

$$m_3 \sim (\Delta m_{\text{atm}}^2)^{1/2} \sim 5 \cdot 10^{-2} \text{ eV}$$

$$m_2 \sim (\Delta m_{\text{sol}}^2)^{1/2} \sim 8 \cdot 10^{-3} \text{ eV}$$

- The "theorem" that large  $\Delta m_{32}$  implies small mixing (pert. th.:  $\theta_{ij} \sim 1/|E_i - E_j|$ ) is not true in general: all we need is  $(\text{sub})\det[23] \sim 0$

- Example:  $m_{23} \sim \begin{bmatrix} x^2 & x \\ x & 1 \end{bmatrix}$

Det = 0; Eigenvl's: 0,  $1+x^2$   
Mixing:  $\sin^2 2\theta = 4x^2/(1+x^2)^2$

So all we need are natural mechanisms for  $\det[23]=0$

For  $x \sim 1$   
large splitting  
and large mixing!



## Examples of mechanisms for $\text{Det}[23] \sim 0$

based on see-saw:  $m_\nu \sim m_D^T M^{-1} m_D$

1) A  $\nu_R$  is lightest and coupled to  $\mu$  and  $\tau$

King; Allanach; Barbieri et al.....

$$M \sim \begin{bmatrix} \varepsilon & 0 \\ 0 & 1 \end{bmatrix} \longrightarrow M^{-1} \sim \begin{bmatrix} 1/\varepsilon & 0 \\ 0 & 1 \end{bmatrix} \approx \begin{bmatrix} 1/\varepsilon & 0 \\ 0 & 0 \end{bmatrix}$$

$$m_\nu \sim \begin{bmatrix} a & b \\ c & d \end{bmatrix} \begin{bmatrix} 1/\varepsilon & 0 \\ 0 & 0 \end{bmatrix} \begin{bmatrix} a & c \\ b & d \end{bmatrix} \approx 1/\varepsilon \begin{bmatrix} a^2 & ac \\ ac & c^2 \end{bmatrix}$$

2)  $M$  generic but  $m_D$  "lopsided"

$$m_D \sim \begin{bmatrix} 0 & 0 \\ x & 1 \end{bmatrix}$$

Albright, Barr; GA, Feruglio, .....

$$m_\nu \sim \begin{bmatrix} 0 & x \\ 0 & 1 \end{bmatrix} \begin{bmatrix} a & b \\ b & c \end{bmatrix} \begin{bmatrix} 0 & 0 \\ x & 1 \end{bmatrix} = c \begin{bmatrix} x^2 & x \\ x & 1 \end{bmatrix}$$



## An important property of SU(5)

Left-handed quarks have small mixings ( $V_{CKM}$ ),  
but right-handed quarks can have large mixings (unknown).

In SU(5):  
LH for d quarks  $\longleftrightarrow$  RH for l- leptons

$$\begin{array}{l}
 \bar{5} \quad \swarrow \quad \nwarrow \quad 10 \\
 m_d \sim \bar{d}_R d_L \\
 \\
 10 \quad \swarrow \quad \nwarrow \quad \bar{5} \\
 m_e \sim \bar{e}_R e_L
 \end{array}
 \quad
 \begin{array}{l}
 \bar{5} : (\underbrace{\bar{d}, \bar{d}, \bar{d}}_R, \underbrace{\nu, e^-}_L) \\
 \\
 m_d = m_e^T
 \end{array}$$

cannot be exact, but approx.

Most "lopsided" models are based on this fact. In these models large atmospheric mixing arises (at least in part) from the charged lepton sector.



- The correct pattern of masses and mixings, also including  $\nu$ 's, is obtained in simple models based on

$SU(5) \times U(1)_{\text{flavour}}$

Ramond et al; GA, Feruglio+Masina; Buchmuller et al;  
King et al; Yanagida et al, Berezhiani et al; Lola et al.....

Offers a simple description of hierarchies, but it is not very predictive (large number of undetermined  $o(1)$  parameters)

- $SO(10)$  models could be more predictive, as are non abelian flavour symmetries, eg  $O(3)$

Albright, Barr; Babu et al; Buccella et al; Barbieri et al;  
Raby et al; King, Ross



With suitable charge assignments all relevant patterns can be obtained

Recall:  $u \sim 10 \ 10$   
 $d=e^T \sim \bar{5} \ 10$   
 $\nu_D \sim \bar{5} \ 1; M_{RR} \sim 1 \ 1$

No structure for leptons

No automatic  $\det 23 = 0$

Automatic  $\det 23 = 0$

1st fam.  $\Psi_{10}: (5, 3, 0)$   
 2nd  $\Psi_5: (2, 0, 0)$   
 3rd  $\Psi_1: (1, -1, 0)$

Equal 2,3 ch. for lopsided

Model	$\Psi_{10}$	$\Psi_{\bar{5}}$	$\Psi_1$	$(H_u, H_d)$
Anarchical (A)	(3,2,0)	(0,0,0)	(0,0,0)	(0,0)
Semi-Anarchical (SA)	(2,1,0)	(1,0,0)	(2,1,0)	(0,0)
Hierarchical ( $H_I$ )	(6,4,0)	(2,0,0)	(1,-1,0)	(0,0)
Hierarchical ( $H_{II}$ )	(5,3,0)	(2,0,0)	(1,-1,0)	(0,0)
Inversely Hierarchical ( $IH_I$ )	(3,2,0)	(1,-1,-1)	(-1,+1,0)	(0,+1)
Inversely Hierarchical ( $IH_{II}$ )	(6,4,0)	(1,-1,-1)	(-1,+1,0)	(0,+1)

all charges positive

not all charges positive



All entries are a given power of  $\lambda$  times a free  $o(1)$  coefficient

$$m_u \sim v_u \begin{bmatrix} \lambda^{10} & \lambda^8 & \lambda^5 \\ \lambda^8 & \lambda^6 & \lambda^3 \\ \lambda^5 & \lambda^3 & 1 \end{bmatrix}$$

In a statistical approach we generate these coeff.s as random complex numbers  $\rho e^{i\phi}$  with  $\phi = [0, 2\pi]$  and  $\rho = [0.5, 2]$  (default) or  $[0.8, 1.2]$ , or  $[0.95, 1.05]$  or  $[0, 1]$  (real numbers also considered for comparison)

For each model we evaluate the success rate (over many trials) for falling in the exp. allowed window:

(boundaries  $\sim 3\sigma$  limits)

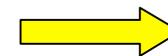
Maltoni et al, hep-ph/0309130

$$r \sim \Delta m^2_{\text{sol}} / \Delta m^2_{\text{atm}}$$



$$\begin{aligned} 0.018 < r < 0.053 \\ |U_{e3}| < 0.23 \\ 0.30 < \tan^2 \theta_{12} < 0.64 \\ 0.45 < \tan^2 \theta_{23} < 2.57 \end{aligned}$$

for each model the  $\lambda, \lambda'$  values are optimised



The optimised values of  $\lambda$  are of the order of  $\lambda_C$  or a bit larger (moderate hierarchy)

model	$\lambda(= \lambda')$
$A_{SS}$	0.2
$SA_{SS}$	0.25
$H_{(SS,II)}$	0.35
$H_{(SS,I)}$	0.45
$IH_{(SS,II)}$	0.45
$IH_{(SS,I)}$	0.25



## Example: Normal Hierarchy

G.A., Feruglio, Masina

Note: not all charges positive  
 $\rightarrow$  det23 suppression

1st fam.      2nd      3rd

$$\begin{aligned} q(10): & (5, 3, 0) \\ q(\bar{5}): & (2, 0, 0) \\ q(1): & (1, -1, 0) \end{aligned}$$

$$\begin{aligned} q(H) &= 0, \quad q(\bar{H}) = 0 \\ q(\theta) &= -1, \quad q(\theta') = +1 \end{aligned}$$

In first approx., with  $\langle \theta \rangle / M \sim \lambda \sim \lambda' \sim 0.35 \sim o(\lambda_C)$

$10_i 10_j$

$$m_u \sim v_u \begin{pmatrix} \lambda^{10} & \lambda^8 & \lambda^5 \\ \lambda^8 & \lambda^6 & \lambda^3 \\ \lambda^5 & \lambda^3 & 1 \end{pmatrix},$$

$10_i \bar{5}_j$

$$m_d = m_e^T \sim v_d \begin{pmatrix} \lambda^7 & \lambda^5 & \lambda^5 \\ \lambda^5 & \lambda^3 & \lambda^3 \\ \lambda^2 & 1 & 1 \end{pmatrix}$$

"lopsided"

$\bar{5}_i 1_j$

$$m_{\nu D} \sim v_u \begin{pmatrix} \lambda^3 & \lambda & \lambda^2 \\ \lambda & \lambda' & 1 \\ \lambda & \lambda' & 1 \end{pmatrix},$$

$1_i 1_j$

$$M_{RR} \sim M \begin{pmatrix} \lambda^2 & 1 & \lambda \\ 1 & \lambda'^2 & \lambda' \\ \lambda & \lambda' & 1 \end{pmatrix}$$

Note: coeffs. 0(1) omitted, only orders of magnitude predicted



$$\bar{5}_{i1_j} \swarrow \quad \mathbf{m}_{\nu D} \sim v_u \begin{bmatrix} \lambda^3 & \lambda & \lambda^2 \\ \lambda & \lambda & 1 \\ \lambda & \lambda & 1 \end{bmatrix}, \quad \mathbf{1}_{i1_j} \swarrow \quad \mathbf{M}_{RR} \sim M \begin{bmatrix} \lambda^2 & 1 & \lambda \\ 1 & \lambda^2 & \lambda \\ \lambda & \lambda & 1 \end{bmatrix}$$

see-saw  $\mathbf{m}_\nu \sim \mathbf{m}_{\nu D}^T \mathbf{M}_{RR}^{-1} \mathbf{m}_{\nu D}$

$$\mathbf{m}_\nu \sim v_u^2/M \begin{bmatrix} \lambda^4 & \lambda^2 & \lambda^2 \\ \lambda^2 & \boxed{1} & \boxed{1} \\ \lambda^2 & \boxed{1} & \boxed{1} \end{bmatrix},$$

$$\det_{23} \sim \lambda^2$$

The 23 subdeterminant is automatically suppressed,  
 $\theta_{13} \sim \lambda^2, \theta_{12}, \theta_{23} \sim 1$

This model works, in the sense that all small parameters are due to various degrees of suppression.

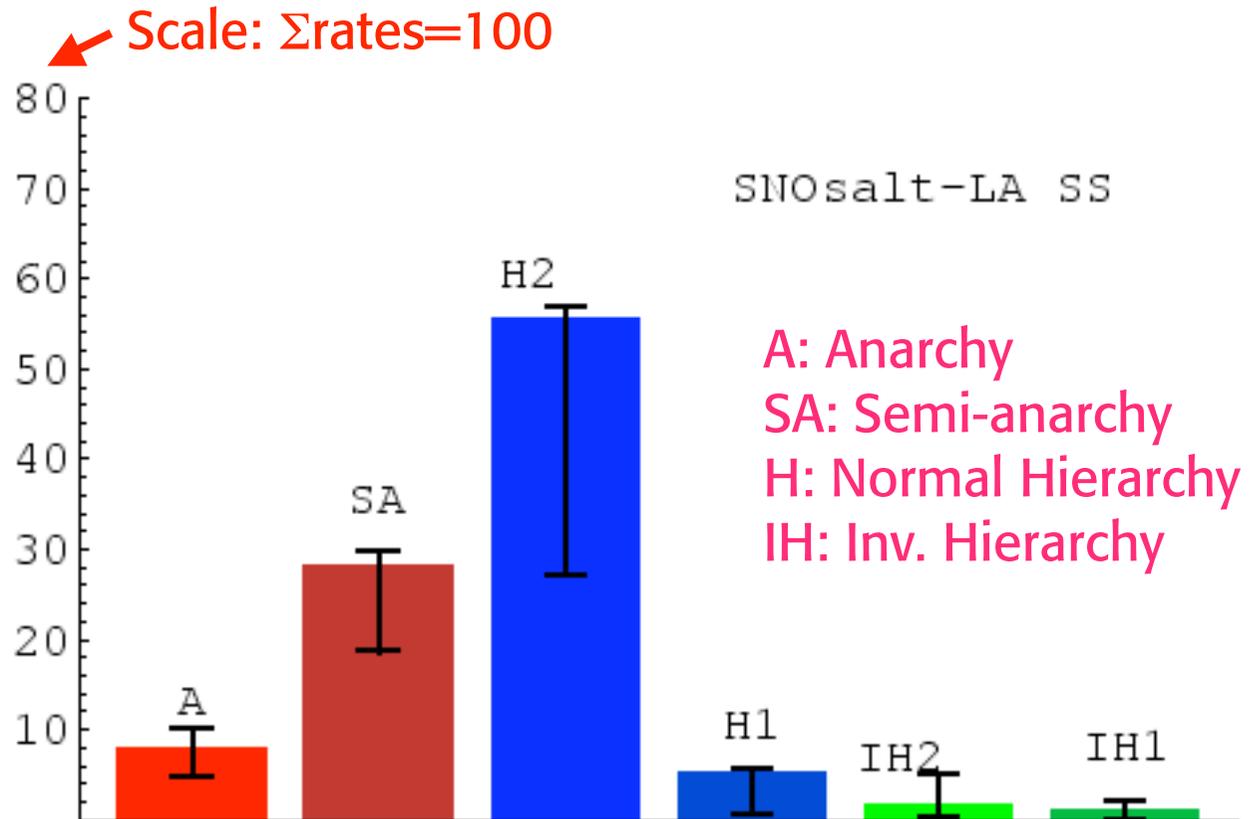


But too many free parameters!!

# Results with see-saw dominance (updated in Nov. '03):

1 or 2 refer to models with 1 or 2 flavons of opposite ch.

With charges of both signs and 1 flavon some entries are zero



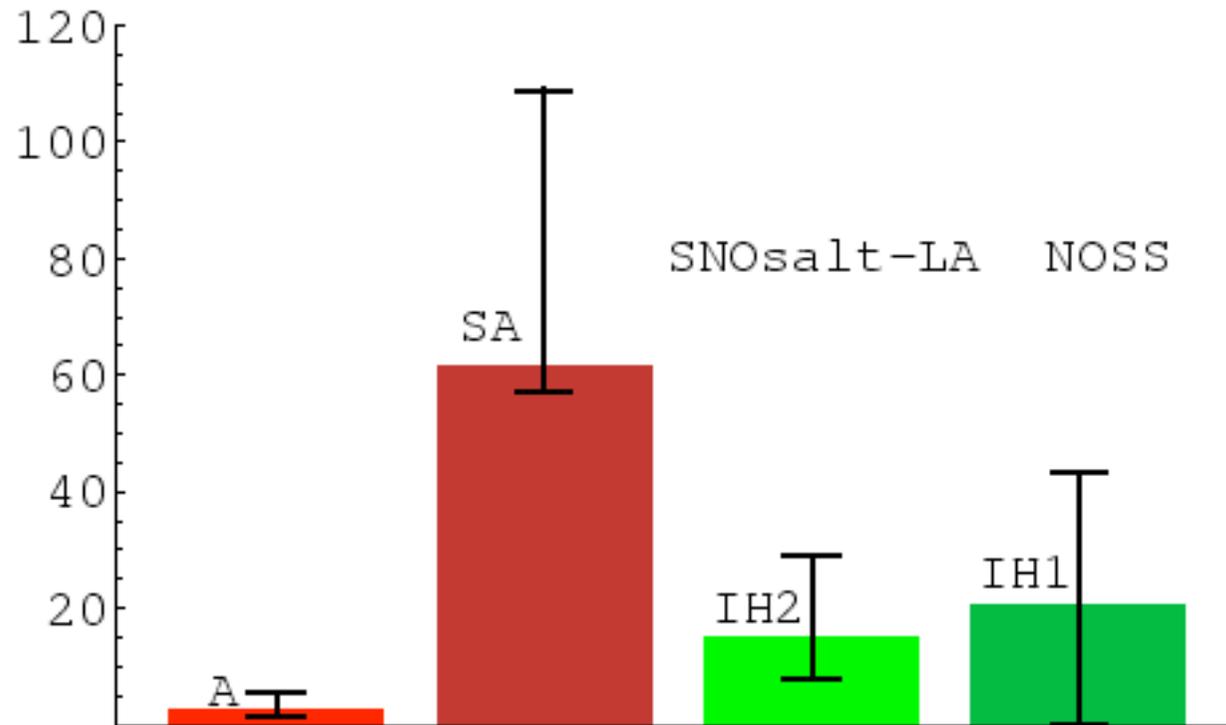
Errors are linear comb. of stat. and syst. errors (varying the extraction procedure: interval of  $\rho$ , real or complex)

H2 is better than SA, better than A, better than IH



With no see-saw ( $m_\nu$  generated directly from  $L^T m_\nu L \sim \bar{5} \bar{5}$ ) IH is better than A

[With no-see-saw H coincide with SA]



Note: we always include the effect of diagonalising charged leptons

