### ILC Detector Requirements for Smuon Analysis

Joseph Proulx

University of Colorado at Boulder

### Outline

- The Colorado SUSY group
- SUSY parameter points and masses
- Simulation tools
- Smuon studies
- Detector requirements

Many undergraduates contribute to SUSY and calorimetry studies at CU with funding fron NSF, UROP, and other sources. These students include:

Chris GeraciPaul SteinbrecherElliot SmithMaria GuldaKyle MillerJack GillKeith DrakeJesse SmockMatthew PhillipsSarah MollSarah Moll

http://hep-www.colorado.edu/SUSY
http://hep-www.colorado.edu/~nlc

Our task:

 Attempt reconstruction of SUSY masses using fast MC detector simulation

Our task:

- Attempt reconstruction of SUSY masses using fast MC detector simulation
- Determine whether measurements are attainable

Our task:

- Attempt reconstruction of SUSY masses using fast MC detector simulation
- Determine whether measurements are attainable
- Determine whether changes in detector or accelerator design might help with measurements

Nine separate parameter points chosen for studies: SPS1-9

- 6 mSUGRA parameter points
- 2 GMSB points
- 1 AMSB point

Nine separate parameter points chosen for studies: SPS1-9

- 6 mSUGRA parameter points
- 2 GMSB points
- 1 AMSB point

Similarities include Higgs mass of about 115GeV Differences include variations in relative masses of SUSY particles representative of various regions of parameter space

Nine separate parameter points chosen for studies: SPS1-9

- 6 mSUGRA parameter points
- 2 GMSB points
- 1 AMSB point

Similarities include Higgs mass of about 115GeV Differences include variations in relative masses of SUSY particles representative of various regions of parameter space

We have focused on SPS1 and several other mSUGRA points

	mSUGRA								
	SPS1	SPS2	SPS3	SPS4	$\mathbf{SPS5}$	SPS6			
$M_0$	100	1450	90	400	150	150			
$M_{1/2}$	250	<b>300</b>	400	<b>300</b>	<b>300</b>	300			
$A_0$	-100	0	0	0	-1000	0			
$tan(\beta)$	10	10	10	50	5	10			
$\mu$	352.39	124.77	508.59	377.03	639.80	393.09			

	GN	/ISB		AMSB
	SPS7	SPS8		SPS9
Λ	40,000	100,000	$M_0$	400
$M_{mes}$	80,000	200,000	$M_{3/2}$	60000
$N_5$	3	1	,	
$\tan(\beta)$	15	15	$tan(\beta)$	10
$\mu$	300.03	398.31	$\mu$	869.90

### **The SUSY Masses**

minimal SUGRA						GMSB		AMSB	
	$\mathbf{SPS1}$	SPS2	SPS3	SPS4	$\mathbf{SPS5}$	SPS6	SPS7	SPS8	SPS9
$\widetilde{\chi}_1^0$	96.05	79.54	160.55	118.66	119.51	117.50	161.65	137.19	175.51
$\widetilde{\chi}^0_2$	176.82	135.34	296.95	218.14	226.33	215.54	260.06	252.33	549.03
$\widetilde{\chi}^0_3$	358.81	140.84	512.87	383.91	642.83	398.70	306.26	404.00	874.37
$\widetilde{\chi}_4^0$	377.81	269.45	529.57	401.08	652.95	418.06	379.94	426.28	875.97
$\widetilde{\chi}_1^+$	176.38	104.03	296.85	218.06	226.33	215.20	256.33	252.03	175.67
$\widetilde{\chi}_2^+$	378.23	269.03	529.51	402.28	652.68	418.19	379.45	426.47	877.22
$h^0$	113.97	115.71	116.95	115.39	119.79	114.71	113.57	114.83	114.83
$H^0$	394.15	1444.10	573.03	404.63	694.03	457.84	378.37	515.01	912.56
$A^0$	393.63	1442.95	572.42	404.43	693.86	457.26	377.89	514.49	911.74
$H^+$	401.77	1446.18	578.30	416.28	698.49	464.40	386.70	521.17	915.83
$\widetilde{ u}_e$	186.00	1454.17	275.99	441.22	244.52	243.25	249.06	347.61	309.71
$\widetilde{e}_R^-$	142.97	1451.69	178.33	416.54	191.45	191.30	127.43	175.87	303.01
$\widetilde{e}_L$	202.14	1456.33	287.11	448.40	256.30	255.81	261.47	356.61	319.66
$\widetilde{\tau}_1^-$	133.22	1439.46	170.59	267.61	180.67	184.34	120.45	169.42	271.28
$\widetilde{\tau}_2^-$	206.13	1450.38	289.22	414.91	257.86	258.31	263.40	357.59	322.54
$\widetilde{t_1}$	379.11	1003.88	623.83	530.58	220.74	474.12	779.09	957.65	1005.17
$\widetilde{t_2}$	574.71	1307.41	819.54	695.88	644.65	659.73	863.00	1058.68	1128.80
$\widetilde{b_1}$	491.91	1296.56	757.50	606.86	535.86	589.80	822.17	1021.90	1112.07
$\widetilde{b_2}$	524.59	1520.09	791.35	706.45	622.99	623.42	843.35	1048.26	1232.88
$\widetilde{u_R}$	520.45	1530.08	791.78	715.10	624.49	621.87	830.54	1033.16	1227.35
$\widetilde{u_L}$	537.25	1532.70	816.57	730.24	641.82	638.97	859.66	1080.25	1218.09
$\widetilde{g}$	595.19	784.37	914.26	721.03	710.31	708.58	926.04	820.50	1275.18

### **The SUSY Masses**

minimal SUGRA						GMSB		AMSB	
	$\mathbf{SPS1}$	SPS2	SPS3	SPS4	$\mathbf{SPS5}$	SPS6	SPS7	SPS8	SPS9
$\widetilde{\chi}_1^0$	96.05	79.54	160.55	118.66	119.51	117.50	161.65	137.19	175.51
$\widetilde{\chi}^0_2$	176.82	135.34	296.95	218.14	226.33	215.54	260.06	252.33	549.03
$\widetilde{\chi}^0_3$	358.81	140.84	512.87	383.91	642.83	398.70	306.26	404.00	874.37
$\widetilde{\chi}_4^0$	377.81	269.45	529.57	401.08	652.95	418.06	379.94	426.28	875.97
$\widetilde{\chi}_1^+$	176.38	104.03	296.85	218.06	226.33	215.20	256.33	252.03	175.67
$\widetilde{\chi}_2^+$	378.23	269.03	529.51	402.28	652.68	418.19	379.45	426.47	877.22
$h^0$	113.97	115.71	116.95	115.39	119.79	114.71	113.57	114.83	114.83
$H^0$	394.15	1444.10	573.03	404.63	694.03	457.84	378.37	515.01	912.56
$A^0$	393.63	1442.95	572.42	404.43	693.86	457.26	377.89	514.49	911.74
$H^+$	401.77	1446.18	578.30	416.28	698.49	464.40	386.70	521.17	915.83
$\widetilde{ u}_e$	186.00	1454.17	275.99	441.22	244.52	243.25	249.06	347.61	309.71
$\widetilde{e}_R^-$	142.97	1451.69	178.33	416.54	191.45	191.30	127.43	175.87	303.01
$\widetilde{e}_L$	202.14	1456.33	287.11	448.40	256.30	255.81	261.47	356.61	319.66
$\widetilde{\tau}_1^-$	133.22	1439.46	170.59	267.61	180.67	184.34	120.45	169.42	271.28
$\widetilde{\tau}_2^-$	206.13	1450.38	289.22	414.91	257.86	258.31	263.40	357.59	322.54
$\widetilde{t_1}$	379.11	1003.88	623.83	530.58	220.74	474.12	779.09	957.65	1005.17
$\widetilde{t_2}$	574.71	1307.41	819.54	695.88	644.65	659.73	863.00	1058.68	1128.80
$\widetilde{b_1}$	491.91	1296.56	757.50	606.86	535.86	589.80	822.17	1021.90	1112.07
$\widetilde{b_2}$	524.59	1520.09	791.35	706.45	622.99	623.42	843.35	1048.26	1232.88
$\widetilde{u_R}$	520.45	1530.08	791.78	715.10	624.49	621.87	830.54	1033.16	1227.35
$\widetilde{u_L}$	537.25	1532.70	816.57	730.24	641.82	638.97	859.66	1080.25	1218.09
$\widetilde{g}$	595.19	784.37	914.26	721.03	710.31	708.58	926.04	820.50	1275.18

### **Simulation Tools**

**Event Generation:** 

- ISAJET SUSY and Standard Model background
- HERWIG  $e^+e^- \rightarrow e^+e^-\gamma^*\gamma^*$
- WHIZARD multiple parton final state events

### **Simulation Tools**

**Event Generation:** 

ISAJET - SUSY and Standard Model background

• HERWIG - 
$$e^+e^- \rightarrow e^+e^-\gamma^*\gamma^*$$

WHIZARD - multiple parton final state events

**Detector Simulation** 

Locally written toy MC simulation; some bells and whistles

### **Simulation Tools**

**Event Generation:** 

ISAJET - SUSY and Standard Model background

• HERWIG - 
$$e^+e^- \rightarrow e^+e^-\gamma^*\gamma^*$$

WHIZARD - multiple parton final state events

**Detector Simulation** 

Locally written toy MC simulation; some bells and whistles

Analysis



 $e^+e^- \to X\bar{X}$ 

 $e^+e^- \to X\bar{X}$ 

 $X \to Y + Z$ 

- $e^+e^- \to X\bar{X}$
- $X \to Y + Z$

### E(X) = 250 GeV

$$e^+e^- \to X\bar{X}$$

$$X \to Y + Z$$

$$E(X) = 250 GeV$$

$$E(Z) = \gamma(E_{Z,cm} + \beta P_{Z,cm} cos(\theta))$$

$$e^+e^- \to X\bar{X}$$

$$X \to Y + Z$$

E(X) = 250 GeV

$$E(Z) = \gamma(E_{Z,cm} + \beta P_{Z,cm} cos(\theta))$$

Energy of visible (particle Z) follows a uniform distribution between an  $E_{max}$  and  $E_{min}$ .

# **Smuon Energy Distribution at SPS1**



Two distributions corresponding to two different smuons add to form energy distribution with four endpoints.

Number of events is proportional to cross section and thus polarization.

#### No cut on angle





Joseph Proulx - University of Colorado - ILCWS05 - p.12/27









Angular distribution of smuon production is peaked at  $cos(\theta) = 0$ .

Therefore angular cuts have only a minor effect, mostly at the low end of energy spectrum.



# W pair production

For t-channel events, (gauginos, e.g.) angular cuts affect the high end of the energy spectrum, and are much more pronounced.















#### Joseph Proulx - University of Colorado - ILCWS05 - p.24/27

# **Two-photon background**







- Extremely high cross section
- Peaked at low energy
- Peaked at low angle

# **Two-photon background**

 $\gamma \gamma \rightarrow \mu^+ \mu^- \mathbf{P}_t$  distribution



- Detector needed in highly forward region to detect scattered electrons
- Background is mostly eliminated with missing Pt cut
- Remaining background (from taus) is polarization-independent

### Conclusions

**Requirements for Smuon Studies:** 

- Muon id
- Good forward electron detector
- Reasonable angular coverage
- Good energy (momentum) resolution

### Conclusions

**Requirements for Smuon Studies:** 

- Muon id
- Good forward electron detector
- Reasonable angular coverage
- Good energy (momentum) resolution

Smuon studies do not drive muon detector requirements.