Can the Relic Dark Matter Density be determined at the ILC?

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Contributions by J. Alexander, E. Baltz, B. Dutta, J. Feng, T. Kamon, M. Peskin 2005 International Linear Collider Physics and Detector Workshop August 24, 2005 Snowmass (CO) USA CMB data from WMAP provides precise determination of relic Dark Matter density, result further corroborated by data on supernovas and galaxy clusters;

EGRET data from Compton Gamma Ray Observatory show excess of γ emission in Inner Galaxy, which may be interpreted as signal from DM annihilation;

Analysis of subtracted WMAP spectrum at different frequencies shows excess microwave emission interpretable as **synchrotron emission** from energetic e⁺e⁻ pairs produced in DM annihilation near galactic center:

Data consistent with ~100 GeV particle annihilation with $\langle \sigma v \rangle = 2 \times 10^{-26} \text{ cm}^3 \text{ s}^{-1}$







Systematic study of ILC reach promoted by White Paper on ILC-Cosmo Connections (M.B., J.Feng, N.Graf, M.Peskin, M.Trodden Editors)

Point	m_0	$m_{1/2}$	aneta	A_0	M(t)	$M(\chi_1^0)$
LCC 1	100	250	10	-100	178.	96.1
LCC 2	3280	300	10	0	175.	107.7
LCC 3	210	360	40	0	178.	142.5
LCC 4	380	420	53	0	178.	169

Compute RGEs with Isajet 7.69 and estimate dark matter density from Isajet spectrum and couplings with MicrOMEGAS 1.3 and DarkSUSY 4.0

Point	DarkSUSY 4.0	MicroMEGAS 1.3
LCC 1	0.193	0.193
LCC 2	0.108	0.110
LCC 3	0.059	0.057
LCC 4	0.113	0.106

SUSY models analysis simplified within cMSSM: dimensionality of parameter space reduced by one $(m_{1/2} \leftrightarrow m_0)$: four regions emerge:

Cosmologically interesting cMSSM Regions and Benchmark points for the ILC-Cosmo White Paper



B.L. Ms. Cotton Nero D IV



m_{1/2}



LHC reach limited towards high end of Focus Point Region where strongly interacting superpartners become too heavy to be produced.



LHC Measurements

Availability of decay chains with multi-leptons, lepton+jets topologies allows to determine masses from kinematical endpoints (but significant correlations from sensitivity to mass differences):

 $\tilde{q} \to q \tilde{\chi}_2^0 \to q \tilde{\ell}^{\pm} \ell^{\mp} \to q \ell^{\pm} \ell^{\mp} \tilde{\chi}_1^0$



Predictions of DM relic density can be obtained, in a model-dependent way, to good numerical accuracy in some regions by reconstructing cMSSM parameters from observed endpoints: SPS1a: $\delta\Omega/\Omega = 0.025$ (stat.) 300 fb⁻¹

$$M_{\ell^+\ell^-}^{\max} = \frac{\sqrt{\left(m_{\chi_2^0}^2 - m_{\tilde{\ell}}^2\right)\left(m_{\tilde{\ell}}^2 - m_{\chi_1^0}^2\right)}}{m_{\tilde{\ell}}}$$
$$M_{\ell_1q}^{\max} = \frac{\sqrt{\left(m_{\chi_2^0}^2 - m_{\tilde{\ell}}^2\right)\left(m_{\tilde{q}}^2 - m_{\chi_2^0}^2\right)}}{m_{\chi_2^0}}$$





Bulk Point LCC1

In bulk region LSP mostly bino and DM density controlled by annihilation to leptons via slepton exchange: need to determine LSP and slepton masses but also ensure no other mechanisms contribute.

0,25 c ¹ 0.2	0.25	0.25			1.1.1.0	
0.15	0.15-	0.15		value	$\delta(LHC)$	$\delta(ILC)$
0.1	0.1	0.1	$M(ilde{\chi}_1)$	96.1	\pm 4.8	\pm 0.05
0.05-	0.05	0.05	$M(ilde{e}_R)$	143.0	\pm 4.8	\pm 0.05
90 60 70 80 90 100 110 120 130 14M	900 120 140 180 180 200 220 240 280 200	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$M(ilde{\mu}_R)$	143.0	\pm 4.8	\pm 0.2
0.25	0.25	.25	$M(ilde{ au}_R)$	133.2	± 5-8	\pm 0.3
0.2	0.2	0.2	$M(ilde{e}_L)$	202.1	\pm 5.0	\pm 0.2
0.15	0.15	1.15	$M(ilde{\mu}_L)$	202.1	\pm 5.0	\pm 0.5
0.1	0.1	0.1-	$M(ilde{ au}_L)$	206.1	?	\pm 1.1
0.05	0.05	1.05	$M(\tilde{\chi}_2) - m(\tilde{\chi}_1)$	80.3	\pm 0.08	± 2
0 200 250 300 350 400 450 50 µ	Abo 200 300 400 500 600 700 800 900 8000 M(A)	0 200 300 400 500 600 700 800 Mg3	M(A)	393.6	$\overline{M}(A) > 200$	M(A) > 220



Focus Point LCC2

In focus point DM density controlled by LSP annihilation to WW and ZZ, large mass splitting between gauginos and sfermions:





ILC Measurements at LCC2

Study of Focus Point at 0.5 TeV is based on five main reactions: $e+e- \rightarrow \chi^+_1 \chi^-_1, \chi^+_1 \chi^-_2, \chi^0_1 \chi^0_3, \chi^0_2 \chi^0_3, \chi^0_3 \chi^0_4$

Determine mass differences from endpoint of 11 and jj distributions and use kinematics to fix masses:



Availability of polarised beams provides additional observables for establishing properties of gauginos;

	value	δ (ILC)
$M(ilde{\chi}_1)$	107.7	\pm 0.7
$M(ilde{\chi}_2) - M(ilde{\chi}_1)$	58.6	\pm 0.4
$M(ilde{\chi}_3) - M(ilde{\chi}_1)$	82.3	\pm 0.3
$M(\tilde{\chi}_1^+) - M(\tilde{\chi}_1)$	143.0	\pm 0.3
$\sigma(e^-e^+ \to \tilde{\chi}_1^+ \chi_1^-)$ (-0.8/+0.6)		\pm 7.7%
$\sigma(e^-e^+ \to \tilde{\chi}_1^+ \chi_1^-)$ (+0.8/-0.6)		$\pm \ 10.5\%$
$\sigma(e^-e^+ ightarrow ilde{\chi}^0_2\chi^0_3)$ (-0.8/+0.6)		\pm 3.8%
$\sigma(e^-e^+ o ilde{\chi}^0_2 \chi^0_3)$ (+0.8/-0.6)		\pm 4.5%
$M(ilde{\ell})$	3270	
$M(ilde{q})$	3300	
M(A)	3242.2	>220



co-Annihilation Point LCC3

DM density controlled by stau-LSP mass splitting and μ : sensitivity to small ΔM depends on $\gamma\gamma$ background rejection:





	value	$\delta(ILC)$
$M(\tilde{\tau}_1) - M(\tilde{\chi}_1)$	9.5	\pm 1.0
$M(ilde{ au}_1)$	151	\pm 0.5
$M(ilde{\chi}^0_1)$	142	\pm 0.1
$M(\tilde{\chi}_2^0) - M(\tilde{\chi}_1^0)$	80.3	\pm 0.5
$M(ilde{\chi}_3^0) - M(ilde{\chi}_1^0)$		\pm 2.0
$M(\tilde{\chi}_2^+) - M(\tilde{\chi}_1^+)$		± 2.0
$M(\tilde{\chi}+_1)$, $M(\mu_R)$	274	\pm 0.7
$M(ilde{e}_R)$, $M(ilde{\mu}_R)$	252	\pm 1.0
$M(\tilde{e}_R) - M(\tilde{\chi}_1^0)$		\pm 1.0
$M(\tilde{\tau}_2) - M(\tilde{\chi}_1)$		\pm 1.1
M(A)		M(A) > 450



ILC Measurements at LCC3

At 0.5 TeV production of $\tau_1 \tau_1$ and $\chi_1 \chi_2$ resulting in $\tau \tau E_{\text{missing}}$ final state; Important to reject $\gamma \gamma$ bkg ee $\rightarrow \text{ee} \tau \tau$ by low angle electron tagging:





Very Fwd. calorimetric coverage controls minimum reachable ΔM:

 Δ M accuracy at ILC = 10% At LHC worst accuracy and feasibility critically depends on fake jet rate.



Determine $M(\tau_1) - M(\chi_1^0)$ from distribution of $M(j_1j_2E_{missing})$





A⁰ Funnel Point LCC4

DM density controlled by $M(A)/2M(\chi)$, $\Gamma(A)$ and μ requires intensive program of measurements from 0.35 TeV to 1.0 TeV:





ILC Measurements at 0.5 TeV

Determine $M(\tau_1)$ and $M(\tau_1) - M(\chi_1^0)$ from stau threshold scan and stau decays;

Estimate $\Gamma(A^0)$ from precise determination of BR(h⁰ \rightarrow bb) at 0.35/0.5 TeV;

$$\Gamma(A^0) = \frac{\mathrm{BR}(h^0 \to b\bar{b})}{\mathrm{BR}(A^0 \to b\bar{b})} \times \Gamma(h^0) \times \tan^2\beta$$



Stau Threshold Scan











Determine $M(\chi_3)$ - $M(\chi_1)$ from Z energy distribution in $\chi_3 \rightarrow \chi_1 Z$ decays in $\chi_3 \chi_2$ events to fix μ value; At LHC M(A) measurable to 2 GeV but difficuly to control $\Gamma(A)$ and μ .







Constraining tan β at 1 TeV

Points at large tan β , such as LCC3 and LCC4 and EGRET compatible region have large sensitivity on tan β ;





e+e- \rightarrow H+H- \rightarrow tbtv sensitive to tan β process produced with typical cross section of ~ 2 fb at 1 TeV giving BRs accuracy of O(3-6%).





Flat Scans

Perform model-independent MSSM scans around each LCC point;

Extract MSSM parameters from cMSSM inputs and vary MSSM parameters in uncorrelated way over ranges consistent with $> 3\sigma$ from anticipated accuracies;

Compute DM relic density at each MSSM point and constructed *pdf* by weighting by $\Pi_J erf(\frac{\delta_j}{\sqrt{\sigma_j}})$

Extract uncertainty on DM density prediction by width of resulting *pdf* distribution.

Parameter	LCC 1	LCC 2	LCC 3	LCC 4
aneta	\pm 10	\pm 5	\pm 4	\pm 10
M_1	\pm 30	± 10	\pm 5	± 10
M_2	\pm 30	\pm 10	\pm 5	\pm 5
M_3	\pm 30	\pm 10	± 5	\pm 5
A_m	± 150	± 150	± 150	\pm 150
A_l	± 150	± 150	± 150	\pm 150
A_b	± 150	± 150	± 150	\pm 150
A_t	± 150	± 150	± 150	\pm 150
Ml1	\pm 15	± 10	\pm 4	\pm 5
Mr1	\pm 15	\pm 10	\pm 4	\pm 5
Ml3	\pm 15	\pm 10	\pm 4	\pm 5
Mr3	\pm 15	\pm 10	\pm 4	\pm 5
Mqu	\pm 10	\pm 10	\pm 10	\pm 10
Mqd	\pm 10	\pm 10	\pm 10	\pm 10
Mq3	± 10	\pm 10	± 10	\pm 10
μ	± 150	± 150	± 150	\pm 100
M_A	\pm 100	\pm 50	\pm 5	\pm 4
M_{top}	± 1	± 1	± 1	\pm 1



Markov Chain Scans

Scan MSSM multi-parameter phase space using Markov Chain Monte Carlo technique: given a point *i* advance to new point *i*+1 if i) $\mathcal{L}(i+1)/\mathcal{L}(i)$ or ii) > rndm() where \mathcal{L} defined by SUSY measurements and anticipated accuracies; (Berg, cond-mat/0410490 and Baltz, Gondolo hep-ph/0407039)

Markov Chain technique is more effcient and has better statistical weight of relevant regions; but reaching into topologically disconnected regions may be problematic;





Selected Parameter Regions at LHC and ILC



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A Comparison on DM density accuracy at LHC and ILC in Bulk Region







ILC Accuracy on Dark Matter (Preliminary)



Relative Accuracy $\delta\Omega/\Omega$







If Lightest Neutralino responsible for observed Dark Matter density in the Universe, **expect important signals to be detected at LHC**;

But to fully understand the role of the newly discovered particles in determining the Dark Matter and its impact on the history of the Universe, the **accuracy provided by the ILC** in studying its microscopic properties and those of the other relevant particles is **crucial**;

A sample of scenarios, widely different in terms of phenomenology and requirements shows that the ILC has the capabilities to promote the study of SUSY Dark Matter to an accuracy competitive to that of present and future satellite CMB data.





