

Understanding LIGHT

Why we need a
terascale photon collider

Zack Sullivan

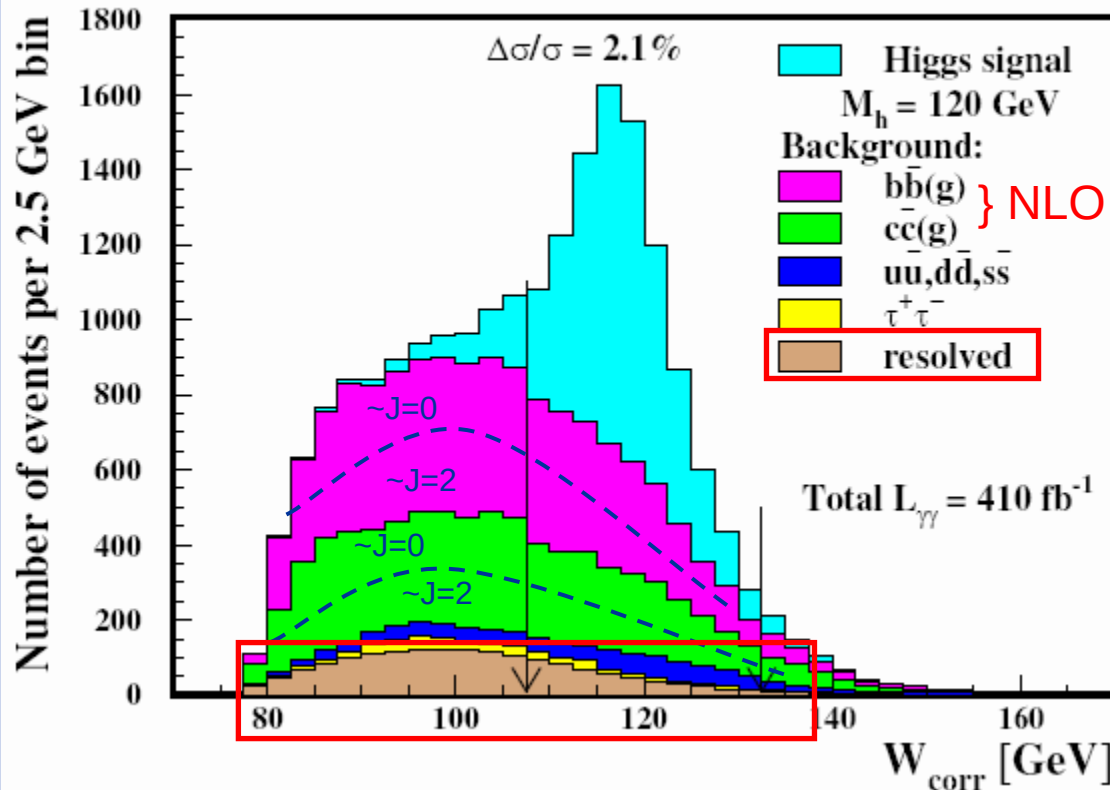
Argonne National Laboratory

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Outline

1. Precision Higgs — the usual suspect
2. $b\bar{b}$ production as a window on light
3. What can we say about it now?
4. Conclusions

Precision Higgs



Measure:

$$\Gamma(h \rightarrow \gamma\gamma) \sim 2\%$$

$$\Gamma_{\text{tot}}(h) \sim 10\%$$

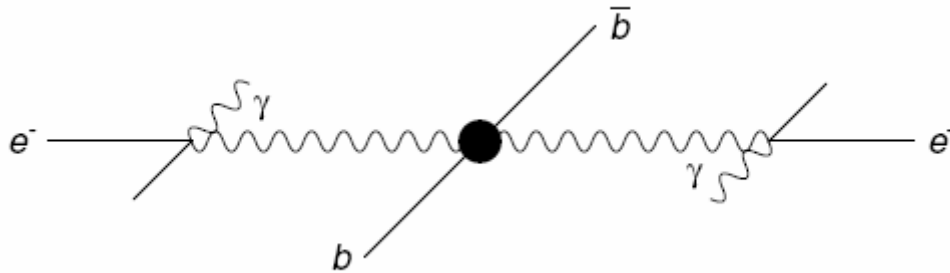
- NLO $b\bar{b}/c\bar{c}$:
 $J=0 \approx J=2$
 – $J=0$ really LO
- How big is the resolved term?

P. Nieżurawski, A. F. Żarnecki, M. Krawczyk
 presented by J. Ciborowski

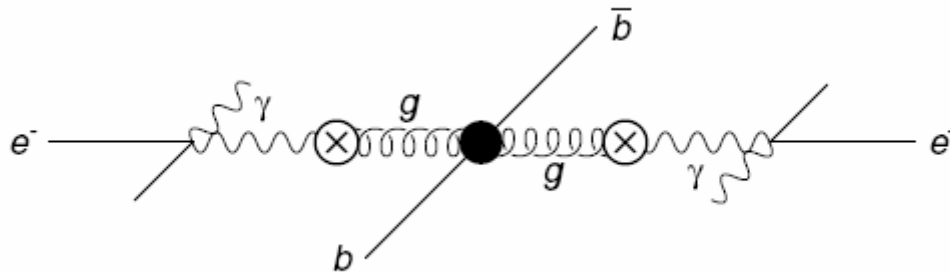
**WE DO NOT
UNDERSTAND
LIGHT**

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UNDERSTAND**
(the ^{apparent} structure of)
LIGHT

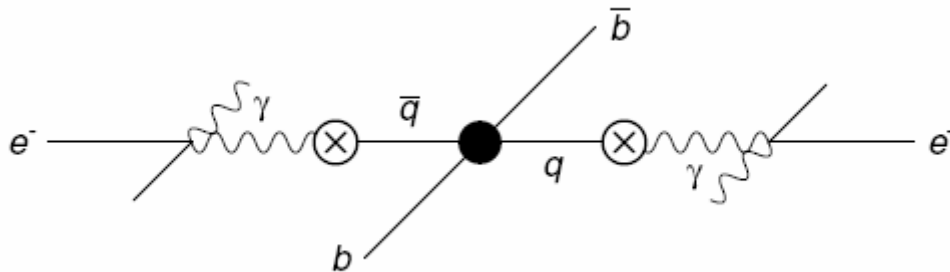
$b\bar{b}$ production through direct and resolved photons



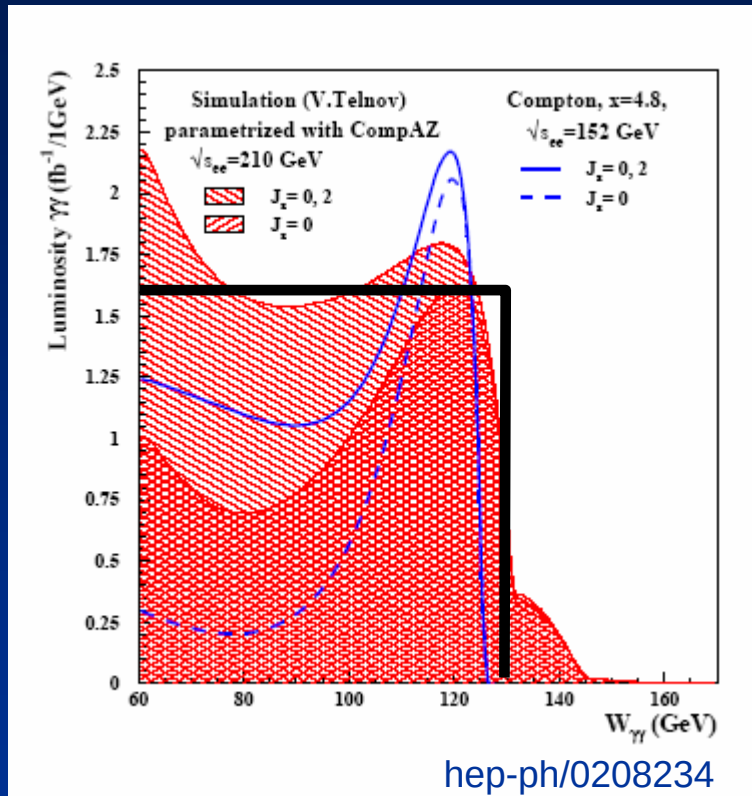
Direct



Resolved



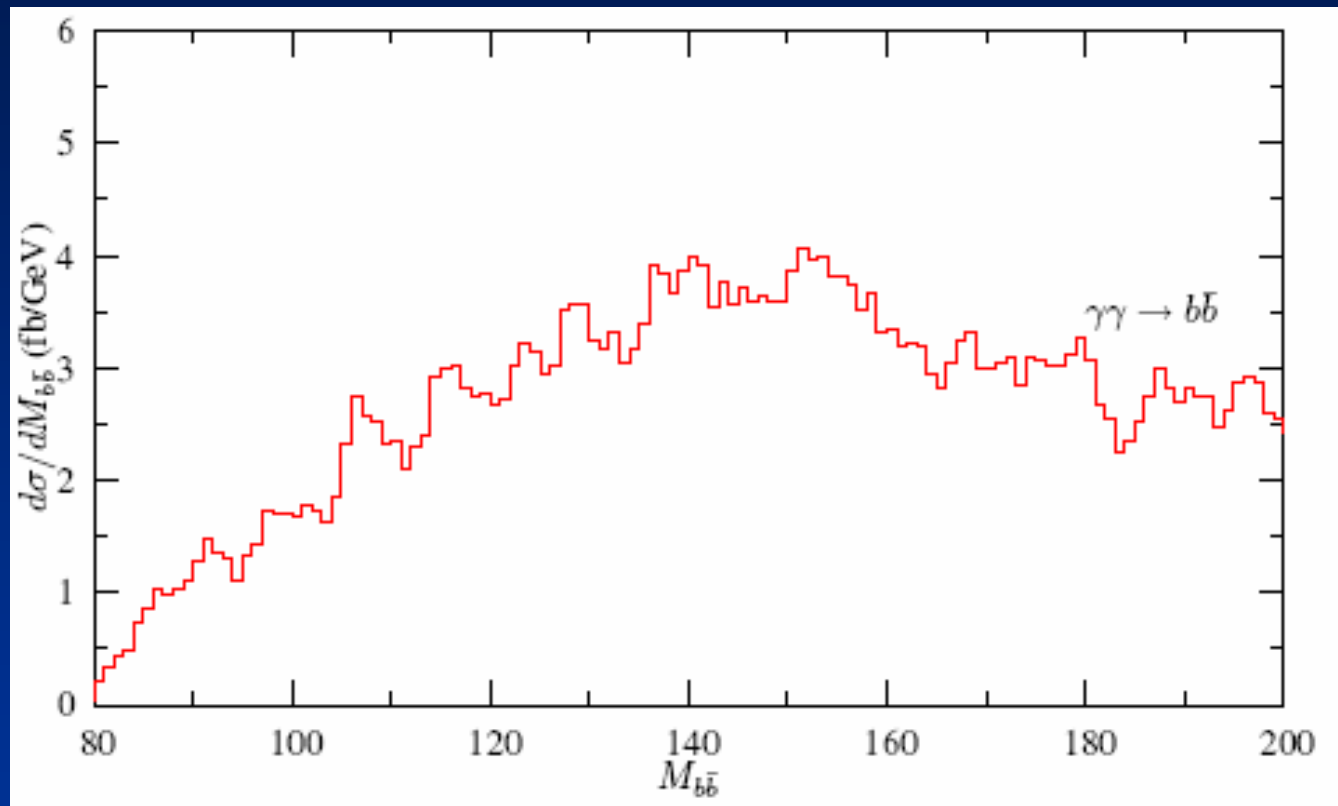
Photon spectrum and simulation



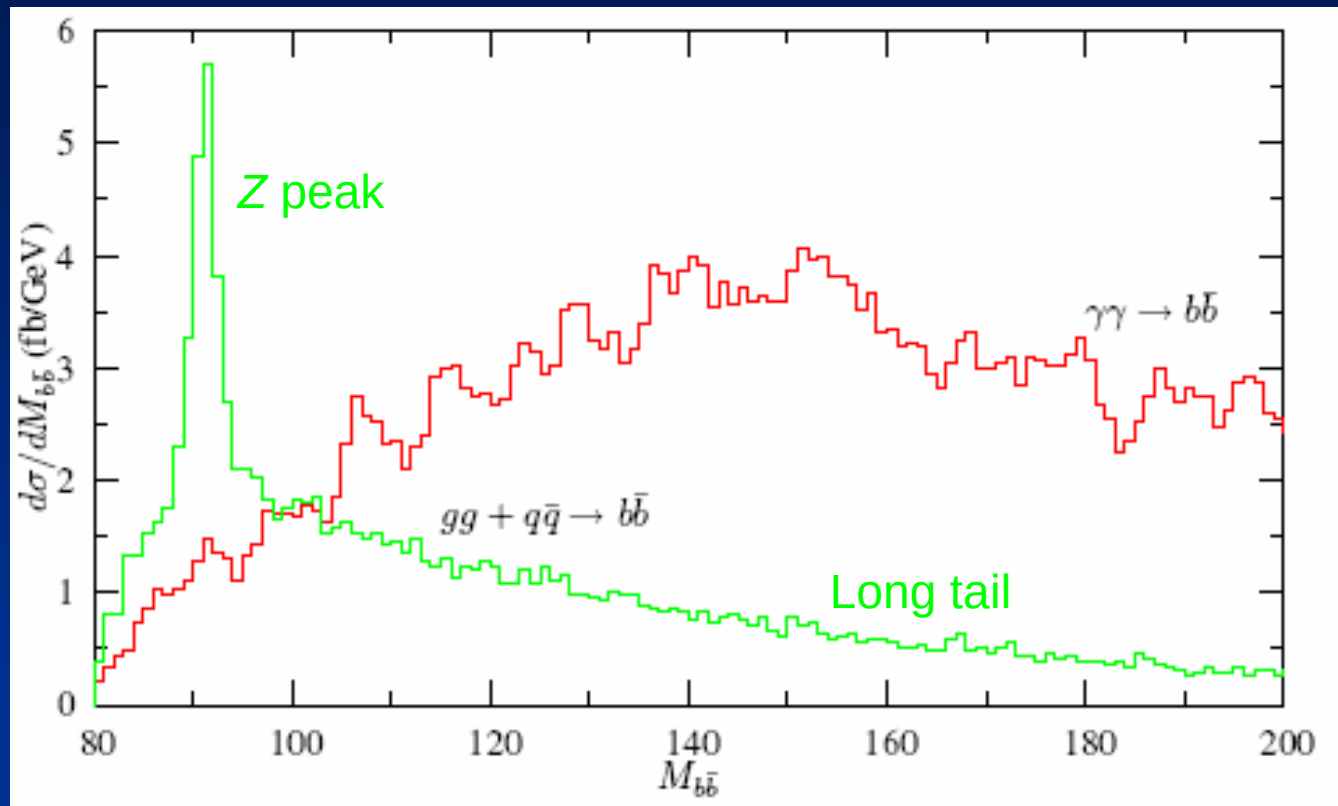
Full $\gamma\gamma$ spectrum is flatter than predicted by Compton scattering alone.

- $\sqrt{s_{ee}} = 500$ GeV
 - $W_{\gamma\gamma} = 50\text{-}400$ GeV
 - $E_\gamma = 25\text{-}200$ GeV, using a flat spectrum
 - Results insensitive to actual shape
- Cuts
 - $E_{Tb} > 40$ GeV, $|\eta_b| < 4$
 - $\Delta R_{iso} = 0.1$
 - $M_{bb} > 80$ GeV

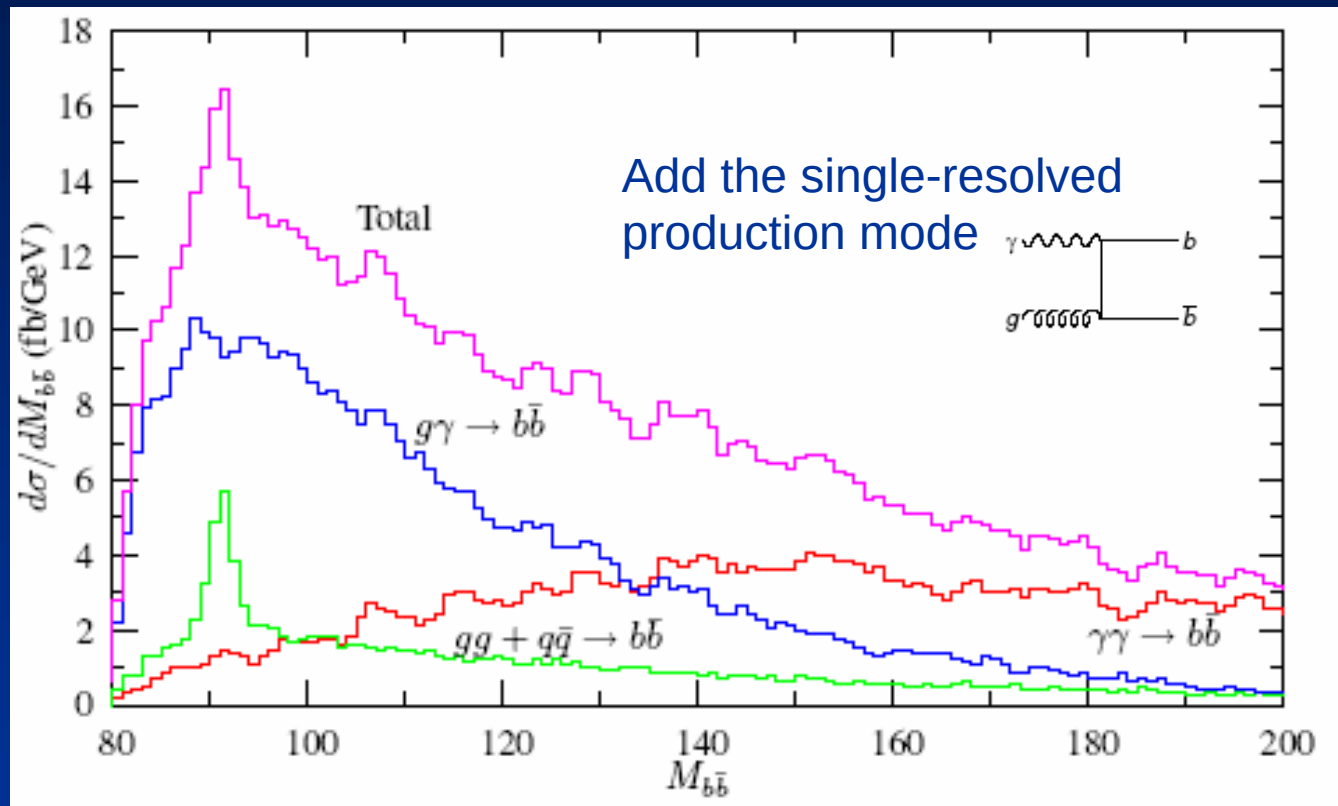
bb cross section



bb cross section



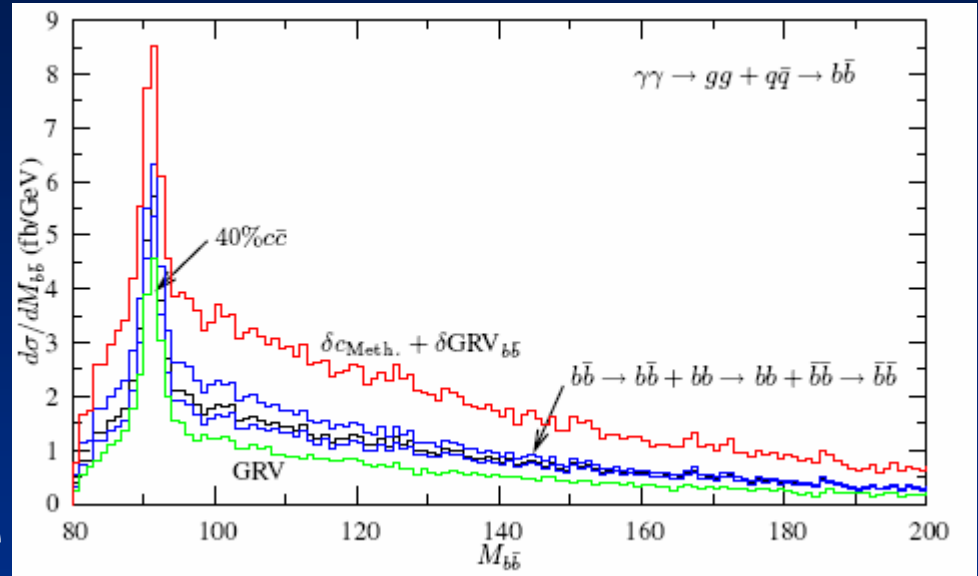
bb cross section



The single-resolved mode completely dominates the cross section below $\sim \frac{1}{3}W_{\gamma\gamma}^{\max}$.

Uncertainty in resolved-resolved cross section

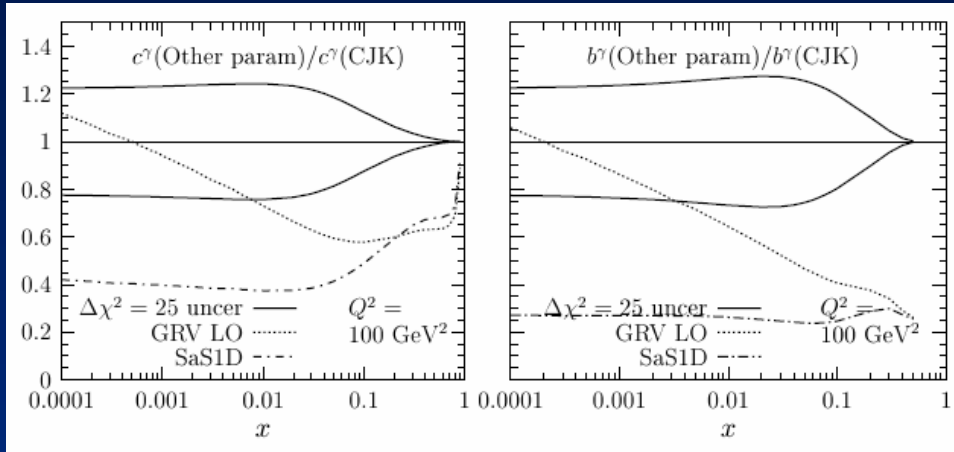
- CJK2 LO PDFs
 - Error calculation is based on Hessian matrix method
- For observables use a “modified tolerance method” ZS, PRD 66, 075011 (02)



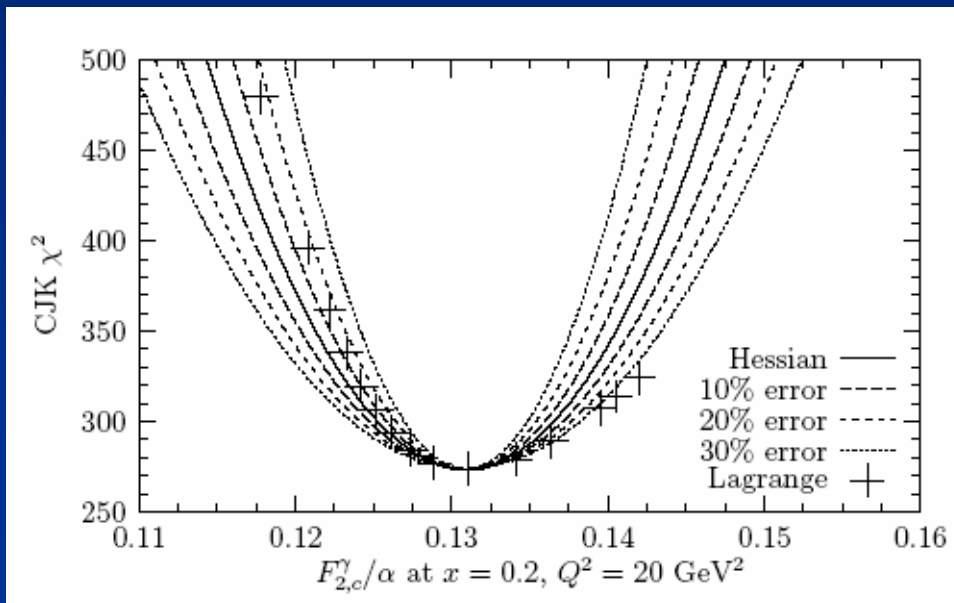
$$\delta\sigma_{\pm} = \frac{T}{10} \sqrt{\sum_i (\sigma_i - \sigma_0)^2 |_{\sigma_i > \sigma_0}}$$

- Assumes fit parameters can be mapped to a hypersphere – Hessian distribution.

CJK2 vs. GRV / Hessian vs. Lagrange



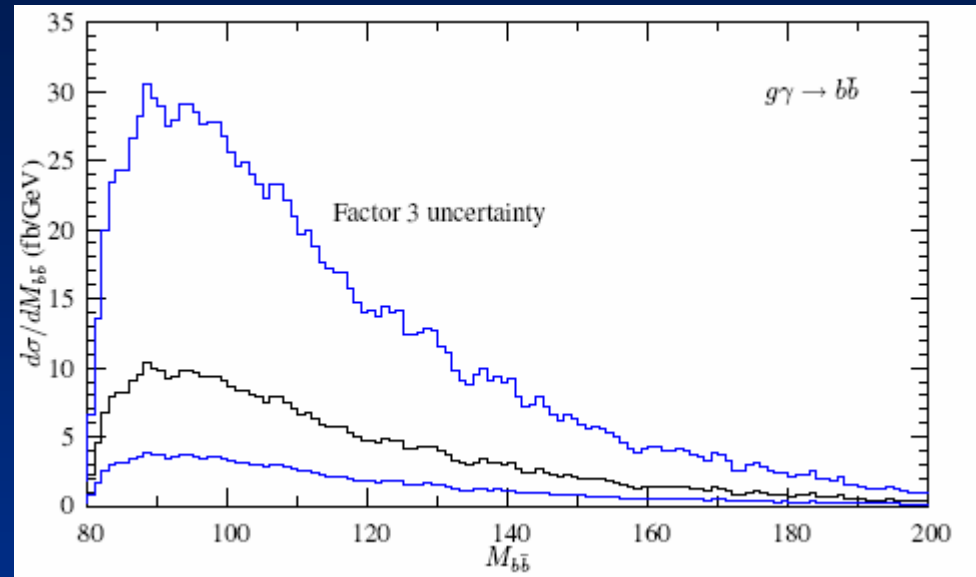
- c , b PDFs differ significantly at large x . Well beyond estimated uncertainties.



- Already at 4 GeV, the c (b ?) are not really Hessian any longer.

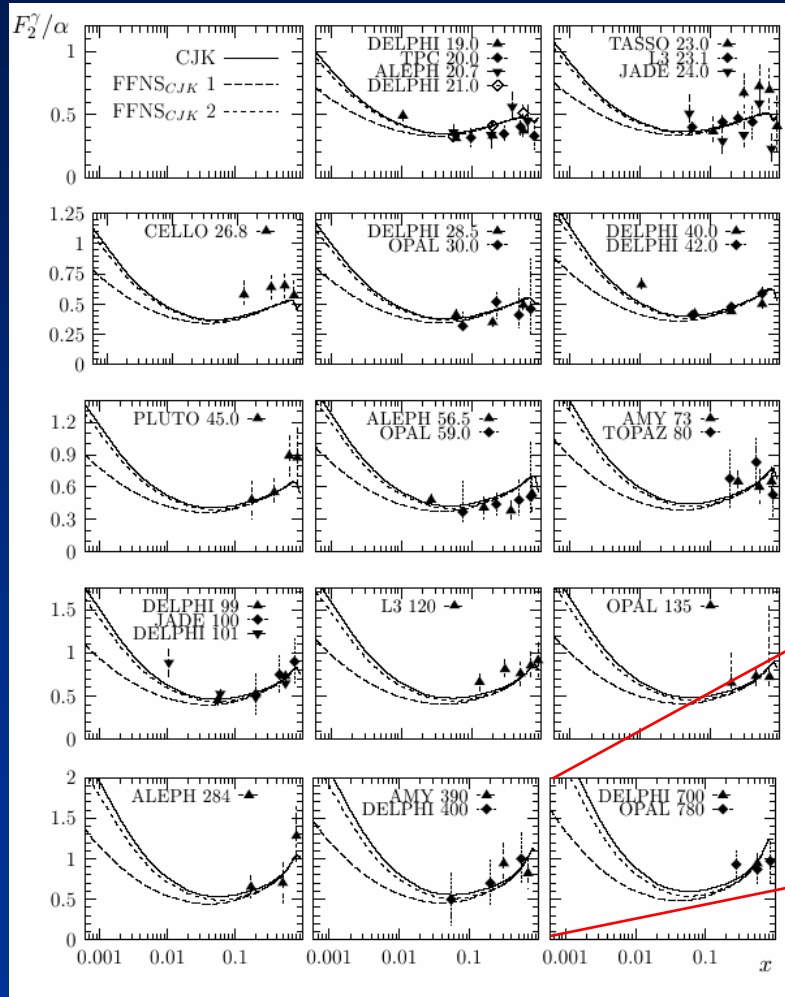
Uncertainty in $g\gamma \rightarrow b\bar{b}$

- Huge uncertainty in the gluon PDF.
 - Very little data to constrain it.
- Tolerance T is much larger for the gluon.



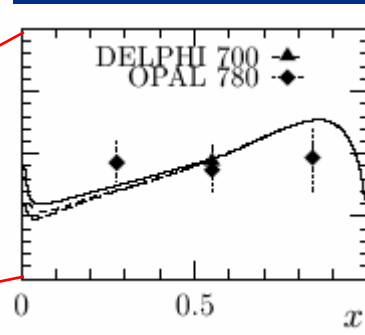
	$T(G\gamma)$	$T(d\gamma)$	$T(u\gamma)$	$T(s\gamma)$	$T(F_2^\gamma)$
$1 \leq Q^2 \leq 100 \text{ GeV}^2$	4.5	7.0	7.0	3.4	8
$1 \leq Q^2 \leq 1000 \text{ GeV}^2$	14.0	7.0	7.0	3.4	10.5
$1 \leq Q^2 \leq 200000 \text{ GeV}^2$	138.0	7.0	7.0	3.4	20.0

Summary of fit data

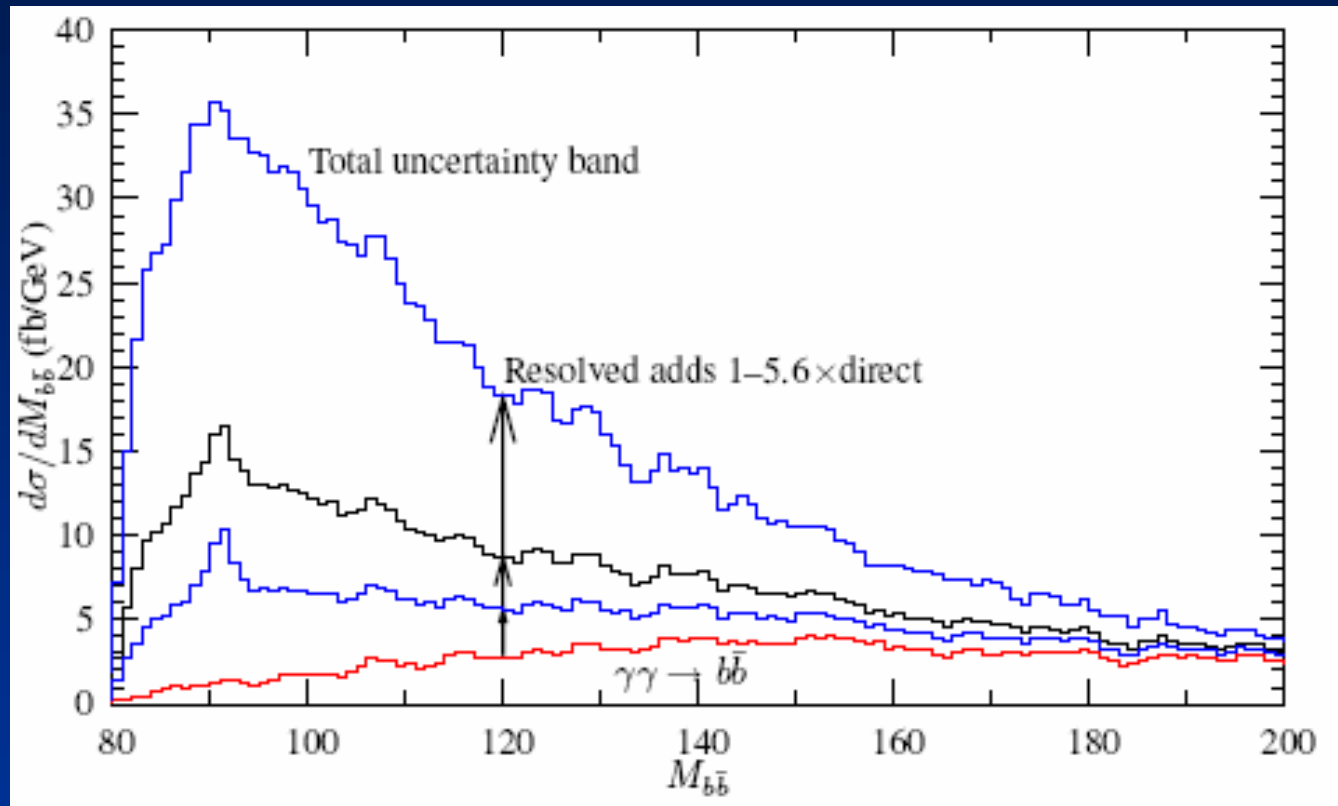


model	χ^2 (182 pts)	χ^2/DOF
CJK	273.7	1.537

- Full F_2^γ fit is <1% consistent
 - At ~ 90 GeV, only 1 point actually fits
- Mostly u, c quarks here



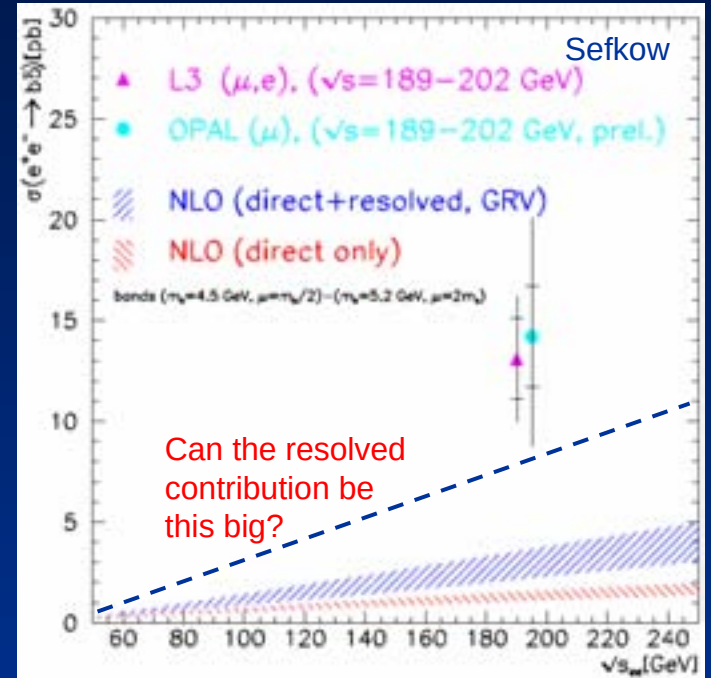
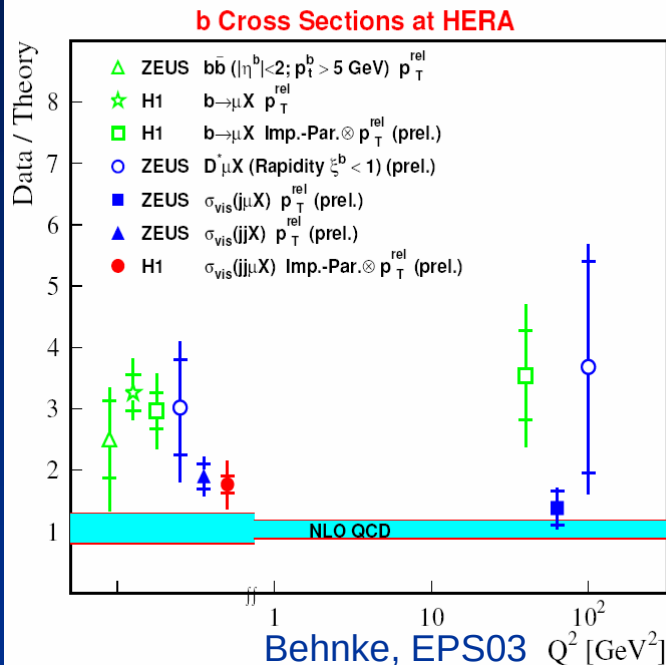
Total uncertainty for bb



- This will only be improved by direct measurement at a real photon collider.

$b\bar{b}$ at HERA / LEP

Beauty at HERA: Summary plot



- Theory underestimates the cross section at both HERA and LEP by factors of 2-3!
- It seems likely this is partially due to additional resolved photons.

Conclusions

- Below $\sim \frac{1}{3}W_{\gamma\gamma}^{\max}$, theorists should think of $\gamma\gamma$ colliders as clean hadron colliders.
- In order to predict cross sections to better than factors of 5, we must measure the structure of the photon *in situ*.
- Today we do not understand LIGHT,
but with a terascale photon collider,
WE WILL

Extra Slides

$b\bar{b}$ cross section with lower cuts

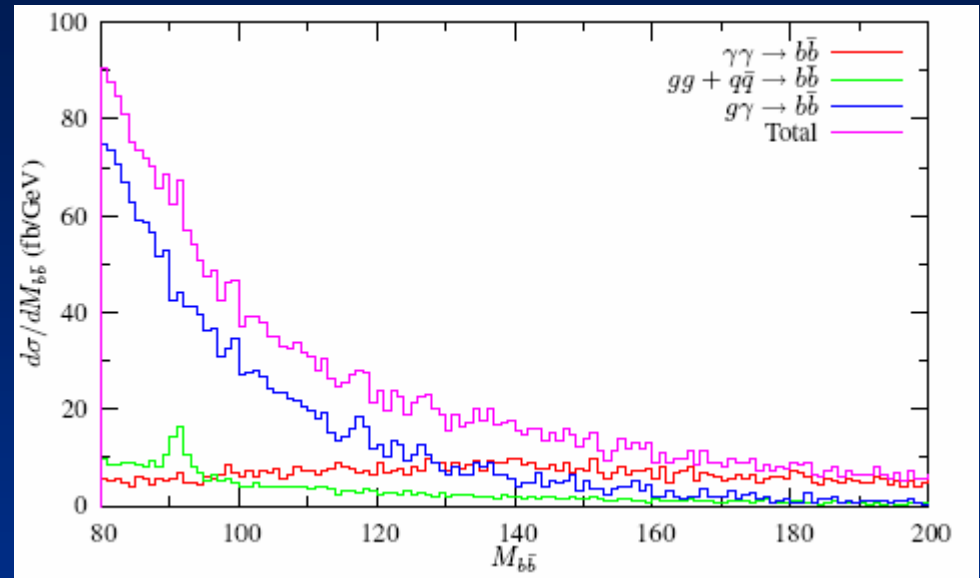
- $S_{ee} = 500$ GeV
 - $W_{\gamma\gamma} = 50\text{-}400$ GeV
 - $E_\gamma = 25\text{-}200$ GeV, using a flat spectrum

- Cuts

$$E_{Tb} > 15 \text{ GeV}, |\eta_b| < 4$$

$$\Delta R_{iso} = 0.1$$

$$M_{bb} > 80 \text{ GeV}$$



- Looser cuts fill in direct production near $M_{bb} \sim 80$ GeV, but $g\gamma \rightarrow b\bar{b}$ grows much more quickly.