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# ***Sensitivities to the Scale of Supersymmetry from Precision Observables***

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Based on collaboration with J. Ellis, S. Heinemeyer, K. Olive  
[hep-ph/0411216 + updates for  \$m\_t = 172.7 \pm 2.9\$  GeV](https://arxiv.org/abs/hep-ph/0411216)

1. Introduction
2. Electroweak precision observables
3. Combined sensitivity: present situation and ILC precision
4. Conclusions

# 1. Introduction

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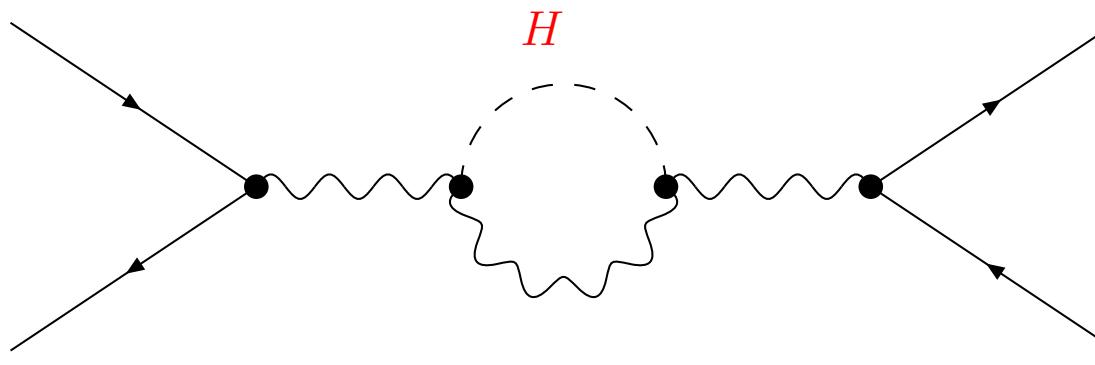
EW precision data:

$M_Z, M_W, \sin^2 \theta_{\text{eff}}^{\text{lept}}, \dots$

Theory:

SM, MSSM, ...

Test of theory at quantum level: sensitivity to loop corrections



Indirect constraints on unknown parameters:  $M_H, m_{\tilde{t}}, \dots$

Effects of “new physics”?

# ***Constrained MSSM (CMSSM) with restrictions from dark matter relic density***

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CMSSM characterised by five parameters:

$m_{1/2}$ ,  $m_0$ ,  $A_0$  (GUT scale),  $\tan \beta$ ,  $\text{sgn}(\mu)$  (weak scale)

⇒ Low-energy spectrum from renormalisation group running  
lightest SUSY particle:  $\tilde{\chi}_1^0$

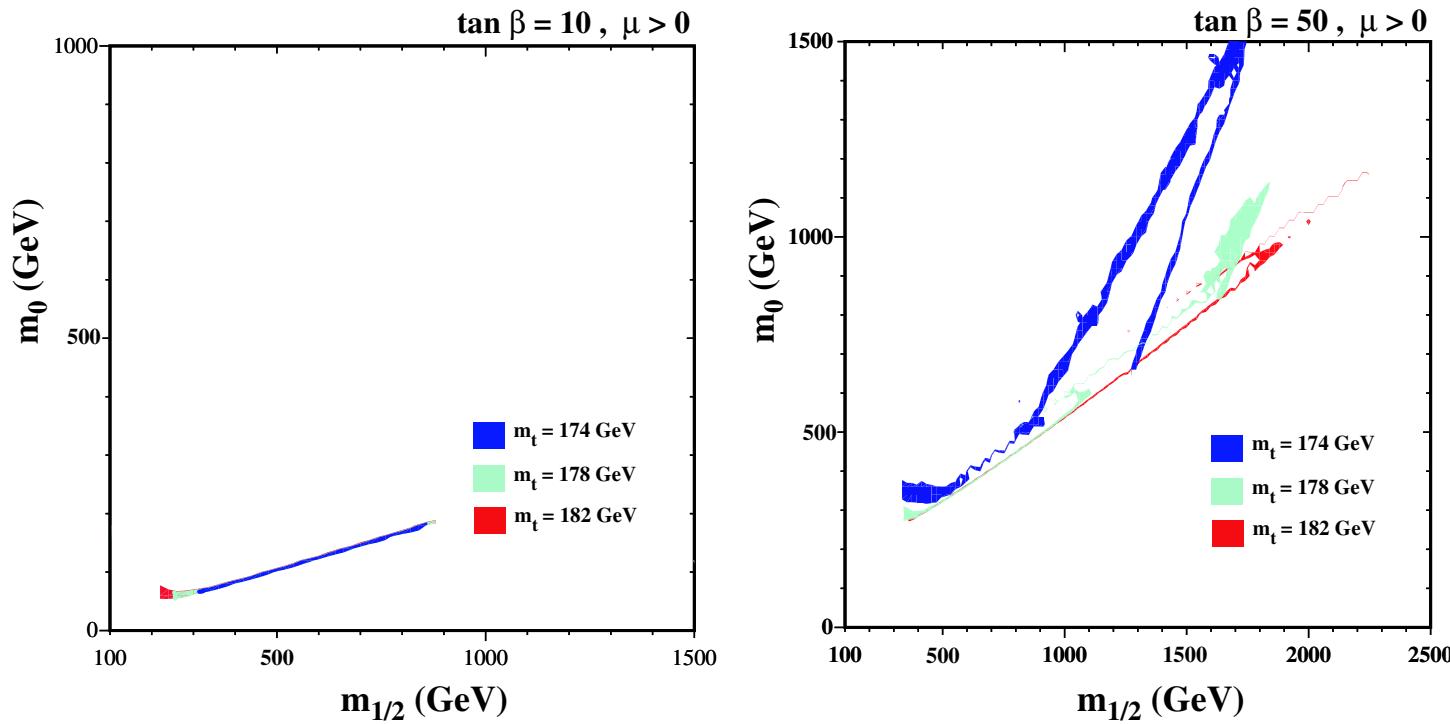
Cold dark matter (CDM) density (WMAP, . . .):

$$0.094 < \Omega_{\text{CDM}} h^2 < 0.129$$

⇒ Constraints on SUSY parameter space

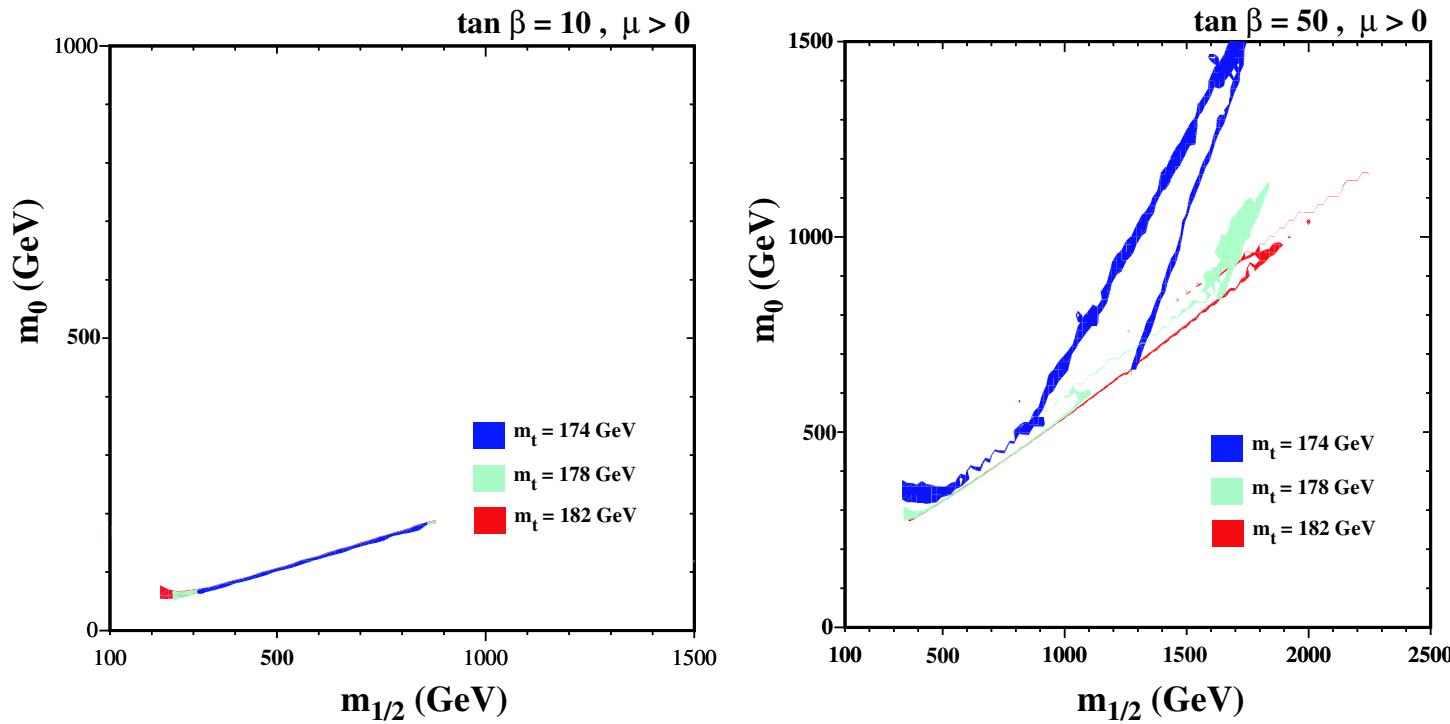
# *Constrained MSSM (CMSSM) with restrictions from dark matter relic density*

Allowed region in  $(m_{1/2}, m_0)$  plane (fixed  $A_0$ , different  $m_t$ ):



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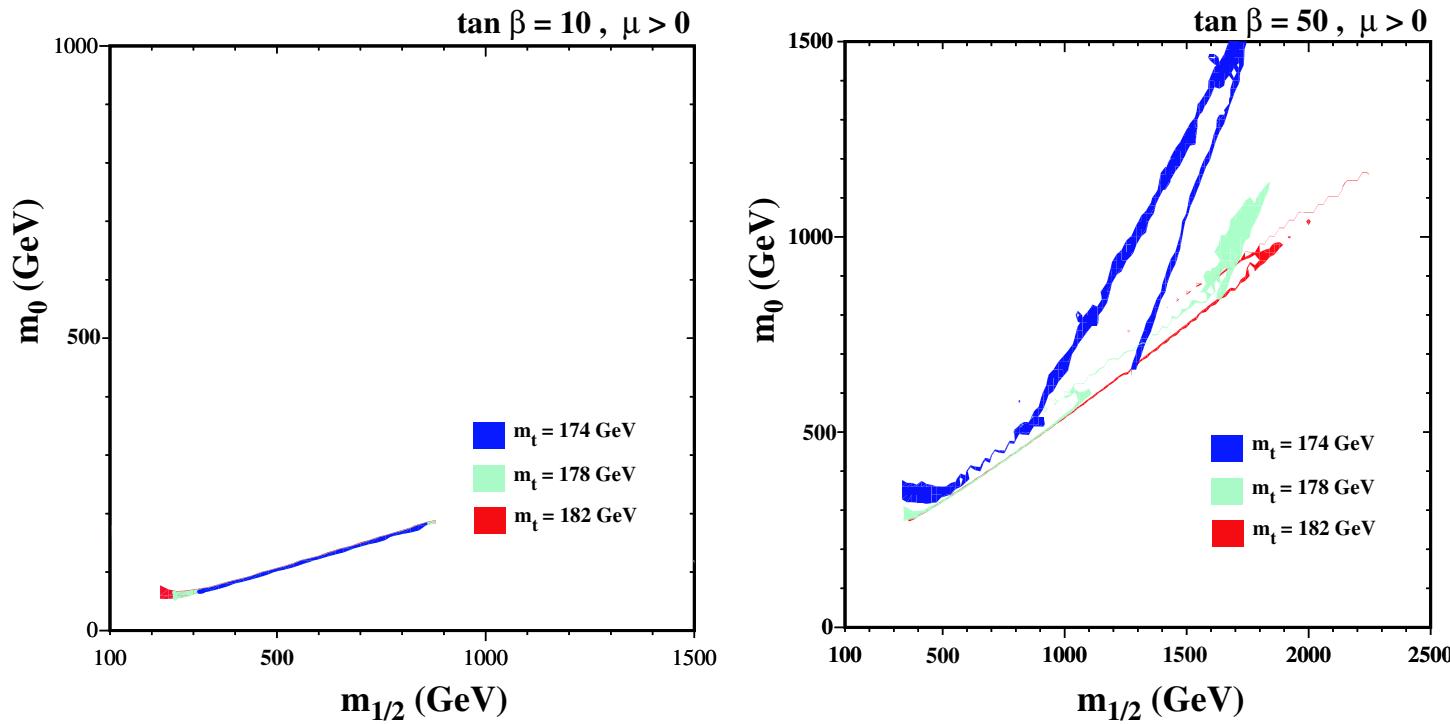
Allowed region in  $(m_{1/2}, m_0)$  plane (fixed  $A_0$ , different  $m_t$ ):



- ⇒ narrow “WMAP strips”
- ⇒ effectively reduces dimensionality of parameter space

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- ⇒ narrow “WMAP strips”
- ⇒ effectively reduces dimensionality of parameter space
- ⇒ analyse CMSSM along WMAP strips

## **2. Electroweak precision observables**

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Observables taken into account:

Present:

$M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $(g - 2)_\mu$ ,  $b \rightarrow s\gamma$ ,  $B_s \rightarrow \mu^+ \mu^-$

Current experimental errors + estimate of current theoretical uncertainties (from unknown higher-orders + experimental errors of input parameters)

## 2. Electroweak precision observables

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Current experimental errors + estimate of current theoretical uncertainties (from unknown higher-orders + experimental errors of input parameters)

Future (ILC):

$M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $(g - 2)_\mu$ ,  $b \rightarrow s\gamma$ ,  $B_s \rightarrow \mu^+ \mu^-$ ,  
 $M_h$ ,  $\text{BR}(h \rightarrow b\bar{b})/\text{BR}(h \rightarrow WW^*)$

Experimental precision at the ILC + estimate of future theoretical uncertainties

## **Theoretical predictions for $M_W$ , $\sin^2 \theta_{\text{eff}}$ :**

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Comparison of prediction for muon decay with experiment (Fermi constant  $G_\mu$ )

$$\Rightarrow M_W^2 \left( 1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi \alpha}{\sqrt{2} G_\mu} (1 + \Delta r),$$

$\Updownarrow$   
**loop corrections**

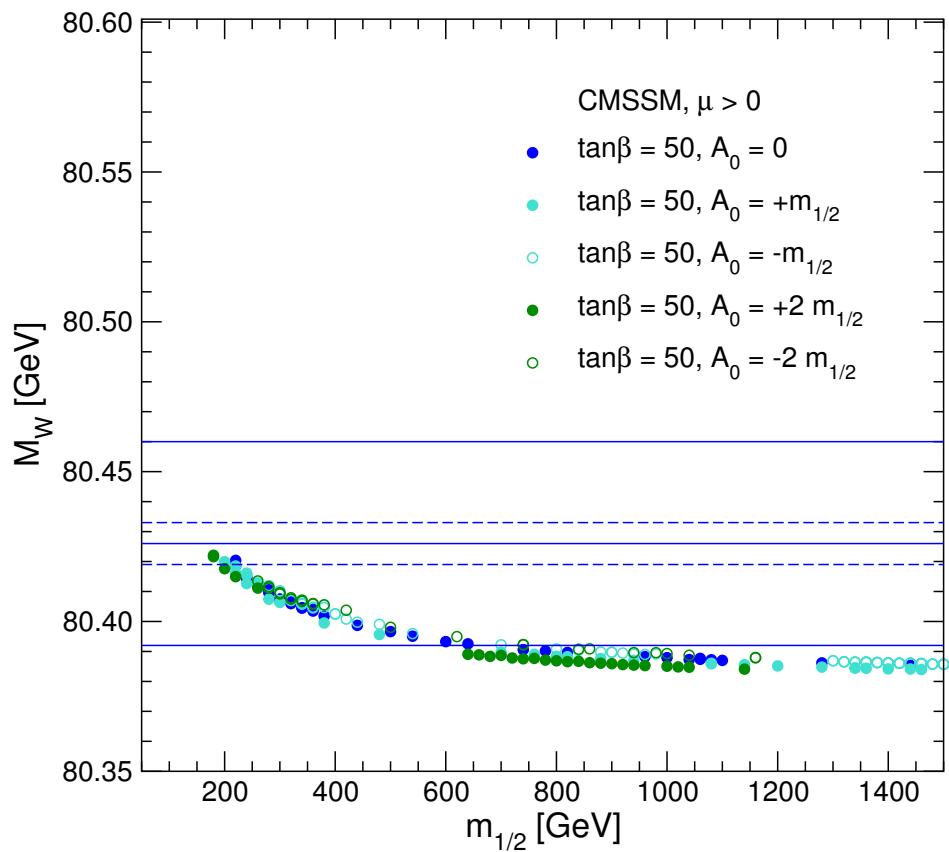
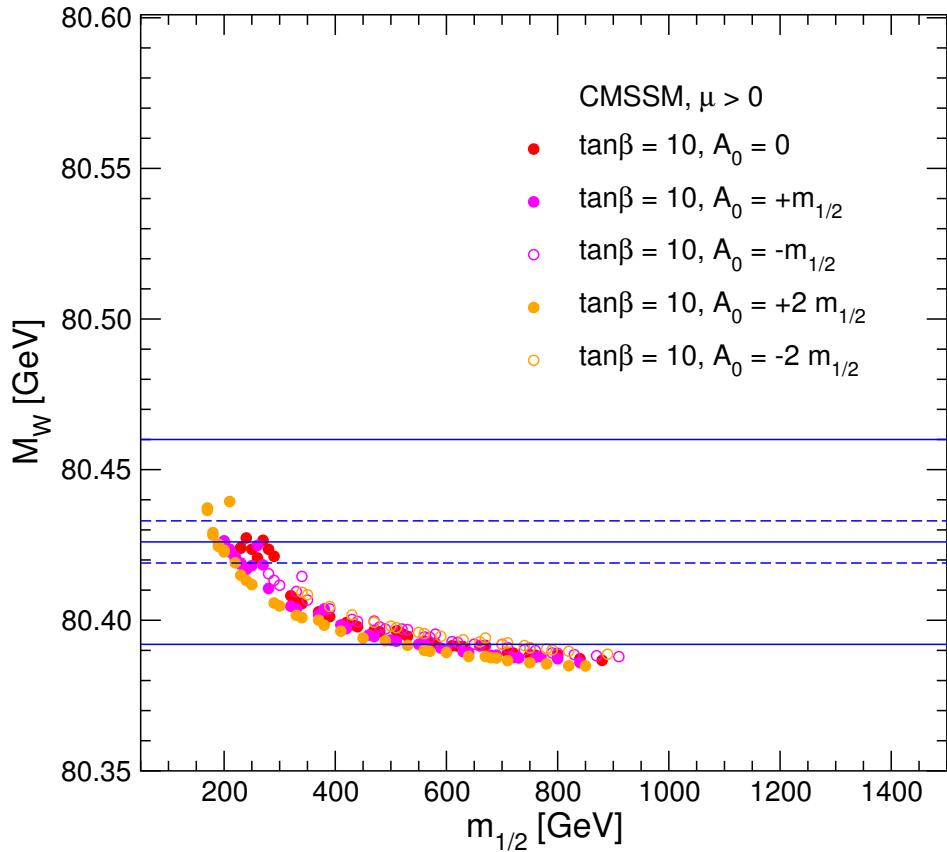
$\Rightarrow$  Theo. prediction for  $M_W$  in terms of  $M_Z$ ,  $\alpha$ ,  $G_\mu$ ,  $\Delta r(m_t, m_{\tilde{t}}, \dots)$

Effective couplings at the Z resonance:

$$\Rightarrow \sin^2 \theta_{\text{eff}} = \frac{1}{4} \left( 1 - \text{Re} \frac{g_V}{g_A} \right) = \left( 1 - \frac{M_W^2}{M_Z^2} \right) \text{Re} \kappa_l(s = M_Z^2)$$

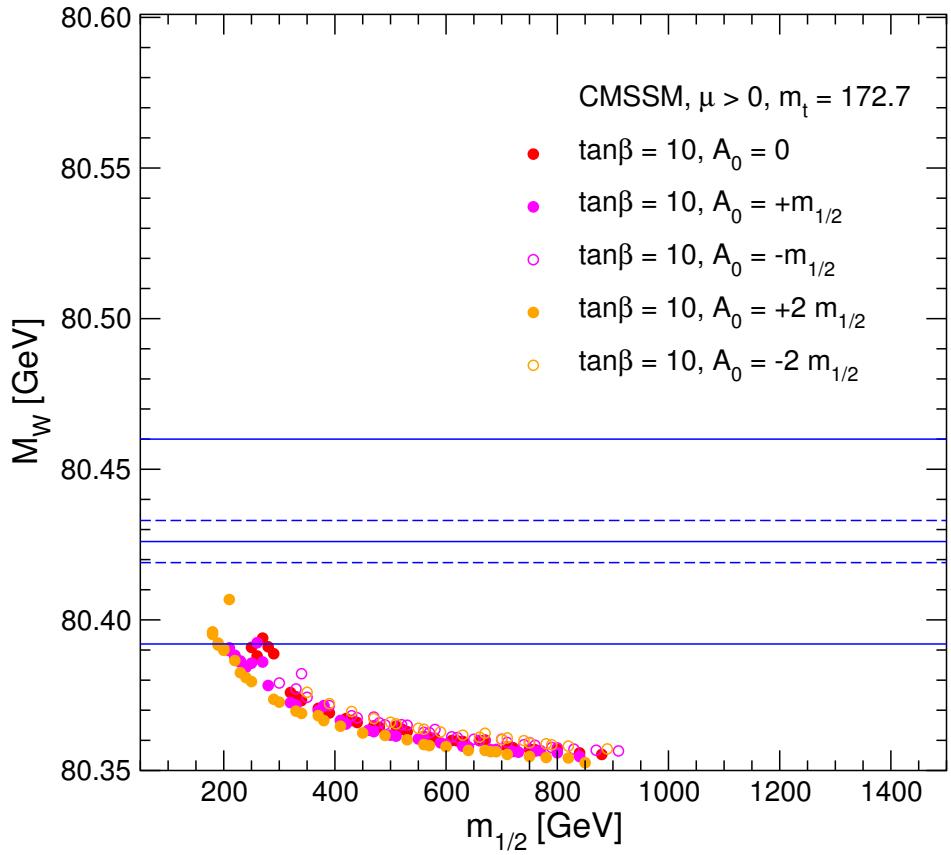
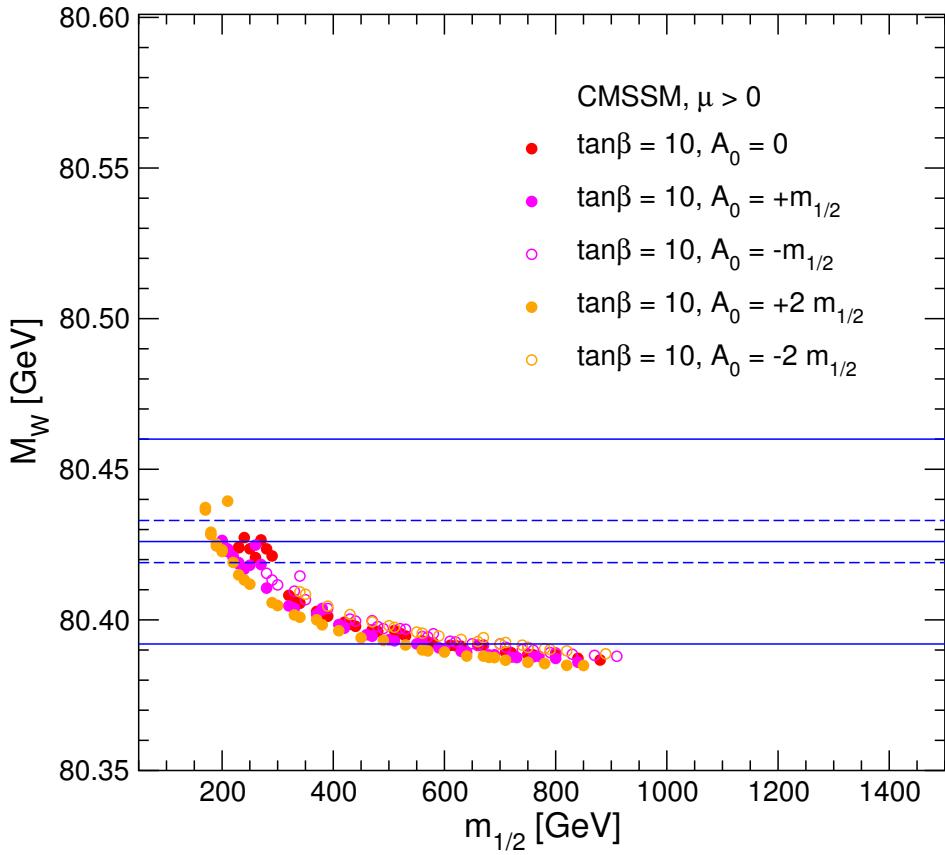
# **CMSSM prediction for $M_W$ vs. current precision and ILC (MegaW) for $m_t = 178.0$ GeV**

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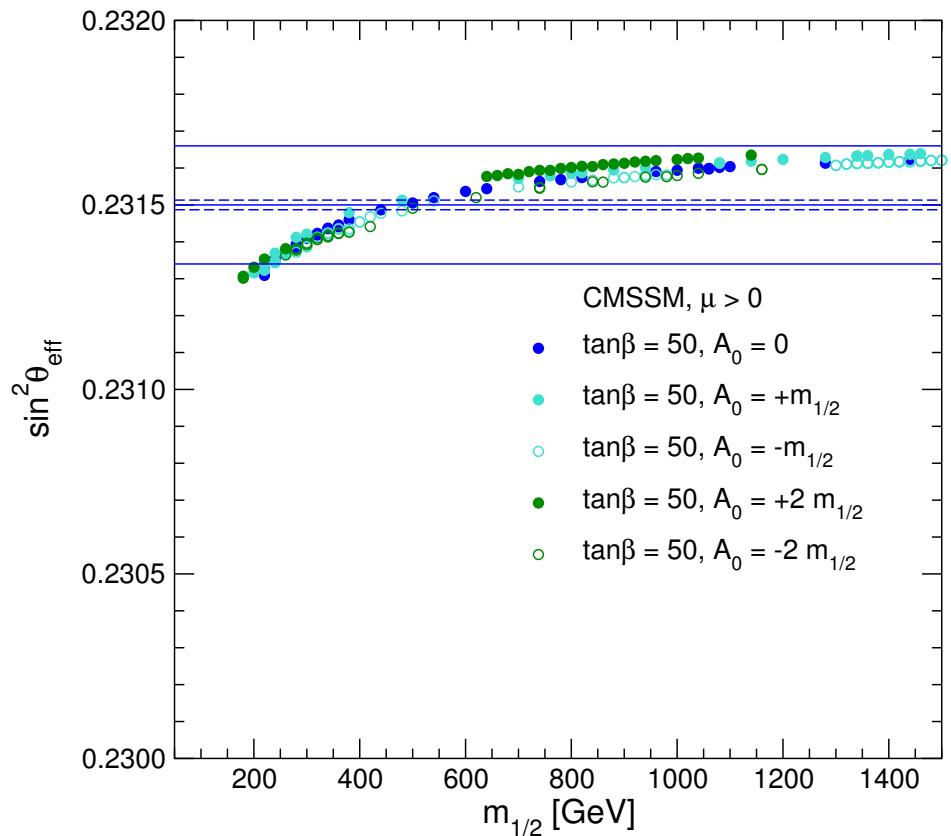
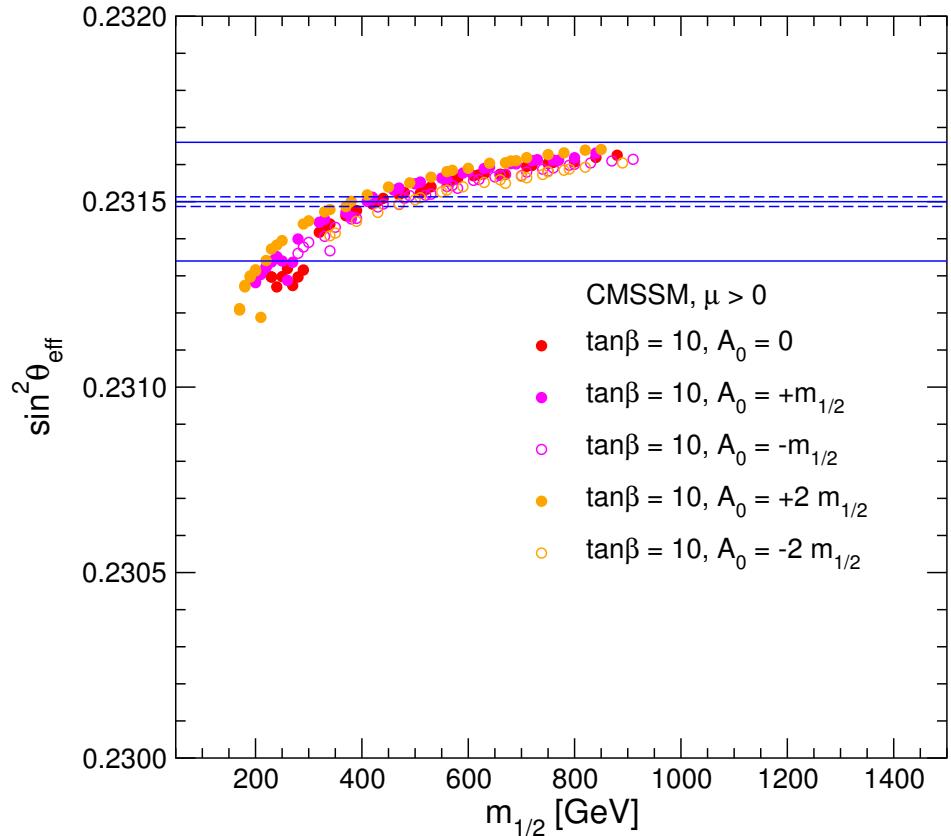
⇒ Relatively small values of  $m_{1/2}$  favoured  
great improvement at the ILC

# **Comparison: $m_t = 178.0 \text{ GeV (left) vs. } m_t = 172.7 \text{ GeV (right)}$**



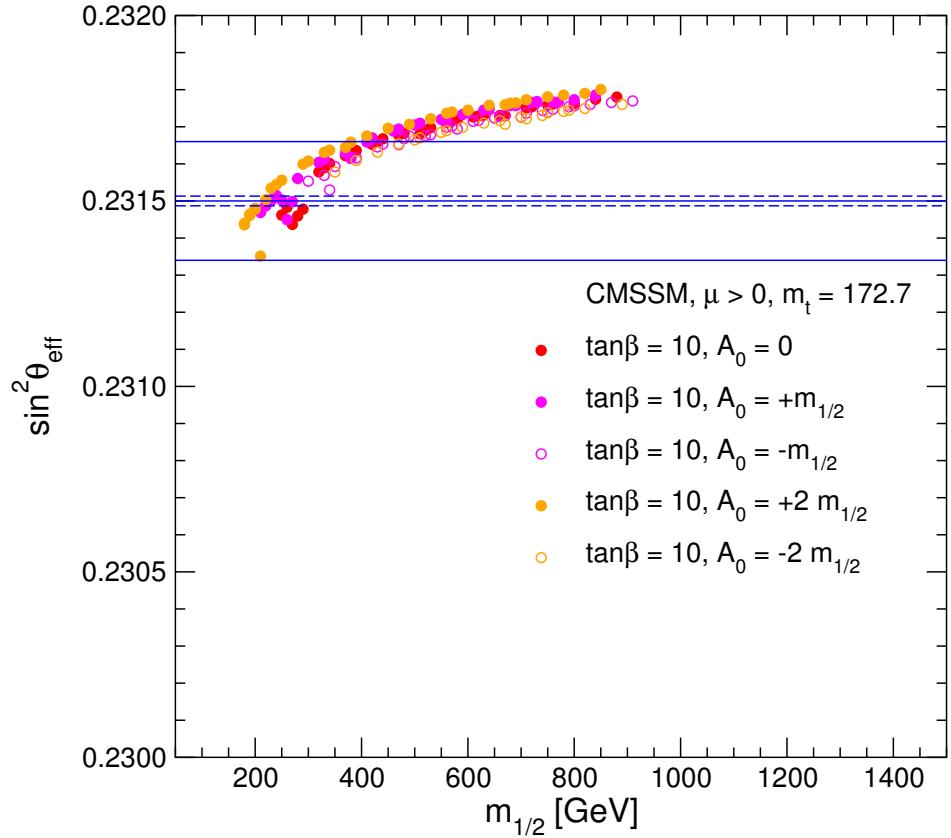
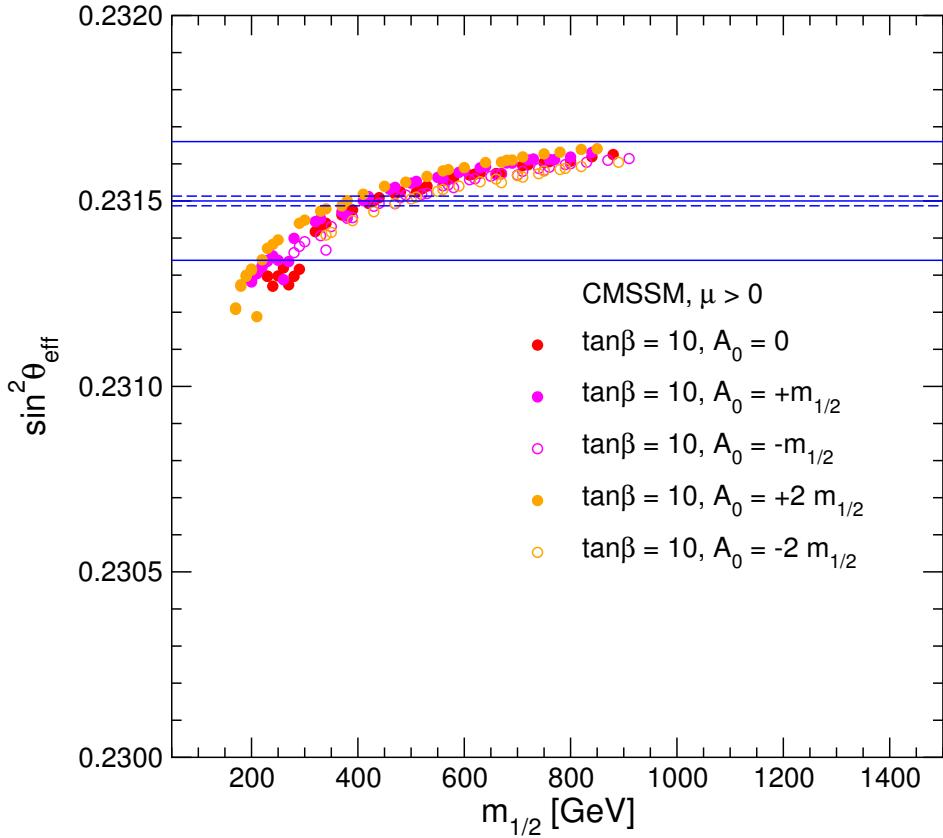
⇒ Lower  $m_t$  yields lower  $M_W$   
 preference for small values of  $m_{1/2}$  becomes more pronounced

# ***CMSSM prediction for $\sin^2 \theta_{\text{eff}}$ vs. current precision and ILC (GigaZ)***



⇒ Relatively small values of  $m_{1/2}$  favoured  
great improvement at the ILC

# **Comparison: $m_t = 178.0 \text{ GeV (left) vs. } m_t = 172.7 \text{ GeV (right)}$**



⇒ Lower  $m_t$  yields higher  $\sin^2 \theta_{\text{eff}}$   
 preference for small values of  $m_{1/2}$  becomes more pronounced

# **The anomalous magnetic moment of the muon:**

$$(g - 2)_\mu \equiv 2a_\mu$$

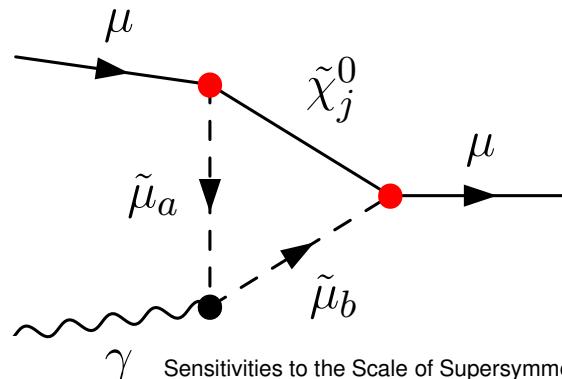
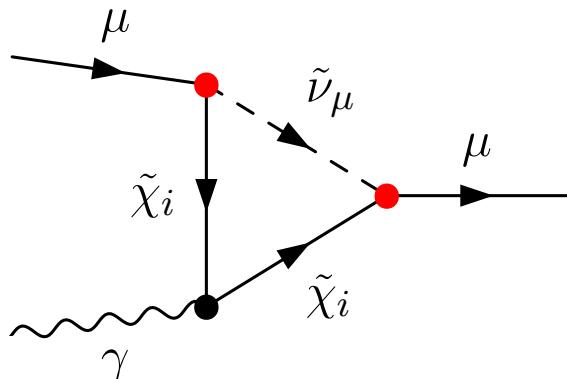
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Experimental result for  $a_\mu$  vs. SM prediction (using  $e^+e^-$  data for hadronic vacuum polarization contributions):

$$a_\mu^{\text{exp}} - a_\mu^{\text{theo}} = (25.2 \pm 9.2) \times 10^{-10} : 2.7\sigma .$$

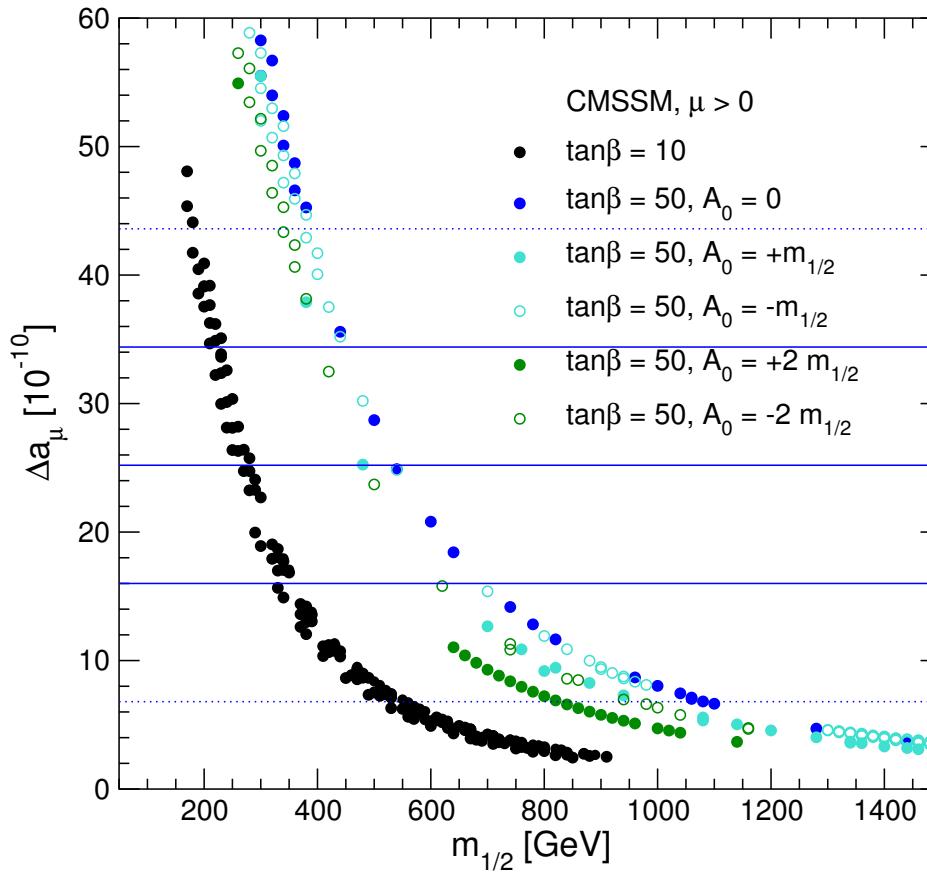
Better agreement between theory and experiment possible in models of physics beyond the SM

Example: one-loop contributions of superpartners of fermions and gauge bosons



# **CMSSM prediction for $\Delta a_\mu$ vs. current precision ( $1\sigma$ and $2\sigma$ bands)**

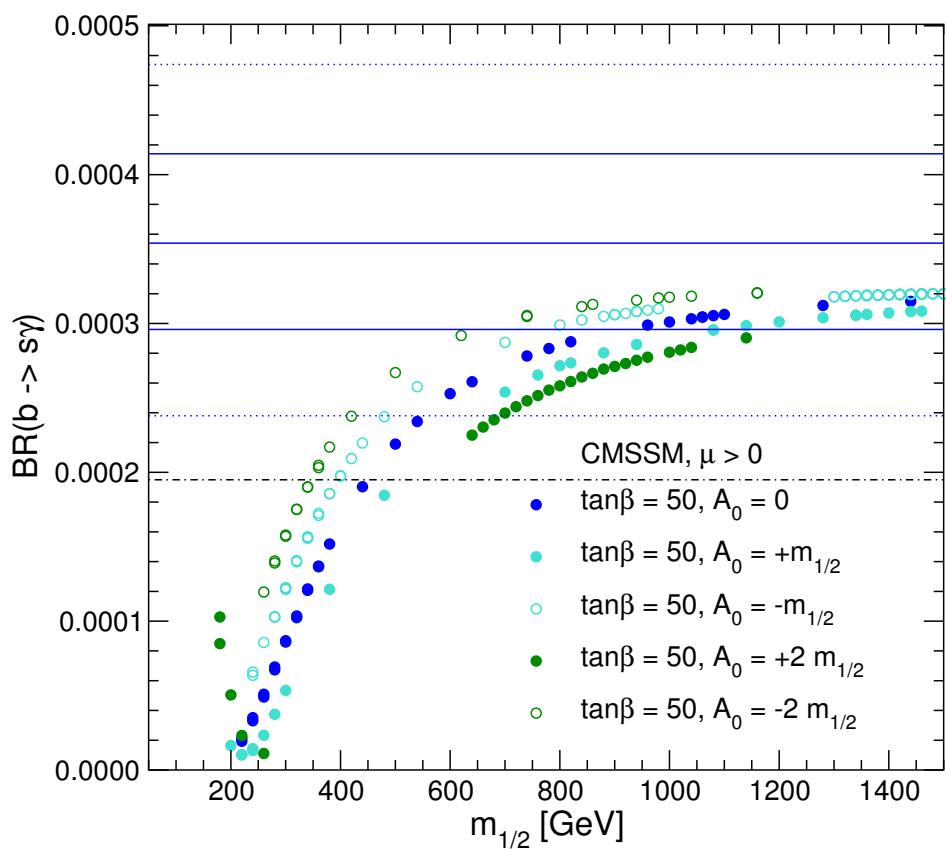
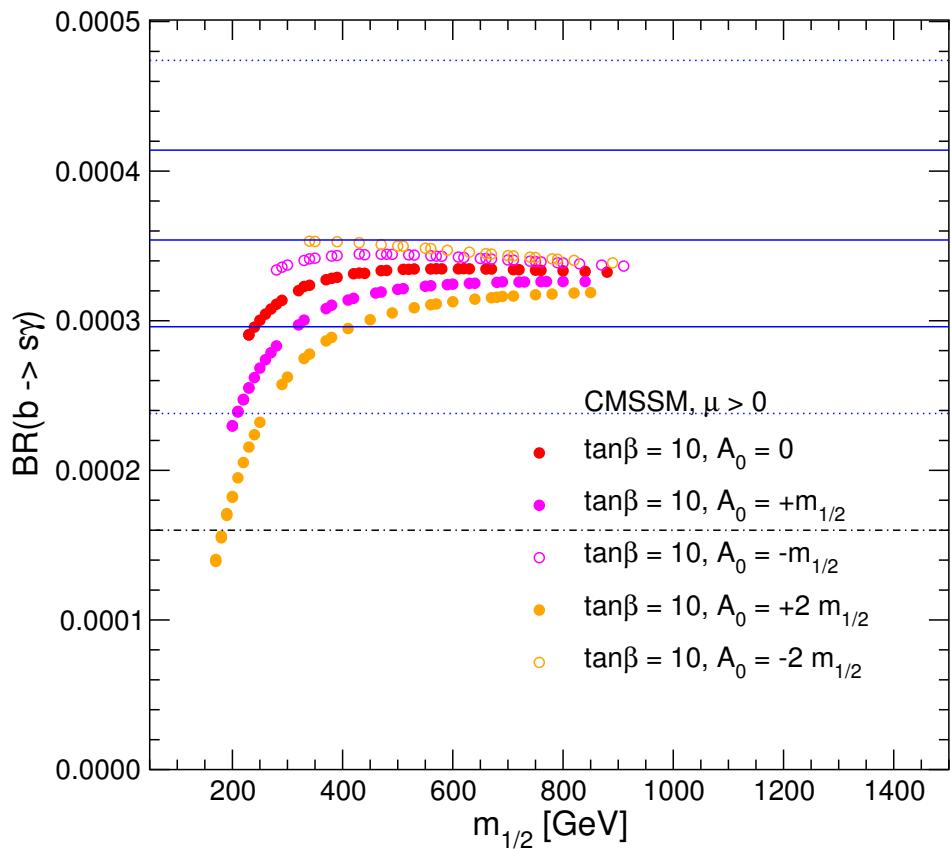
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⇒ For  $\tan\beta = 10$ : relatively small values of  $m_{1/2}$  favoured

# **CMSSM prediction for $\text{BR}(b \rightarrow s\gamma)$ vs. current precision (1 $\sigma$ and 2 $\sigma$ bands)**

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⇒ For  $\tan\beta = 10$ : relatively small values of  $m_{1/2}$  allowed

# ***Higgs mass prediction in the MSSM:***

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Prediction for  $M_h$ ,  $M_H$ , ...

Tree-level result for  $M_h$ ,  $M_H$ :

$$m_{H,h}^2 = \frac{1}{2} \left[ M_A^2 + M_Z^2 \pm \sqrt{(M_A^2 + M_Z^2)^2 - 4M_Z^2 M_A^2 \cos^2 2\beta} \right]$$

$\Rightarrow M_h \leq M_Z$  at tree level

MSSM tree-level bound (gauge sector): excluded by LEP!

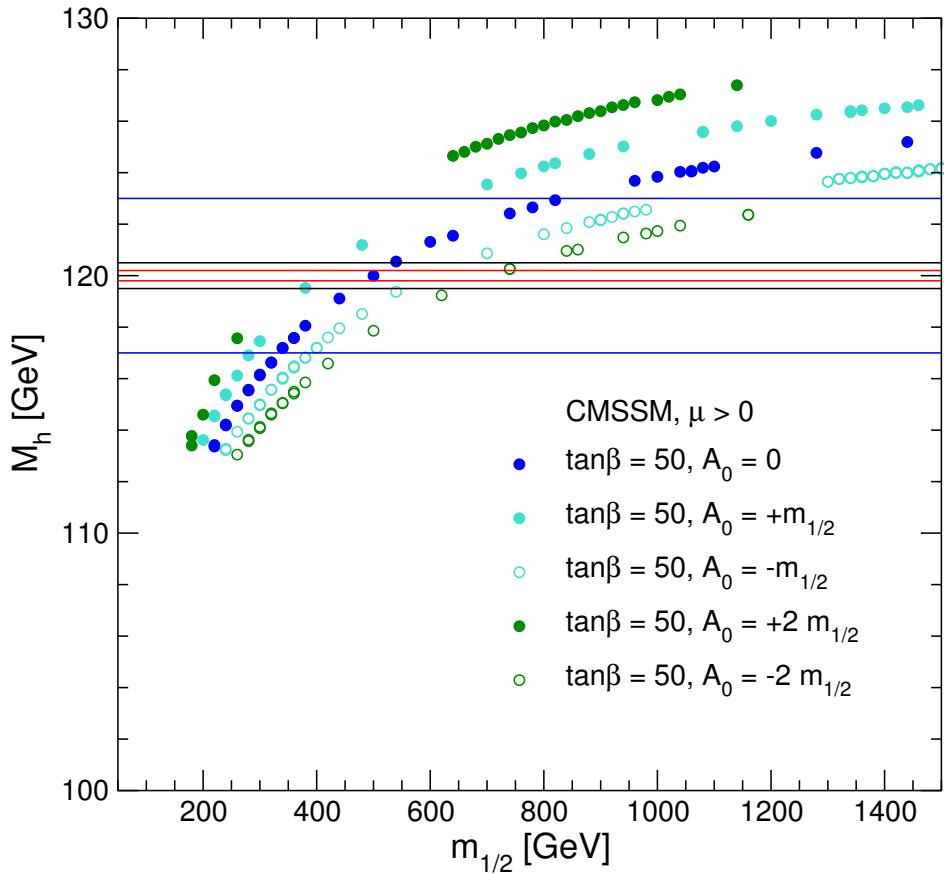
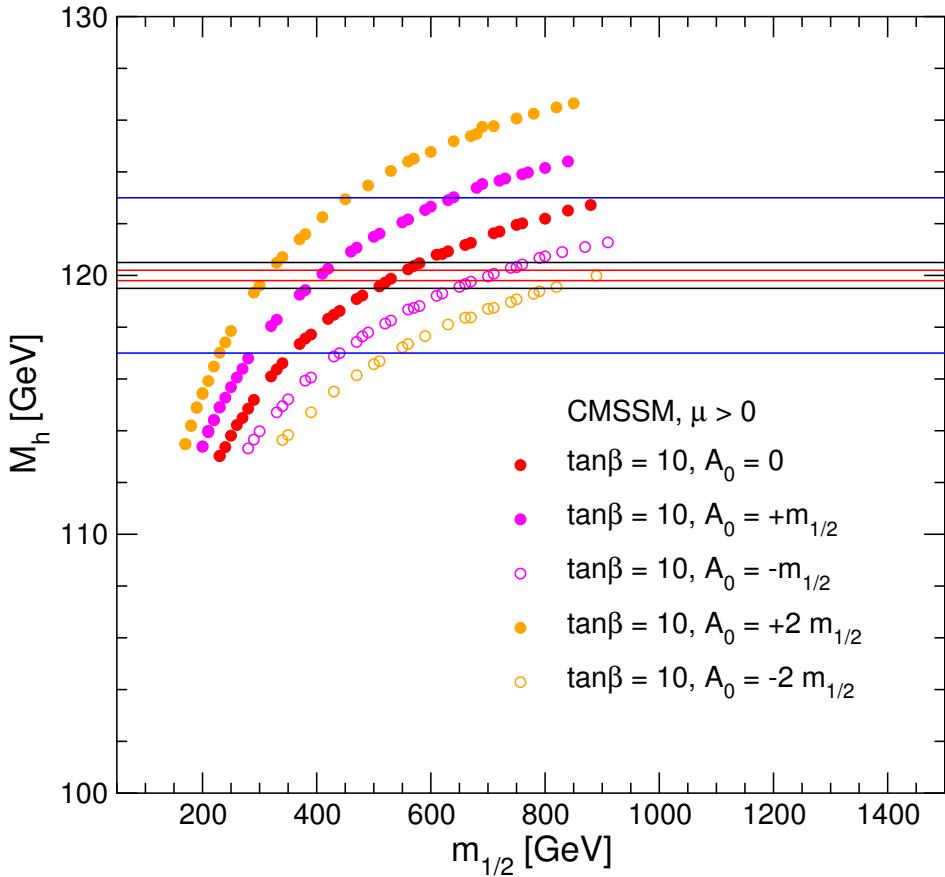
Large radiative corrections (Yukawa sector, ...):

Yukawa couplings:  $\frac{e m_t}{2 M_W s_W}$ ,  $\frac{e m_t^2}{M_W s_W}$ , ...

$\Rightarrow$  Dominant one-loop corrections:  $G_\mu m_t^4 \ln \left( \frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{m_t^2} \right)$ ,  $\mathcal{O}(100\%)$  !

# **CMSSM prediction for $M_h$ vs. assumed experimental value for current and different future theoretical uncertainties**

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- ⇒ High sensitivity to variations of  $m_{1/2}, A_0$
- ⇒ constraints on SUSY parameter space

### **3. Combined sensitivity: present situation and ILC precision**

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Combined sensitivity investigated for  
present situation + ILC precision

Two kinds of fits:

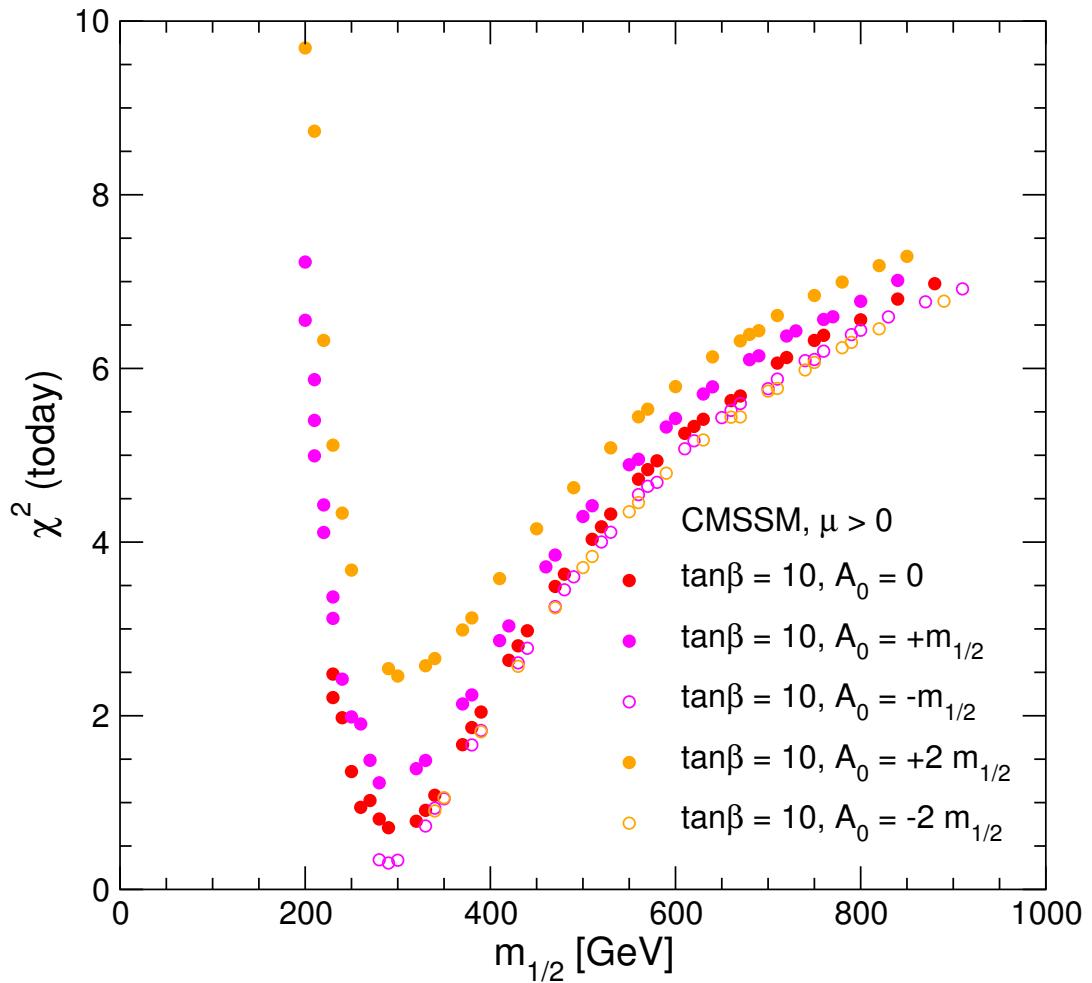
- $\chi^2$  fit for fixed  $A_0$
- $\chi^2$  fit in  $(m_{1/2}, A_0)$  plane  
(scan of CMSSM parameter space)

# $\chi^2$ fit in CMSSM with dark matter constraints:

$M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $(g - 2)_\mu$ ,  $\text{BR}(b \rightarrow s\gamma)$ , present situation

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$\tan \beta = 10$ :



⇒ very good description  
of the data

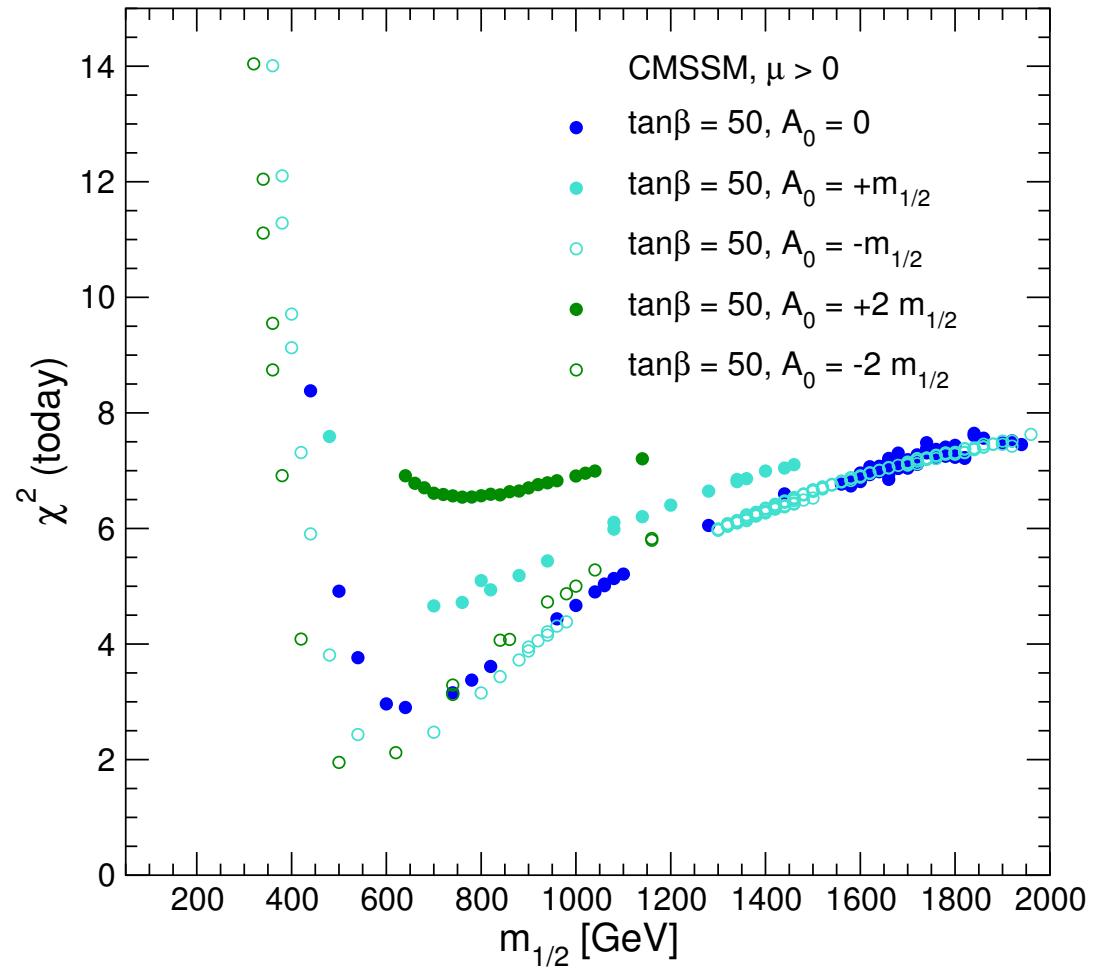
preference for relatively  
small mass values

# $\chi^2$ fit in CMSSM with dark matter constraints:

$M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $(g - 2)_\mu$ ,  $\text{BR}(b \rightarrow s\gamma)$ , present situation

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$\tan \beta = 50$ :

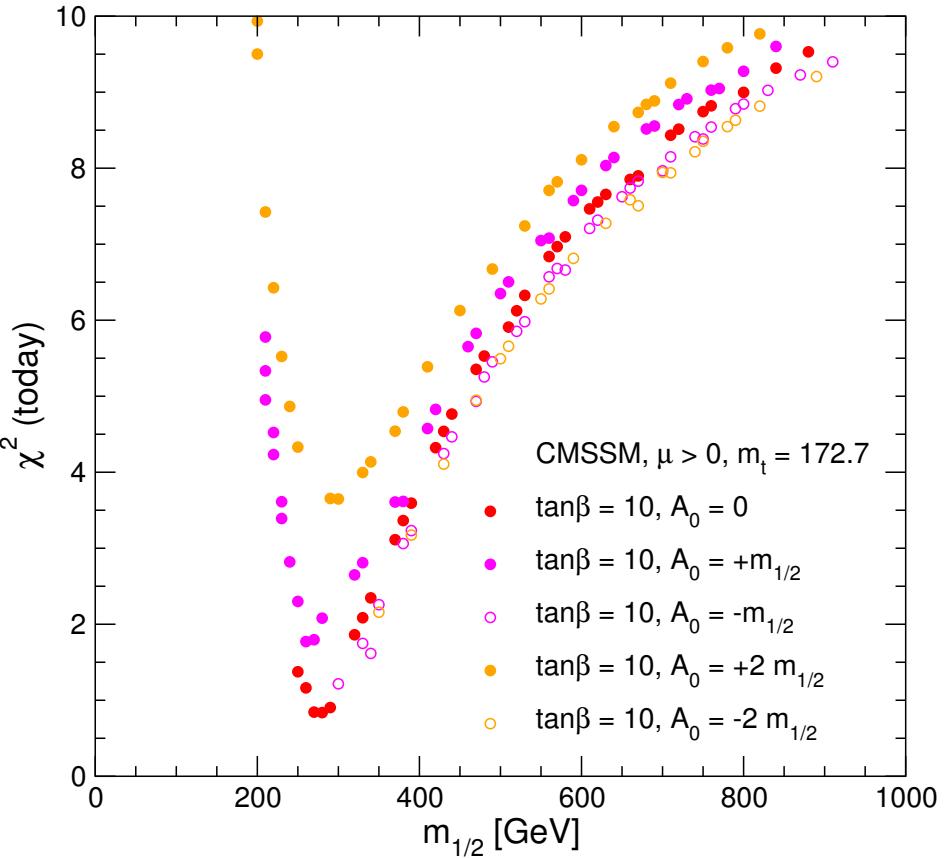
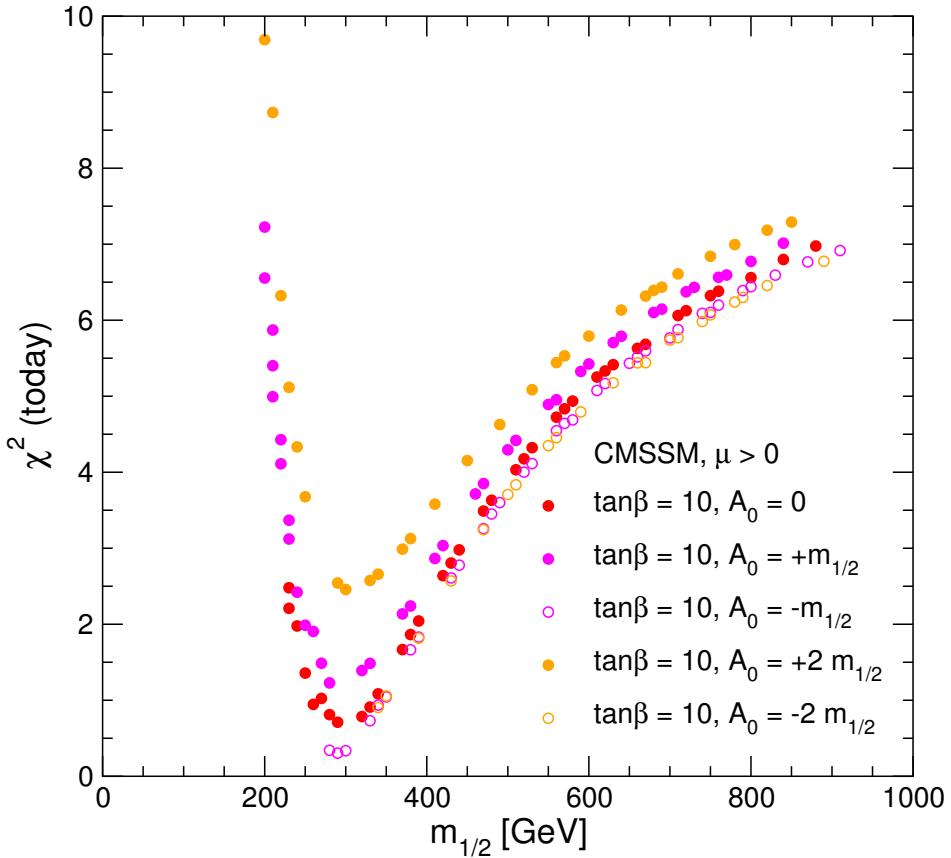


⇒ worse fit quality

preferred  $m_{1/2}$  values  
larger by 200–300 GeV  
compared to  $\tan \beta = 10$   
case

**Comparison:**  $m_t = 178.0 \pm 4.3$  GeV (**left**)

**vs.**  $m_t = 172.7 \pm 2.9$  GeV (**right**)



⇒ best fit value shifted to slightly smaller  $m_{1/2}$

slightly larger minimum  $\chi^2$  value

higher precision on  $m_t$  yields tighter bounds on  $m_{1/2}$

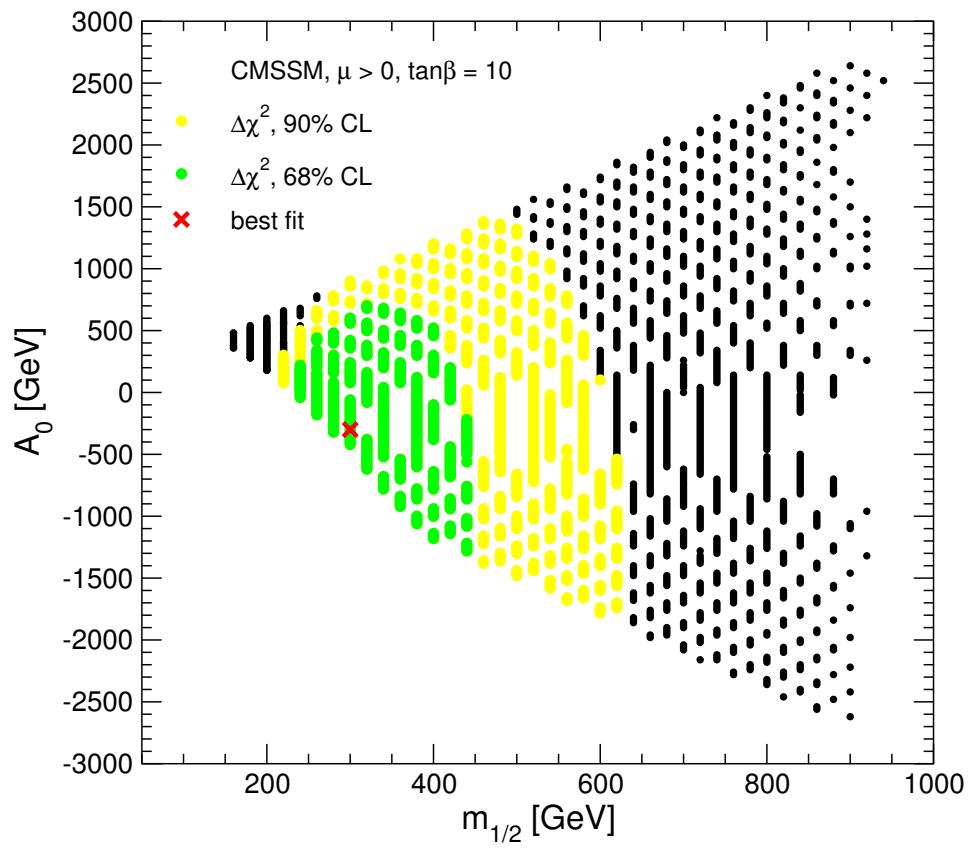
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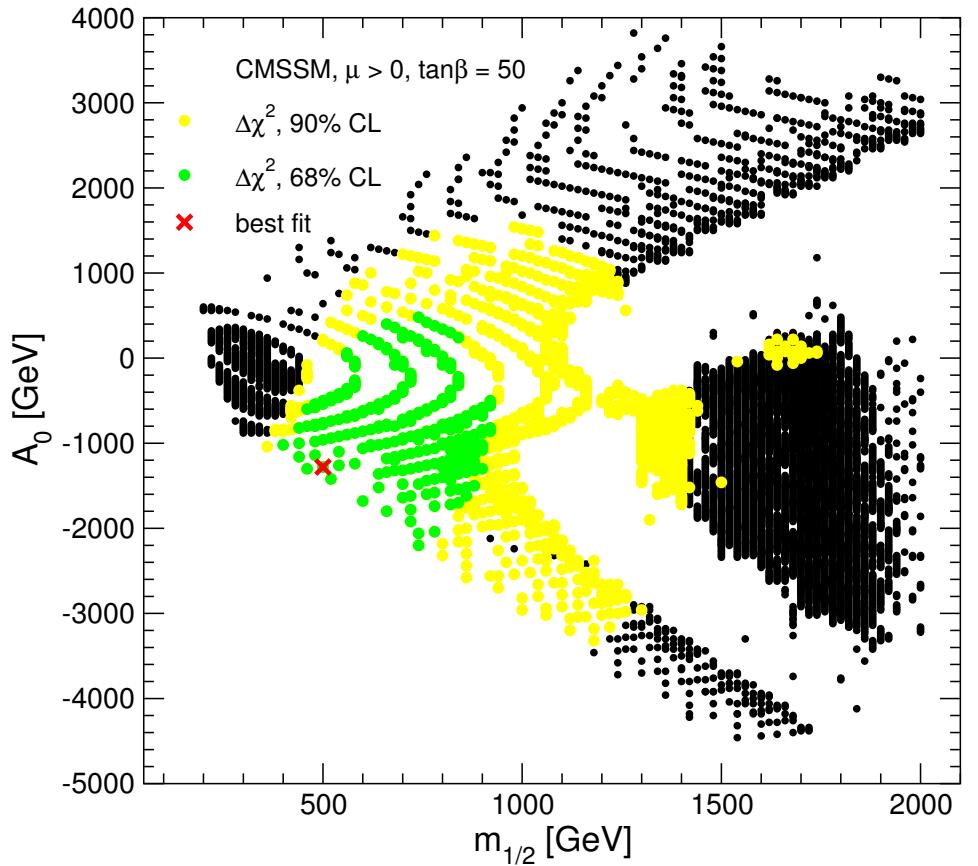
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68% and 90% C.L. regions in  $m_{1/2}$ – $A_0$  plane:

$\tan \beta = 10$ :



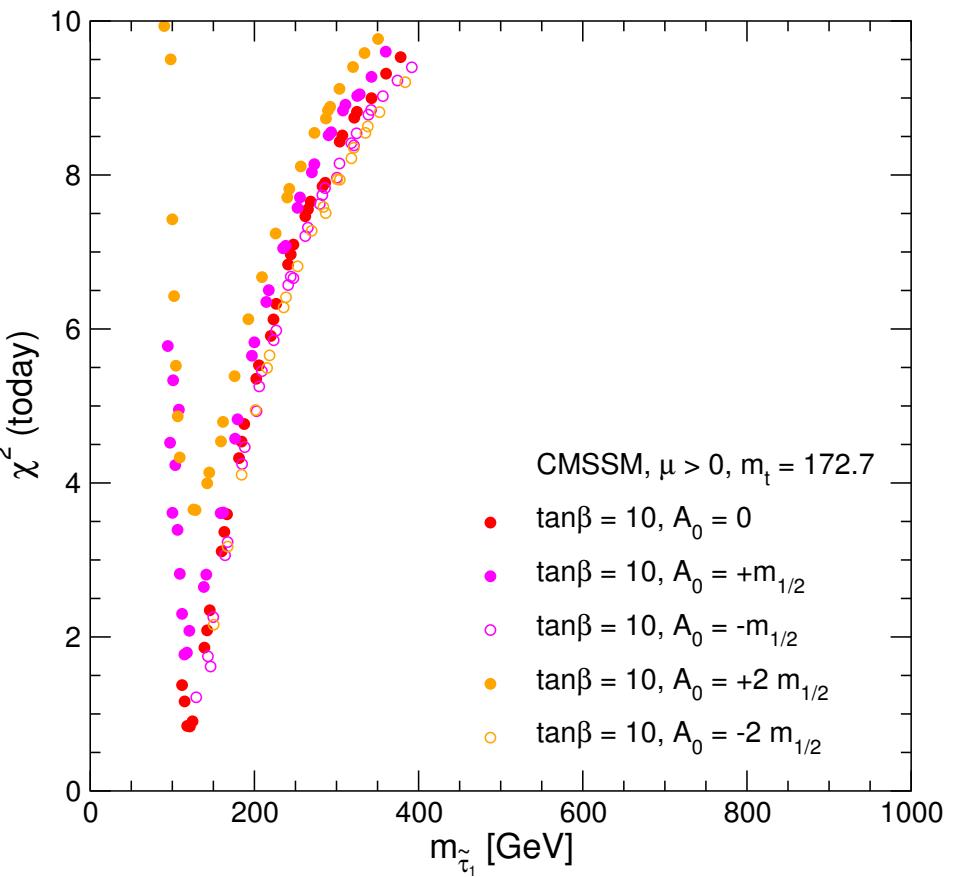
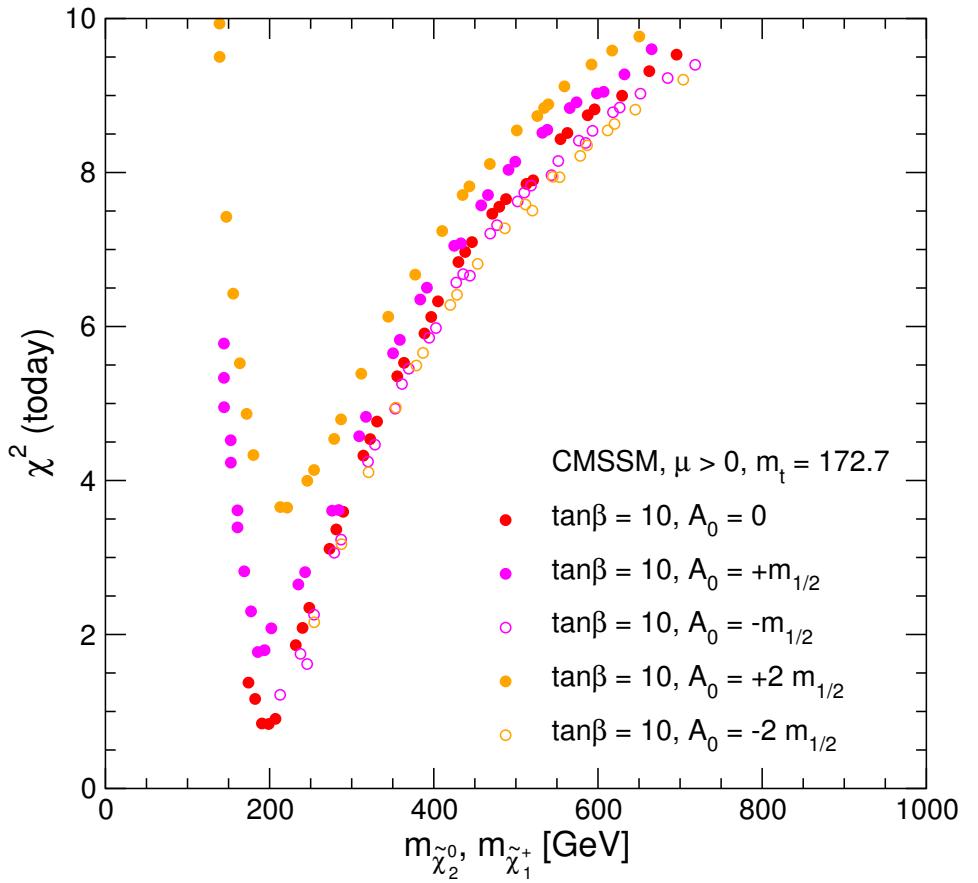
$\tan \beta = 50$ :



# **Fit results (present) for particle masses,**

$\tan \beta = 10$ :  $m_{\tilde{\chi}_1^+} \approx m_{\tilde{\chi}_2^0}$ ,  $m_{\tilde{\tau}_1}$ ,  $m_t = 172.7 \pm 2.9$  GeV

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⇒ Good prospects for the LHC and ILC

## ***Comparison with other recent analysis***

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Fit involving all CMSSM parameters, but omitting electroweak precision observables  $M_W$  and  $\sin^2 \theta_{\text{eff}}$

[*B. Allanach, C. Lester '05*]

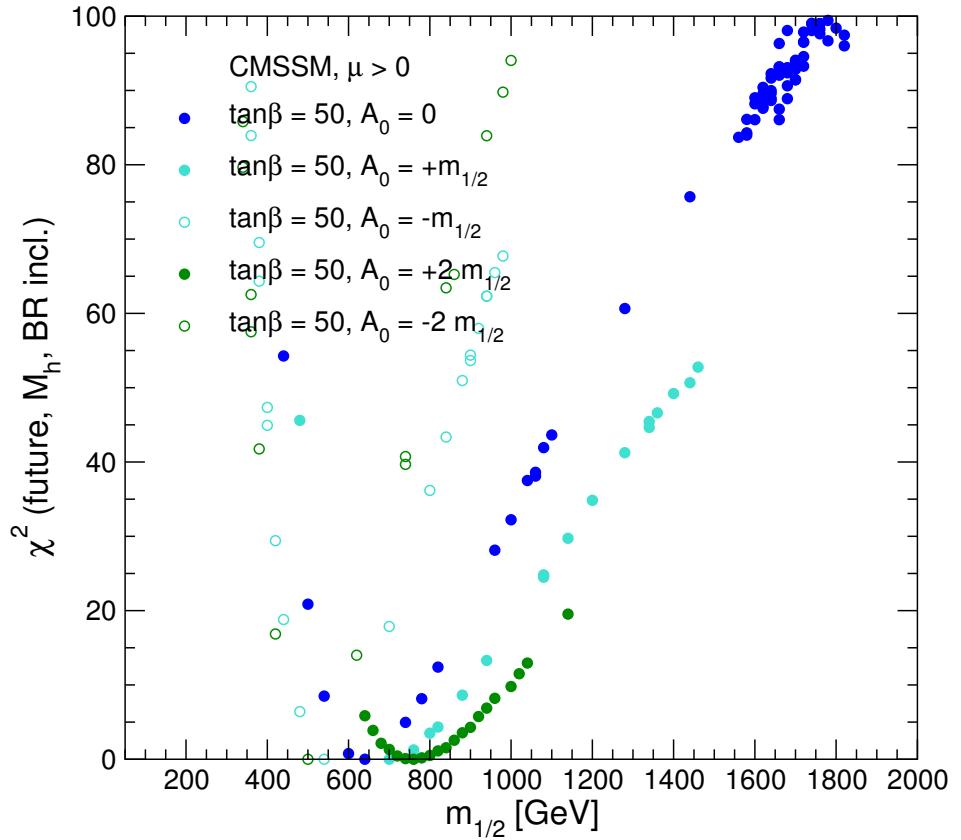
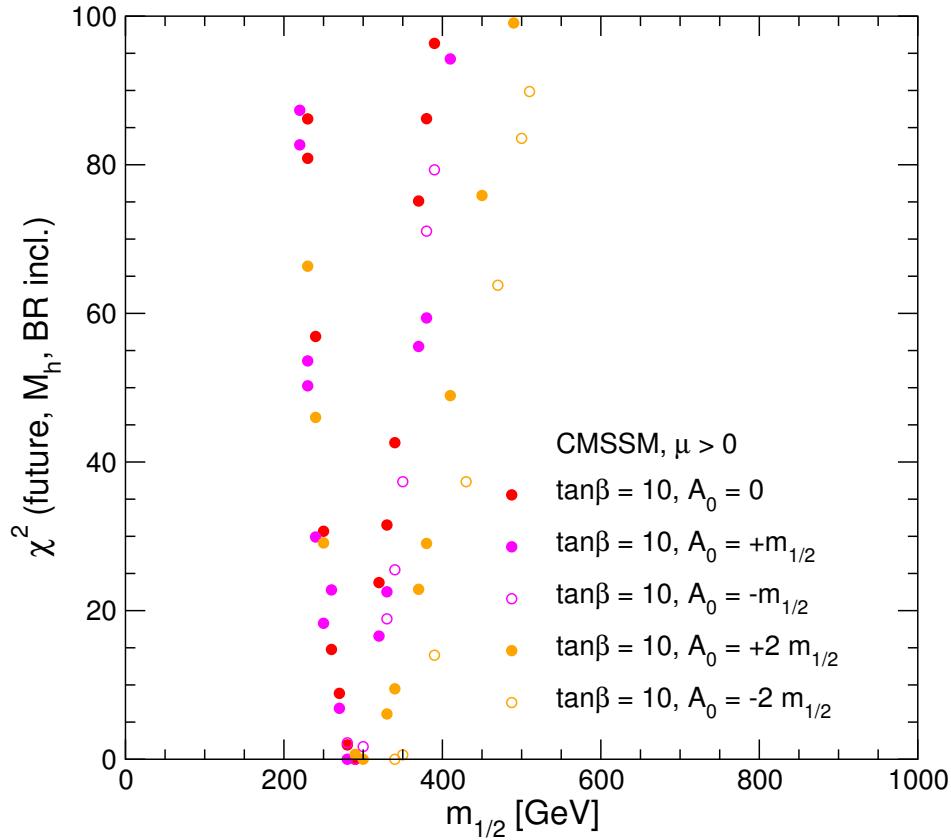
⇒ preference for larger values of  $\tan \beta$  and somewhat larger  $m_{1/2}$

⇒  $M_W$  and  $\sin^2 \theta_{\text{eff}}$  are crucial for the fit,  
give rise to preference for relatively small  $m_{1/2}$

$\chi^2$  fit in CMSSM with CDM constraints:  $M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $(g - 2)_\mu$ ,

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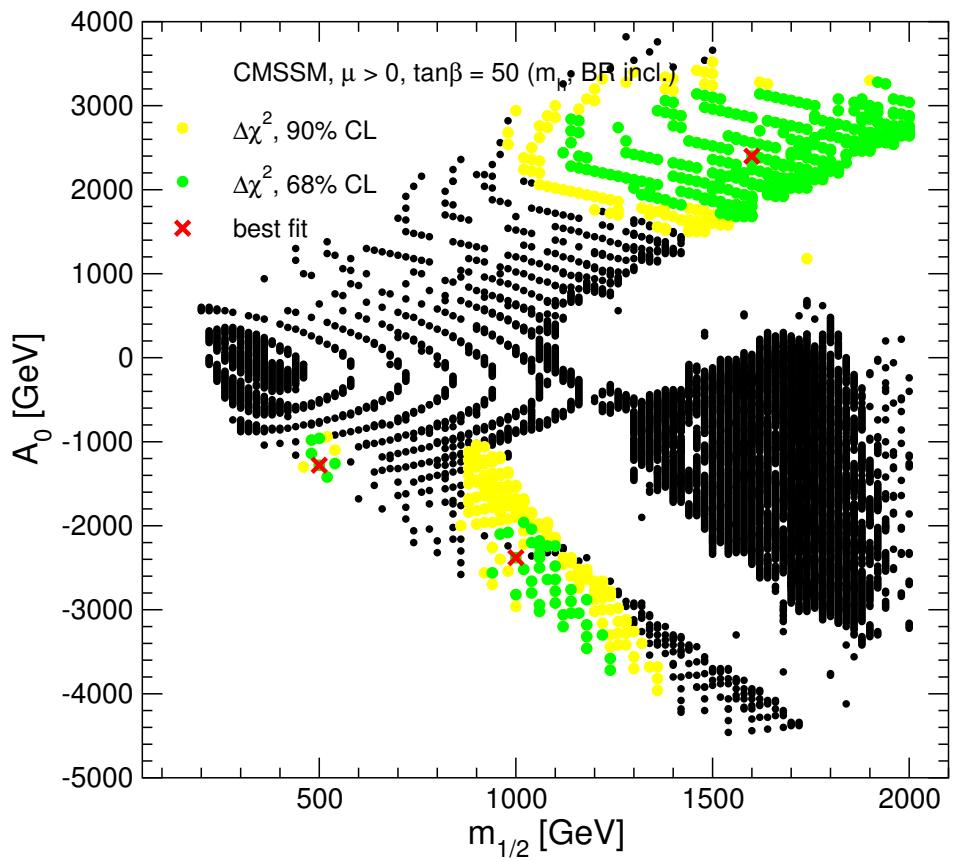
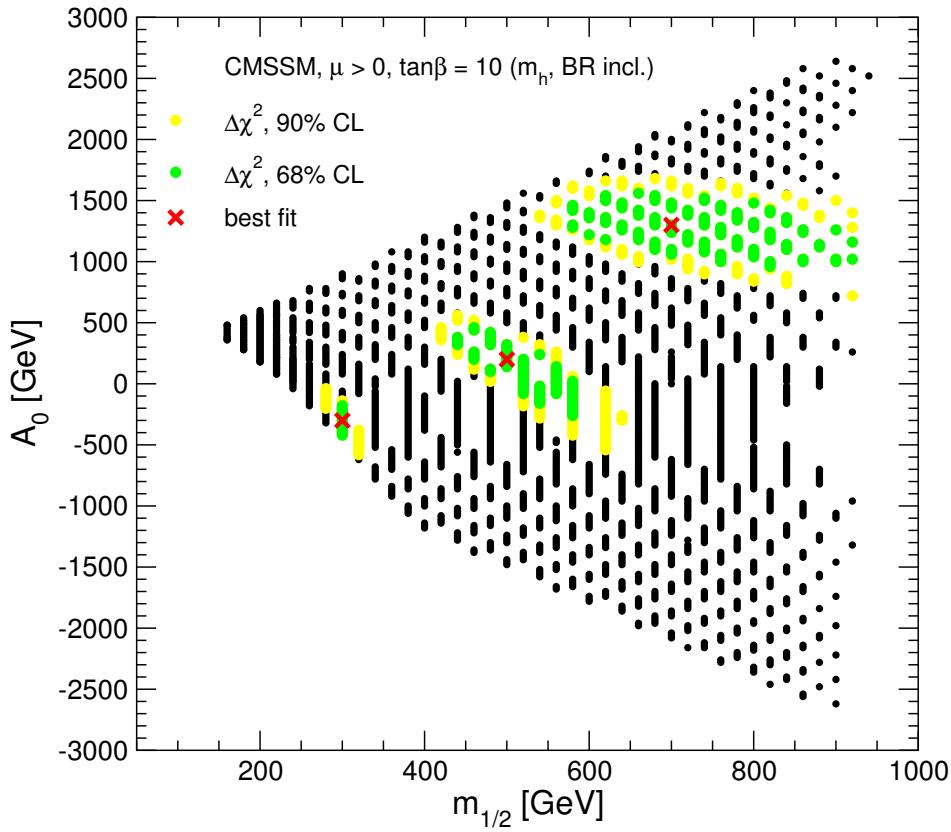


⇒ Great increase in sensitivity

⇒ tight constraints on particle masses

$\chi^2$  fit in CMSSM with CDM constraints:  $M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $(g - 2)_\mu$ ,

$\text{BR}(b \rightarrow s\gamma)$ ,  $M_h$ ,  $\text{BR}(h \rightarrow b\bar{b})/\text{BR}(h \rightarrow WW^*)$ , ILC precision



⇒ ILC precision greatly improves sensitivity to  $m_{1/2}$ ,  $A_0$   
high indirect sensitivity up to  $m_{1/2} \lesssim 1$  TeV

## ***4. Conclusions***

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**Comparison with direct experimental information  $\Rightarrow$  test of CMSSM at the loop level**