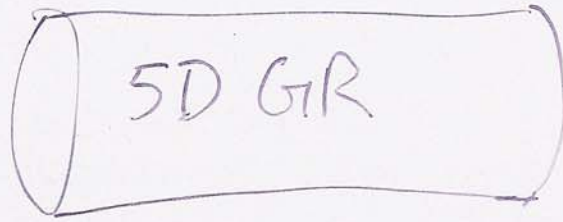


# XD General Relativity (GR)



$$M_{MN} \longrightarrow G_{MN}(x^M, \theta)$$

$$\downarrow E \ll R$$

$$G_{MN}^{(0)}(x) / \sqrt{2\pi R}$$

4D  
Massless fields  
on plugging into  
5D Einstein action

$$\equiv \frac{G_{\mu\nu}^{(0)}(x)}{\sqrt{2\pi R}}, \quad \frac{G_{\mu 5}^{(0)}(x)}{\sqrt{2\pi R}}, \quad \frac{G_{55}^{(0)}(x)}{\sqrt{2\pi R}}$$

effective 4D metric      4D vector      4D scalar

4D vector  $\equiv$  4D gauge field for KK charge  $n$ .

Can this KK mechanism be origin of SM gauge fields?

Idea has not met with success in field theory, but mechanism can be seen to survive in construction of heterotic string theory (with some stringy enhancements).

4D Gravity:  $G_{N_{d+1}} = G_5 / 2\pi R$

# The RADION

4D scalar  $\equiv G_{55}^{(0)}(x)$  has no potential

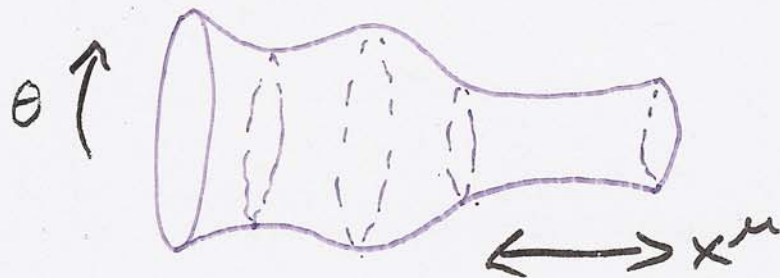
(all terms in 5D Einstein action

$$\Rightarrow \partial_\mu \text{ or } \frac{\delta}{\delta G_{MN}^{(0)}(x)} G_{MN}^{(0)}(x) \rightarrow 0)$$

Geometrically, these fluctuations  $\equiv$

$$ds^2 \ni G_{55}^{(0)}(dx^5)^2 = G_{55} R^2 d\theta^2$$

$\Rightarrow G_{55}^{(0)}(x) \equiv$  dynamical XD radius



$\therefore \langle R^{(x)}_{\text{physical}} \rangle$  undetermined, & yet  
 $\equiv \langle \sqrt{G_{55}} \rangle R$  different VEVs are  
 physically distinct

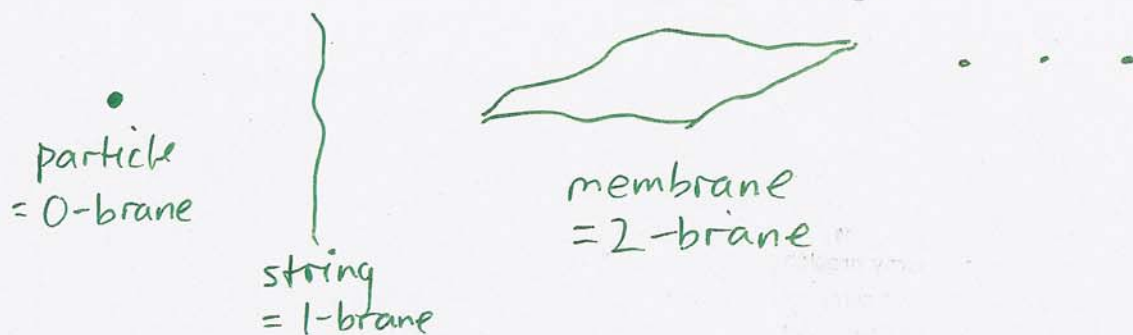
Eg. Adding 5D scalar,

$$\lambda_4 \ll \lambda_5 = \frac{\lambda_5}{2\pi \langle R_{\text{phys}} \rangle} \quad \text{undetermined}$$

Such massless scalars are MODULI & greatly limit predictivity until VEVs determined.

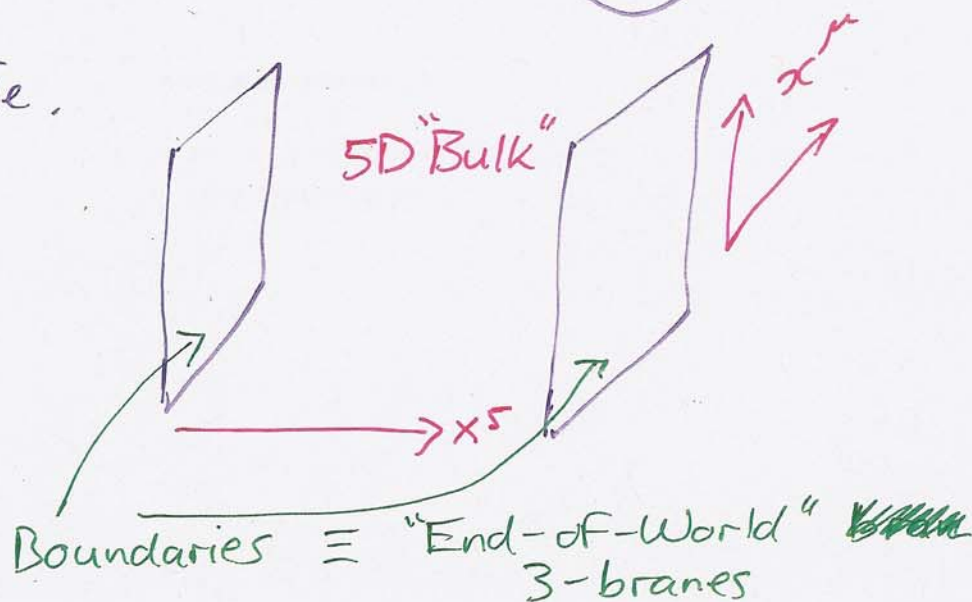
# BRANES

$\equiv$  localized defects in higher dimensions



Eg. Suppose  $XD$    $\rightarrow$  

ie.

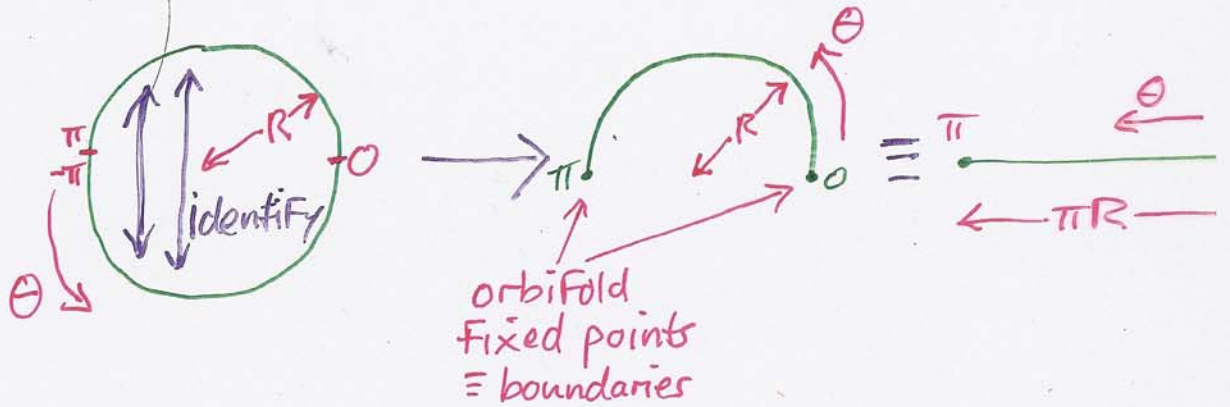


5D fields must  $\therefore$  have boundary conditions prescribed.



# Orbifold boundary conditions

PRETEND still  
live on XD



$$\phi(x, \theta) = (-1)^P \phi(x, -\theta)$$

orbifold parity  
respected by action

## Brane localized interactions

allow different dynamics at branes  
than in bulk.

$$\text{Eg. } S = \int d^4x \int_{-\pi}^{\pi} d\theta R \left\{ \frac{1}{2} (\partial_M \phi)^2 - \frac{m_s^2}{2} \phi^2 - \kappa \phi \delta(x_5) - \kappa \phi \delta(x_5 - \pi R) \right\}$$

$$\Rightarrow \text{Equations } -\partial_M \partial^M \phi + \frac{\partial_\theta^2}{R^2} \phi - m_s^2 \phi = \frac{\kappa \delta(\theta)}{R} + \frac{\kappa \delta(\theta - \pi)}{R}$$

⇒ Non-trivial Vacuum

$$\langle \phi(x, \theta) \rangle = \bar{\phi}(\theta),$$

$$\frac{1}{R^2} \partial_\theta^2 \bar{\phi} - m_s^2 \bar{\phi} = \frac{\kappa \delta(\theta)}{R} + \frac{\kappa \delta(\theta - \pi)}{R}$$

Ansatz:  $\bar{\phi} = A \left[ e^{m_s R \theta} + e^{m_s R(\pi - \theta)} \right]$

N.B. correct symmetries  
&  $2\pi$ -periodicity.

Matching  $\delta$ -Functions,

$$2A(1 - e^{m_s R \pi}) \frac{m_s}{R} = \frac{\kappa}{R}$$

Plugging  $\bar{\phi}$  back into action,

$$\begin{aligned} S &= \int d^4x \int_{-\pi}^{\pi} d\theta R \left\{ -\frac{\kappa \bar{\phi} \delta(\theta)}{R} - \frac{\kappa \bar{\phi} \delta(\theta - \pi)}{R} \right\} \\ &= -\int d^4x \frac{\kappa^2}{2m_s} \frac{1 + e^{m_s \pi R}}{1 - e^{m_s \pi R}} \end{aligned}$$

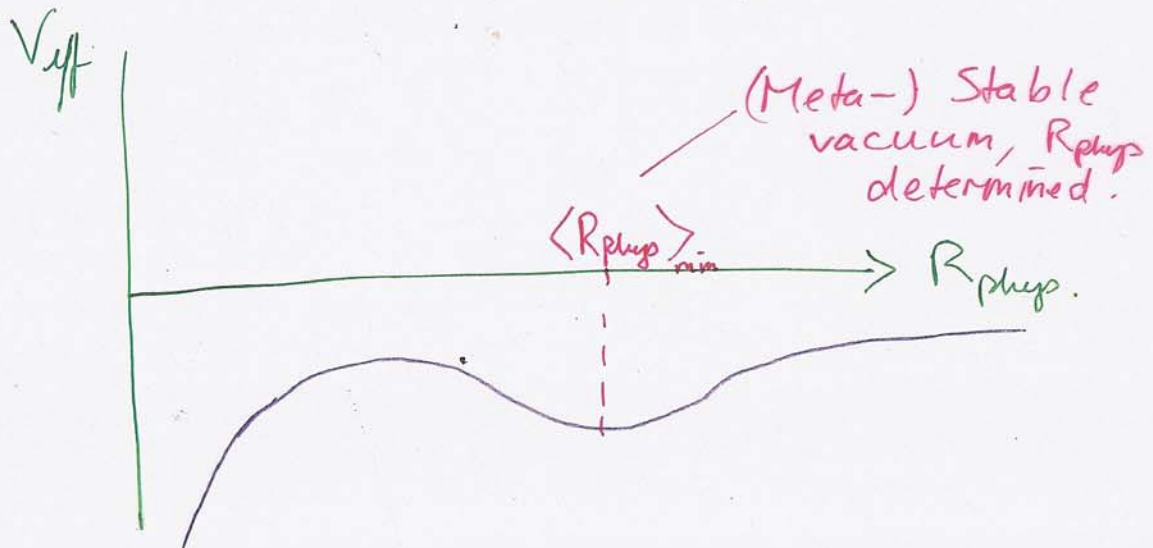
Looks like irrelevant constant, BUT  
with 5D GR added...

# RADION POTENTIAL

$$V_{\text{eff}}(R_{\text{phys}}^{(x)}) \equiv \frac{\kappa^2}{2m_5} \frac{1 + e^{m_5 \pi R_{\text{phys}}}}{1 - e^{m_5 \pi R_{\text{phys}}}}$$

Adding contribution of a second scalar, with opposite sign but equal magnitude brane terms,

$$V_{\text{eff}}(R_{\text{phys}}^{(x)}) = \frac{\kappa^2}{2m_5} \frac{1 + e^{m_5 \pi R_{\text{phys}}}}{1 - e^{m_5 \pi R_{\text{phys}}}} + \frac{\tilde{\kappa}^2}{2\tilde{m}_5} \frac{1 - e^{\tilde{m}_5 \pi R_{\text{phys}}}}{1 + e^{\tilde{m}_5 \pi R_{\text{phys}}}}$$

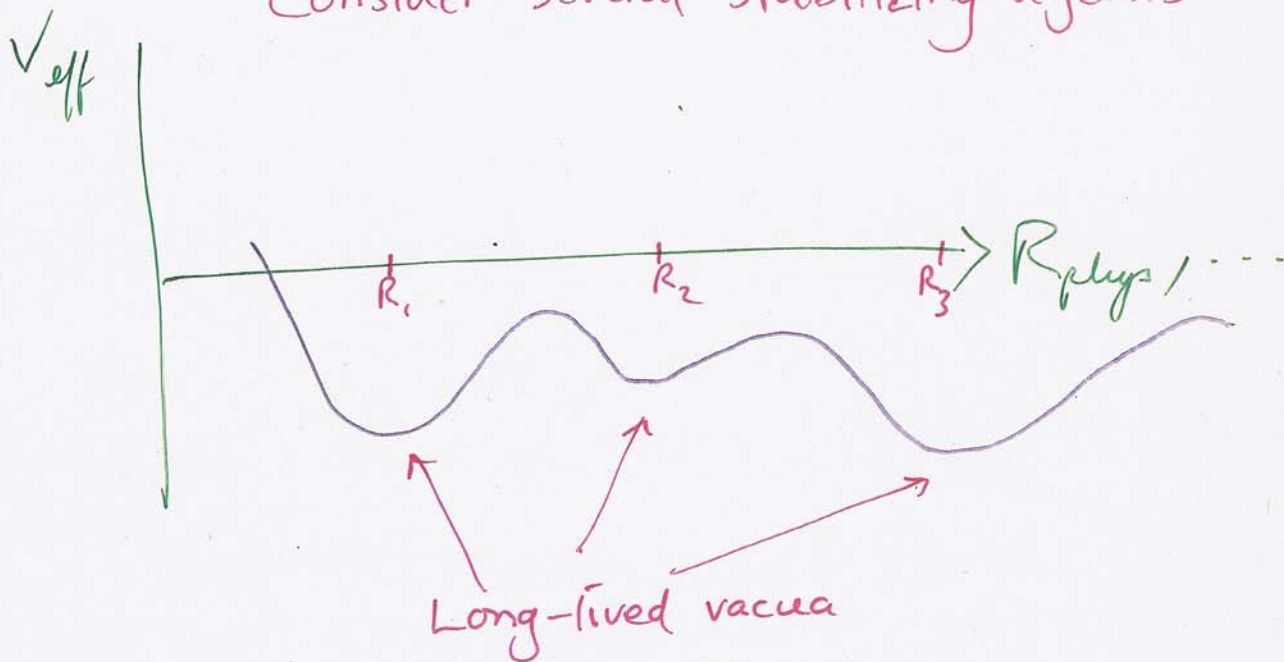




# THE "LANDSCAPE"

## OF XD SHAPES & SIZES

Consider several stabilizing agents



$\Rightarrow$   $\infty$  different possible low-energy worlds

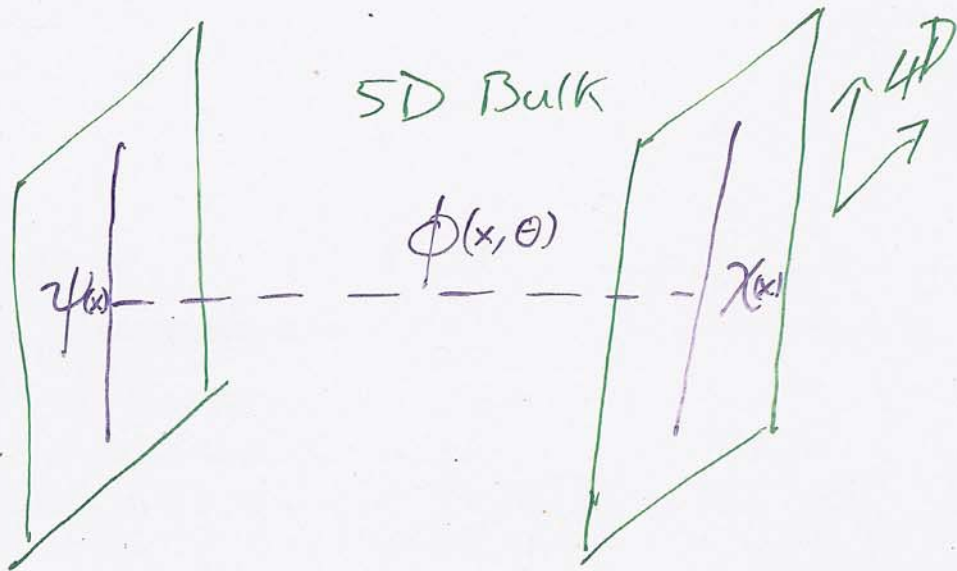
$$\lambda_{4d}(R_i)$$

??

String theory has run into this loss of predictivity.

# BRANEWORLD

Some degrees of freedom may reside (localized) entirely on the defects,



Eg.  $\psi \equiv$  Standard Model ~~(+ gravity)~~  
 $\phi \equiv$  gravity  
 $\chi \equiv$  "Hidden/Dark" sector.

Large Extra Dimensions

Arkani-Hamed, Dimopoulos, Dvali '98

Gravity appears weak because of XD "dilution"

$$G_N \sim \frac{1}{M_{pl}^2}$$

$$G_{\text{higher dimensions}} \sim \frac{1}{(2\pi R)^n}$$

rough generalization to arbitrary

number of XD

Perhaps  $G_{\text{higher}} \sim G_{\text{Fermi}}$  & gravity strongly coupled above weak scale!



# PARALLEL UNIVERSE FORM-FACTORS

Arkani-Hamed,  
Schmaltz ~ '99  
-100

Consider  $\psi + \chi \xrightarrow{\phi\text{-exchange}} \psi + \chi$   
all massless

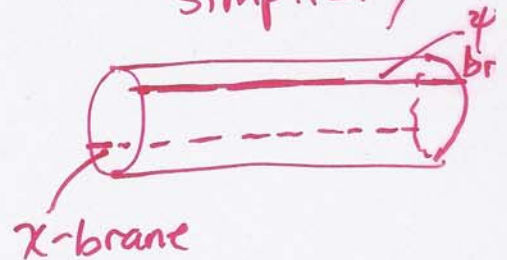
$\propto$   $\phi$ -propagator (spacelike 4-momentum,  
 $\Delta x_5 = \text{brane separation}$ )

$\propto$  assuming no brane-localized terms for  $\phi$  alone.

$$\sum_{P_5 = \frac{n}{R}} \frac{e^{i P_5 \pi R}}{\underbrace{P_\mu P^\mu - P_5^2}_{< 0} + i\epsilon}$$

$\leftarrow ie \pm 1$

$\leftarrow$  neglecting orbifold boundary conditions for simplicity



$$\propto \int_C dp_5 \frac{e^{i P_5 \pi R}}{(P_5 + i\sqrt{-P^2})(P_5 - i\sqrt{-P^2})} \frac{1}{e^{2i P_5 \pi R} - 1}$$

$$\propto \int_{C_2 + C_3} dp_5 \frac{e^{ip_5 \pi R}}{(p_5 + i\sqrt{-p^2})(p_5 - i\sqrt{-p^2})} \frac{1}{e^{2ip_5 \pi R} - 1}$$

$$\propto \frac{e^{+\sqrt{-p^2} \pi R}}{\sqrt{-p^2} [e^{2\sqrt{-p^2} \pi R} - 1]}$$

$$\widetilde{-p^2 \gg 1/R}$$

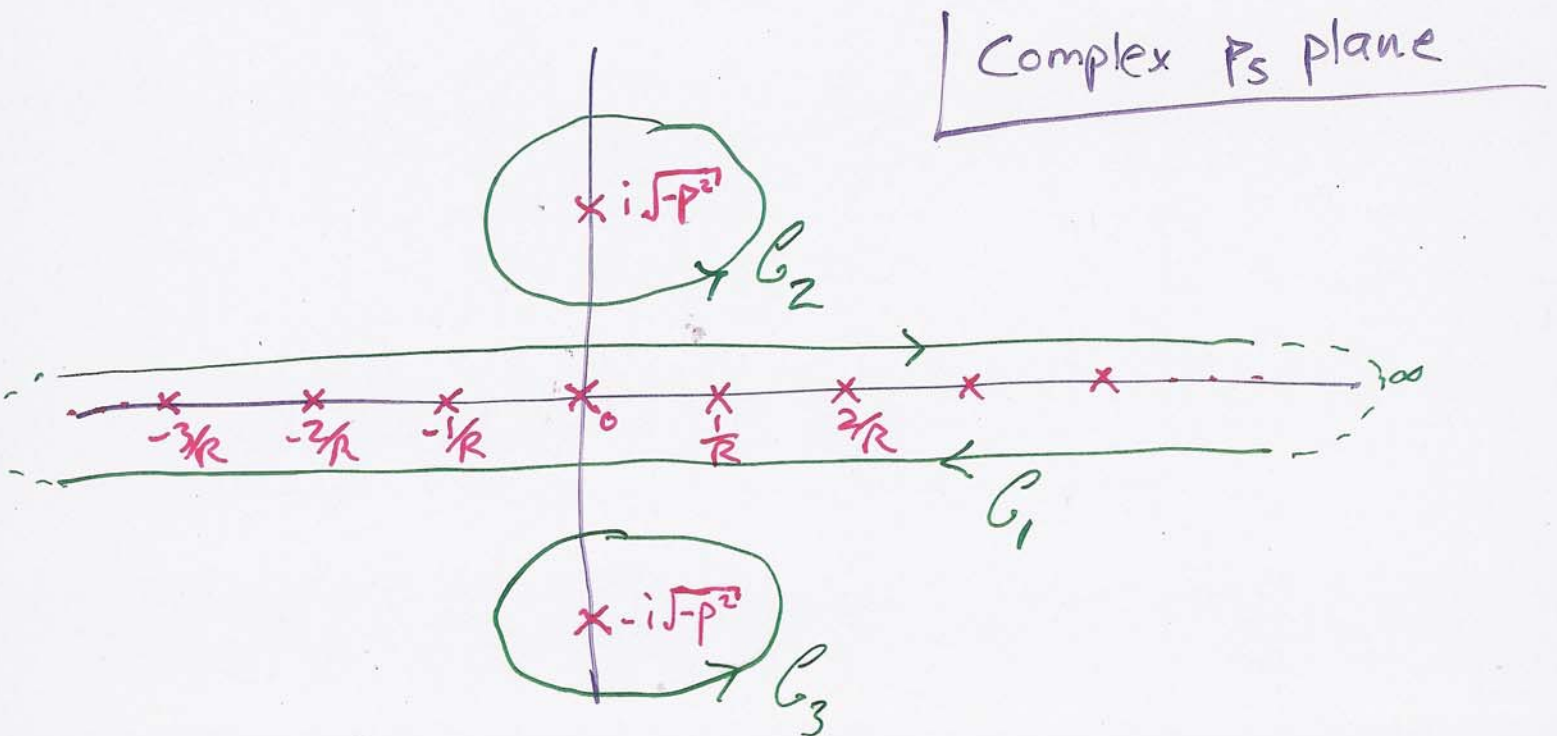
$$\frac{e^{-\sqrt{-p^2} \pi R}}{\sqrt{-p^2}}$$

Weird form-factor.

$$\widetilde{-p^2 \ll 1/R}$$

$$\frac{1}{p^2}$$

Usual 4D scattering



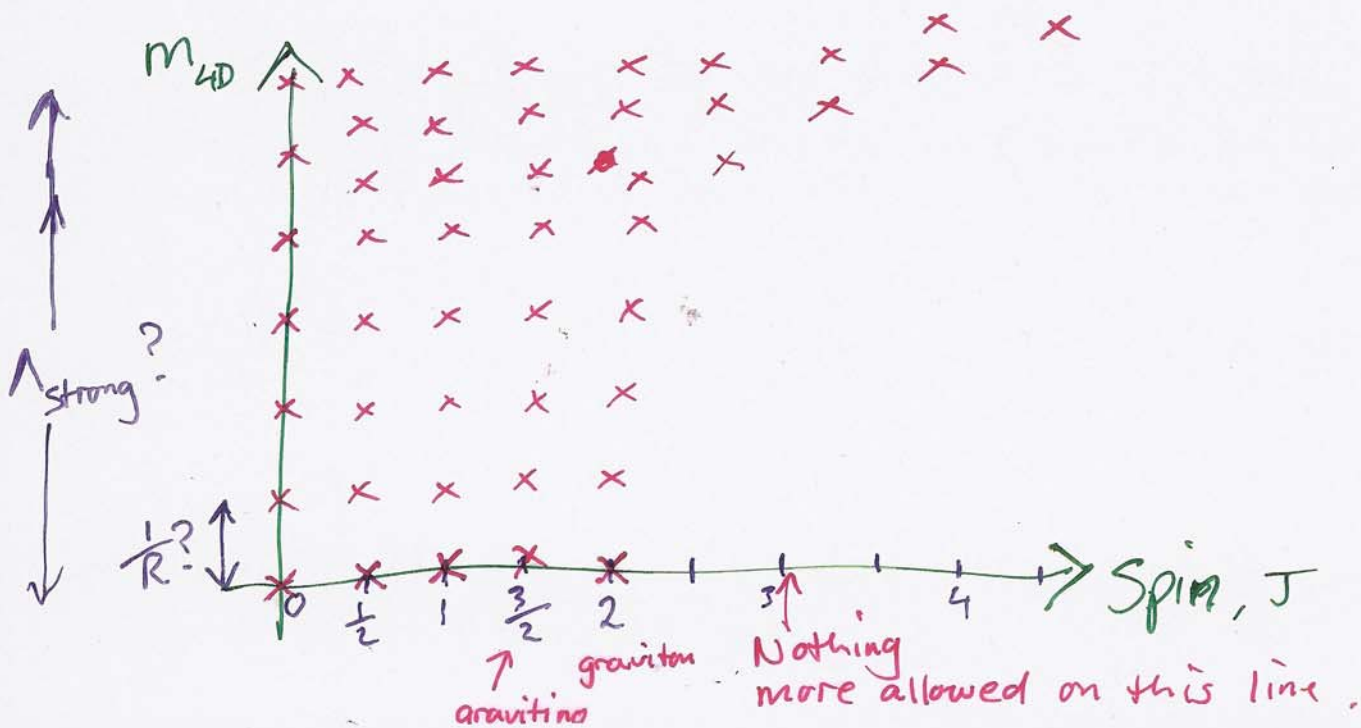
# When is it XD?

The only known UV completion of XD and/or GR (preserving extended spacetime symmetry in UV) is string theory. When stringiness is relevant, there are Regge trajectories (proliferation of high spin excitations).



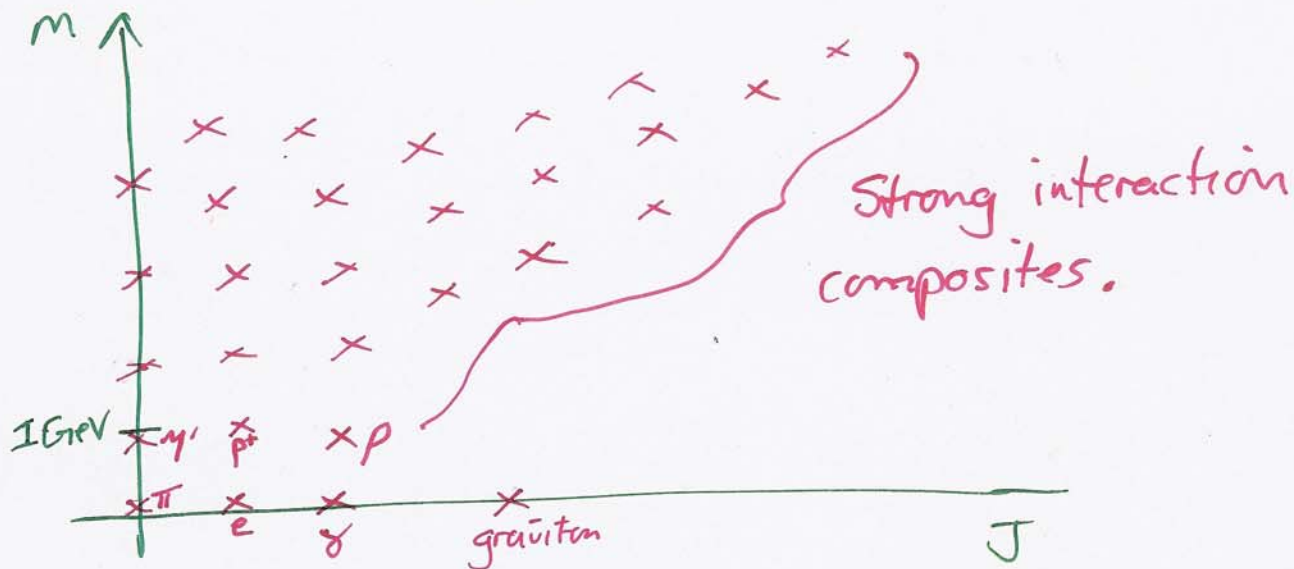
open string stretches by its orbital motion.

∴ look for this:





# THE PARTICLE DATA BOOK



How to distinguish XD from discovery of new strong interaction ~~scale~~ ~~in sense~~ composites of new high scale variant of QCD ("Technicolor" in its loosest sense)

- (i) XD can be weakly coupled (if  $m_{\text{string}} < \Lambda_{\text{strong}}$ ) but new strong dynamics will be "strong".
- (ii) In XD spectrum, high spins appear after many low spin ~~low~~ recurrences.
- (iii) At v. high energies strong dynamics becomes asymptotically free, revealing constituent partons.

BUT

- (i) In gauge theories with  $N_{\text{color}} \gg 1$  the strong dynamics can just confine quarks into hadrons, the residual interactions between hadrons can be  $\sim \frac{1}{N_{\text{color}}}$ .
- (ii) There are strongly interacting 4D dynamics where high spins are delayed to much higher mass than low spin recurrences.
- (iii) There are strongly interacting 4D theories where coupling does not run & remains strong at high energy (conformal field theories).

In fact there are examples where

Weakly-coupled  $X$ D theory  $\equiv$  Strongly-coupled 4D gauge theory  
in every physical consequence      Eg. see Strassler TASI lectures

Analogy: Wave-Particle duality in quantum mechanics



- Central module of such "duality" is the AdS/CFT<sub>4D</sub> Correspondence.

"Anti-de Sitter" <sup>5D</sup>

review Aharony, Gubser, Maldacena, Ooguri, Oz '99

- The catch in this duality is

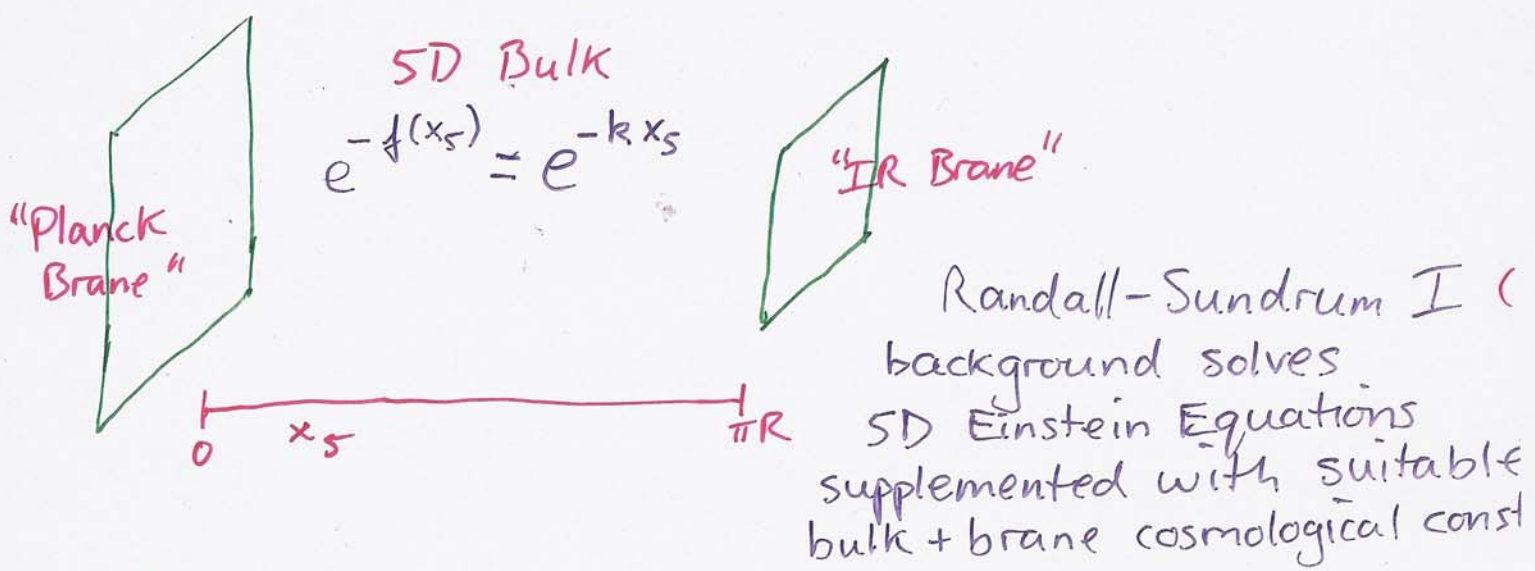
that

$$\eta_{MN} \longrightarrow \begin{pmatrix} \eta_{\mu\nu} & e^{-\frac{2f(x_5)}{\Lambda}} & 0 \\ 0 & & -1 \end{pmatrix}$$

"warp factor" preserves 4D Poincaré symmetry

which makes calculations harder.

- Simplest warped compactification, used in phenomenological effective field theories



String realization: ... Polchinski



Analogy — Inflating universe (dS spacetime)

$$\eta_{\mu\nu} \rightarrow g_{\mu\nu} = \begin{pmatrix} +1 & 0 \\ 0 & -e^{Ht} \delta_{ij} \end{pmatrix}$$

solve 4D Einstein Equations with positive cosmological constant. Space inflates in time. "scale factor"

ie.  $AdS_5 \equiv$  "inflation" of 4D spacetime in XI

KK decomposition of bulk fields is more subtle (not sinusoidal).

General features:

- Low-lying KK excitations localize near IR brane
- Lowest mode shape is v. sensitive to boundary conditions & 5D spin, mass