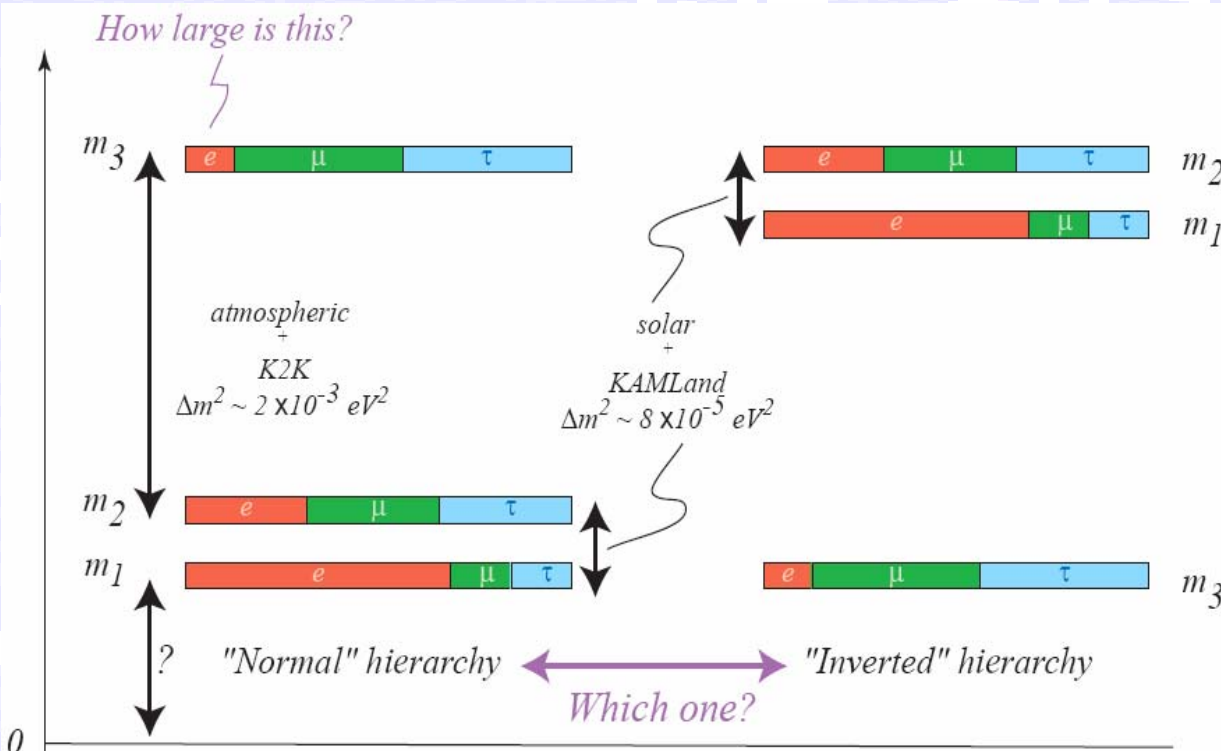


A Program of Long Baseline Neutrino Exploration at Fermilab

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Fermilab

April 8, 2005

What we know, what we would like to know...



Would like to have more precise knowledge of mixing. Do ν_e 's participate in oscillations at atmospheric scale?

Is $\Delta m^2 > 0$ or < 0 ?

Is CP violated?

The Fermilab long-range plan for long baseline experiments addresses all of these issues in a step-by-step program of detectors and beamline upgrades.

MINOS (Main Injector Neutrino Oscillation Search)

Long baseline oscillation experiment designed to:

- Demonstrate oscillation behavior
 - Confirm and describe flavour oscillations
 - Provide *high statistics* discrimination against alternative models (decoherence, ν decay, extra dimensions, etc.)
- Precise Measurement of Δm_{23}^2 to $\sim 10\%$
- Search for $\nu_{\mu} \rightarrow \nu_e$ oscillations (θ_{13})
- First direct measurement of ν vs $\bar{\nu}$ oscillations from atmospheric neutrino events
 - MINOS is the first large deep underground detector with a B-field



The NuMI beam

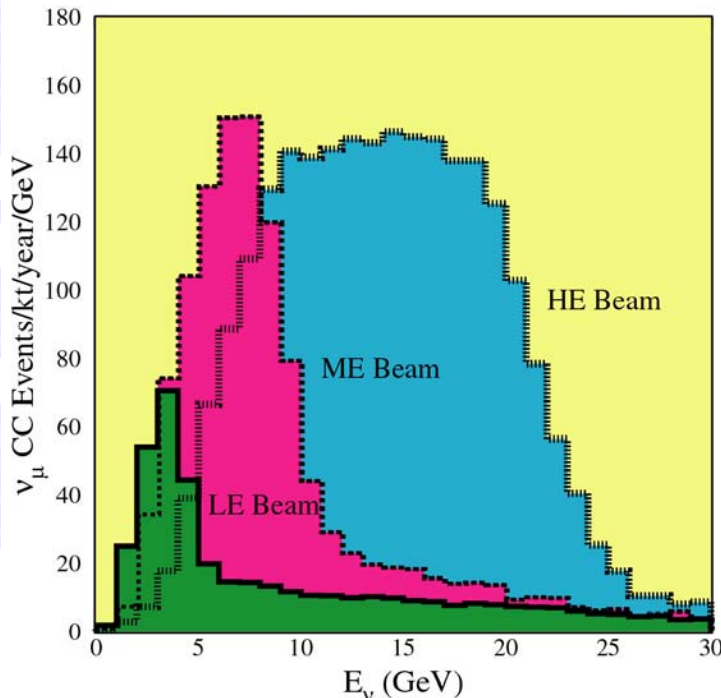
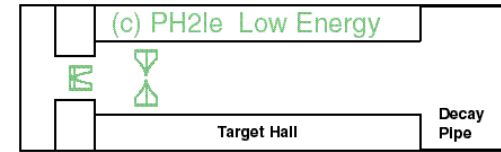
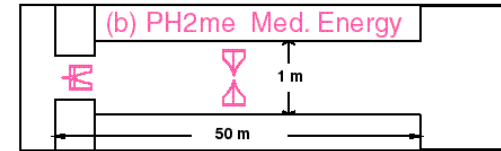
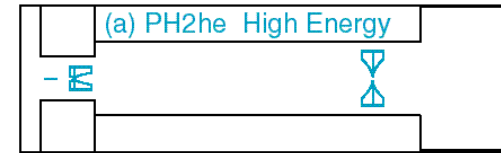


FERMILAB #98-765D

- 120 GeV protons extracted from the Main Injector in a single turn (8.7ms)
- 1.9 s cycle time *i.e.* ν beam 'on' for 8.7 ms every 1.9 s
- 2.5×10^{13} protons/pulse initially
- 2.5×10^{20} protons/year initial intensity
- 0.25 MW on target !
- 0.4 MW at 4×10^{13} p/pulse!

Tuneable Beam

- Relative position of target and horns allows tuning of beam energy. Act like a pair of highly achromatic lenses.
- MINOS starts with LE beam - best for $\Delta m^2 \sim 0.002 \text{ eV}^2$
- Can run neutrinos or antineutrinos



LE BEAM:

ν_μ CC Events Observed/yr:

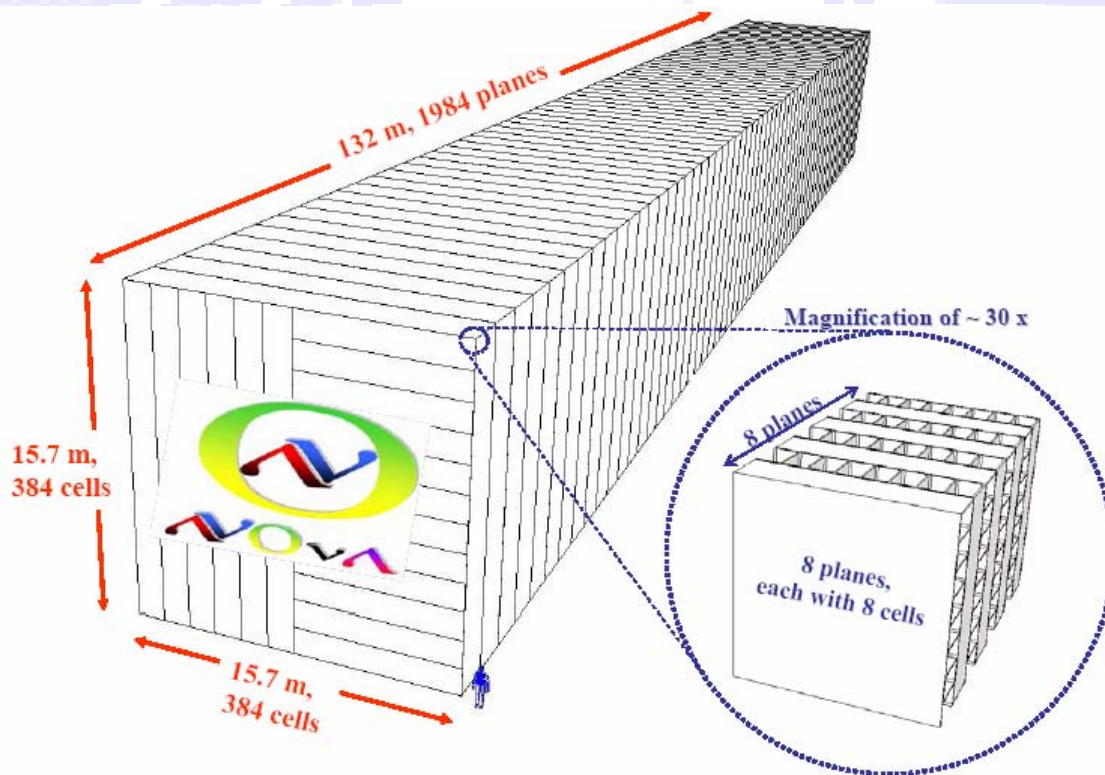
Low	Medium	High
1600	4300	9250

(2.5×10^{20} protons on target/year)

NO ν A

NuMI Off-Axis ν_e Appearance Experiment

Argonne, Athens, Caltech, UCLA, Fermilab, College de France, Harvard, Indiana, ITEP, Lebedev, Michigan State, Minnesota/Duluth, Minnesota/Minneapolis, Munich, Stony Brook, Northern Illinois, Ohio, Ohio State, Oxford, Rio de Janeiro, Rutherford, South Carolina, Stanford, Texas A&M, Texas/Austin, Tufts, Virginia, Washington, William & Mary, Wisconsin



Goals of the NO ν A Experiment

- Observe ν_e appearance
- Sensitivity to $\text{Sin}^2(2\theta_{13})$ a factor of 10 below CHOOZ sensitivity, i.e. down to ~ 0.01
- $\text{Sin}^2(2\theta_{23})$ measurement to 2% accuracy
- Resolve or contribute to determination of mass hierarchy via matter effects
- Begin to study CP violation in lepton sector

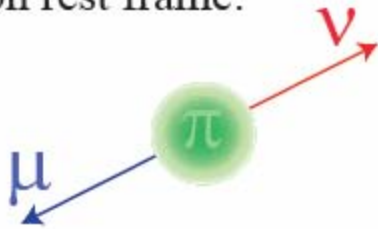
How NO_vA Will Meet its Goals

- Reduce backgrounds to ν_e appearance search by going **off the NUMI beam axis** for a narrow-band beam. Will use Medium Energy configuration.
- Increase flux/POT at oscillation max by ~ 2 by going off-axis
- Increase detector mass a factor of 6 over MINOS while reducing cost/kiloton by a factor of 3
- **80% active detector design** (compared to $1.5 X_0$ sampling in MINOS)
 - electron showers appear as “fuzzy” tracks with 1-4 hits/plane/view
 - allow better separation of γ 's from π^0 decays
 - good energy resolution to focus on signal energy region
- Choose long baseline to enhance matter effects

Off-Axis Neutrino Beams

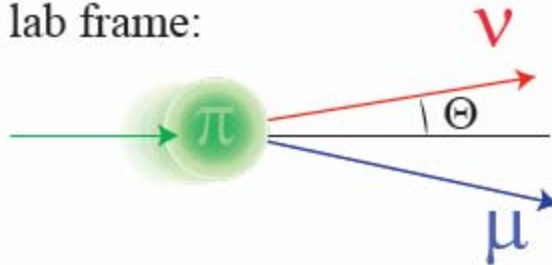
First proposed by BNL E-889

In pion rest frame:

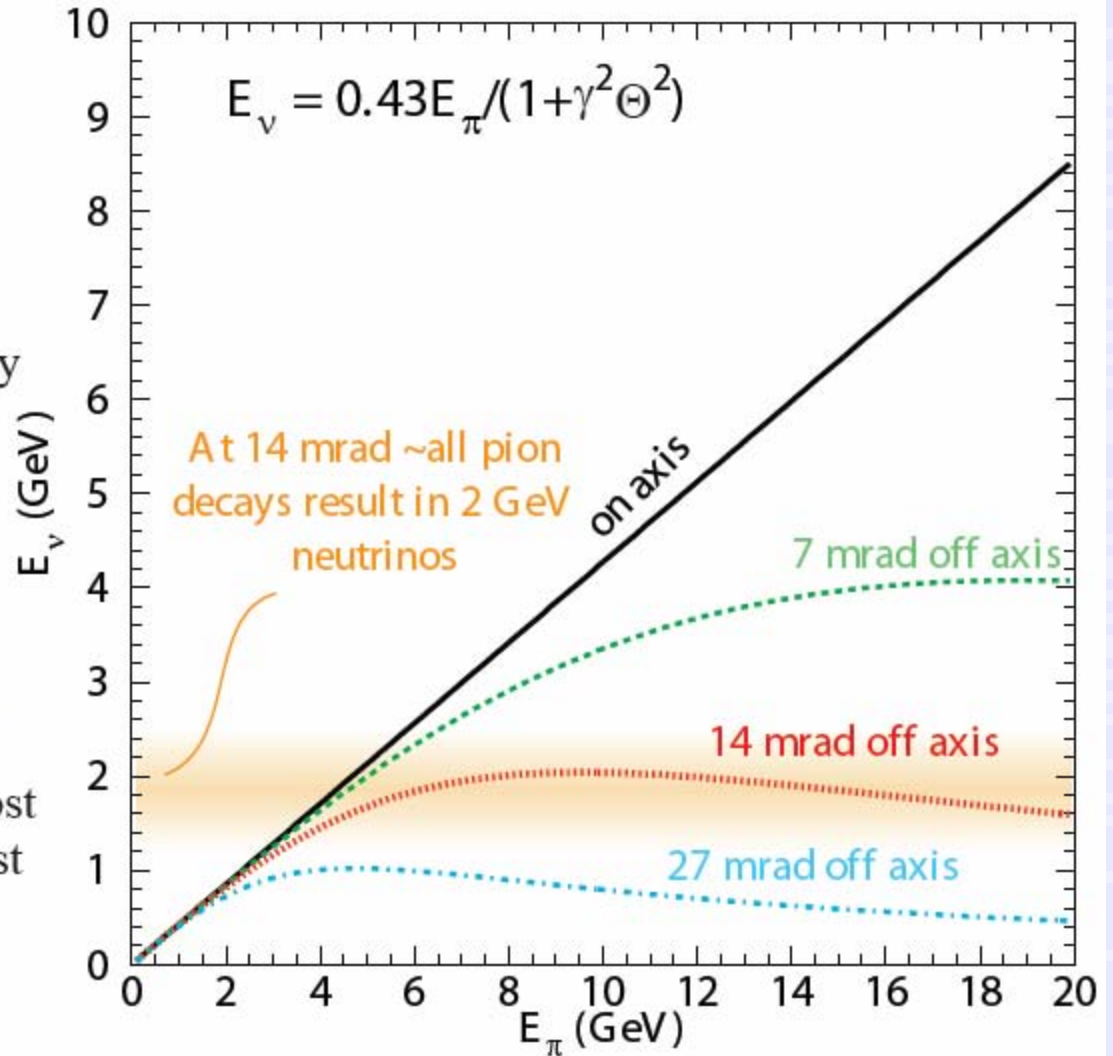


Neutrino and muon energy completely determined by energy conservation

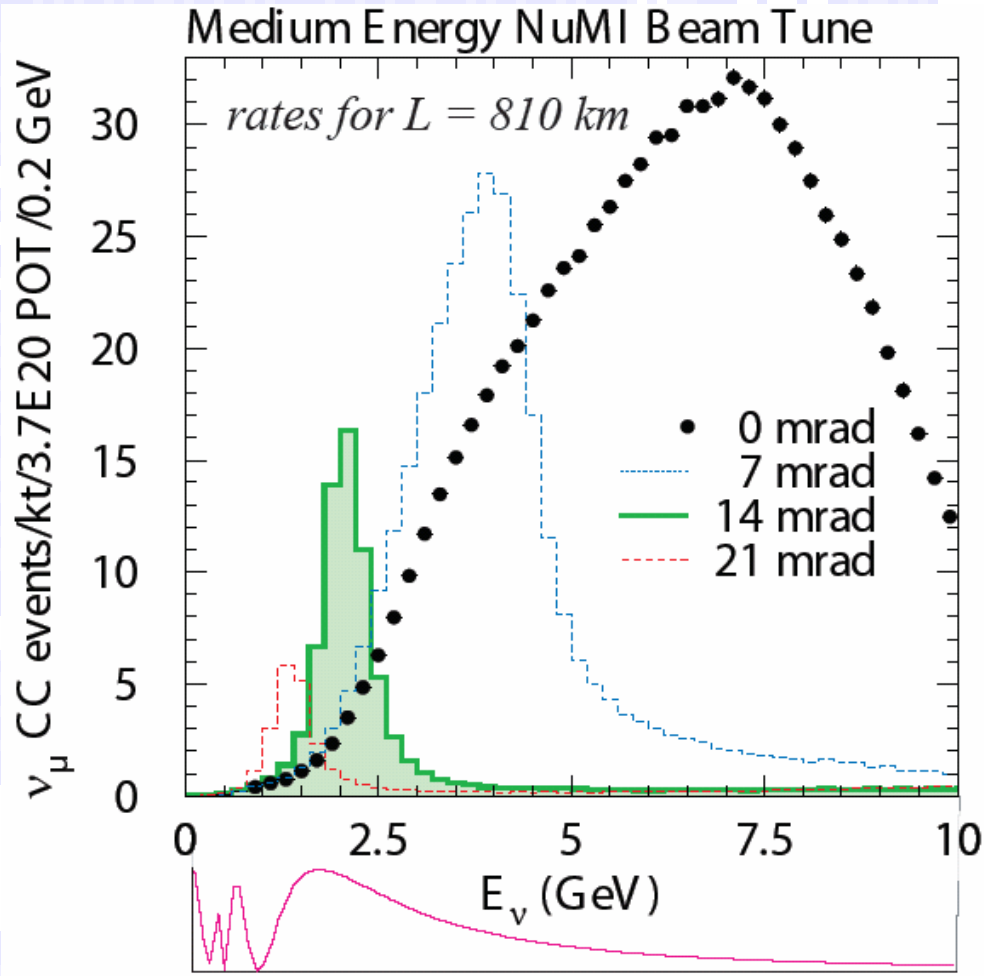
In lab frame:



Neutrino energy depends on boost and angle between neutrino boost direction



NUMI Neutrino Spectra



- 14 mrad off-axis beam peaks just above oscillation max at ~ 2 GeV with $\sim 20\%$ width
- High energy tail suppressed
 - Reduces NC and τ backgrounds
- Main peak from π decays. K decay ν at much wider angles.
 - Spectrum prediction insensitive to knowledge of k/π ratio

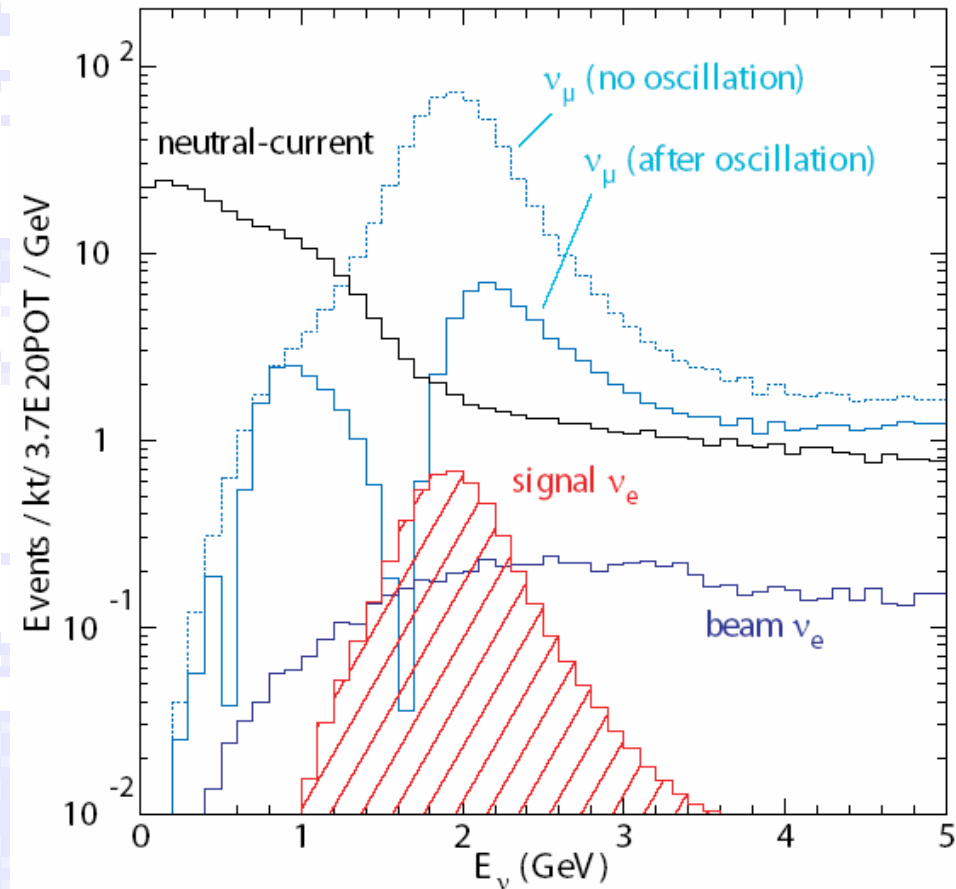
Event Rates

Event rates calculated for

- $L=810$ km, 12 km off-axis
- $\Delta m_{23}^2 = 2.5 \times 10^{-3} \text{ eV}^2$
- $\text{Sin}^2 2\theta_{23} = 1$
- $\text{Sin}^2 2\theta_{13} = 0.01$

To Reject Background:

- 50:1 rejection of ν_μ CC required
 ⇒ Easy!
- Need 100:1 NC rejection
 ⇒ fine grained, low density
- Good energy resolution
 ⇒ reject beam ν_e

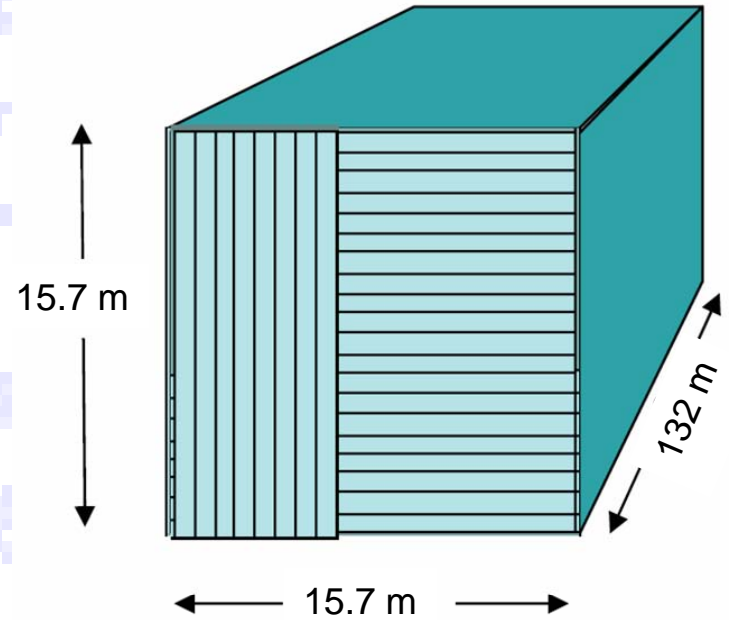
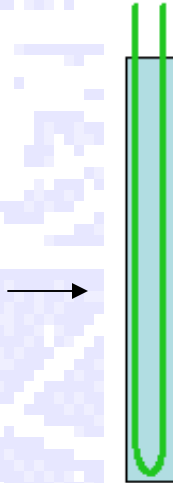


The NOvA Far Detector

- 30 kT, low Z tracking calorimeter
- 80% active material (by weight).
- Optimized for detecting 2 GeV electrons.
- PVC extrusions filled with Liq. Scint.

- Cell size of 3.87cm x 6.0 cm x 15.7 m
- 12 extrusions/plane
- 32 cells/extrusion
- 1984 planes
- = 23,808 extrusions
- = 761,856 channels

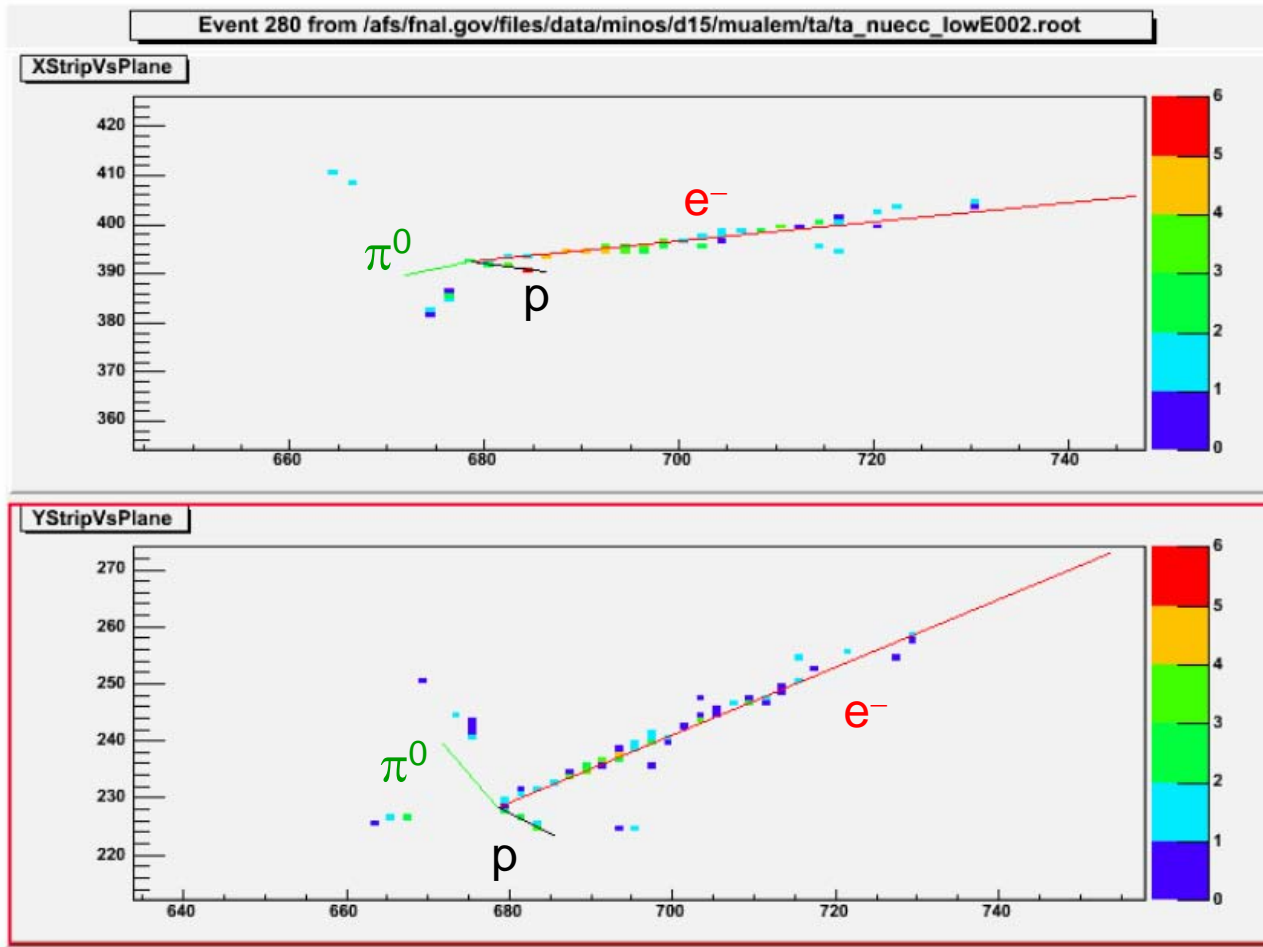
- 0.8 mm looped WLS fiber into APD readout



APD Readout

- Cooled to -15°C
- Q.E. 85%
- 22 p.e. at far end
- 250 e noise
- S/N 10:1

Typical $\text{NO}_{\nu A}$ Event: $\nu_e A \rightarrow p e^- \pi^0, E_\nu = 1.65 \text{ GeV}$



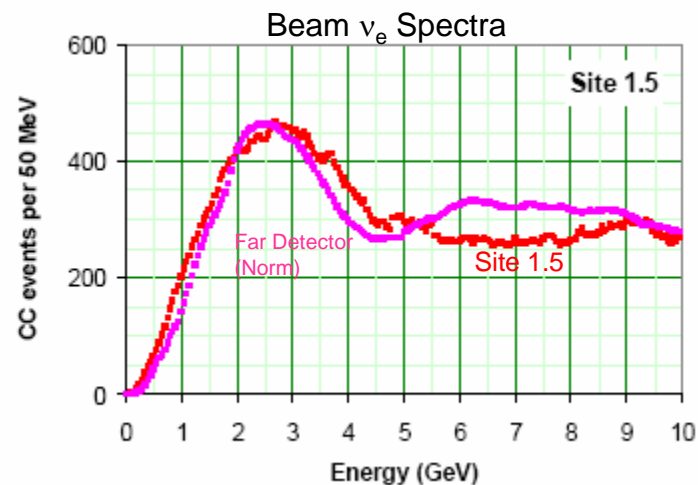
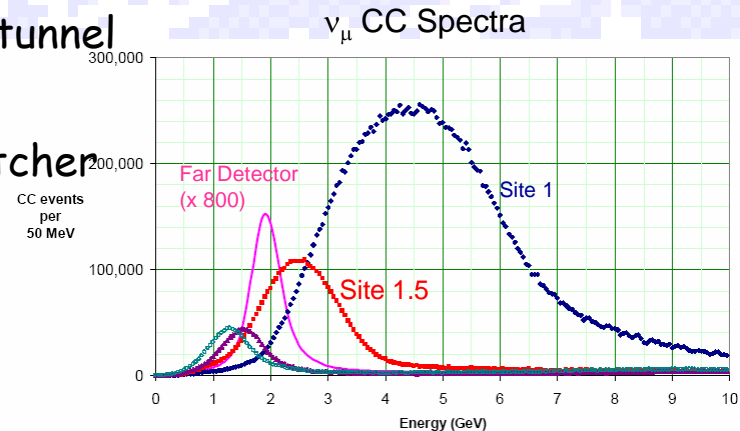
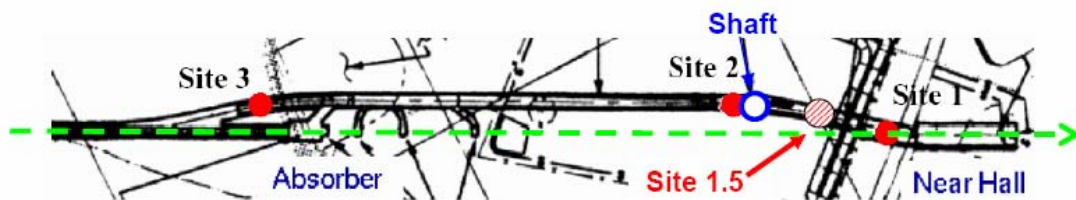
Signal efficiency 24%

signal/background 7.3

signal/sqrt(bg.) 32

NO_vA Near Detector

- ~ 1 km from NUMI target
- Fits in several existing locations in NUMI access tunnel
 - 3.5 m x 4.8 m x 9.6 m
 - Includes veto, shower containment, muon catcher
- No single location optimizes all parameters
 - Make movable or build more than 1



NOvA Milestones

Milestone	Date (in months relative to Project Start)	Proposed Calendar Date	FY
Project Start	t_0	October, 2006	07
Order extrusions and fiber	t_0+1	Nov., 2007	07
Start extrusion module factories	t_0+12	October, 2007	08
Start operation of Near Detector	t_0+21	July, 2008	08
Far building complete	t_0+31	May, 2009	09
Start Construction of Far detector	t_0+31	May, 2009	09
First kiloton operational	t_0+36	Oct., 2009	10
First 15 kilotons operational	t_0+47	June 2010	10
Full 30 kilotons operational	t_0+57	July, 2011	11

Fermilab Proton Plan

	Booster Batch Size	Main Injector Load (AP + NuMI)	Cycle Time (sec)	MI Intensity (protons)	Booster Rate* (Hz)	Total Proton Rate (p/hr)	Annual Rate at end of Phase	
							NuMI	BNB
Actual Operation								
July, 04	5.0E+12	1+0	2.0	0.5E+13	5.1	0.8E+17	0	3.3E+20
Proton Plan								
Phase I	5.10E+12	2+1→2+5	2.0	3.6E+13	6.3	1.0E+17	2.0E+20	1.5E+20
Phase II	5.3E+12	2+5	2.0	3.7E+13	7.5	1.2E+17	2.2E+20	2.8E+20
Phase III	5.50E+12	2+9	2.2	6.0E+13	8.3	1.5E+17	3.4E+20	2.2E+20
Beyond Scope of Present Plan								
11 Hz	5.50E+12	2+9	2.2	6.1E+13	11.0	2.0E+17	3.4E+20	5.0E+20

2008

TABLE 6: Performance parameters at the completion of each phase of operation.

* Booster rate is limited by radiation levels, except for the 11 Hz case

http://www.fnal.gov/directorate/program_planning/Nov2004PACPublic/Draft_Proton_Plan_v2.pdf

Fermilab Proton Plan after 2009

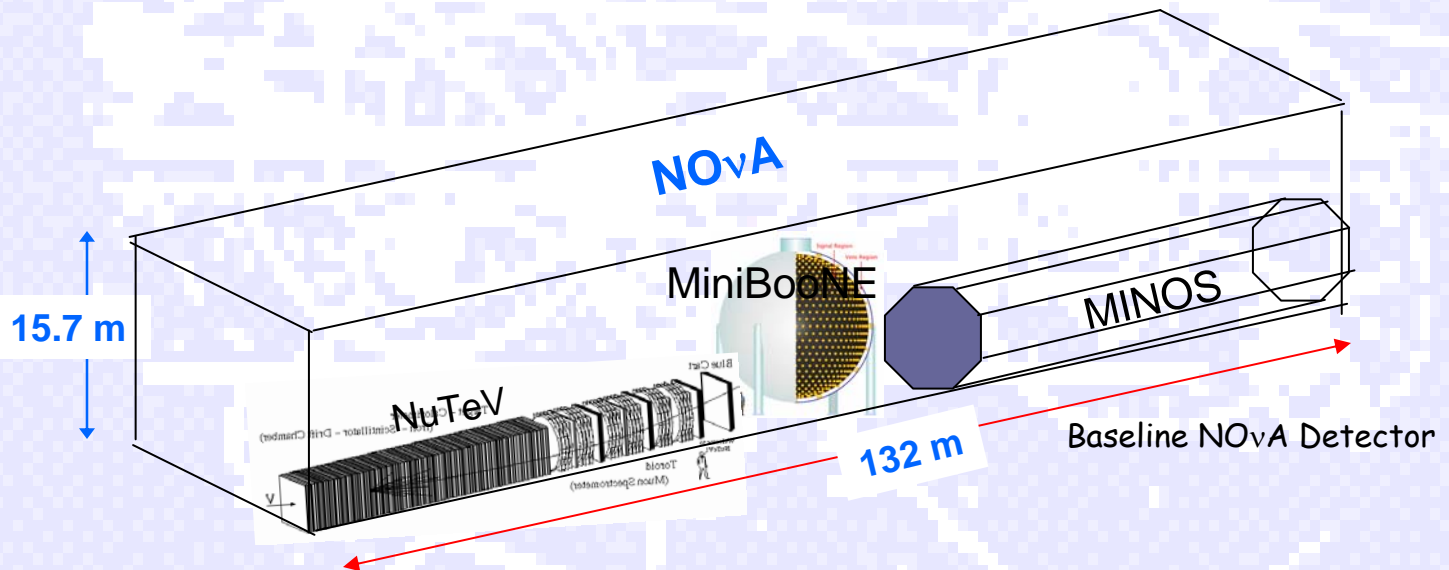
Collider operations end in 2009

- Proton bunches in MI used to \bar{p} now available to NuMI $\rightarrow \times 11/9$
- No NuMI downtime due to shot setup (10%) or antiproton transfers to Recycler (5%) $\rightarrow \times 1.176$
- Load 11 booster batches into Recycler and transfer from Recycler to MI in a single booster cycle. MI cycle time reduced from 2.2 s to 1.467 s $\rightarrow \times 1.5$
- $(1.22)(1.176)(1.5)(3.4 \times 10^{20} \text{ p/yr}) = 7.3 \times 10^{20} \text{ p/yr}$

Assume 90% $\rightarrow 6.5 \times 10^{20} \text{ p/yr}$

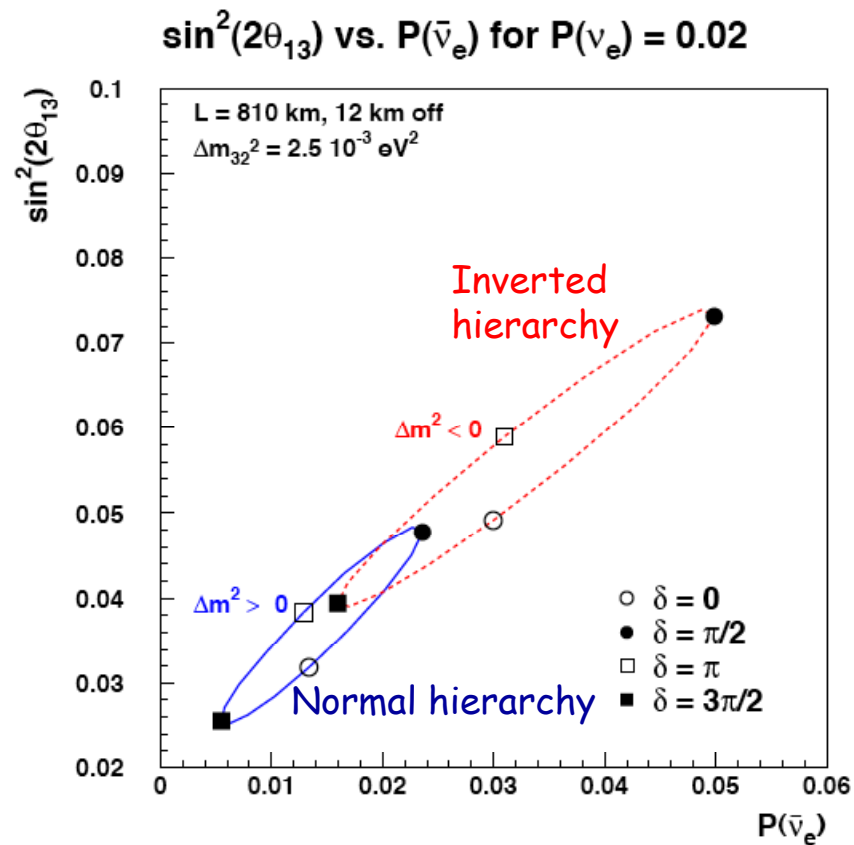
Statistically Limited

- NOvA will be statistically limited. Thus, the power of the experiment is proportional to mass times the neutrino flux.
- A **Fermilab Proton Driver** would provide 25×10^{20} pot/yr, a factor of $\times 4$.
- Same effect as building 4 NOvA's which would cost \$500M more and be truly enormous:



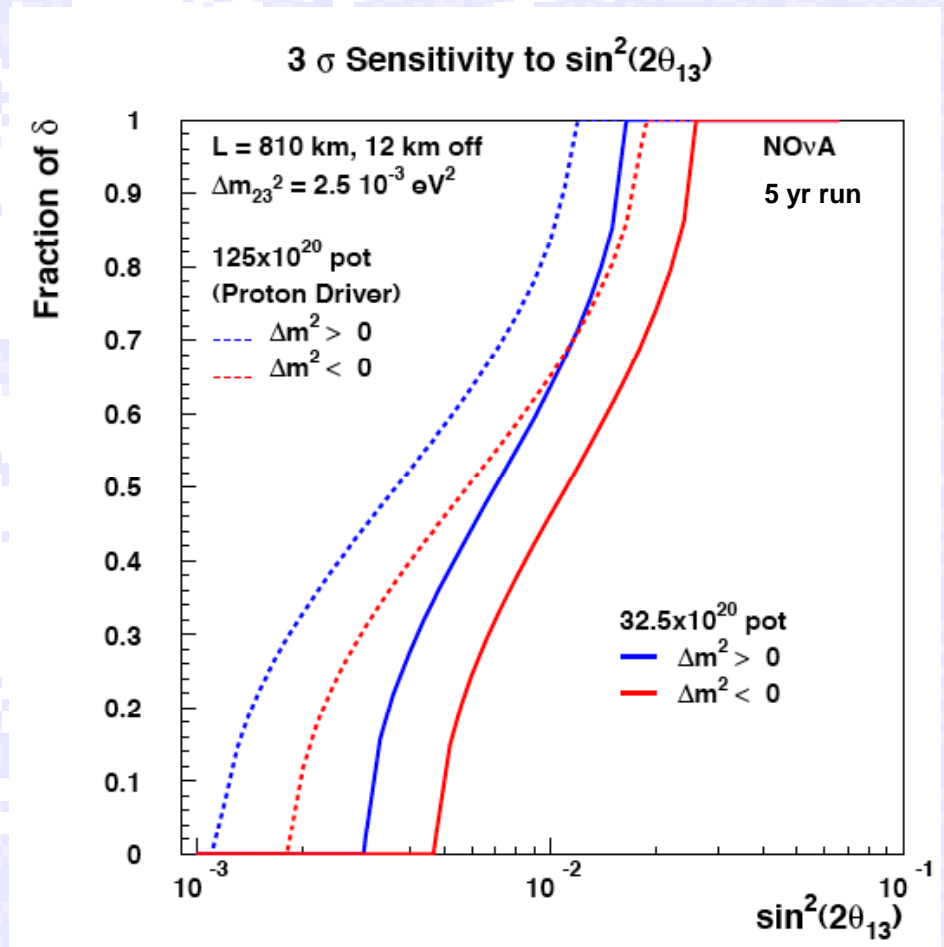
Interpreting what we measure

- Experiments measure oscillation probabilities
- Ambiguities in $\sin^2(2\theta_{13})$ due to CP phase δ and mass hierarchy
- Comparison of NO ν A and T2K at different baselines can break ambiguities
- Possibly use a 2^d NUMI off-axis detector at the 2^d oscillation maximum
- Sensitivity varies with CP phase
- Quote sensitivities vs the fraction of the CP ellipse covered

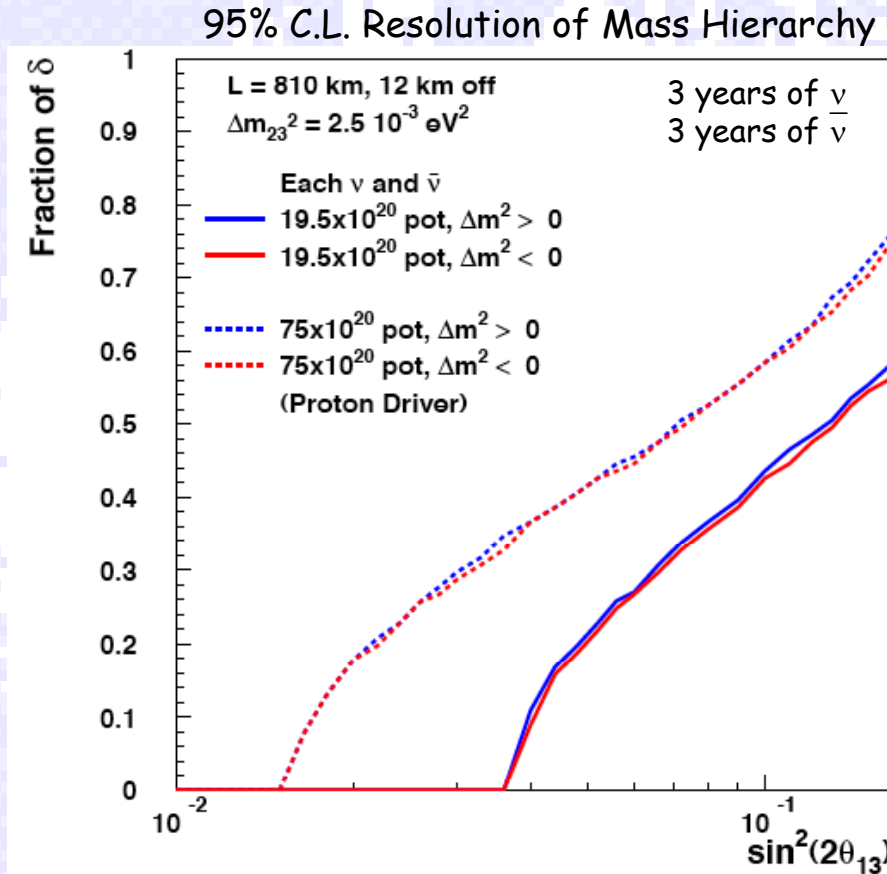


$\sin^2(2\theta_{13})$ Sensitivity

- Vertical axis is the fraction of possible δ values for which a 3σ discovery could be made.
- At large values of $\sin^2(2\theta_{13})$ a 3σ discovery can be made for all values of δ .
- At lower values of $\sin^2(2\theta_{13})$ a 3σ discovery is only possible for a range of δ .
- 5% systematic error on background determination included.



Resolving the Mass Hierarchy



There is a reasonable region of parameter space for which NO ν A can resolve hierarchy. Proton Driver extends reach by factor of 2.

Resolving the Mass Hierarchy (cont.)

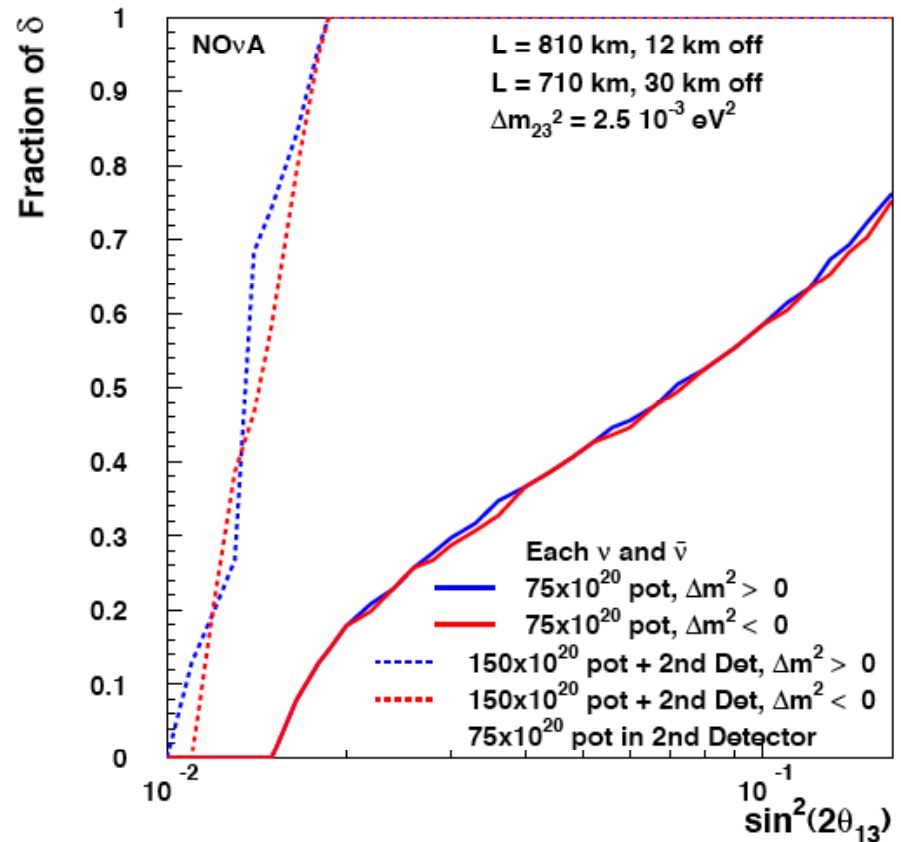
NOvA alone and with an additional off-axis detector at the 2^d maximum

At 2^d oscillation maximum

- L=710 km, 30 km off-axis
- Energy lower by x 3
 - ⇒ Matter effect smaller by x 3
 - ⇒ CP violation larger by x 3

Mass hierarchy resolved for all δ for $\sin^2(2\theta_{13}) > \sim 0.015$

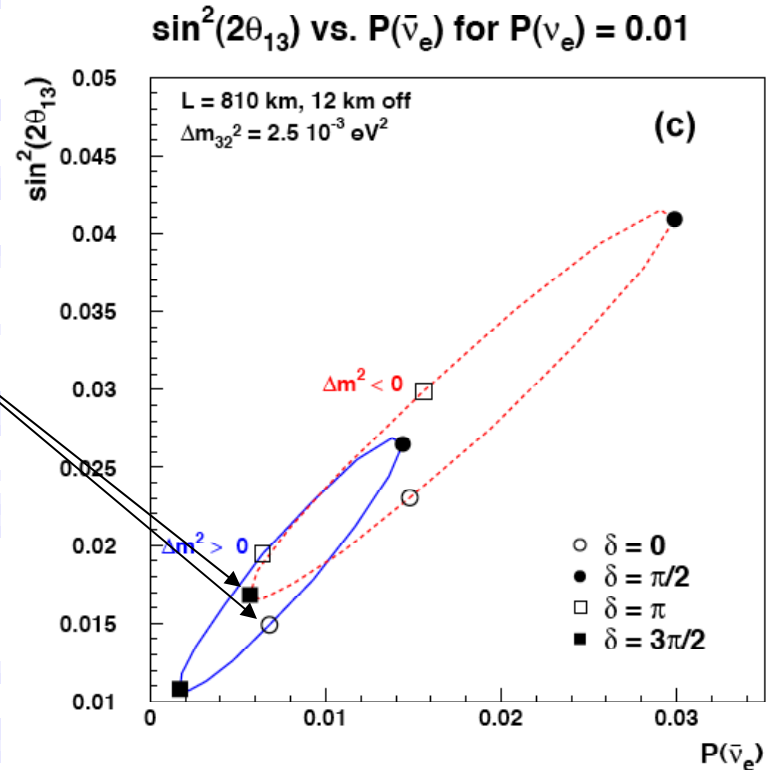
95% C.L. Resolution of Mass Hierarchy



12 years with proton driver
6 years with 2 detectors

Sensitivity to CP Violation

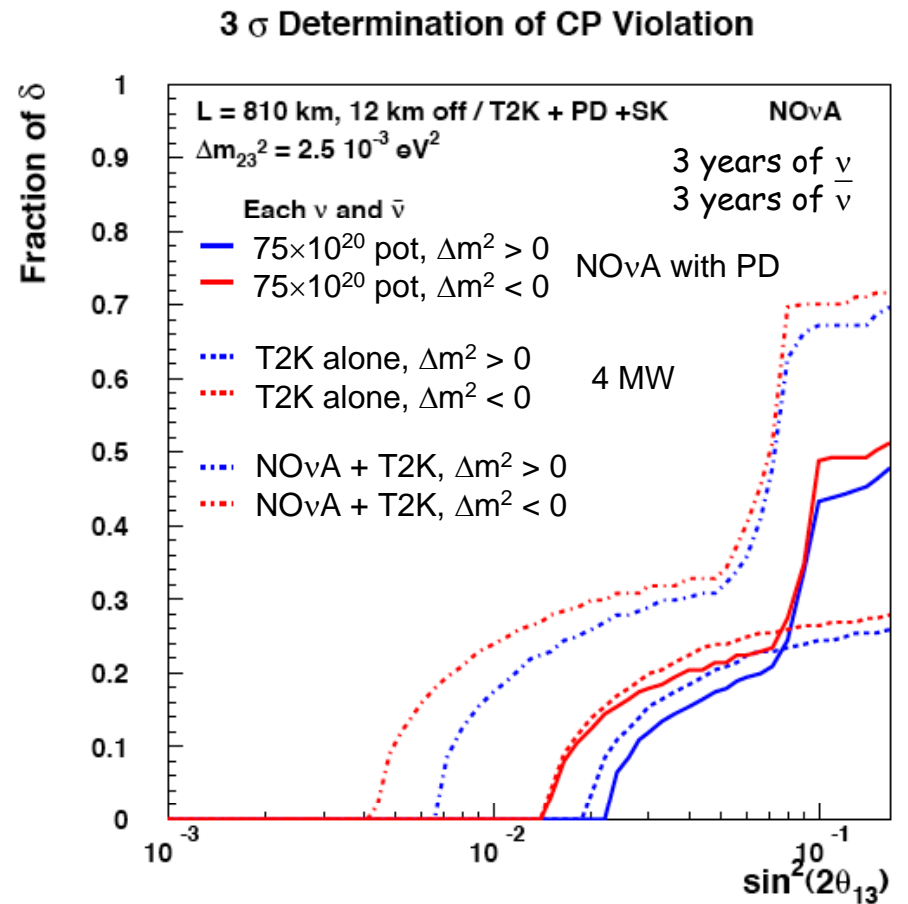
- Long baseline experiments generally need to know the hierarchy to measure the CP phase
- Maximal CPV for one mass ordering can have ν and $\bar{\nu}$ probabilities corresponding to no CPV for the other mass ordering



- Neither NOvA nor T2K can demonstrate CP violation in 6 years of running without enhanced proton sources

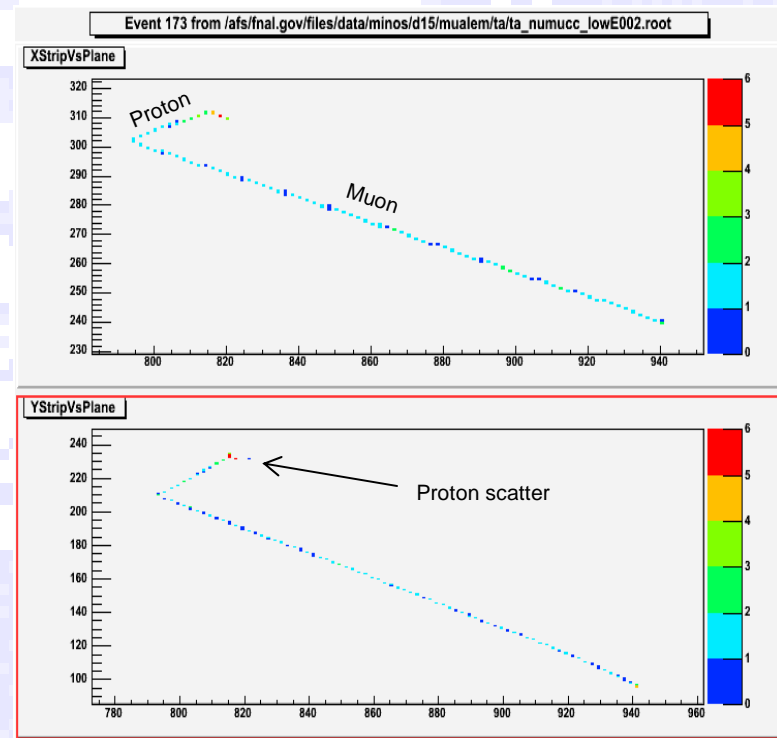
Sensitivity to CP Violation (cont.)

- Fraction of possible δ values for which there is a 3σ demonstration of CP violation
 - i.e. δ is neither 0 nor π for both mass orderings.



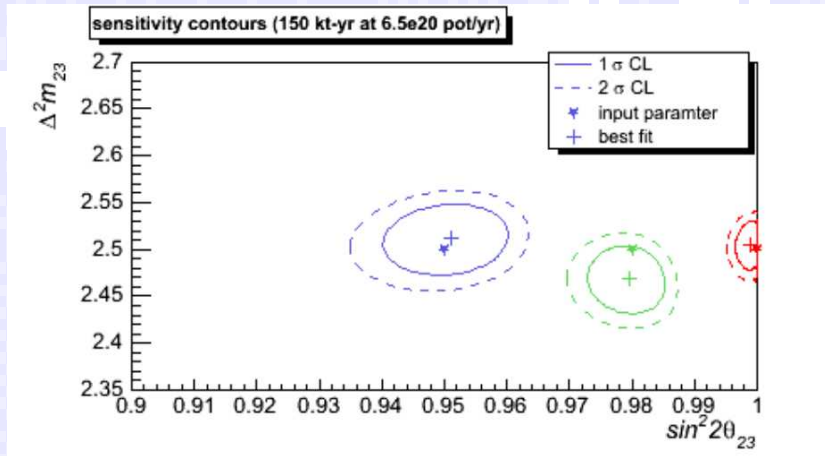
Precise Determination of $\sin^2(2\theta_{23})$

- Important because:
 - If mixing is maximal, could be due to unknown symmetry
 - $\nu_\mu \rightarrow \nu_e$ oscillation is proportional to $\sin^2(\theta_{23})\sin^2(2\theta_{13})$.
 - If mixing is not maximal, this leads to an ambiguity in comparing reactor and accelerator results.
- Precision measurement requires
 - good statistics
 - excellent ν energy resolution
 - good control of systematics
- Use totally contained quasi-elastic events
 - Very clean, essentially no NC background
 - Can measure $\sin^2(2\theta_{23})$ to $\sim 1\text{-}2\%$ level

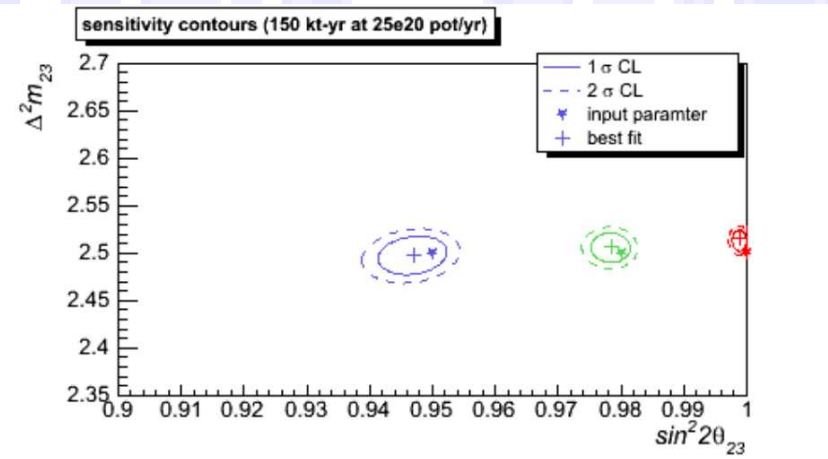


Precise Determination of $\sin^2(2\theta_{23})$

1σ and 2σ contours for simultaneous measurement of Δm_{32}^2 and $\sin^2(2\theta_{23})$ for a 5 year ν run without a Proton Driver.



5 year ν run with Proton Driver



For maximal mixing, error on $\sin^2(2\theta_{23})$ is about 0.004 without Proton Driver and 0.002 with a Proton Driver.

Summary

- Fermilab's long-range plan includes an ongoing program of long baseline neutrino experiments.
- MINOS is the first step in this program and is just now underway
(see talk by Mary Bashi at this workshop).
- **NO_vA would be the next step**
 - Presented current design to Fermilab PAC yesterday
 - Hoping for rapid consideration, by June at the latest
 - NuSAG review by funding agencies to report in June
 - **NO_vA and Fermilab are very open to new collaborators**
- A Proton Driver is being considered to augment the neutrino program and to support a wide range of other physics programs
(see talk by John Ellis at this workshop)
- In some scenarios a 2^d NUMI off-axis detector at the 2^d maximum is helpful