

Large TPCs for low energy rare event detection

NNN05 Next Generation of Nucleon Decay and Neutrino Detectors 7-9 April 2005 Aussois, Savoie, France

- Highlights from the Paris TPC workshop
- Spherical TPC project and motivation

SECOND WORKSHOP ON LARGE TPC FOR LOW ENERGY RARE EVENT DETECTION

LPNHE - Paris VI and VII Universities

Place Jussieu Tour 33 Rdc PARIS, France 20 - 21 December 2004

Gaseous TPCs :

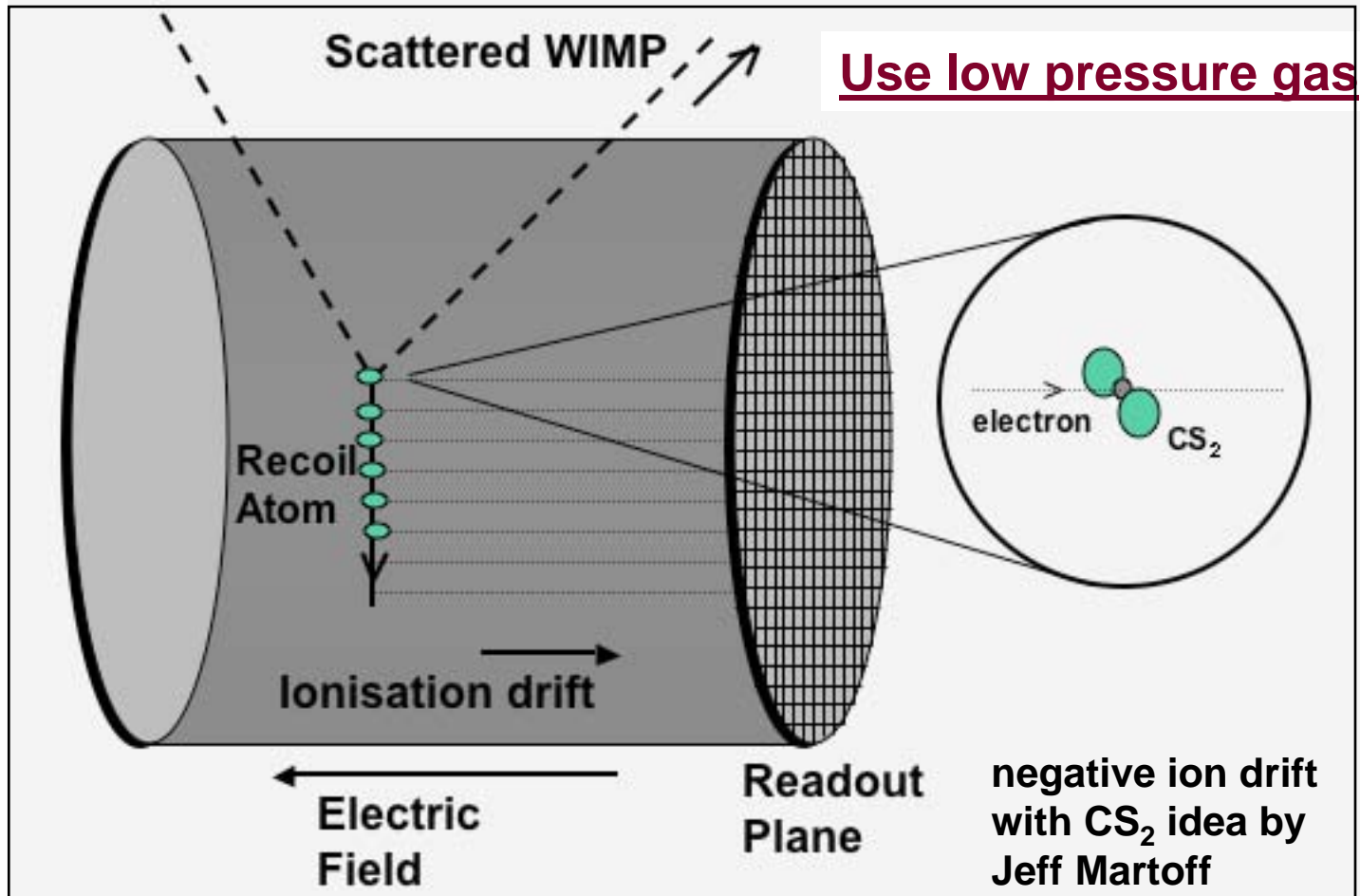
- 1) **Low energy neutrino detection** (neutrino oscillations, solar neutrinos, double beta decay, magnetic moment, supernova), I. Vergados, G. Gounaris, I. Irastorza, Ph. Gorodetzky, G. Bonvicini, Z. Daraktchieva, M. Green, M. Zito
- 2) **Axion search**, Th. Dafni, B. Beltran
- 3) **WIMP search with recoil direction**, B. Sadoulet, N. Spooner, D. Santos

Liquid TPCs,

A. Rubbia, E. Aprile, N-J-T. Smith, Ph. Lightfoot, V. Peskov
I. Giomataris

DRIFT and Prospects for a Large Scale Directional WIMP TPC

N. Spooner



Main motivation : drifting ions instead of electrons reduces the diffusion effect

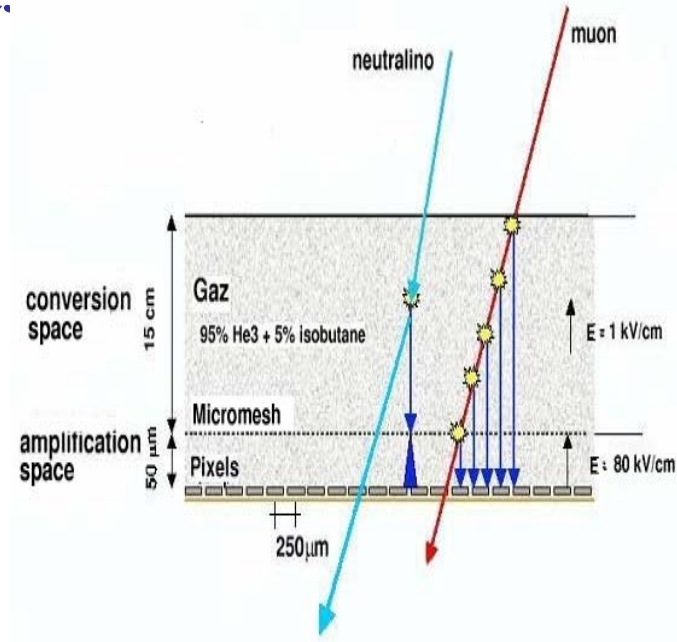
I. Giomataris

MIMAC-He3 : Micro-tpc Matrix of Chambers of He³

(D. Santos)

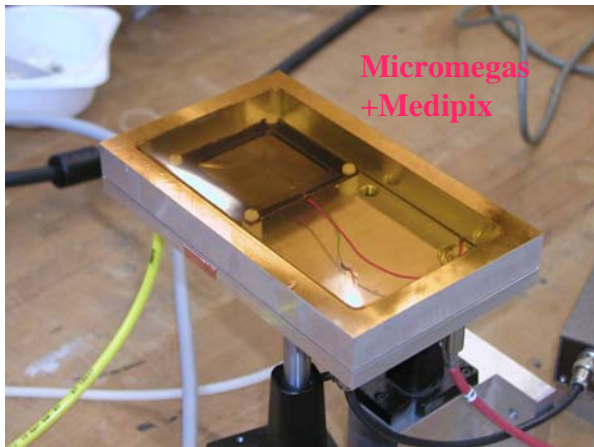
³He for axial detection of non-baryonic dark matter

High spatial temporal resolution
recoil track projection
⇒ energy threshold < 1 keV
⇒ electron/recoil discrimination

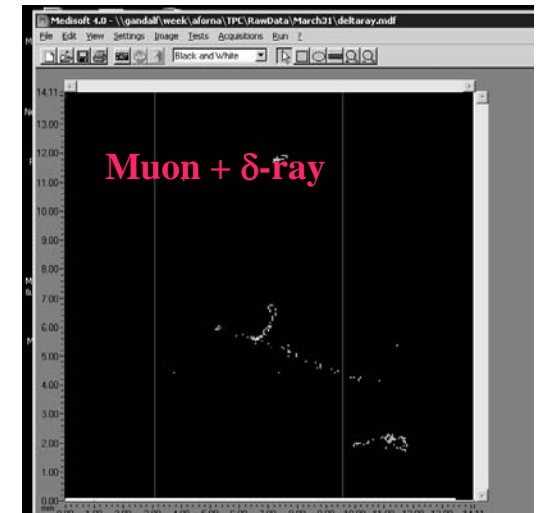


Last refinement: CMOS integrated pixel anodes (H. Van der Graaf)

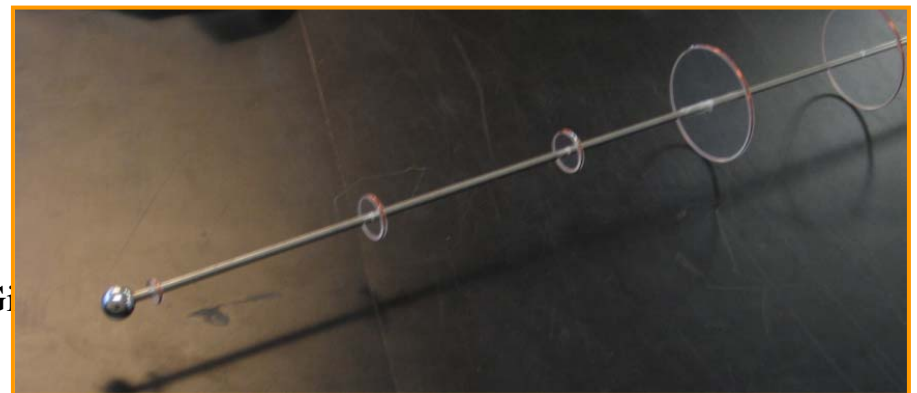
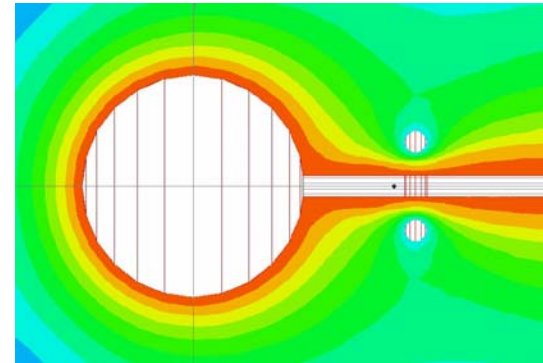
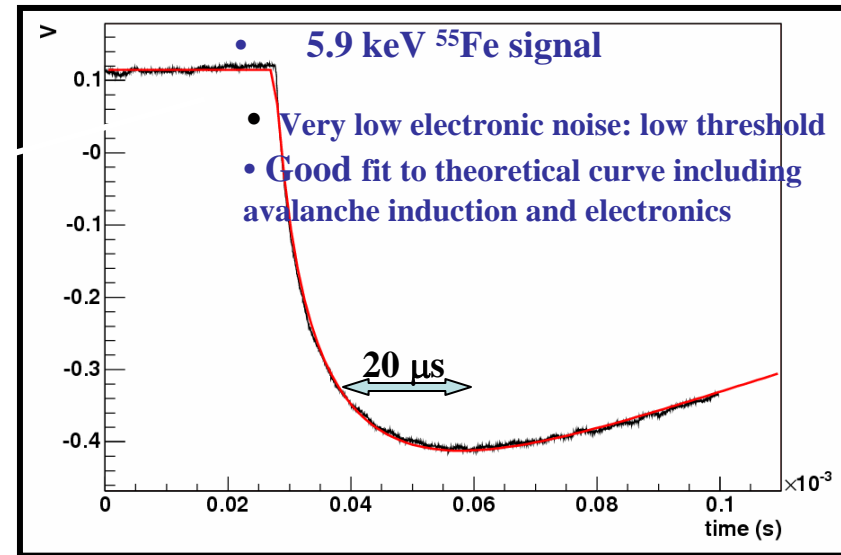
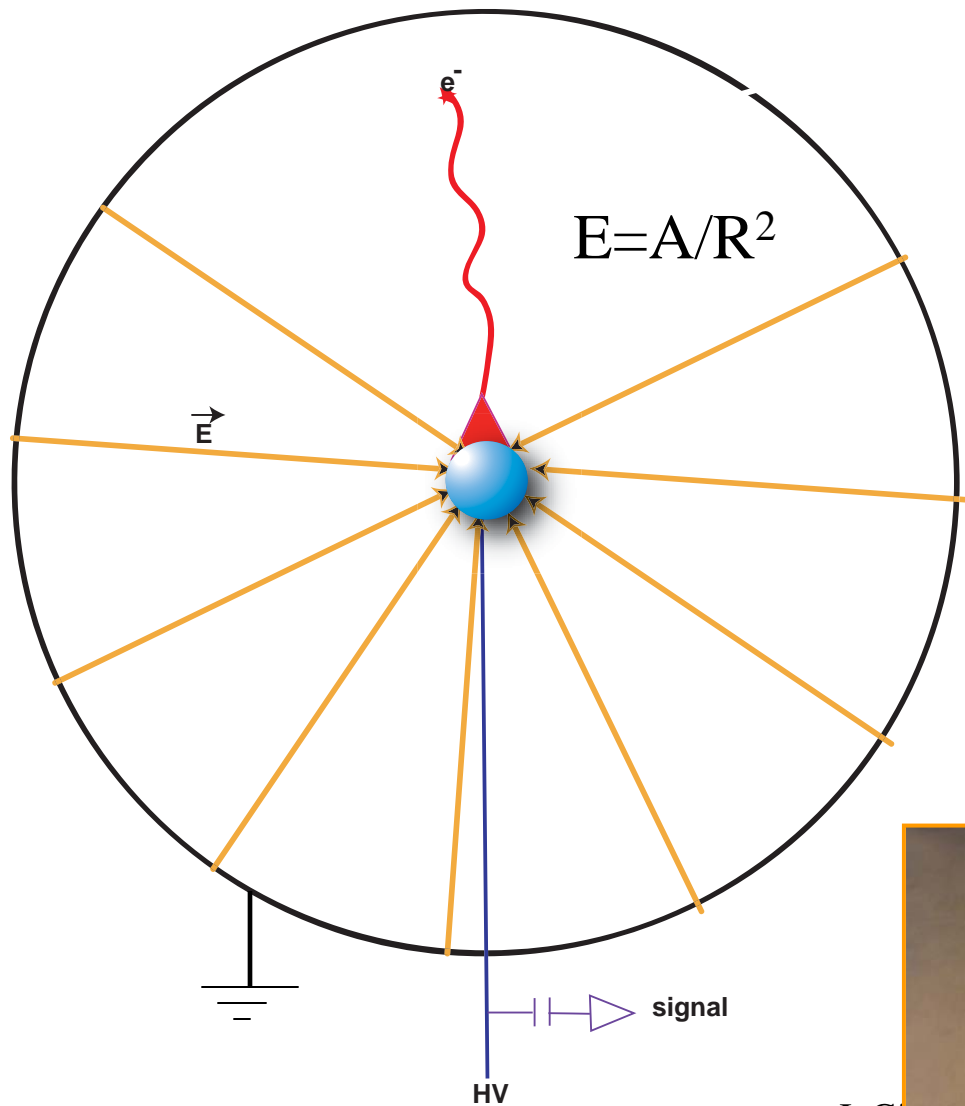
Idea : Combine micro-pad CMOS with high accuracy MPGD like Micromegas



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Spherical TPC with spherical proportional counter read-out



The spherical TPC concept:

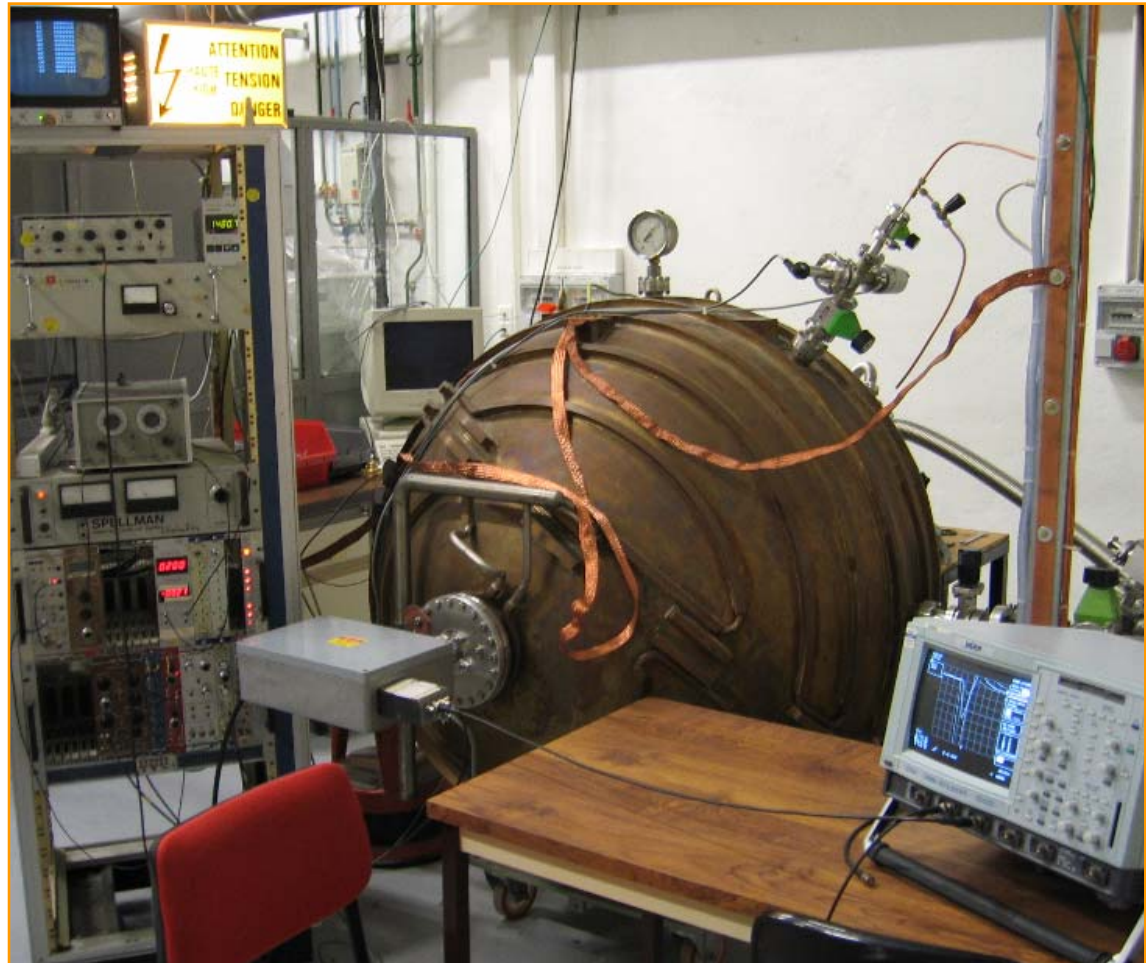
Advantages

- Natural focusing:
 - large volumes can be instrumented with a small readout surface and few (or even one) readout lines
- 4π coverage: better signal
- Still some spatial information achievable:
 - Signal time dispersion
- Other practical advantages:
 - Symmetry: lower noise and threshold
 - Low capacity
 - No field cage
- Simplicity: few materials. They can be optimized for low radioactivity.
- Low cost

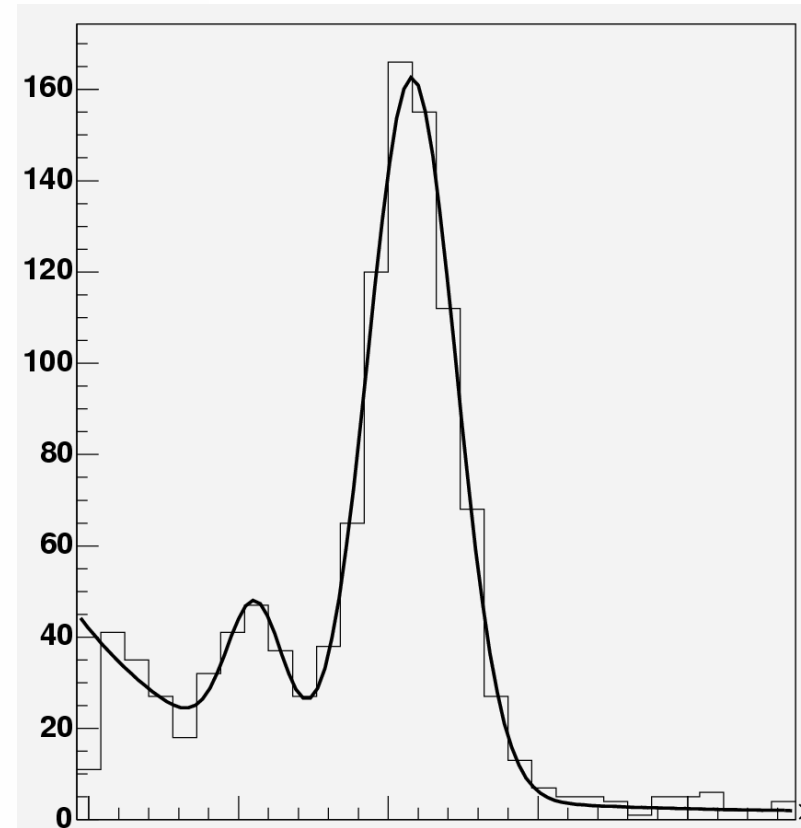
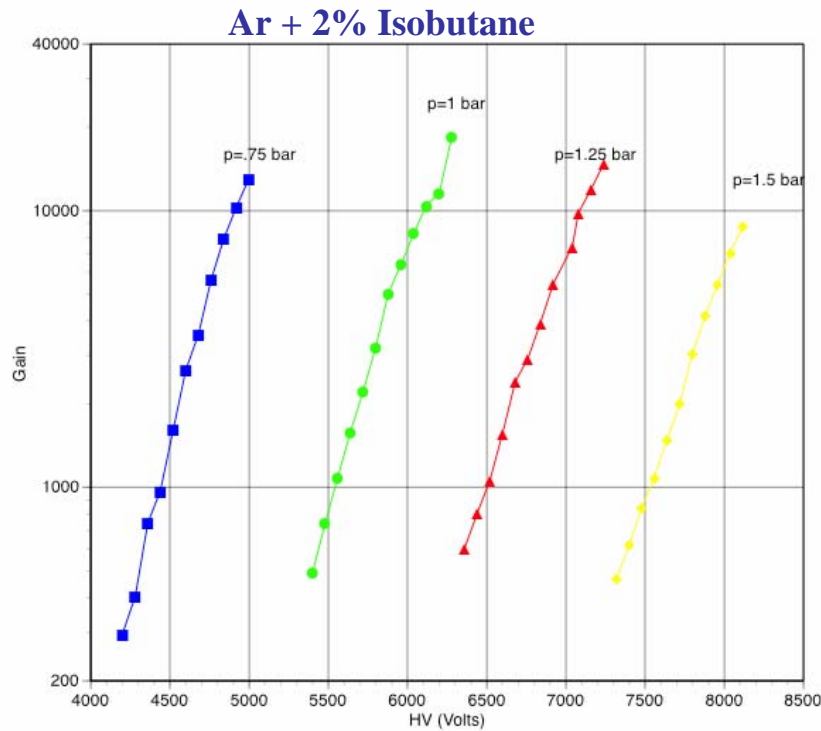
The way to obtain large detector volumes keeping low background and threshold

First prototype: the Saclay sphere

- $D=1.3$ m
- $V=1$ m³
- Spherical vessel made of Cu (6 mm thick)
- P up to 5 bar possible (up to 1.5 tested up to now)
- Vacuum tight: $\sim 10^{-6}$ mbar (outgassing: $\sim 10^{-9}$ mbar/s)



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■ Stability:

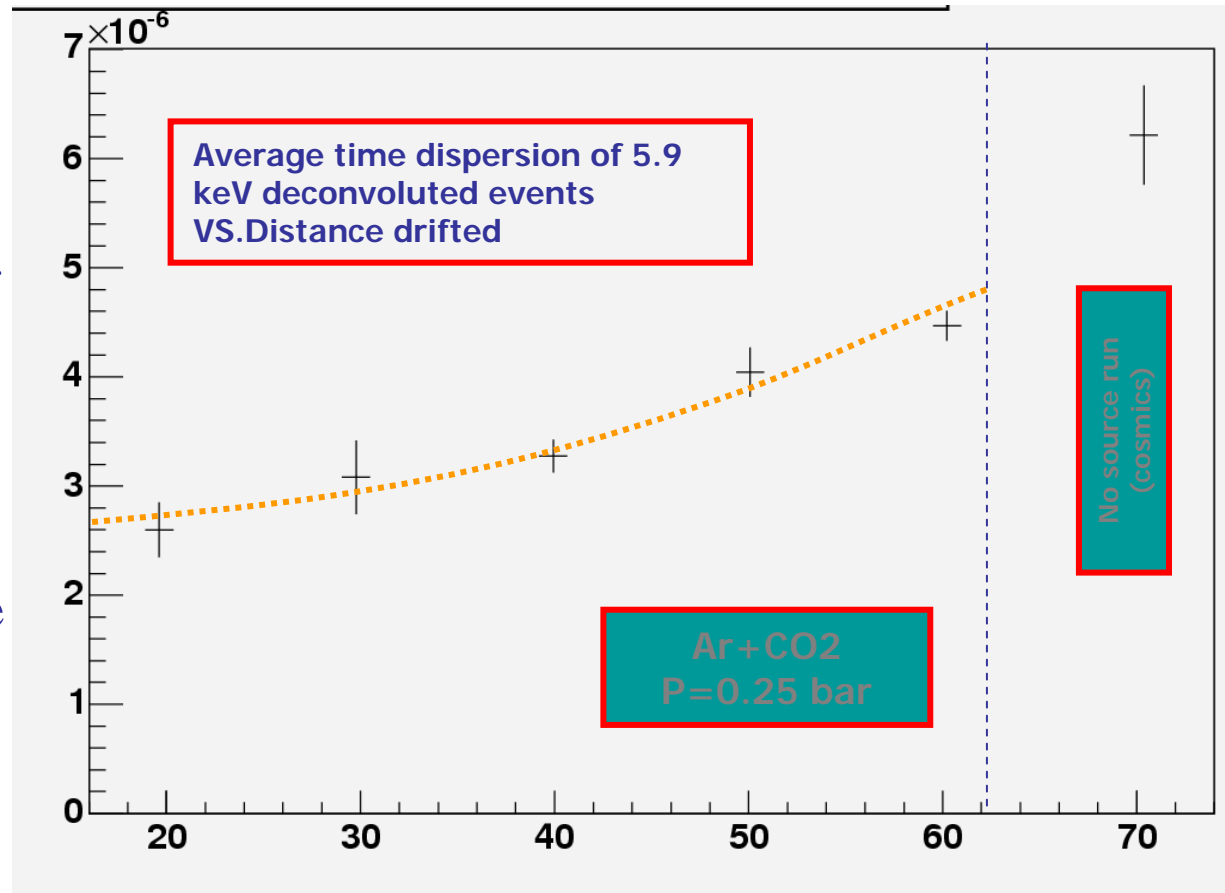
- tested up to ~3 months.
- No circulation of gas. Detector working in sealed mode. (1 pass through an oxysorb filter)

■ No absorption observed

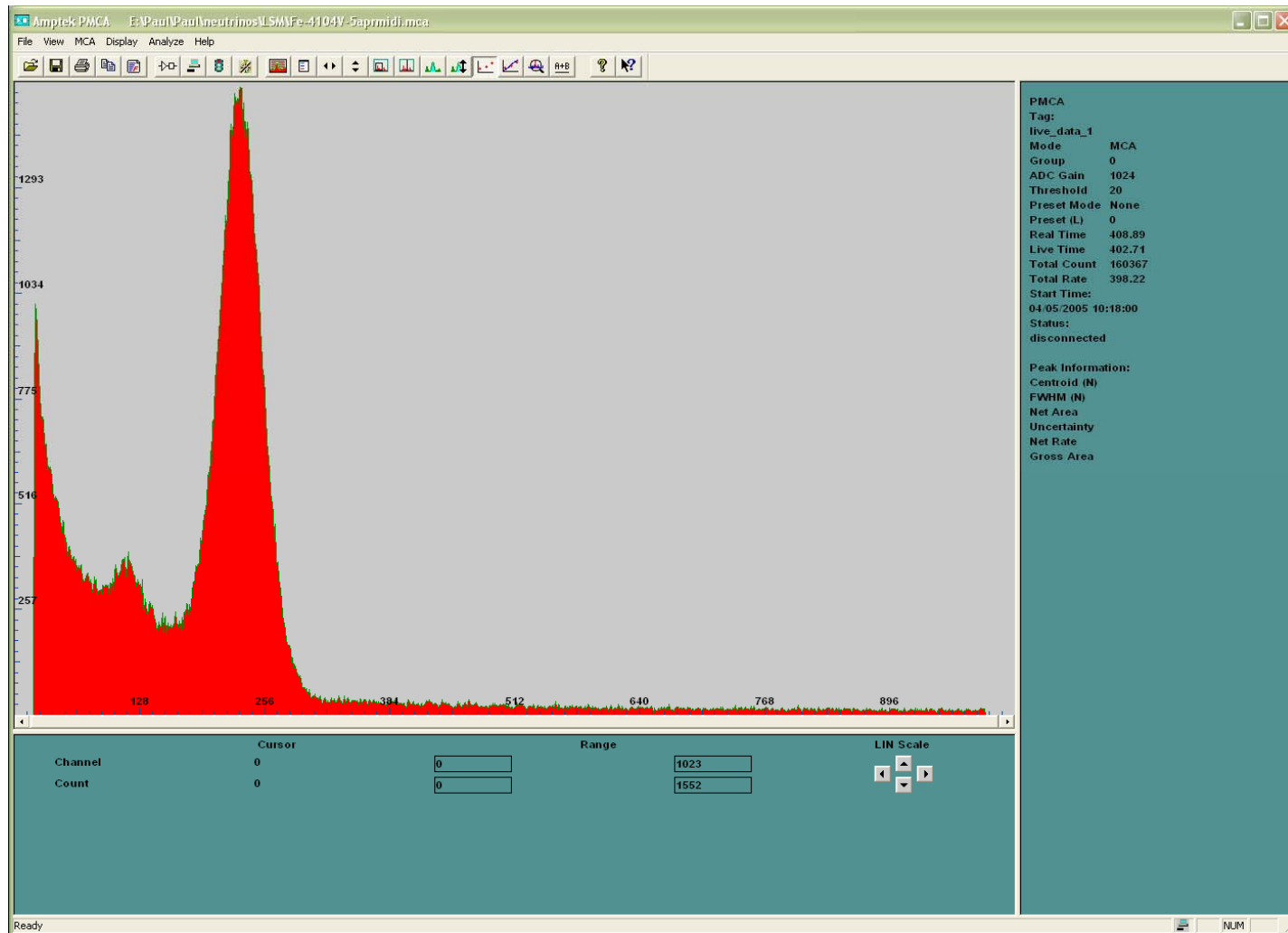
- Signal integrity preserved after 60 cm drift.
- Not high E needed to achieve high gain.

First results

- Even with a very simple (and slow) readout, we have proved the use of dispersion effects to estimate the position of the interaction (at least at ~ 10 cm level).
- Further test are under preparation to better calibrate (external trigger from Am source)



First underground tests in LSM 5-4-2004



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Long term program

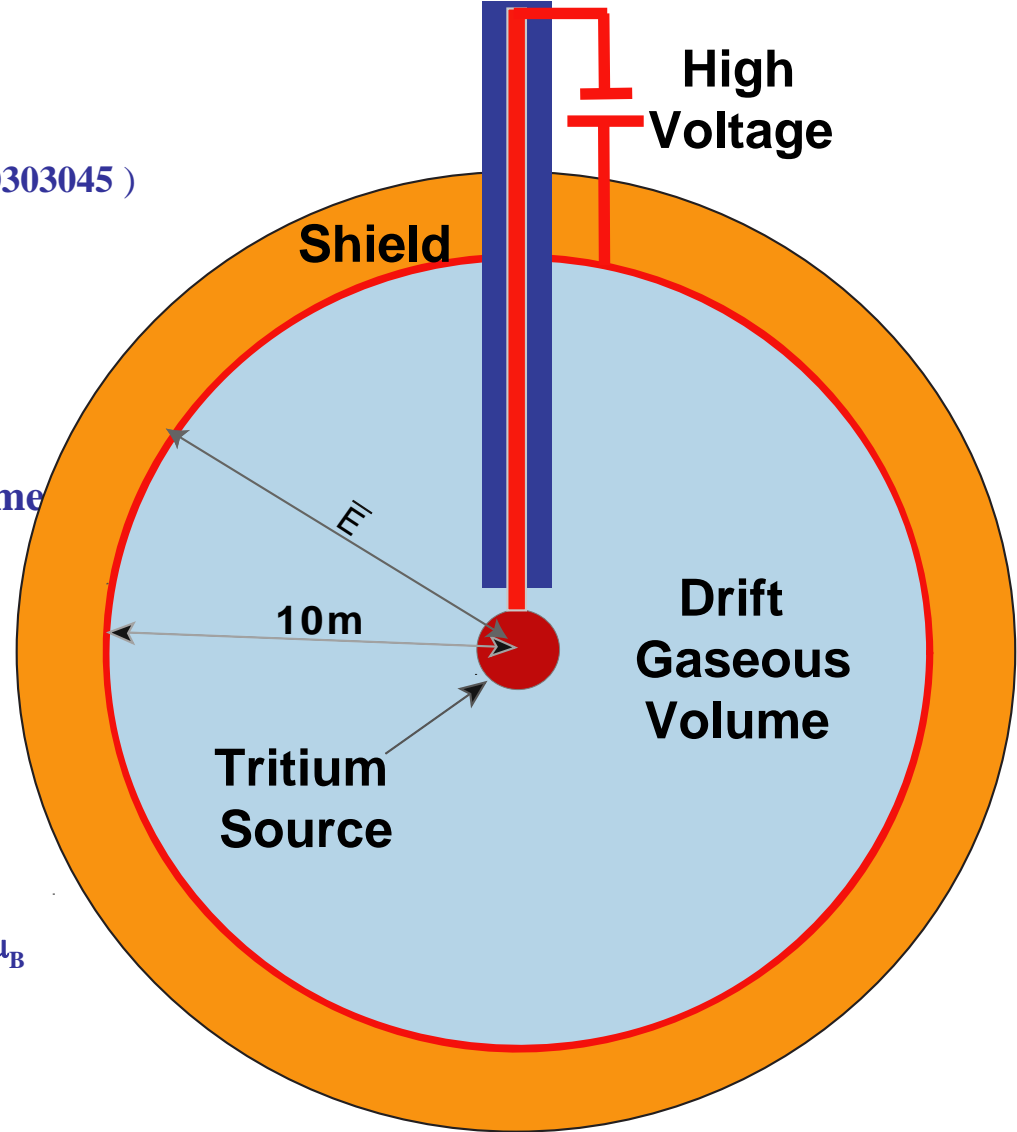
NOSTOS I. Giomataris, J. Vergados, hep-ex/0303045)

- Large Spherical TPC 10 m radius
- 200 MCi tritium source in the center
- Neutrinos oscillate inside detector volume
 $L_{23}=13$ m

$$P(\nu_e \rightarrow \nu_{\mu,\tau}) = \sin^2 2\theta_{13} \sin^2 \pi \frac{L}{L_{23}}$$

Objectives

- Measure θ_{13} (systematic free)
- Neutrino magnetic moment studies $\ll 10^{-12} \mu_B$
- Measurement of the Weinberg angle at low energy



Short term (3 year program) Neutrino-nucleus coherent elastic scattering

$$\sigma \approx N^2 E^2, \text{ D. Z. Freedman, Phys. Rev.D, 9(1389)1974}$$

1. Nuclear reactor measurement sensitivity with present prototype after 1 year run (2×10^7 s), assuming full detector efficiency:

- Xe ($\sigma \approx 2.16 \times 10^{-40} \text{ cm}^2$), 2.2×10^6 neutrinos detected, $E_{\text{max}} = 146 \text{ eV}$
- Ar ($\sigma \approx 1.7 \times 10^{-41} \text{ cm}^2$), 9×10^4 neutrinos detected, $E_{\text{max}} = 480 \text{ eV}$
- Ne ($\sigma \approx 7.8 \times 10^{-42} \text{ cm}^2$), 1.87×10^4 neutrinos detected, $E_{\text{max}} = 960 \text{ eV}$

Challenge : Very low energy threshold
We need to calculate and measure the quenching factor

2. Spallation source measurement with present prototype

3. Supernova neutrino detection with a 2nd demonstrator (4 m)

For $E_\nu = 10 \text{ MeV}$ $\sigma \approx N^2 E^2 \approx 2.5 \times 10^{-39} \text{ cm}^2$, $T_{\text{max}} = 1.500 \text{ keV}$

For $E_\nu = 25 \text{ MeV}$ $\sigma \approx 1.5 \times 10^{-38} \text{ cm}^2$, $T_{\text{max}} = 9 \text{ keV}$

Expected signal : 100 events (Xenon at $p=10 \text{ bar}$) per galactic explosion (including detector threshold and quenching factor)

Idea : A European or world wide network of several (tenths or hundreds) of such simple (one channel), robust and low cost detectors ($T_{\text{life time}} \gg 1 \text{ century}$)

Conclusions

- **Large volume TPCs are already used for rare event detection**
- **Combined with new MPGD precise detector can provide low energy threshold and recoil directionality**
- **A novel detector based in the spherical geometry with spherical proportional counter read-out has been successfully tested and it is under development.**
- **Many applications in low energy neutrino physics are open**