

Physics Case for the ILC Options

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at LCWS05, SLAC

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ILC-I (base design)

$$\sqrt{s} = 200-500 \text{ GeV}$$

$$L > 100 \text{ fb}^{-1}/\text{year}$$

$$P(e^-) \sim 80\%$$

$$\Sigma h, \nu\bar{\nu}h, t\bar{t}, W^+W^-$$

$$> 10^4 h, m_t, W\text{-couplings}$$

quantum #'s of new particles
EW tests

Physics case

independent of LHC

Needed as early as possible for

ILC-LHC synergy.

* next pages.

ILC-II, options

$$\sqrt{s} = 1 \text{ TeV}$$

$$e^-e^-, e^-e^+, \tau\tau$$

Giga Z

e^+ polarization

fixed target (beam dump)

tunnel length, MV/m

2nd int. region, crossing angle

e^+ source, pol. scheme

additional detector

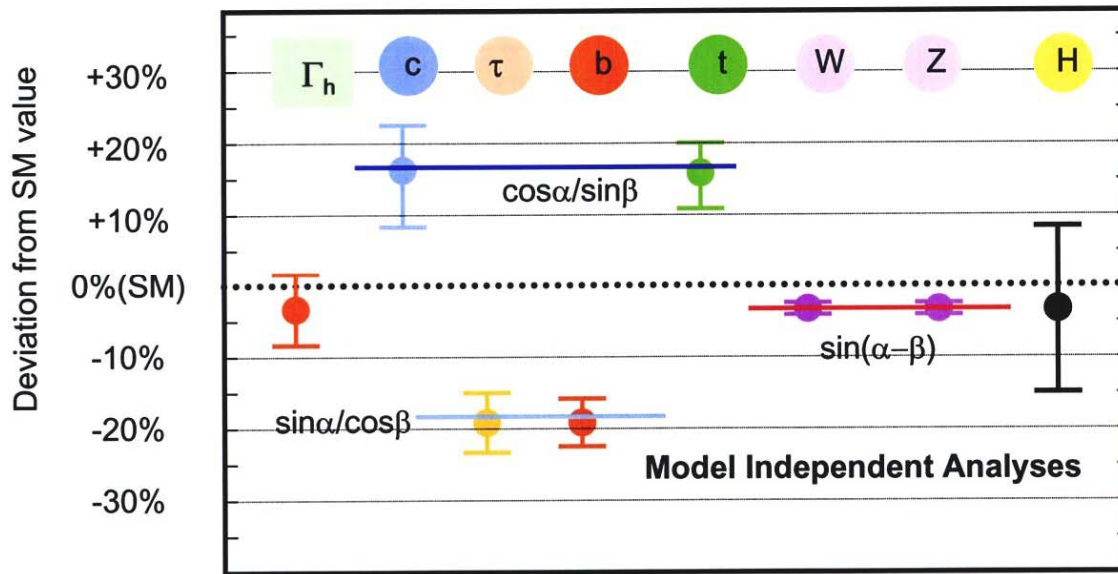
Inputs
from
LHC

cf. Kanemura et al

$eN \rightarrow \tau X$ measures $e\tau$ FCNC
 10^3 times better than $\tau \rightarrow \mu \nu, \mu \gamma, \dots$

SUSY or 2HDM

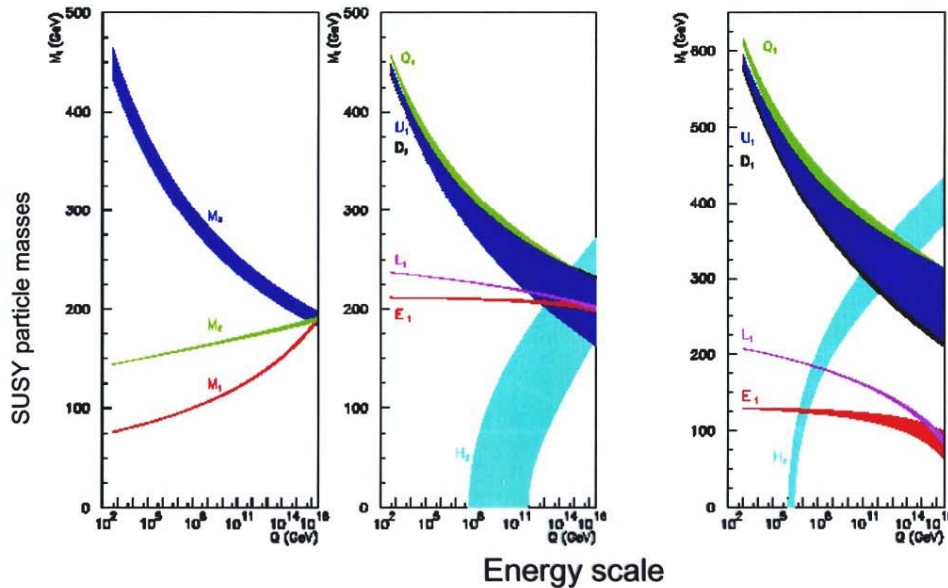
ILC



Determining SUSY breaking mechanism

LHC+ILC
Combined analysis

SUSY breaking scenario

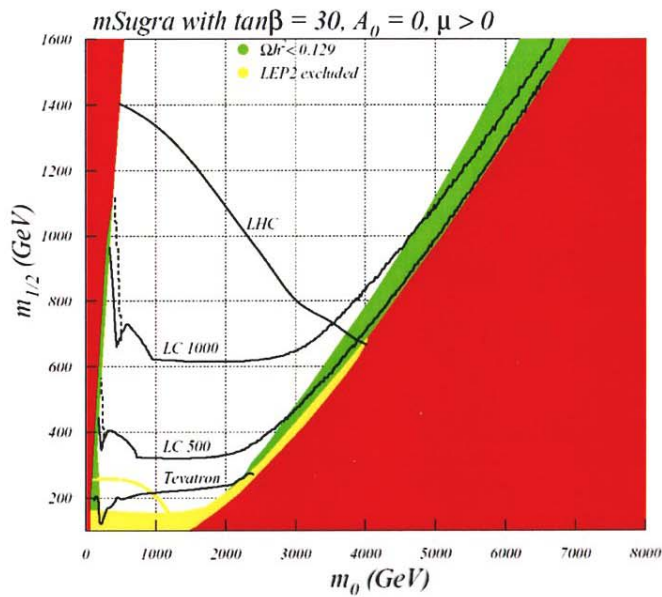


Super Gravity (mSUGRA)

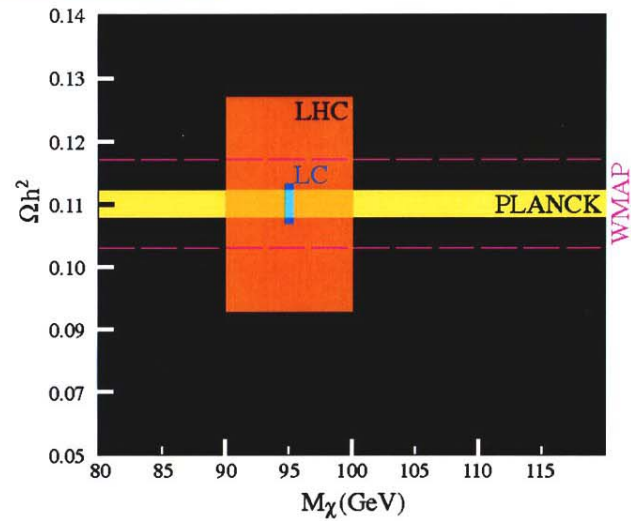
Gauge Mediation

G.A.Blair, W.Porod, and P.M.Zerwas

Cosmology: Dark Matter = LSP ?!



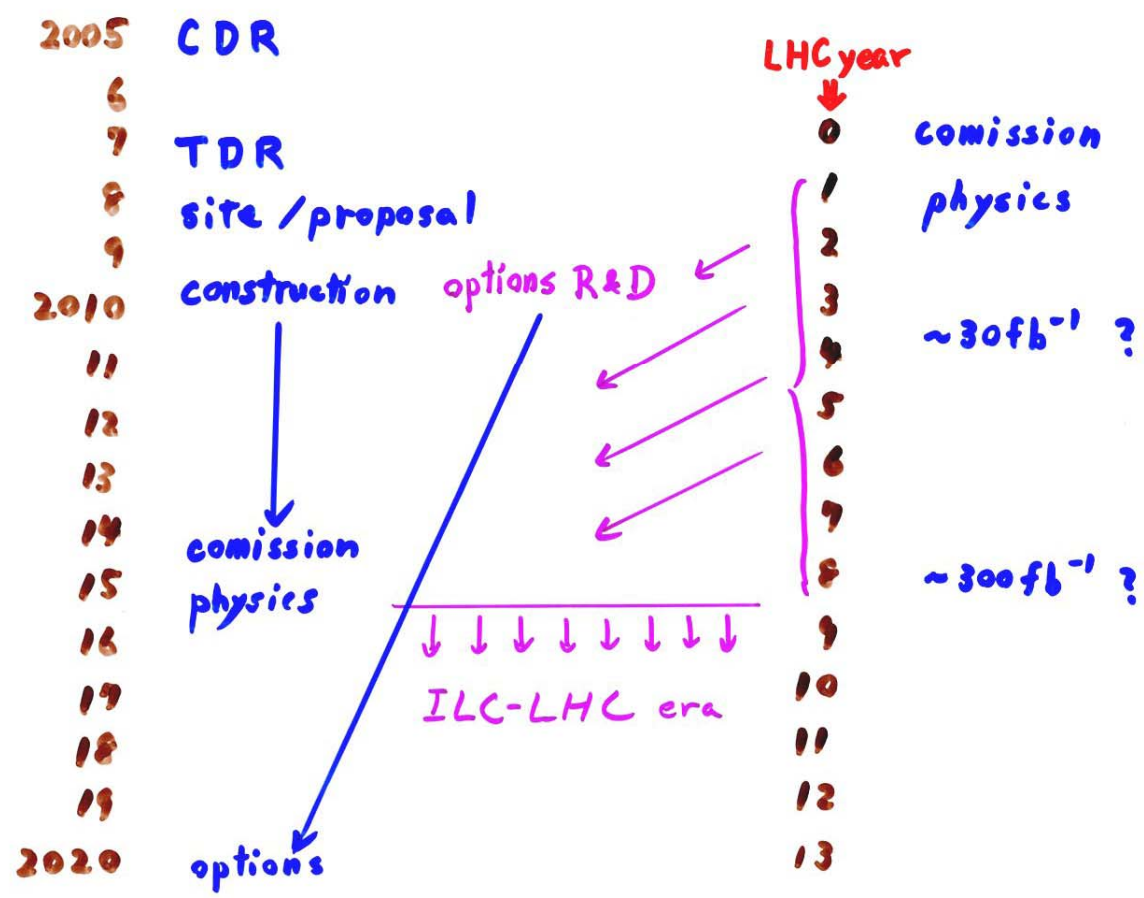
WMAP $.094 < \Omega h^2 < .128$ (2 sigma)



'WMAP'	7 %
LHC	~15 %
'Planck'	~2 %
ILC	~3 %

ILC

LHC



e^-e^-

easy to realize

$$L_{e^-e^-} \sim (0.1 - 0.3) L_{e^+e^-}$$

3 physics channels

$$e_L^- e_L^- \rightarrow \nu_L \nu_L W^- W^- \dots \dots W^- W^- \text{ scattering}$$

$$\rightarrow \nu_L \nu_L H^{--} \dots \dots$$

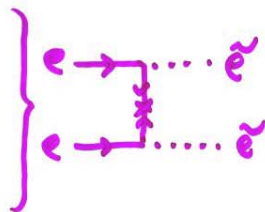
$$\rightarrow W^- W^-$$

$\dots \dots$ heavy Majorana ν 's @ TeV

$$\rightarrow \tilde{e}_L^- \tilde{e}_L^-$$

$$e_L^- e_R^- \rightarrow \tilde{e}_L^- \tilde{e}_R^-$$

$$e_R^- e_R^- \rightarrow \tilde{e}_R^- \tilde{e}_R^-$$



Majorana neutralino's

S-wave threshold

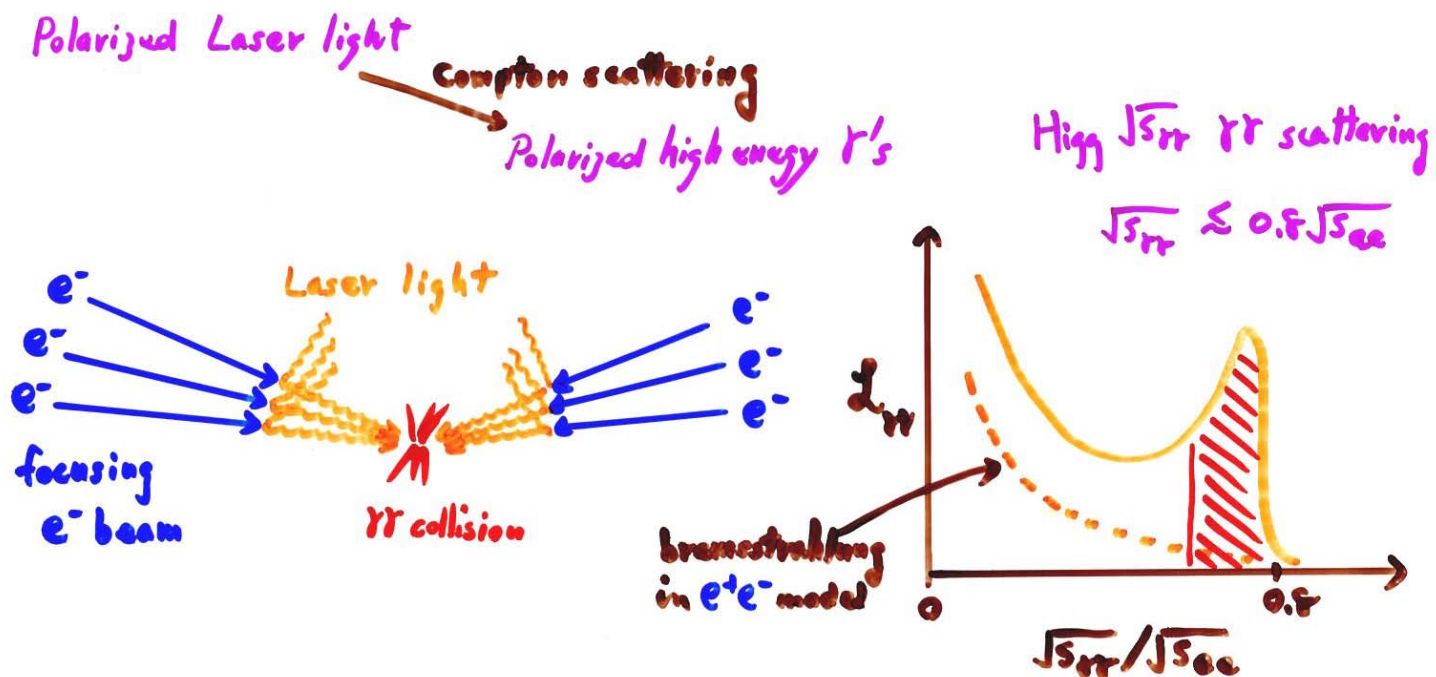
$$\Rightarrow m_{\tilde{e}}$$

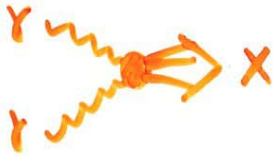
γγ

Compton back scattering of laser light off beam e^-

Large crossing angle. $\Rightarrow$ 2nd int. region.

Laser cavity for long beam bunch.



$\gamma\gamma$ 

$$J_X = 0$$

s-channel production of h, H, A and their mixture
if $\cancel{CP}$

circular pol. $\Rightarrow J_Z = 0$

enhances S/N of h, H, A signals

$$J_Z = \pm 2$$

spin 2 effect search

linear pol. $\Rightarrow CP = \pm$

select h, H vs A
aid $\cancel{CP}$ search

$$J^{PC} = 0^{++} \Rightarrow \text{s-wave threshold for } \tilde{l}^+ \tilde{l}^-, \tilde{\tau}, \tilde{\bar{\tau}}, \dots$$

Large cross sections : $\sigma_{WW}, \sigma_{\tilde{l}\tilde{l}}, \dots$

$$\mathcal{L}_{\gamma\gamma}(\sqrt{s_{\gamma\gamma}} > 0.8 \sqrt{s_{\gamma\gamma, \text{max}}}) \sim (0.1 - 0.3) \mathcal{L}_{e^+e^-}$$

We need $\gamma\gamma$ because

if light h : $\Gamma(h \rightarrow \gamma\gamma) / \Gamma(h \rightarrow b\bar{b}) \sim 2\%$

$\Rightarrow \mathcal{P}(h \rightarrow \gamma\gamma) \sim 2\%$ sensitive to new particles



$\cancel{CP}$ by $\gamma\gamma$ polarization far more sensitive than $e^+e^- \rightarrow h \rightarrow e^+e^-$ because $h\gamma\gamma$ is a loop amp.

if $m_{H,A} < 0.8\sqrt{s_{e^+e^-}}$: Possible discovery

CP & $\cancel{CP}$

$\rightarrow$ mixing (ϵ) and direct (ϵ')



interference allows us to measure the phase of $\phi_{\gamma\gamma}$ vertex.

in addition: $\gamma\gamma \rightarrow W^+W^-, t\bar{t}$
 $J=0, 2$

to supplement

$e^+e^- \rightarrow W^+W^-, t\bar{t}$ precision phys.
 $J=1$

e^+ Polarization

e^+ source $\sim 150 \text{ GeV } e^-$ undulator $\left. \begin{array}{l} \sim 5 \text{ GeV } e^- \text{ laser back-scattering} \end{array} \right\} \text{pol. } \gamma \rightarrow \text{pol. } e^+$

T. Omori et al., K. Mönig

The roles of e^+ polarization in e^+e^- collisions

e^+e^- annihilation $\rightarrow e_L \bar{e}_L$ & $e_R \bar{e}_R$ only (e-chirality)

W-exchange processes $\rightarrow e_L \bar{e}_L$ only (survive)

$\Rightarrow P_{e^+} \sim 0.4 \left\{ \begin{array}{l} \text{helps } \left\{ \begin{array}{l} \text{enhancing of } \sigma(e_L \bar{e}_L \rightarrow X) \\ \text{reducing b.g. from } e_L \bar{e}_L \rightarrow u^+ w^- \text{ etc.} \end{array} \right. \leftarrow \star \\ \text{increases } P_{\text{eff}} \text{ from } P_{e^-} = 0.2 \text{ to } P_{\text{eff}} = 0.95 \end{array} \right.$

Transverse Pol. of both $e^\pm$ beams could be useful $\star \text{ ex.) } e^+e^- \rightarrow \tilde{\tau}^+ \tilde{\tau}^-$

ex: If $\left| \begin{array}{c} \text{diagram with } W_3, B \\ e_L, \bar{e}_L \end{array} \right|^2 \gg \left| \begin{array}{c} \text{diagram with } B \\ e_R, \bar{e}_R \end{array} \right|^2$ Then $\left(\begin{array}{c} \text{diagram with } W_3, B \\ e_L, \bar{e}_L \end{array} \right) \left(\begin{array}{c} \text{diagram with } B \\ e_R, \bar{e}_R \end{array} \right)$ can be probed by $P_\perp(e^+)$

Giga Z

10^9 polarized Z's $\Rightarrow$ A_{LR} precision measurement

• LEP1 A_{FB}^b suffers from hadron/QCD uncertainty

but $\frac{\delta A_{LR}}{A_{LR}} \sim \frac{\delta P_{e^-}}{P_{e^-}} \sim 0.5\%$

With e^+ Polarization we can measure $\sigma_{LR}, \sigma_{RL}, P_{e^-}, P_{e^+}$ from

$$\left. \begin{aligned} \sigma_{++} &= \frac{1-P}{2} \frac{1+\bar{P}}{2} \sigma_{LR} + \frac{1+P}{2} \frac{1-\bar{P}}{2} \sigma_{RL} \\ \sigma_{+-} &= \frac{1-P}{2} \frac{1-\bar{P}}{2} \sigma_{LR} + \frac{1+P}{2} \frac{1+\bar{P}}{2} \sigma_{RL} \\ \sigma_{-+} &= \frac{1+P}{2} \frac{1+\bar{P}}{2} \sigma_{LR} + \frac{1-P}{2} \frac{1-\bar{P}}{2} \sigma_{RL} \\ \sigma_{--} &= \frac{1+P}{2} \frac{1-\bar{P}}{2} \sigma_{LR} + \frac{1-P}{2} \frac{1+\bar{P}}{2} \sigma_{RL} \end{aligned} \right\} \Rightarrow A_{LR} = \frac{\sigma_{LR} - \sigma_{RL}}{\sigma_{LR} + \sigma_{RL}}$$

with 4×10^{-5} error

$\Rightarrow \delta \sin^2 \theta_w^{\text{eff}} \sim 1 \times 10^{-5} !!$

Note $\frac{\delta G_F}{G_F}, \frac{\delta m_Z}{m_Z} \sim 2 \times 10^{-5}, \frac{\delta \alpha}{\alpha(m_Z^2)} \sim 25 \times 10^{-5}$ Powerful EW constraints!

Physics case for ILC-I (base)

$$\sqrt{s} = 200 - 500 \text{ GeV}$$

$$L > 100 \text{ fb}^{-1}/\text{yr}$$

$$P(e^-) \sim 80\%$$

is clear indep. of LHC.

but is much stronger if its commission is in the 1st 10y's of LHC



Proposed Time Schedule should be respected as much as possible

- ① Technological reliability
- ② Cost ③ Time Schedule
- ④ Flexibility to accommodate options later.

★ Longer tunnel for TeV

★ 2nd Int. Sect. for $\gamma\gamma$

LHC scenarios

A: finds nothing (not even a Higgs)

→ n-page

⇒ ILC-I Zh search, $t\bar{t}$, W^+W^- prec. exp.

⇒ Giga-Z + $P(e^+)$ check our starting point.

⇒ TeV search

⇒ $\gamma\gamma$ $t\bar{t}$, W^+W^- prec. exp.

B: finds a light Higgs only

⇒ ILC-I precision Higgs, $t\bar{t}$, WW

⇒ $\gamma\gamma$ $\Gamma(h \rightarrow \gamma\gamma)$, $S\bar{S}$, heavy H.A search

⇒ TeV $t\bar{t}h$, hhh , search

C: finds a heavy Higgs only

⇒ ILC-I confirm nothing lighter, prec.

⇒ TeV, $\gamma\gamma$ ZH , $\nu\bar{\nu}H$, $\gamma\gamma \rightarrow H$

⇒ Giga-Z + $P(e^+)$ hint of new phys. signal

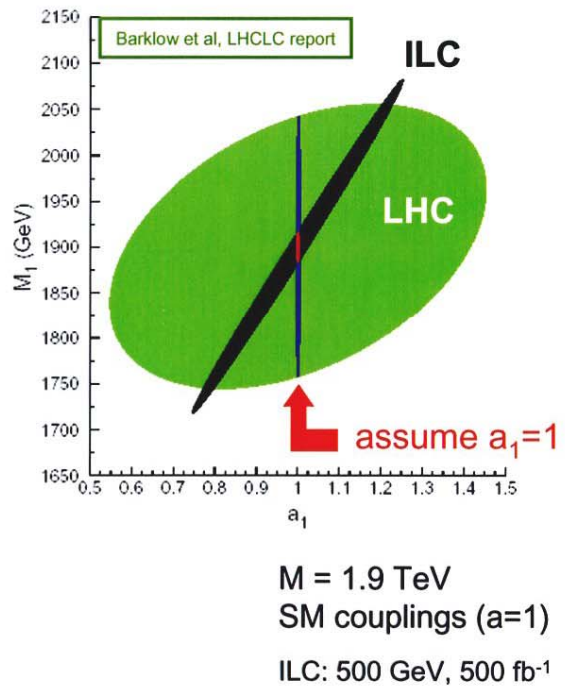
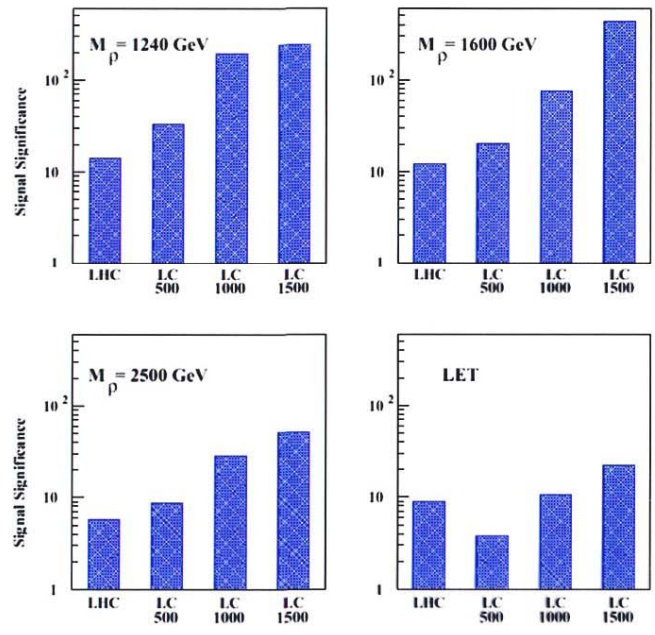
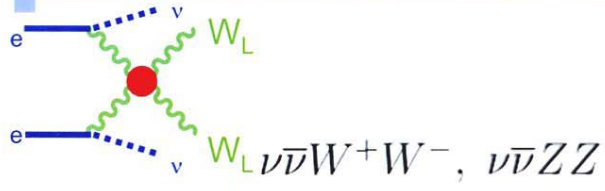
D: finds many new particles, non-SM effects

⇒ ILC-I prec. spectroscopy

⇒ TeV " "

⇒ $\gamma\gamma$ $J=0$ channel, scalar sector

sensitivity to new heavy resonances in $ee \rightarrow \nu\nu WW$



ILC is very sensitive to $\nu\nu WW, \nu\nu ZZ$

We are all most grateful for the **ITRP** members

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V. Soskol, M. Sugawara, D. Plane

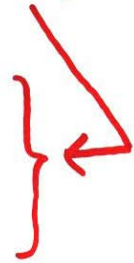
for making the toughest decision. I feel that

the most important decision for the **ILC** has already been made!

Our task is to identify and pursue the best strategy to optimize the physics outputs with a reasonable cost. I think

ILC-I should start as early as possible $\Rightarrow$ **ILC-LHC** synergy
" should leave the flexibility for

- * TeV upgrade
- * 2'nd Int. Sect. for $\gamma\gamma$ collider
- * Giga-Z with e^+ polarization.



Let us work together !

to meet the deadline.