

SLAC - PUB - 3742

July 1985

(T/E)

STRONG INTERACTION CONTRIBUTIONS
TO ONE LOOP LEPTONIC PROCESSES

B. W. LYNN*

*Stanford Linear Accelerator Center
Stanford University, Stanford, California, 94305*

and

G. PENSO

*Dipartimento di Fisica, Università di Roma, Italy
INFN, Sezione di Roma*

and

C. VERZEGNASSI[†]

*Dipartimento di Fisica Teorica, Università di Trieste, Italy
INFN, Sezione di Trieste
ISAS, Trieste*

Submitted to *Physical Review D*

* Work supported by the Department of Energy, contract DE - AC03 - 76SF00515.

† Work supported by Ministry of Public Instruction, Italy

ABSTRACT

We classify all strong interaction contributions to all four-lepton processes through one loop in electroweak $SU_2 \times U_1$ and to all orders in strong interactions. We show that those parts which are not reliably calculated in perturbative QCD are all related to a certain integral over the *total* cross section for $e^+e^- \rightarrow$ hadrons at low energies. We evaluate this integral for the most recent data and find that, for most four-lepton processes of interest, it is dominated by the time-like $|q'^2|$ region from 1 to 100 GeV^2 . We show that the associated theoretical strong interaction uncertainty is a factor of ~ 2 smaller than previously published estimates. We give the strong-interaction contribution and associated theoretical uncertainty for future SLAC, CERN and FNAL precision experiments and show that the theoretical uncertainty is quite small, thus allowing precision tests of the electroweak theory at the one-loop level.

1. Introduction

One of the most attractive features of the new generation of high energy accelerators will be their ability to study leptonic processes with great precision, thus gaining access to information about new currents and one loop electroweak radiative corrections. These corrections depend intimately on the gauge structure of the theory and, even within the context of $SU_2 \times U_1$, vary considerably depending on which representations of particles, even very heavy ones, are included in the model. Thus, by studying radiative corrections to leptonic processes, we can hope to see effects of new particles, even if they are too heavy to be produced directly. For example, there are measurable corrections to the various asymmetries in $e^+e^- \rightarrow \mu^+\mu^-$, especially the initial state longitudinal polarization asymmetry A_{LR} , on Z^0 resonance (where statistics will be high at LEP/SLC) within the context of the standard model of Glashow, Salam and Weinberg (GSW).¹ There are also measurable shifts from new particles (extra quarks and leptons, SUSY, technicolor, etc.) from beyond GSW which enter at the one-loop level. Some generic values for the shifts due to various one-loop effects are displayed in Table I for various precision measurements.²

There is one problem with this scenario. At the one-loop level there are hadronic effects due to the presence of strongly interacting particles in the various vacuum polarization amplitudes and thus there are strong interaction contributions even to leptonic processes and to the masses and widths of the W^\pm and Z^0 . Any theoretical uncertainty induced by strong interactions must be understood before the one-loop effects of new physics can be de-convoluted from the leptonic data. In this paper we study the effects of familiar quarks and hadrons on all four-lepton processes to one loop. We show that the strong interaction uncer-

tainties induced in the various precision asymmetries and mass measurements are smaller than most contributions of new particles listed in Table I and show that by remeasuring the total cross section for $e^+e^- \rightarrow$ hadrons in the time-like energy region $1 \text{ GeV}^2 \leq |q^2| \leq 100 \text{ GeV}^2$ with greater accuracy (to say 5%) it could be reduced much further.

Let us write down² the most general neutral and charged current four fermion matrix elements, including all one-loop electroweak corrections and strong interactions to all orders, in electroweak $SU_2 \times U_1$ where the internal symmetry breaking is done primarily by Higgs doublets. If external fermion masses are neglected, all external fermion vertices are helicity conserving and all cross sections may be written in terms of effective matrix elements where the initial state left-handed isospin I_3 and electric charge Q as well as the final state I'_3, Q'_3 are specified. We choose a renormalization scheme³ where α , the muon decay constant G_μ (i.e. the two best known electroweak constants of Nature) and M_Z , the Z^0 mass (expected to be measured very precisely by SLC/LEP), are used as precise input data. Then, in *Euclidean metric* ($q^2 = \bar{q}^2 - q_0^2$), the neutral current matrix element (normalized to 1 for photon exchange in $e^+e^- \rightarrow \mu^+\mu^-$) is

$$\begin{aligned}
\mathcal{M}_{pp'} = & Q \frac{1}{1 - \Delta_\alpha(q^2) - i \text{Im} \Pi'_{AA}(q^2)} \left(\frac{-s}{q^2} \right) Q' \\
& + \left[\frac{I_3 - Q(s_\theta^2 + \Delta_p(q^2) - i s_\theta c_\theta \text{Im} \Pi'_{ZA}(q^2))}{s_\theta c_\theta} \right]_p \\
& \times \frac{-s}{(q^2 + M_Z^2) (1 - \Delta_\rho(q^2) - 0.06) - i \text{Im} \Pi_{ZZ}(q^2)} \\
& \times \left[\frac{I'_3 - Q'(s_\theta^2 + \Delta_p(q^2) - i s_\theta c_\theta \text{Im} \Pi'_{ZA}(q^2))}{s_\theta c_\theta} \right]_{p'} + X^{NC}
\end{aligned} \tag{1}$$

while the charged current matrix element is

$$\begin{aligned}
M^{CC}(q^2) = & \frac{1}{2 \sin^2 \theta_w} (-s) \left[(1 - 0.06) [(q^2 + \cos^2 \theta_w M_Z^2)(1 - \Delta_W(0)) \right. \\
& \left. + \cos^2 \theta_w M_Z^2 \cdot \Delta_W(q^2)] - i \text{Im} \Pi_{WW}(q^2) \right]^{-1} + X^{CC} .
\end{aligned} \tag{2}$$

Here s is the Mandelstam variable ($q^2 = -s$ in the s channel) while Δ_α , Δ_ρ , Δ_p and Δ_W are certain finite combinations of the one-loop vacuum polarization amplitudes defined in Fig. 1 to be discussed later. Here, the quantities X^{NC} and X^{CC} represent the one-loop 1PI vertex, box and fermion self-energy contributions, the so-called “direct” coupling corrections. These do not suffer strong interaction effects for leptonic processes and we drop them from further consideration. We define the weak mixing angle used throughout the calculation:

$$\sin^2 \theta_w \equiv s_\theta^2 = \frac{1}{2} - \frac{1}{2} \left[1 - \frac{4\pi\alpha}{\sqrt{2}G_\mu M_Z^2 (1 - 0.06)} \right]^{1/2} \tag{3}$$

to include the largest part of the QED corrections to the renormalization of α from $q^2 = 0$ to $-M_Z^2$ by light quarks and leptons. There is no strong interaction uncertainty in the Born terms; the constant 0.06 is chosen to establish a convention in which s_θ^2 is directly calculable from α , G_μ and M_Z .

The “oblique correction” functions are finite combinations of electroweak one-loop 1PI vector boson self-energies as defined in Fig. 1 with $i, j = W^\pm, Z, A$ (photon) or SU_2 and QED currents $i, j = 1, 2, 3, Q$, and are to contain strong

interactions to all orders. We have

$$\begin{aligned}
\Delta_\alpha &= \text{Re} \left[\Pi'_{AA}(q^2) - \Pi'_{AA}(0) \right] \\
\Delta_\rho &= \text{Re} \left[\frac{\Pi_{ZZ}(-M_Z^2)}{M_Z^2} - \frac{\Pi_{WW}(0)}{M_W^2} + \frac{\Pi_{ZZ}(q^2) - \Pi_{ZZ}(-M_Z^2)}{q^2 + M_Z^2} \right] \\
\Delta_p &= \text{Re} \left[-s_\theta c_\theta \left[\Pi'_{ZA}(q^2) - \Pi'_{ZA}(-M_Z^2) \right] + \frac{s_\theta^2 c_\theta^2}{1 - 2s_\theta^2} (\Delta_\alpha(-M_Z^2) - 0.06) \right. \\
&\quad \left. + \frac{s_\theta^2 c_\theta^2}{1 - 2s_\theta^2} \left[-\Pi'_{AA}(-M_Z^2) + \frac{\Pi_{WW}(0)}{M_W^2} - \frac{\Pi_{ZZ}(-M_Z^2)}{M_Z^2} \right] - s_\theta c_\theta \Pi'_{ZA}(-M_Z^2) \right] \\
\Delta_W &= \text{Re} \left[\frac{s_\theta^2}{1 - 2s_\theta^2} (\Pi'_{AA}(0) + 0.06) \right. \\
&\quad \left. - \frac{\Pi_{WW}(q^2)}{M_W^2} + \frac{c_\theta^2}{1 - 2s_\theta^2} \frac{\Pi_{ZZ}(-M_Z^2)}{M_Z^2} - \frac{s_\theta^2}{1 - 2s_\theta^2} \frac{\Pi_{WW}(0)}{M_W^2} \right].
\end{aligned} \tag{4}$$

It is easy to see that there can only be four such functions; in $SU_2 \times U_1$, there are only four vector self-energies, $Z - Z$, $A - A$, $Z - A$ and $W - W$, and these self-energies will of course appear at one-loop. The quantities $\Pi'_{AA}(0)$, $\Pi_{ZZ}(-M_Z^2)$ and $\Pi_{WW}(0)$ appear because we used α , M_Z and G_μ respectively as physical input data and the experimental values of these quantities already include some radiative corrections.

Now it is simple task to track down the strong interaction contributions in the $\text{Im} \Pi_{ij}$ and the Δ_i . Let us assume that the influence of strong interactions is entirely due to the presence of quarks in the various vector self energies. Then, forgetting for the moment the top quark, we assume that we may calculate $\text{Im} \Pi_{ZZ}(q^2)/M_Z^2$, $\text{Im} \Pi_{WW}(q^2)/M_W^2$, $\text{Im} \Pi'_{ZA}(q^2)$ and $\text{Im} \Pi'_{AA}(q^2)$ for $|q^2| \gg m_q^2$ (where m_q^2 is a generic quark mass) dropping terms of order $\alpha_{em} m_q^2/q^2$ and using

perturbative QCD. For example,

$$\text{Im } \Pi'_{AA}(q^2) \simeq \text{Im } \Pi'_{AA}(q^2) \Big|_{\text{free field theory}} \left[1 + \frac{\alpha_{\text{QCD}}(q^2)}{\pi} + c \frac{\alpha_{\text{QCD}}^2(q^2)}{\pi^2} \right]. \quad (5)$$

We will, in fact, be interested in low energy neutral current neutrino scattering and so would have to evaluate $\text{Im } \Pi'_{ZA}(q^2)$ but this is in the t channel and so vanishes.

Similarly, we assume perturbative QCD to be valid for the calculation of

$$\text{Re } \frac{\Pi_{ZZ}(q^2)}{M_Z^2}, \quad \frac{\Pi_{WW}(0)}{M_W^2}, \quad \frac{\text{Re } \Pi_{WW}(q^2)}{M_W^2} \quad (6)$$

for all q^2 and $\text{Re } \Pi'_{AA}(q^2)$, $\text{Re } \Pi'_{ZA}(q^2)$ for $|q^2| \gg \Lambda_{\text{QCD}}^2$.

There are though two warnings to be made here. The first is the possibility of toponium $t\bar{t}$ resonances which could destroy our ability to use perturbative QCD for $|q^2| \sim m_{t\bar{t}}^2$. If toponium has some substantial mixing with the Z^0 this effect would need to be included in the analysis. The second warning is that in order to calculate absolute cross sections near the Z^0 or W^\pm poles or the widths of these particles to 1% accuracy we should properly include the 2-loop contributions to their respective propagators' imaginary parts. This we regard as beyond the scope of this paper. Note, however, that it is still safe to form asymmetries to one loop near Z^0 resonance because Z^0 propagator effects (and also the luminosity) cancel there. Either of these two effects could give strong interaction contributions to experiments at SLC and LEP and will be included in a further analysis.⁴

The hadronic contributions not suppressed by powers of m_q^2/M_Z^2 or calculable

in perturbative QCD, enter via the two finite combinations $\Delta_\alpha(q^2)$ and

$$\Delta_{ZA}(-M_Z^2, q^2) = \text{Re} [\Pi'_{ZA}(-M_Z^2) - \Pi'_{ZA}(q^2)] \quad (7)$$

for low q^2 . These two combinations, then, give all of the non-perturbative strong interaction effects for four-lepton processes at one loop and we will concentrate on these for the remainder of this paper.

We will show in Section 2 that strong interaction effects in Δ_{ZA} can be related to those in Δ_α , which in turn can be related to low energy data in $e^+e^- \rightarrow$ hadrons. We will use the most recent available data to evaluate the hadronic contribution to Δ_α . In Section 3 we will use this to give the hadronic contributions to and bounds on the hadronic uncertainties in various precision measurements to be performed by the CHARM II collaboration and by experimental groups at SLC, LEP and FNAL in the near future.

2.

As we saw in the previous section, all complicated strong interaction effects in one loop leptonic processes are contained in the two quantities $\Delta_\alpha(q^2)$ and Δ_{ZA} . Here, q^2 represents a four-momentum square which is typically small, $|q^2| \ll M_Z^2$ which prevents us from relying on free field theory (FFT) particularly when light quark contributions are involved.

In the case of $\Delta_\alpha(q^2)$, the problem can be circumvented since this quantity is directly related to $e^+e^- \rightarrow$ hadrons data. The situation is less simple in the case of Δ_{ZA} , where a more detailed analysis of flavor dependent effects is required. Π'_{ZA} is defined from the vacuum expectation value of the product of $J^{e.m.} = eJ^Q$

and $J_{\text{vector}}^{Z^0} = \frac{e}{s_\theta c_\theta} (J_{\text{vector}}^3 - s_\theta^2 J^Q)$ where J_{vector}^3 is the vector part of the third component of the weak isospin current. Thus Δ_{ZA} contains the flavor dependent term $\langle J_{\text{vector}}^3 J^Q \rangle$. Now write

$$J_{\text{vector}}^3 = \frac{1}{2} J^Q + \left(J_{\text{vector}}^3 - \frac{1}{2} J^Q \right)$$

and note that the hadronic part of the last term can be written

$$\begin{aligned} J_{\text{vector}}^3 - \frac{1}{2} J^Q &= -\frac{1}{12} [\bar{d}\gamma_\mu d + \bar{u}\gamma_\mu d + \bar{s}\gamma_\mu s + \bar{c}\gamma_\mu c \\ &\quad + \bar{b}\gamma_\mu b + \bar{t}\gamma_\mu t + \dots] \\ &= -\frac{1}{2} J^\omega + \frac{1}{4} J^\phi + \text{heavy quarks} \\ J^\omega &= \frac{1}{6} [\bar{d}\gamma_\mu d + \bar{u}\gamma_\mu u] \end{aligned} \tag{8}$$

$$J^\phi = -\frac{1}{3} \bar{s}\gamma_\mu s$$

$$J^\rho = \frac{1}{2} [\bar{u}\gamma_\mu u - \bar{d}\gamma_\mu d]$$

$$J_Q = J^\rho + J^\omega + J^\phi + \text{heavy quarks} .$$

Thus $J_{\text{vector}}^3 - \frac{1}{2} J^Q$ does not contain the (dominant) ρ meson current and is entirely of weak isospin $\vec{I} = 0$. Taking care with the various isospin components we have

$$\begin{aligned} \Pi'_{ZA} &= \frac{1}{s_\theta c_\theta} \left(\frac{1}{2} - s_\theta^2 \right) \Pi'_{AA} \\ &\quad + \frac{1}{4s_\theta c_\theta} \left[\Pi'_{AA}(\phi\phi) - 2\Pi'_{AA}(\omega\omega) \right] \\ &\quad + \frac{1}{4s_\theta c_\theta} \left[\Pi'_{AA}(\rho\phi) - 2\Pi'_{AA}(\rho\omega) - \Pi'_{AA}(\omega\phi) \right] \\ &\quad + \text{heavy quark terms} \end{aligned} \tag{9}$$

where $\langle J_\mu^a J_\nu^b \rangle = \delta_{\mu\nu} q^2 \Pi'_{AA}(ab)$ gives the relevant flavor contribution to the photon vacuum polarization with *vector currents* only.

Let us start our examination of (9) with the heavy quark components of Π'_{AA} . We will assume that all flavor mixing terms involving heavy quarks vanish. If we are in the spacelike q^2 region or even at $q^2 = 0$ we know from QCD sum rules⁵ that free field theory (FFT) plus calculable QCD should be a reliable approximation for the remainder of the heavy quark contribution. So this contribution can be straightforwardly evaluated; in fact, we find it to be very small.

The second term in the bracket contains all possible non-vanishing interference terms $(\rho\omega, \omega\phi, \rho\phi)$, e.g. $\sim \langle J_\mu^{(\rho)} J_\nu^{(\omega)} \rangle$. These are familiar quantities,⁵ their expression is well known and given by nonperturbative QCD condensates, isospin-conservation breaking terms or flavor mixing terms suppressed by factors $\alpha_{e.m.}(m_u^2 - m_d^2)/(GeV)^2$ or $\alpha_{e.m.}^2$ or suppressed via Zweig's flavor mixing rule. These contributions are orders of magnitude smaller than the leading one in Eq. (9). Thus, we see that, for those q^2 values which are relevant for $\bar{\nu}_\mu e, \nu_\mu e$ scattering, we can safely "reduce" $\Delta_{ZA}(Q_0^2, q^2) \equiv \text{Re} [\Pi'_{ZA}(Q_0^2) - \Pi'_{ZA}(q^2)]$ (where Q_0^2 is a suitable subtraction point where FFT can be used) to a sum:

$$\begin{aligned} \Delta_{ZA}(Q_0^2, q^2) &= \frac{(1/2 - s_\theta^2)}{s_\theta c_\theta} [\Delta_\alpha(Q_0^2) - \Delta_\alpha(q^2)] \\ &+ \frac{1}{4s_\theta c_\theta} \left[(\Delta_\alpha^{(\phi\phi)}(Q_0^2) - \Delta_\alpha^{(\phi\phi)}(q^2)) - 2(\Delta_\alpha^{(\omega\omega)}(Q_0^2) - \Delta_\alpha^{(\omega\omega)}(q^2)) \right] \\ &+ \text{"small" terms with negligible errors.} \end{aligned} \tag{10}$$

To be more precise, let us consider the specific value $q^2 = 0$. We find in this case:

$$\begin{aligned}
\Delta_{ZA}(Q_0^2, 0) &= \frac{(1/2 - s_\theta^2)}{s_\theta c_\theta} [\Pi'_{AA}(Q_0^2) - \Pi'_{AA}(0)] \\
&\quad - \frac{1}{4s_\theta c_\theta} \left[\frac{\alpha}{9\pi} \left(\ln \frac{|Q_0^2|}{m_c^2} - \frac{5}{3} \right) \right] \\
&\quad + \frac{1}{4s_\theta c_\theta} \left[\left(\Pi'^{(\phi\phi)}_{AA}(Q_0^2) - \Pi'^{(\phi\phi)}_{AA}(0) \right) \right. \\
&\quad \left. - 2 \left(\Pi'^{(\omega\omega)}_{AA}(Q_0^2) - \Pi'^{(\omega\omega)}_{AA}(0) \right) \right] \\
&\quad + \text{calculable small terms} .
\end{aligned} \tag{11}$$

Q_0^2 must be such that we can safely use FFT for $\Pi'_{AA}(Q_0^2)$, $\Pi'_{ZA}(Q_0^2)$.

The evaluation of the last parenthesis in the r.h.s. of Eq. (11) could, in principle, be performed if precise flavor isospin tagging data in $e^+e^- \rightarrow$ hadrons existed; we could compute it phenomenologically from its definition:

$$\Delta_\alpha^{(\phi\phi)}(Q_0^2) - 2\Delta_\alpha^{(\omega\omega)}(Q_0^2) = \frac{\alpha Q_0^2}{3\pi} \int_{-\infty}^0 \frac{dq'^2}{q'^2(q'^2 - Q_0^2)} (R^{(\phi)}(q'^2) - 2R^{(\omega)}(q'^2)) \tag{12}$$

for $|Q_0^2| \ll M_Z^2$ where

$$R^{(\phi)} = \frac{\sigma_{e^+e^- \rightarrow \phi, \phi', \dots}}{\sigma_{e^+e^- \rightarrow \mu^+ \mu^-}}, \quad R^{(\omega)} = \frac{\sigma_{e^+e^- \rightarrow \omega, \omega', \dots}}{\sigma_{e^+e^- \rightarrow \mu^+ \mu^-}} . \tag{13}$$

In practice, the data are not to our knowledge available at the moment. We will still be able to give a reasonable estimate because the asymptotic part ($|q'^2| \gg m_s^2$) of the numerator of (12) vanishes exactly in FFT. Thus, it is only the region of small q'^2 values, $|q'^2| \lesssim 1 \text{ GeV}^2$, which can effectively contribute. For this

region it is certainly a good approximation to consider the ω, ϕ contribution as due to the dominant resonances treated in the narrow width representation. Thus we can write this contribution as:

$$\begin{aligned} \frac{\alpha Q_0^2}{3\pi} \int_{-\infty}^0 \frac{dq'^2}{q'^2(q'^2 - Q_0^2)} \left[R^{(\phi)}(q'^2) - 2R^{(\omega)}(q'^2) \right] \\ \simeq \frac{3Q_0^2}{\alpha} \left[\frac{\Gamma_{\phi \rightarrow e^+e^-}}{m_\phi} \frac{1}{(Q_0^2 + m_\phi^2)} - \frac{2\Gamma_{\omega \rightarrow e^+e^-}}{m_\omega} \frac{1}{(Q_0^2 + m_\omega^2)} \right]. \end{aligned} \quad (14)$$

Numerically, this turns out to be $\simeq -0.00025$ at the spacelike point $Q_0^2 = 79 \text{ GeV}^2$; we shall assume that the possible error on this estimate is equal to the estimate itself; although based on QCD sum rules we feel that our approximation should not be that bad. A final comment on this $(\phi - 2\omega)$ term is that if we had used the (a priori unjustified) FFT evaluation of the l.h.s. of Eq. (12), we would have obtained a result $\simeq -\frac{2\alpha}{9\pi} \ln \frac{m_s}{m}$ (where m denotes the common value of $m_{u,d}$) which for any reasonable choice of the m_s/m ratio turns out to be numerically very close to our estimate in Eq. (14). A similar narrow width estimate can be given when q^2 increases from zero to spacelike values, with minor modification, and we shall not discuss it further.

We now give an explicit evaluation of Δ_α from the most recent available $e^+e^- \rightarrow$ hadrons data and discuss in some detail the related experimental errors. Figure 2 shows the results of our evaluation of the relevant expression (*written in Euclidean metric*):

$$\Delta_\alpha(Q_0^2) = \frac{\alpha}{3\pi} Q_0^2 \int_{-\infty}^0 \frac{dq'^2 R(q'^2)}{q'^2(q'^2 - Q_0^2)} ; \quad (15)$$

$$R = \frac{\sigma_{e^+e^- \rightarrow \text{hadrons}}}{\sigma_{e^+e^- \rightarrow \mu^+\mu^-}}$$

when Q_0^2 varies in the range $-200 \text{ GeV}^2 \leq Q_0^2 \leq 200 \text{ GeV}^2$. For comparison, we have also included an older estimate by Paschos.⁶ One notices that the two tend to differ somewhat in the very low q^2 region, i.e. in that dominated by the very low energy e^+e^- data, for which we have taken the most recent results, quoting an accurate estimate of the systematic error.⁷ This was not available for the earlier estimate. Assuming for the remaining higher energy data⁸ a realistic systematic error of 10% (see later), we have computed what we consider to be a realistic error for $\Delta_\alpha(Q_0^2)$ in the Q_0^2 range above. To get a deeper understanding of the details of our evaluation, we have divided the integration region of Eq. (14) into six parts, i.e.:

region a: $|q'^2| \leq (0.8 \text{ GeV})^2$, where one expects⁷ the systematic error to be no larger than $\sim 4\%$

region b: $(0.8 \text{ GeV})^2 \leq |q'^2| \leq m_\psi^2$, where the systematic error⁸ is expected to range from $\sim 4\%$ to 15%

region c: $(m_\psi)^2 \leq |q'^2| \leq (m_\gamma)^2$

region d: $(m_\gamma)^2 \leq |q'^2| \leq (46 \text{ GeV})^2$

region e: $(46 \text{ GeV})^2 \leq |q'^2| \leq (80 \text{ GeV})^2$

region f: $(80 \text{ GeV})^2 \leq |q'^2|$.

In the last four regions, we accepted the quoted⁸ systematic error of 10%.

Having divided the integration range in this way, we can now see how much of the overall error, at variable q^2 , comes from the different regions. Considering, e.g. the specific spacelike value $Q_0^2 = 79 \text{ GeV}^2$ which corresponds to the “optimal” subtraction point to be discussed later, we have listed in Table I the individual contributions coming from the six regions a-f. As one sees, the overall result is

$$\Delta_\alpha(79 \text{ GeV}^2) = 0.0145 \pm 0.0013 . \quad (16)$$

Note that the error is ~ 2 times smaller than the previous estimate ± 0.002 by Sirlin³ and, consequently will lead to smaller errors in A_{LR} than previously estimated.¹ Of the ± 0.0013 error, only ± 0.0001 comes from region (a), while ± 0.0010 comes from the two regions (b) and (c). Thus, our result Eq. (16) is *not* dominated by the very low energy e^+e^- data, but rather from those data approximately in the timelike $|q'^2|$ region from 1 to 100 GeV^2 . More precise measurements of the threshold region would consequently not be of great help for $Q_0^2 = 79 \text{ GeV}^2$. Note however that if the experimental error in the region $1 \leq |q'^2| \leq 100 \text{ GeV}^2$ were reduced to 5%, the error on $\Delta_\alpha(79 \text{ GeV}^2)$ would be ± 0.0008 , a substantial reduction. We urge experimentalists to reexamine this region in order to make the hadronic uncertainties in future SLAC/CERN/FNAL precision experiments completely negligible.

These same conclusions would in general apply to the range of spacelike Q_0^2 investigated, i.e. $|Q_0^2| \leq 400 \text{ GeV}^2$. In Fig. 3 we have shown in more detail the values of $\Delta_\alpha(Q_0^2)$ in this range, together with some error bars. Table III contains the contributions of the different regions to the overall results at a number of Q_0^2 values. As one sees, the contribution from the threshold region becomes less and

less relevant as Q_0^2 increases. As a general rule, in the whole range $|Q_0^2| \gtrsim 1 \text{ GeV}^2$ the main contribution, giving the largest fraction of the error, to the relevant quantities comes from the region of the data $(1 \text{ GeV})^2 \leq |q'^2| \lesssim (10 \text{ GeV})^2$.

These same conclusions apply for timelike $|Q_0^2|$ larger than approximately 150 GeV^2 . For smaller timelike Q_0^2 values, as can be seen in Fig. 1, the quantity $\Delta_\alpha(Q_0^2)$ is subject to oscillations due to contributions from regions a-f of various signs. As a consequence, the overall error, which is of the order of $\sim 8\%$ in the more favorable spacelike Q_0^2 or large timelike Q_0^2 cases, becomes somewhat larger ($\sim 20\%$).

As a final comment to motivate our choice of the optimal spacelike subtraction point $\tilde{Q}_0^2 = 79 \text{ GeV}^2$ for Eq. (16), we would like to point out that it is possible to derive bounds on the quantity $\Delta_\alpha(Q_0^2)$ in the spacelike region⁹ which are a consequence of the experimental value of the muon anomaly and of the assumptions that QED is correct and that QCD gives respectable predictions for the photon vacuum polarization in the spacelike region, in the spirit suggested by SVZ.⁵ In particular, it was shown in a previous paper¹⁰ that these bounds become optimal, i.e. most strict, at the point $Q_0^2 = 79 \text{ GeV}^2$, where one obtains the general result:

$$0.0115 \leq \Delta_\alpha(79 \text{ GeV}^2) \leq 0.0157 . \quad (17)$$

As one notices, the upper limit of this general bound (coming from theoretical considerations of a vastly different sort) is exactly saturated by the purely phenomenological evaluation based on e^+e^- data Eq. (15) and (16). Actually, the two different estimates are consistent over the whole spacelike region $|Q_0^2| \leq 150 \text{ GeV}^2$ where the general bounds can be derived. This strengthens our

belief in the correctness of the result Eq. (16) to be used in what follows.

Inserting Eqs. (16) and (14) in Eq. (11) we obtain (assuming $s_0^2 = 0.215$, i.e. $M_Z = 94 \text{ GeV}$):

$$\begin{aligned} \Delta_{ZA}(79 \text{ GeV}^2, 0) &= [0.0101 \pm 0.0009] - [0.0003] - [0.0002 \pm 0.0002] \\ &= 0.0096 \pm 0.0011 \end{aligned} \tag{18}$$

where the numbers on the r.h.s. represent the contributions to the overall quantity coming from the three pieces in Eq. (11). Thus, we see that the bulk of the result and of its error comes from the same e^+e^- data which determined the photon vacuum polarization Eq. (16), and the same considerations of that case still apply.

Having completed our numerical analysis of the two quantities, Eqs. (16) and (18), which are affected by strong interaction uncertainties, we are now ready to discuss what the effects of these uncertainties on a number of measurable quantities will be. This will be done in Section 3.

3.

Let us begin our analysis with a discussion of the contribution due to strong interactions to the theoretical prediction for the initial state polarization asymmetry A_{LR} for the process $e^+e^-_{L,R} \rightarrow \mu^+\mu^-$ at the Z_0 resonance, which will soon be measured at SLC. This is given by the expression:^{1,2}

$$A_{LR}(-M_Z^2) \equiv \left[\frac{\sigma(e^+e^-_L \rightarrow \mu^+\mu^-) - \sigma(e^+e^-_R \rightarrow \mu^+\mu^-)}{\sigma(e^+e^-_L \rightarrow \mu^+\mu^-) + \sigma(e^+e^-_R \rightarrow \mu^+\mu^-)} \right]_{q^2=-M_Z^2} \tag{19}$$

The contribution to this quantity due to u, d, s, c, b , and t quarks is easily written

from Eqs. (1) and (4)

$$\begin{aligned}
\delta A_{LR}^{(u \rightarrow t)}(-M_Z^2) &= \frac{-64 c_\theta^2 s_\theta^4}{(1 + v_\theta^2)^2} \Delta_\alpha^{(u \rightarrow t)}(-M_Z^2) + \text{const.} \\
&= \frac{-64 c_\theta^2 s_\theta^4}{(1 + v_\theta^2)^2} [(\Delta_\alpha(-m_Z^2) - \Delta_\alpha(79 \text{ GeV}^2)) \\
&\quad + \Delta_\alpha(79 \text{ GeV}^2)]^{(u \rightarrow t)} + \text{const.}
\end{aligned} \tag{20}$$

with $v_\theta = 4s_\theta^2 - 1$.

The first term in the bracket is evaluated using FFT while the second can be gotten from Table II. Collecting the various terms (remember that since $m_t \gg m_b$ there is a contribution from the ρ parameter $\Delta\rho(0)$) we have for $M_Z = 94 \text{ GeV}$, $m_b = 3m_c = 4.5 \text{ GeV}$ and $m_t = 30 \text{ GeV}$

$$\delta A_{LR}^{(u \rightarrow t)} = -0.0615 \pm 0.0029 . \tag{21}$$

Note that this hadronic contribution is a substantial fraction of the prediction from the *complete* standard model of Glashow-Salam and Weinberg^{‡1} through one-loop¹ including leptons, quarks and vector and scalar bosons in internal loops for $M_Z = 94$, $m_t = 30$ and $m_{\text{Higgs}} = 100 \text{ GeV}$

$$A_{LR}^{GSW}(q^2 = -M_Z^2) = 0.2692 \tag{22}$$

but that the *hadronic uncertainty* is quite small compared to the total radiative correction from GSW listed in Table I or the contributions to A_{LR} from beyond the standard model. The hadronic uncertainty in (21) is a factor ~ 2 smaller

‡1 All one loop GSW predictions quoted in this paper specifically exclude *only* the detector dependent QED contributions from graphs in Fig. 4.

than previous estimates.^{3,1} We conclude that the GSW prediction Eq. (22) for A_{LR} is theoretically “clean” since the uncertainties from strong interaction effects of light quarks can be safely controlled. Thus any shifts from this value greater than say, 0.005, must be attributed to new physics from beyond the standard model. Some candidates are listed in Table I.

We have also considered other possible asymmetries in e^+e^- annihilation. At the Z_0 resonance and including one-loop effects, their expressions and the related strong interaction contributions and uncertainties are simply related to those of the longitudinal asymmetry A_{LR} , as has been extensively discussed elsewhere.² Table I contains the relevant uncertainties, which one can evaluate straightforwardly with Eq. (16).

As a next application, assuming M_Z to be very accurately measured at SLC and LEP, we consider the theoretical prediction for the W mass (to be measured at LEP2 and FNAL), related to Eq. (16) through Sirlin’s formula³ :

$$M_W^2 \left[1 - \frac{M_W^2}{M_Z^2} \right] = \frac{(37.281 \text{ GeV})^2}{1 - \Delta r} \quad (23)$$

$$\Delta r = \frac{2s_\theta^2 - 1}{s_\theta^2} \Delta_W(-M_W^2) + \text{const.}$$

which can be gotten, alternatively, by examination of the pole structure of Eq. (2). The contribution from u, d, s, c, b and t quarks is for $M_Z = 94 \text{ GeV}$, $m_t = 30 \text{ GeV}$

$$\begin{aligned} \Delta r^{(u \rightarrow t)} &= \Delta_\alpha^{(u \rightarrow t)}(-M_Z^2) + \text{const}' \\ &= 0.0333 \pm 0.0013 . \end{aligned} \quad (24)$$

This gives a contribution to W^\pm mass

$$\delta M_W^{(u \rightarrow t)} = -484 \pm 19 \text{ MeV} . \quad (25)$$

A glance at Table I shows that the strong interaction uncertainty is smaller than the possible effects coming from physics beyond the GSW model. The ($u \rightarrow t$) contribution is to be compared to the prediction for M_W including all one loop standard model contributions^{11,3,2,1} for $M_Z = 94$, $m_t = 30$ and $m_H = 100 \text{ GeV}$

$$M_W = 83.33 \text{ GeV} \quad (26)$$

To conclude our analysis, we have considered the following two ratios, soon to be measured by the CHARM II collaboration:

$$R_{\nu\bar{\nu}}(-t) = \frac{\sigma_{\nu_\mu e \rightarrow \nu_\mu e}}{\sigma_{\bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e}} \simeq \left(1 - \frac{\hat{v}_\theta}{1 + \hat{v}_\theta^2}\right) / \left(1 + \frac{\hat{v}_\theta}{1 + \hat{v}_\theta^2}\right) \quad (27)$$

$$R_{NC,CC}(-t) = \frac{\sigma_{\nu_\mu e \rightarrow \nu_\mu e}}{\sigma_{\nu_\mu e \rightarrow \nu_e \mu}} \simeq \frac{1 - \hat{v}_\theta + \hat{v}_\theta^2}{12} (1 - \Delta_\rho(-t))^{-2} \left(1 - \frac{m_\mu^2}{2m_e E_\nu}\right)^{-2} \quad (28)$$

with $\hat{v}_\theta = 4(s_\theta^2 + \Delta_\rho(-t)) - 1$. We evaluated the contribution from $u \rightarrow t$ quarks and the resulting theoretical uncertainties. We find for $M_Z = 94$, $m_t = 30 \text{ GeV}$ and $E_\nu = 70 \text{ GeV}$

$$\delta R_{\nu\bar{\nu}}^{(u \rightarrow t)} = -0.163 \pm 0.0089 \quad (29)$$

$$\delta R_{NC,CC}^{(u \rightarrow t)} = -0.0094 \pm 0.0005 \quad (30)$$

which are to be compared to the prediction from the complete GSW theory to one loop^{12,2} for $M_Z = 94 \text{ GeV}$, $m_t = 30 \text{ GeV}$, $m_{\text{Higgs}'} = 100 \text{ GeV}$ and for

$$E_\nu = 70 \text{ GeV}$$

$$R_{\nu D} = 1.2862 \quad (31)$$

$$R_{NC;CC} = 0.1295 . \quad (32)$$

Note that the $u \rightarrow t$ quarks give quite a large fraction of the $R_{\nu D}$ and $R_{NC;CC}$ but that the theoretical uncertainties are small.

4. Conclusions

We have classified the hadronic contributions to *all* one-loop four-lepton processes. We conclude that the hadronic corrections to most leptonic processes contain a rather small uncertainty, which is mainly due to that of the e^+e^- data in the region of time-like $|q'^2|$ from 1 to 100 GeV^2 . This uncertainty could be further substantially reduced if an experimental effort in this region brought the systematic error below that of the available data (which is of approximately 10%) to say 5%. We have shown however that the theoretical uncertainty with present data is sufficiently small to allow a whole series of future experiments at SLAC, CERN and FNAL to carry on a systematic test of the theory of electroweak forces at one-loop with clean theoretical predictions.

ACKNOWLEDGEMENTS

C.V. would like to thank the SLAC Theory Group for their kind hospitality during the autumn of 1985.

REFERENCES

1. B. W. Lynn and R. G. Stuart, Nucl. Phys. B253 (1985) 216.
2. B. W. Lynn and M. E. Peskin, SLAC-PUB-3724, June 1985; B. W. Lynn, M. E. Peskin and R. G. Stuart, SLAC-PUB-3725, June 1985.
3. A. Sirlin, Phys. Rev. D22 (1980) 285; A. Sirlin in Proceedings of the 1983 Trieste Workshop on Radiative Corrections in $SU_2 \times U_1$, B. W. Lynn and J. F. Wheeler, editors, World Scientific Pubs, Singapore (1984).
4. B. W. Lynn and C. Verzegnassi (in preparation).
5. M. A. Shifman, A. I. Vainshtein and V. I. Zakharov, Nucl. Phys. B147 (1979) 386, 448.
6. E. A. Paschos, Nucl. Phys. B159 (1979) 285.
7. L. M. Kurdadse *et al.*, JETP Lett. 37 (1983) 733.
8. G. P. Murtas; *Proceedings of the XIX International Conference on High Energy Physics* (Tokyo, 1978), p. 269; M. Ambrosio *et al.*, Phys. Lett. B91 (1980) 155; R. Baldini-Celio, *et al.*; Lett. Nuovo Cimento 24 (1979) 324; J. E. Augustine, *e^+e^- physics below J/ψ resonance*, invited talk at the *EPS International Conference on High Energy Physics* (Geneva, 1970). See also G. Wolf; *High energy physics trends in e^+e^- physics*, rapporteur talk at *EPS International Conference on High-Energy Physics* (Geneva, 1979); G. Flügge; *Recent Experiments at DESY*, lectures presented at the *XVIII International Symposium für Kernphysik, Schludming, Austria, 1979*, DESY preprint; SLAC-LBL Collaboration, Phys. Rev. Lett. 36 (1976) 300; DASP Collaboration, Nucl. Phys. B148 (1979) 184; DELCO Collaboration, presented by J. Kirkby, *Proceedings of the International Conference on High-*

Energy Physics (Tokyo, 1978), p. 249; PLUTO Collaboration, Phys. Lett. B66 (1977) 395; G. Feldman, *Proceedings of the XIX International Conference on High Energy Physics* (Tokyo, 1978); p. 777; J. L. Se , et al., SLAC-PUB-2831, LBL 13464 (1981) (T/E). B. H. Wiik, *New e^+e^- physics*, invited talk at the *XX International Conference on High-Energy Physics*, Madison, WI (1980, DESY preprint) and references quoted therein; R. Brankelik, et al., DESY 82-010 (1982).

9. C. Verzegnassi, Phys. Lett. B147 (1984) 455.
10. J. Cole, G. Penso, and C. Verzegnassi, ISAS preprint 19/85 EP.
11. W. J. Marciano and A. Sirlin, Nucl. Phys. B189 (1981) 442; W. J. Marciano, Phys. Rev. D20 (1979) 274; A. Sirlin, Phys. Rev. D29 (1984) 89.
12. S. Sarantakos and A. Sirlin, Nucl. Phys. B217 (1983) 84; M. Bohm *et al.*, DESY preprint DESY-84-067 (1984); D. Yu. Bardin, Nucl. Phys. B246 (1984) 221; K. I. Aoki *et al.*, Prog. Theor. Phys. 65 (1981) 1001; W. J. Marciano and A. Sirlin, Phys. Rev. D22 (1980) 2695.

FIGURE CAPTIONS

1. Vector Self Energies.
2. The graph of $\text{Re } 10^2 \cdot \Delta_\alpha(Q_0^2)$ at timelike (right-hand axis) and spacelike (left-hand axis) Q_0^2 values $|Q_0^2| \leq 200 \text{ GeV}^2$. The squares represent the older evaluation by Paschos.⁶
3. Same as in Fig. 2 for spacelike Q_0^2 values $|Q_0^2| \leq 400 \text{ GeV}^2$. A few error bars have been computed for some low Q_0^2 points, according to the prescription given in Section 2.

TABLE CAPTIONS

- Table I. Responses at one loop of various asymmetries on Z^0 resonance and the W^\pm mass to new one-loop physics. Numbers are generic, calculated using $M_Z = 94 \text{ GeV}$.
- Table II. Contributions to the quantity $10^2 \cdot \Delta_\alpha(79 \text{ GeV}^2)$ in Eq. (16) coming from the different regions a-f and related error.
- Table III. Contributions to the quantity $10^2 \cdot \Delta_\alpha(Q_0^2)$ coming from the different regions a-f at several spacelike Q_0^2 values. The related error can be easily worked out and is pictured in Fig. 3.

Table I

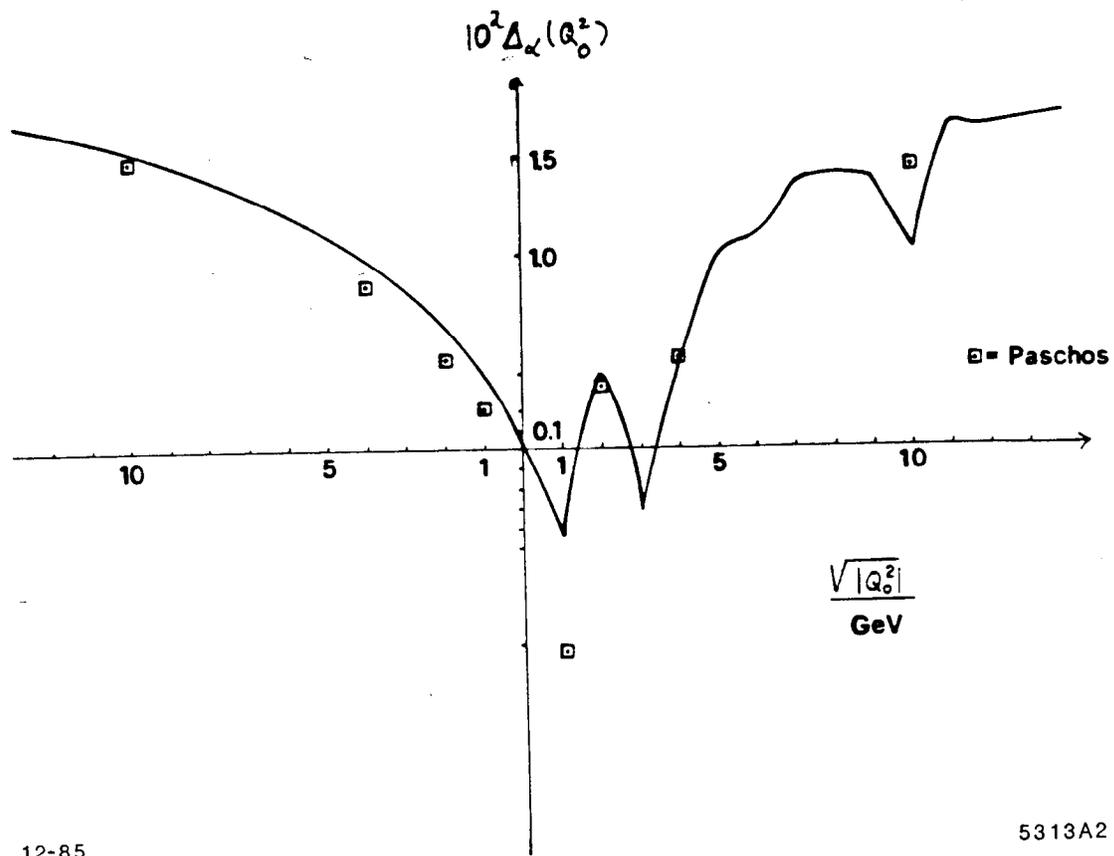
One-Loop Physics	$\delta A_{LR} = \delta A_{\tau pol}$	δA_{FB}	δA_{\perp}	δM_W (MeV)
GSW Weak $m_t = 30$ $m_H = 100$	-0.03	-0.01	.005	-180
Heavy Top Quark $m_t \simeq 180$ GeV	0.03	0.0075	0.004	780
Heavy Higgs ~ 1 TeV	-0.01	-0.0045	-0.003	-160
Heavy Quark Pair a) Large I Splitting b) Degenerate	0.02 -0.004	0.01 -0.002	0.007 -0.001	300 -42
Heavy Lepton Pair a) Large I Splitting $m_{\nu} = 0$ b) Degenerate	0.012 -0.0013	0.006 -0.0006	0.004 -0.0004	300 -14
Heavy Squark Pair a) Large I Splitting b) Degenerate	0.02 0	0.01 0	0.007 0	300 0
Heavy Slepton Pair a) Large I Splitting b) Degenerate	0.012 0	0.006 0	0.004 0	300 0
Winos a) $m_{3/2} \ll 100$ GeV b) $m_{3/2} \gg 100$ GeV	0.005 <0.001	0.0025 <0.001	0.001 $\ll 0.001$	100 <10
Technicolor $SU_8 \times SU_8$ O_{16}	-0.04 -0.07	-0.018 -0.032	-0.012 -0.021	-500 -500
Strong Interaction Uncertainty	± 0.0029	± 0.0014	± 0.001	± 19 MeV

Table II

Region	$10^2 \cdot \Delta_\alpha(79 \text{ GeV}^2)$
a	(0.29 ± 0.01)
b	(0.43 ± 0.05)
c	(0.53 ± 0.05)
d	(0.19 ± 0.02)
e	0.007
f	0.004
Total	(1.45 ± 0.13)

Table III

Q_0^2/GeV^2	a	b	c	d	e	f	Total
1	0.20	0.14	0.03	/	/	/	0.37
4	0.26	0.27	0.11	0.01	/	/	0.65
9	0.27	0.34	0.19	0.03	/	/	0.83
16	0.28	0.37	0.28	0.05	/	/	0.98
25	0.28	0.40	0.35	0.08	/	/	1.11
36	0.29	0.41	0.41	0.10	/	/	1.21
49	0.29	0.42	0.46	0.13	0.005	0.002	1.31
64	0.29	0.43	0.50	0.16	0.006	0.003	1.39
81	0.29	0.43	0.53	0.20	0.007	0.003	1.46
100	0.29	0.43	0.56	0.23	0.009	0.005	1.52



12-85

5313A2

Fig. 2

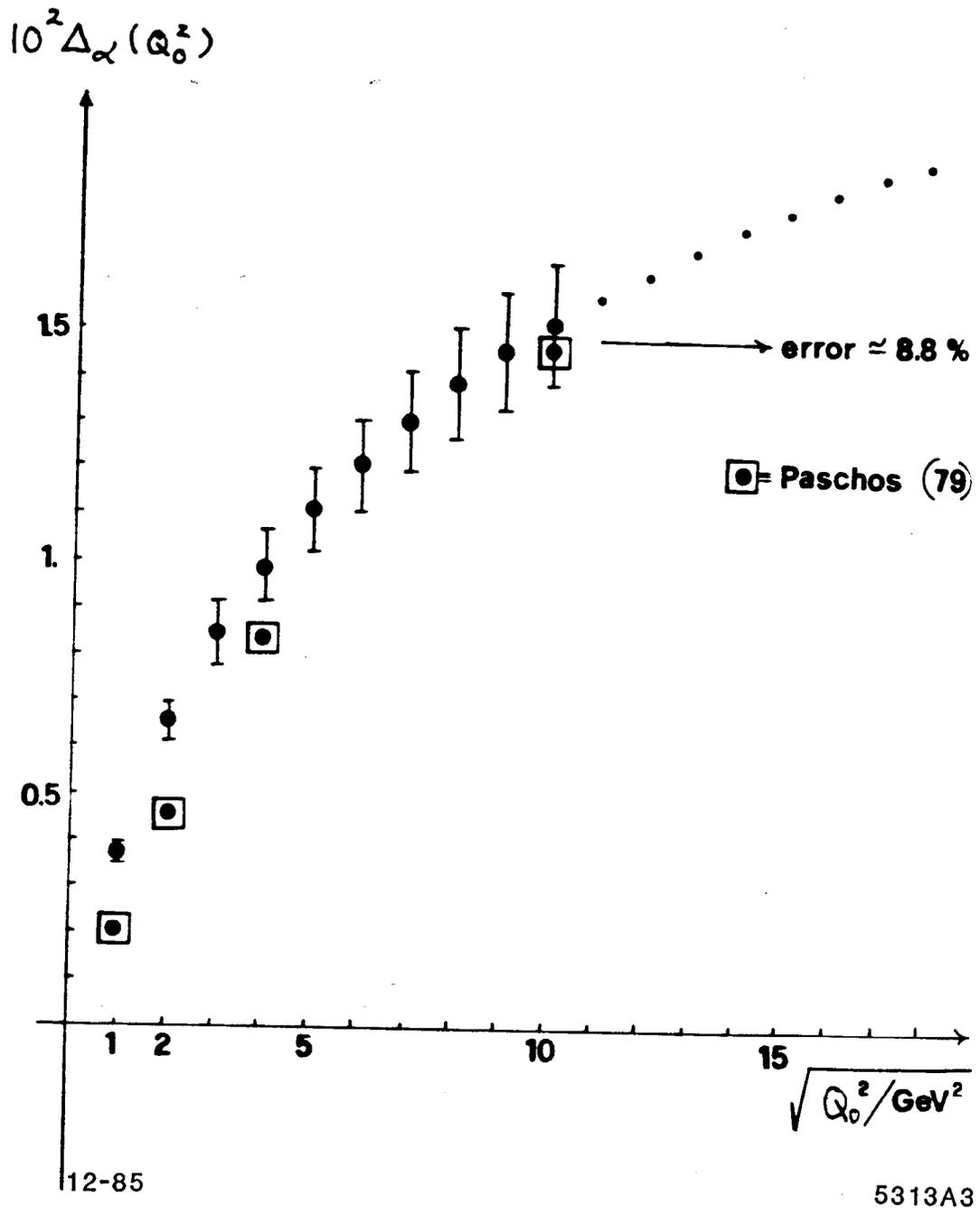
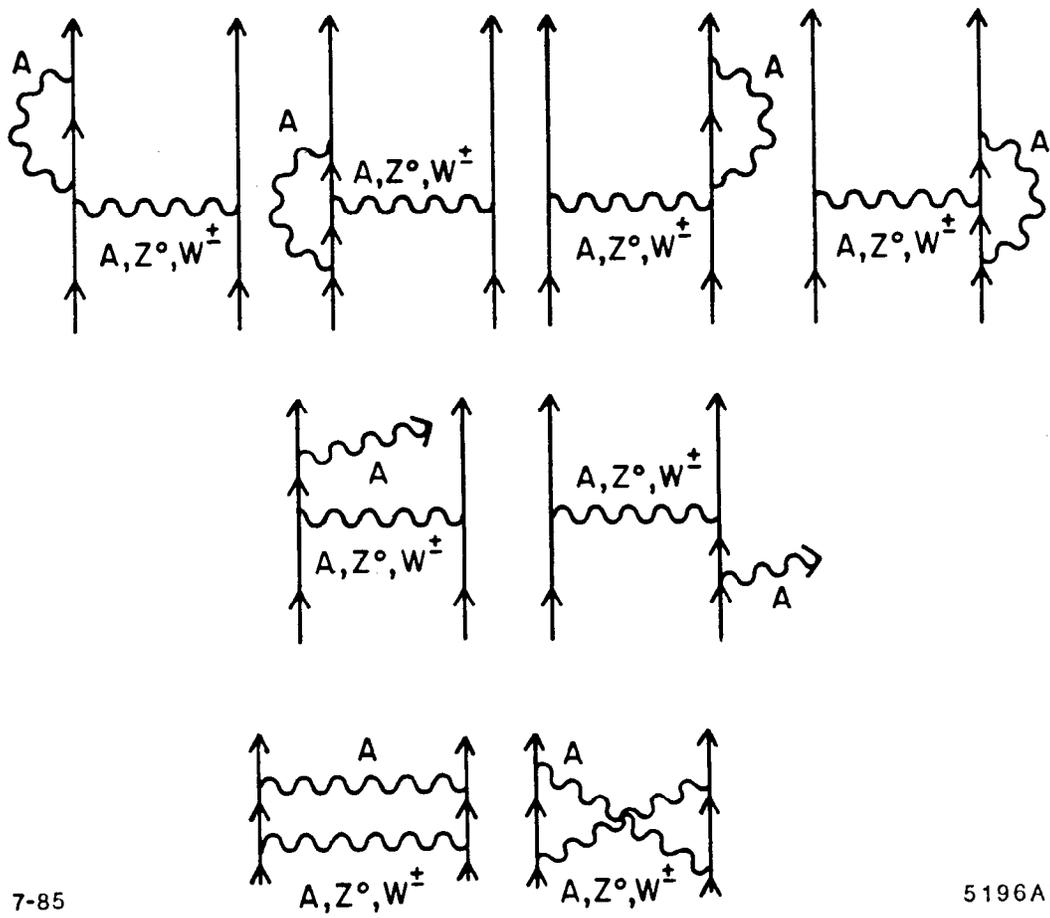


Fig. 3



7-85

5196A2

Fig. 4