EVOLUTION AND COVARIANCE IN CONSTRAINT DYNAMICS*

F. Rohrlich ${ }^{\dagger}$<br>Stanford Linear Accelerator Center<br>Stanford University, Stanford, California 94305

ABSTRACT

Two problems of first class relativistic constraint dynamics of direct interaction are considered. The first is the relation of the gauge motion to the physical motion. In a manifestly covariant formulation they are shown to be exactly equivalent, the latter duplicating the former. The physical motion arises as a consequence of the explicit dependence on an (invariant) time parameter introduced by the gauge fixations (specification of "equal time" surfaces). No simple relation exists in general between the evolution generator and the translation generators. The second problem deals with the covariance of the physical world lines. It is satisfied trivially for Lorentz covariance (since the formalism is covariant) and nontrivially for translations in the equal time surface orthogonal to the total momentum. Necessary conditions on the fixations are given to ensure this.
(Submitted to Physical Review D)

[^0]
## I. BACKGROUND

The attempts to formulate relativistic particle dynamics with interactions that are not necessarily field mediated go back half a century. But the desirability of using more than the minimum number of variables together with suitable constraints is of much more recent vintage. Constraint formulations in this context were suggested independently and to various degrees of generality and explicitness by at least four different authors. ${ }^{1-4}$ The subsequent development resulted in a large number of papers which have recently focussed on the problem of evolution of the dynamical system and on the covariance of its world 1ines. ${ }^{5-9}$

Of special interest is the constraint dynamics with first class constraints because these permit a quantization of this classical system. 10 The elimination of first class constraints requires "fixations" which are time dependent constraints. Such constraints have recently been studied in their own right.ll

The problem of evolution deals with two questions. One is the fixation dependence of the dynamics and the other is the relation of the gauge dynamics (generated by the first class constraints within an equivalence class of points in phase space) and the physical dynamics. The two problems are related.

The problem of covariance is raised by the requirement of relativity which demands that different inertial observers see the same world lines. This condition is not trivial to satisfy and the so-called "world-line condition" has played an important role ever since the no-go theorem of
$1965 .{ }^{12}$
In the present paper we shall study the problems of evolution and covariance from a slightly different point of view than has been done recently. $5,6,8,9$ While our approach is in many respects equivalent to what was done by these authors and our results overlap, the following differences should be noted.

The present work deals with $N$-particle systems ( $N \geq 2$ ) and is throughout manifestly Lorentz covariant. In view of the results by J. N. Goldberg ${ }^{7}$ this covariance ensures Lorentz covariance in both the phase space and configuration space (Minkowski space). As we shall see, even this formulation leads to nontrivial problems of the type mentioned above. Furthermore, use will be made of the Bergmann-Komar reduced variables (star variables). ${ }^{13}$ These are a much more powerful tool than Dirac brackets only. And finally we permit very general fixations which we cast into standard forms that are found to be convenient.

After a brief review of those aspects of first class constraint dynamics that are of interest here (Sec. II) fixations are introduced and standardized (Sec. III). Evolution and covariance are then discussed in Secs. IV and $V$, respectively. The last section summarizes our conclusions.

## II. THE QUOTIENT SPACE

Classical relativistic dynamics in Hamiltonian form is based on an 8N dimensional symplectic space $\Gamma$ for $N$ spinless particles of masses $m_{a}>0(a=1, \ldots, N)$. It is spanned by the canonical variables $q_{a}, p_{a}$
which are fourvectors in $3+1$ dimensional Minkowski space with metric $\eta$ (trace +2 ). They satisfy the covariant Poisson bracket algebra

$$
\begin{equation*}
\left\{q_{a}^{\mu}, p_{b v}\right\}=\delta_{a b} \delta_{v}^{\mu} \tag{1}
\end{equation*}
$$

The Poincaré generators are determined by

$$
\begin{equation*}
P_{\mu}=\sum_{a} P_{a \mu}, \quad M_{\nu}^{\mu}=\sum_{a}\left(q_{a} \wedge p_{a}\right)_{\nu}^{\mu} \tag{2}
\end{equation*}
$$

so that Eq. (1) is a realization of the Poincaré algebra via (2). The N constraints

$$
\begin{equation*}
\mathrm{K}_{\mathrm{a}} \approx 0 \quad(\mathrm{a}=1, \ldots, \mathrm{~N}) \tag{3}
\end{equation*}
$$

(the symbol $\approx$ indicating weak equality, valid in a subspace of $\Gamma$ only) are assumed to be first class,

$$
\begin{equation*}
\left\{K_{a}, K_{b}\right\} \approx 0 \quad \forall a, b \tag{4}
\end{equation*}
$$

since this is the interesting case for the purpose of quantization. The variables $K_{a}$ are the interacting mass shells

$$
\begin{equation*}
K_{a}=p_{a}^{2}+m_{a}^{2}+\phi_{a}(\gamma) \tag{5}
\end{equation*}
$$

with $\gamma \equiv\left(q_{1}, \ldots q_{N}, p_{1}, \ldots p_{N}\right) \in \Gamma$ and $\phi_{a}$ a reasonably well behaved interaction function consistent with Eq. (4).

The constraints (3) restrict $\Gamma(8 \mathrm{~N})$ to a constraint hypersurface of 7 N dimensions, $\mathscr{M}(7 \mathrm{~N})$, which is the system mass shell hypersurface.

Within $\mathscr{M}$ each $K_{a}$ generates a trajectory for every point $\gamma \in \mathscr{M}$, parametrized by $\lambda_{a}$, such that

$$
\begin{equation*}
\frac{d \gamma}{d \lambda}=\left\{\gamma, K_{a}\right\} \tag{6}
\end{equation*}
$$

and these trajectories remain in $\mathscr{M}$ for $a 11 \lambda_{a}$ because the trajectory generators are first class, (4). The $N$ trajectories from any point $\gamma \boldsymbol{\in} \mathscr{M}$ span an $N$ dimensional surface $\Sigma(N) . \mathscr{M}(7 N)$ is thus foliated by these $\Sigma(N)$ and one can define a quotient space $\Phi(6 N)$ by

$$
\begin{equation*}
\Phi(6 \mathrm{~N})=\mathscr{M}(7 \mathrm{~N}) / \Sigma(\mathrm{N}) \tag{7}
\end{equation*}
$$

This quotient space is the physical phase space of 6 N dimensions. The above construction ensures that the reduced 6 N dimensional phase space $\Phi(6 \mathrm{~N})$ carries a nondegenerate symplectic structure, i.e., there exist 6 N canonical variables on it by Darboux's theorem.

The physical motion takes place entirely in $\Phi$ and obeys the constraints (3) which satisfy (4) and which now hold strongly on $\Phi$. A11 points on a given $\Sigma$ then are physically equivalent and correspond to a single point on $\Phi$. The surfaces $\Sigma$ are thus equivalence classes. The "motions" (6) in $\Sigma$ are gauge motions of no physical significance.

How, then, does one determine the physical motion of the system when all the physical information about the interaction between the particles is buried in the constraint functions $K_{a}$, (5)?

## III. THE FIXATIONS

The theory developed so far is still lacking an essential ingredient. It is necessary to specify the observer, i.e., a parametrization by which one can describe the evolution of the system. In Minkowski space Hamiltonian dynamics requires the specification of threedimensional (usually spacelike) hypersurfaces parametrized by some parameter $\tau$, say, so that the infinitesimal evolution takes the system from the surface $\tau$ to the surface $\tau+\delta \tau$. A corresponding specification of hypersurfaces in phase space has not yet been made. In Komar's termino$\log y^{14}$ the syntactic aspect of the theory has been stated but the semantic one is still missing.

If the intended meaning of the $\mathrm{q}_{\mathrm{a}}$ is a position fourvcctor for particle $a$, then $q_{a}^{0}$ is its coordinate time. The hypersurfaces to be specified are therefore expected on physical grounds to relate the $q_{a}^{0}$ to $\tau$.

In a non-covariant way the simplest such specification would thus be

$$
\begin{equation*}
x_{a} \equiv q_{a}^{0}-t \approx 0 \tag{8}
\end{equation*}
$$

In this case $\tau=t$ is not a Lorentz invariant parameter. Since the physical (intended) meaning of the generators $P_{\mu}$, (2), of space-time translations is that of total momentum and energy, a Lorentz covariant statement instead of (8) would be

$$
\begin{equation*}
x_{a} \equiv-q_{a} \cdot \hat{p}-\tau \approx 0 \tag{9}
\end{equation*}
$$

where the unit vector $\hat{P}$ is defined as $P / \sqrt{-P^{2}}$. A still more sophisticated choice would be the one advocated in Ref. 15 which has the form (at least for $N=2$ ),

$$
\begin{equation*}
\chi_{a} \equiv-q_{a} \cdot \hat{P}+p_{a} \cdot \hat{P}_{\tau} \approx 0 \tag{10}
\end{equation*}
$$

A11 these are examples of "fixations", constraints that specify the generalized "equal time surface". In general, they are of the form

$$
\begin{equation*}
x_{a}(\gamma, \tau) \approx 0 \quad(a=1, \ldots, N) \tag{11}
\end{equation*}
$$

but since they are to be monotonic functions of $\tau$ (on physical grounds) it is reasonable to assume that (11) can be solved with respect to $\tau$ and we shall restrict ourselves to that case. The fixations can then be brought to the "first standard form",

$$
\begin{equation*}
x_{a} \equiv \sigma_{a}(\gamma)-\tau \approx 0 \quad(a=1, \ldots, N) \tag{12}
\end{equation*}
$$

and the examples (8) and (9) are of this form. Equation (10) can obviously also be put into the form (12). A "second standard form" is obtained by retaining $X_{N}$ but replacing the other $x_{a}$ by $x_{a}-x_{N}$,

$$
\begin{align*}
& \bar{x}_{a} \equiv \chi_{a}-\chi_{N}=\sigma_{a}(\gamma)-\sigma_{N}(\gamma) \approx 0 \quad(a=1, \ldots, N-1)  \tag{13}\\
& \bar{x}_{N}=x_{N}=\sigma_{N}(\gamma)-\tau \approx 0 \quad .
\end{align*}
$$

This second standard form has the advantage that $\tau$ occurs in only one of the N fixations.

Geometrically, the $N$ fixations specify an $N$ dimensional surface.

When used simultaneously with the constraints (3) they form a set of 2 N second class constraints. We require that the matrix

$$
\begin{equation*}
\Delta=\left(\left\{x_{\alpha}, k_{b}\right\}\right) \quad(\alpha, b=1, \ldots, N) \tag{14}
\end{equation*}
$$

is non-singular,

$$
\begin{equation*}
\operatorname{det} \Delta \neq 0 \tag{15}
\end{equation*}
$$

The phase space $\Gamma(8 \mathrm{~N})$ is thus reduced by the complete set of 2 N second class constraints to a nondegenerate phase space $\Phi^{*}(6 \mathrm{~N})$ of 6 N dimensions. This space is diffeomorphic to the quotient space $\Phi$ for every fixed $\tau$, but gives in general a different diffeomorphism for each value of $\tau$.

In terms of the geometrical picture of a foliated space $\mathscr{M}(7 \mathrm{~N})$ which we had before the introduction of fixations, the $N$-dimensional fixation surface (for a fixed $\tau$ ) can be considered as restricting each surface $\Sigma(N)$ to a single point, which can be characterized by $q_{a}^{*}(\tau)$ and $p_{a}^{*}(\tau)$, $\gamma^{*}(\tau) \equiv\left(q_{1}^{*}, \ldots q_{N}^{*}, p_{1}^{*}, \ldots p_{N}^{*}\right)$. The $\tau$ is necessary because different $\tau$ give different fixation surfaces.

The point $\gamma^{*} \epsilon \Phi^{*}$ can be constructed by a method due to Bergmann and Komar ${ }^{13}$ in which the $\tau$ dependence becomes explicit in $\gamma^{*}$. This construction is simply the reduction of the 8 N component $\gamma$ to that particular 6 N component $\gamma^{*}$ which is compatible with the 2 N constraints. For fixed $\tau$

$$
\begin{equation*}
\gamma^{*}=\gamma-\sum_{m=1}^{2 N} \sum_{n=1}^{2 N}\left\{\gamma, C_{m}\right\}\left(D^{-1}\right)_{m n} C_{n} \tag{16}
\end{equation*}
$$

where

$$
\begin{align*}
& C_{m}=x_{m} \quad(m=1, \ldots, N) \\
& C_{m}=K_{m} \quad(m=N+1, \ldots, 2 N) \tag{17}
\end{align*}
$$

and the $2 \mathrm{~N} \times 2 \mathrm{~N}$ dimensional matrix D consists of f (our $\mathrm{N} \times \mathrm{N}$ submatrices

$$
\begin{align*}
& D=\left(\begin{array}{rr}
\theta & \Delta \\
-\widetilde{\Delta} & 0
\end{array}\right)=\left(\left\{C_{m}, C_{n}\right\}\right)  \tag{18}\\
& \theta=\left(\left\{x_{\alpha}, x_{\beta}\right\}\right) \quad(\alpha, \beta=1, \ldots, N) \tag{19}
\end{align*}
$$

and $\Delta$ was defined earlier in (14). In (18) 0 indicates the $N \times N$ zero matrix and is there because of (4). The inverse matrix $D^{-1}$ which occurs in (16) follows from (18) to be [it exists in view of (15)]

$$
D^{-1}=\left(\begin{array}{lc}
0 & -\tilde{\Delta}^{-1}  \tag{20}\\
\Delta^{-1} & \Delta^{-1} \theta \tilde{\Delta}^{-1}
\end{array}\right)
$$

so that (16) can be written more explicitly as

$$
\begin{align*}
\gamma^{*}=\gamma & -\left\{\gamma, x_{\alpha}\right\}\left(-\tilde{\Delta}^{-1}\right)_{\alpha a} K_{a}  \tag{21}\\
& -\left\{\gamma, K_{a}\right\}\left[\left(\Delta^{-1}\right)_{a \alpha} \chi_{\alpha}+\left(\Delta^{-1} \theta \tilde{\Delta}^{-1}\right)_{a b} K_{b}\right]
\end{align*}
$$

Here and in the following we use the summation convention for simplicity, summing all repeated indices from 1 to $N$.

The reduced variables $\gamma^{*}$ are now functions of $\tau$ via $\chi_{\alpha}$ in the third
term on the right. The $\tau$ dependence disappears in $D$ if we assume one of the standard fixation forms (12) or (13). But for any choice of fixations the $\gamma^{*}$ are thus uniquely determined functions of $\tau$. The reduction of $\gamma$ to $\gamma^{*}$ via (16) ensures that

$$
\begin{equation*}
\left\{r^{*}, x_{\alpha}\right\}=0, \quad\left\{r^{*}, k_{a}\right\}=0 \quad \forall \alpha, a \tag{22}
\end{equation*}
$$

and thus identifies the $\gamma^{*}$ as physical variables.
IV. THE EVOLUTION

Since the fixations are to be preserved throughout the motion, the evolution operator should not change them. Now in first class constraint theory the evolution operator is not known: any combination of the first class constraints

$$
\begin{equation*}
H=\omega_{a} K \tag{23}
\end{equation*}
$$

is acceptable. The $\omega_{a}$ can be almost arbitrary functions of the $q$ and $p$ but will be assumed positive. ${ }^{16}$ The requirement of conservation of fixations,

$$
\frac{d x_{\alpha}}{d \tau}=\frac{\partial x_{\alpha}}{\partial \tau}+\left\{x_{\alpha}, H\right\}=0
$$

now fixes the $\omega_{a}$. From (14) and

$$
\begin{equation*}
\left\{x_{\alpha}, k_{a}\right\} \omega_{a} \approx-\frac{\partial x_{\alpha}}{\partial \tau} \tag{24}
\end{equation*}
$$

follows $\omega_{a}$ uniquely on $\Phi^{*}$,

$$
\begin{equation*}
\omega_{a}=-\left(\Delta^{-1}\right)_{a \alpha} \frac{\partial x_{\alpha}}{\partial \tau} . \tag{25}
\end{equation*}
$$

If we adopt the first standard form, (12), this expression simplifies to

$$
\begin{equation*}
\omega_{\mathrm{a}}=\sum_{\alpha}\left(\Delta^{-1}\right)_{\mathrm{a} \alpha} \tag{26a}
\end{equation*}
$$

and if we adopt the second standard form, (13), of the fixations

$$
\begin{equation*}
\omega_{a}=\left(\Delta^{-1}\right)_{a N} \tag{26b}
\end{equation*}
$$

Thus, $H$ is uniquely determined by the fixations,

$$
\begin{equation*}
\mathrm{H}=\mathrm{K}_{\mathrm{a}}\left(\Delta^{-1}\right)_{\mathrm{aN}} . \tag{27}
\end{equation*}
$$

But this $H$ does not seem to provide for the evolution of the physical variables $\gamma^{*}$. These commute (weakly) with $H$ in view of (22). One does observe though that the original variables $\gamma$ now describe a unique trajectory on $\Sigma$ :

$$
\begin{equation*}
\frac{d \gamma}{d \lambda}=\{\gamma, H\} \approx\left\{\gamma, K_{a}\right\}\left(\Delta^{-1}\right)_{\mathrm{aN}} . \tag{28}
\end{equation*}
$$

The choice of fixations selects exactly one particular trajectory on the N -dimensional surface $\Sigma$. This is, however, a gauge dynamics since it relates equivalent points (points on $\Sigma$ ) to one another, all of which referring to the same point on the quotient space. This is not the evolution of the physical variables. Where then is the evolution of the
physical variables?
The answer to this puzzle lies in the $\tau$ dependence of the $\gamma^{*}$ introduced by the reduction process (21) through the fixations:

$$
\begin{equation*}
\frac{d \gamma^{*}}{d \tau}=\frac{\partial \gamma^{*}}{\partial \tau}+\left\{\gamma^{*}, \mathrm{H}\right\}=\frac{\partial \gamma^{*}}{\partial \tau} \tag{29}
\end{equation*}
$$

Let us for convenience again adopt the second standard form of the fixations so that

$$
\begin{equation*}
\frac{\partial x_{\alpha}}{\partial \tau}=-\delta_{\alpha N} \tag{30}
\end{equation*}
$$

Then (21) inserted into (29) gives, since the only explicit $\tau$ dependence in (21) occurs in the $x_{\alpha}$ that are not in P.B.'s,

$$
\begin{equation*}
\frac{d \gamma^{*}}{d \tau}=-\left\{\gamma, K_{a}\right\}\left(\Delta^{-1}\right)_{a \alpha}\left(-\delta_{\alpha N}\right)=\left\{\gamma, K_{a}\right\} \omega_{a} \tag{31}
\end{equation*}
$$

Thus one obtains from (28) the fundamental result

$$
\begin{equation*}
\frac{d \gamma^{*}}{d \tau}=\left\{\gamma, K_{a}\right\} \omega_{a} \approx\{\gamma, H\}=\frac{d \gamma}{d \lambda} \tag{32}
\end{equation*}
$$

One is led to the important conclusion that the evolution of the physical variables is exactly realized by the gauge dynamics. We can thus identify $\tau$ and $\lambda$, and we have the following picture.

The $\tau$-dependent realizations $\Phi^{*}$ of the quotient space $\Phi$ associate with each point of the trajectory of $\gamma$ on $\Sigma$ (generated by $H$ ) a different set of physical variables $\gamma^{*}$, each set being labelled by a different value of $\tau$. There is (at least locally) a one-to-one map of the
trajectory $\gamma(\lambda)$ to the trajectory $\gamma^{*}(\tau)$, the latter being generated by the reduction process for 2 N second class constraints. This bijection is the fundamental mechanism that permits one to treat the gauge motion as if it were the physical motion. It thus resolves the puzzle created by papers in the literature that treat the gauge motion as physical motion without clear justification.

## V. COVARIANCE

The physical trajectory is a one-dimensional object in a $6 \mathrm{~N}+1$ dimensional space, the direct product space of $\Phi^{*}=\Phi^{*}\left(\gamma^{*}\right)$ and $\tau$. A projection into a $3 \mathrm{~N}+1$ dimensional Minkowski space yields one spacetime trajectory (world line) for the $N$ particles. When $M_{3 N+1}$ is mapped into $M_{3+1}$ N trajectories result and the conventional space-time description emerges: one obtains $N$ world lines in $M_{3+1}$.

It is now necessary to show that the first class constraint dynamics just described is consistent with Poincaré transformations according to the principle of relativity.

The generator of infinitesimal Poincaré transformations ( $\Lambda$, a) is

$$
\begin{gather*}
G=L+T \\
L=\frac{1}{2} \omega_{\mu \nu} M^{\mu \nu} \quad T=a^{\mu} P_{\mu} . \tag{33}
\end{gather*}
$$

Now all our constraints are manifestly Lorentz invariant, if we exclude the type exemplified by (8) which has been discussed before, 8

$$
\begin{equation*}
\left\{x_{\alpha}, L\right\}=0 \quad\left\{\mathrm{~K}_{\mathrm{a}}, \mathrm{~L}\right\}=0 . \tag{34}
\end{equation*}
$$

In particular, $\tau$ is a Lorentz invariant parameter. Thus, the $\gamma^{*}$ are Lorentz covariant. But translation invariance is not so easily dealt with.

The mass shell constraints are chosen to be translation invariant,

$$
\begin{equation*}
\left\{\mathrm{K}_{\mathrm{a}}, \mathrm{~T}\right\}=0 \tag{35}
\end{equation*}
$$

because one is interested only in translation invariant interactions $\phi_{a}$; they can depend only on the position differences $q_{b}-q_{c}$. But the fixations are by nature not translation invariant because at least the variables $q_{a}^{0}$ are involved individually rather than as differences. Otherwise "equal $\tau$ surfaces" could not be specified. Another way of saying this is the following: for spatial distances the interparticle separations are sufficient and no distinguished point needs to be specified; but for the evolution of the system all points in the threespace must agree on the same time $\tau=0$, say. The examples (8) - (10) show this lack of T -invariance explicitly,

$$
\begin{equation*}
\left\{x_{\alpha}, \mathrm{T}\right\} \neq 0 \tag{36}
\end{equation*}
$$

The second standard form (13) does, however, have the advantage that at least in some cases [example (9) but not (10)] $N-1$ of the $N X_{\alpha}$ are T-invariant and only $X_{N}$ is not.

What effect does this have on $\gamma^{*}$ ? One finds easily from (21)

$$
\begin{equation*}
\left\{\gamma^{*}, P_{\mu}\right\} \approx\left\{\gamma, P_{\mu}\right\}-\left\{\gamma, K_{a}\right\}\left(\Delta^{-1}\right)_{a \alpha}\left\{x_{\alpha}, P_{\mu}\right\} \tag{37}
\end{equation*}
$$

because all other terms have vanishing P.B. with $P_{\mu}$.
Now from the general theory of reduced variables one knows that

$$
\left\{A^{*}, B\right\} \approx\left\{A, B^{*}\right\} \approx\left\{A^{*}, B^{*}\right\} \approx\{A, B\}^{*}
$$

where the last bracket is the Dirac bracket. Therefore

$$
\begin{align*}
\left\{\gamma, P_{\mu}^{*}\right\} & \approx\left\{\gamma, P_{\mu}\right\}^{*} \\
& =\left\{\gamma, P_{\mu}\right\}-\left\{\gamma, K_{a}\right\}\left(\Delta^{-1}\right)_{a \alpha}\left\{x_{\alpha}, P_{\mu}\right\}  \tag{38}\\
& =\left\{\gamma, P_{\mu}-K_{a}\left(\Delta^{-1}\right)_{a \alpha}\left\{x_{\alpha}, P_{\mu}\right\}\right\}
\end{align*}
$$

By the same reduction process as used in (16) and (21)

$$
\begin{equation*}
P_{\mu}^{*}=P_{\mu}-K_{a}\left(\Delta^{-1}\right)_{a \alpha}\left\{X_{\alpha}, P_{\mu}\right\} \tag{39}
\end{equation*}
$$

and we see that (38) is in fact a strong equality. Thus, the lack of T-invariance of the fixations is exactly taken into account when one makes translations with the reduced generators $P^{*}$.

The reduced translation generators $P^{*}$ can be simply related to $H$ if the $X_{\alpha}$ satisfy the condition

$$
\begin{equation*}
\left\{X_{\alpha}, P_{\mu}\right\}=C_{\mu} \frac{\partial x_{\alpha}}{\partial \tau} \tag{40}
\end{equation*}
$$

In that case, using (25) in (39)

$$
\begin{equation*}
P_{\mu}^{*}=P_{\mu}-C_{\mu} K_{a}\left(\Delta^{-1}\right)_{a \alpha} \frac{\partial \chi_{\alpha}}{\partial \tau}=P_{\mu}+C_{\mu}^{H} \tag{41}
\end{equation*}
$$

This is a generalization of a result first obtained by Bergmann and Komar. ${ }^{8}$ Since the assumption (40) is indeed satisfied in some cases [e.g., by the popular fixations (9)] it is desirable to consider its consequences. One now finds a relationship between the translations (generated by the $P_{\mu}$ ) and $H$. From the fundamental relations (32) and (41) follows on $\Phi^{*}$

$$
\begin{equation*}
\frac{d \gamma^{*}}{d \tau}=\frac{d \gamma}{d \lambda}=\{\gamma, H\}=\frac{C^{\mu}}{C^{2}}\left\{\gamma, P_{\mu}^{*}-P_{\mu}\right\} \tag{42}
\end{equation*}
$$

In conventional Hamiltonian dynamics one identifies the Hamiltonian with $P^{0}$, the total energy of the system. But in those cases $P^{0}$ is a given function of the $\vec{q}$ and $\vec{p}$, i.e., $P^{0}=P^{0}(\vec{\gamma})$. This is not the case here where $P^{\circ}=\sum_{a} p_{a}^{o}$ and the $p_{a}^{o}$ are independent of the other variables; here

$$
\begin{equation*}
\left\{A, P^{0}\right\}=\sum_{a} \frac{\partial A}{\partial q_{a}^{o}} \tag{43}
\end{equation*}
$$

which by itself implies no dynamics. The relation (42) is therefore of interest. If one uses the fixations (9)

$$
\begin{equation*}
C^{\mu}=\hat{P}^{\mu}, \quad P_{\mu}^{*}=P_{\mu}\left(1+\frac{H}{\sqrt{-P^{2}}}\right) \tag{44}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{d \gamma^{*}}{d \tau}=\hat{P}^{\mu}\left\{\gamma, P_{\mu}-P_{\mu}^{*}\right\} \tag{45}
\end{equation*}
$$

In general. however, (40) is not satisfied as can be seen by the example (10). Then there is no simple relation between the dynamics of the physical variables $\mathrm{d} \gamma^{*} / \mathrm{d} \tau$ and their space-time translations.

Now it is clear from (39) that the reduced generators $P_{\mu}^{*}$ and $M_{\mu \nu}^{*}=M_{\mu \nu}$ satisfy the Poincare algebra on $\Phi^{*}$. Thus, the physical variables must have the appropriate transformation properties with respect to them. For $M_{\mu \nu}^{*}$ this is trivial as indicated above. For $P_{\mu}^{*}$ we have from (37)

$$
\begin{equation*}
\left\{q_{a}^{\lambda^{*}}, P_{\mu}^{*}\right\} \approx\left\{q_{a}^{\lambda}, P_{\mu}\right\}-\left\{q_{a}^{\lambda}, k_{b}\right\}\left(\Delta^{-1}\right)_{b \alpha}\left\{x_{\alpha}, p_{\mu}\right\} . \tag{46}
\end{equation*}
$$

At this point one must recall that each $q_{a}^{*}$ and each $p_{a}^{*}$ although a fourvector, involves only three independent variables, the space they span involving only 6 N dimensions. This results from the fact that the constraint equations $K_{a} \approx 0$ and $X_{a} \approx 0$ are strong equations in $\Phi^{*}$. One can thus use them to eliminate the $q_{a}^{\|} \equiv-q_{a} \cdot \hat{P}$ and the $p_{a}^{\|} \equiv-p_{a} \cdot \hat{P}$ in favor of the $q_{a}^{\perp}$ and the $p_{a}^{\perp}$. Thus the $q_{a}^{*}$ and $p_{a}^{*}$ can be expressed entirely in terms of the $q_{a}^{\perp}$ and the $p_{a}^{\perp}$ and $\hat{P}$.

One can therefore require correct transformation properties under translations only in the hyperplane orthogonal to $P^{\mu}$. Specifically, the $\mathrm{q}_{\mathrm{a}}^{*}$ must transform covariantly, the $\mathrm{p}_{\mathrm{a}}^{*}$ invariantly under such translations,

$$
\begin{align*}
& \left\{q_{a}^{* \lambda}, P_{\mu}\right\} p_{\perp}^{\mu \nu}=P_{\perp}^{\lambda \nu}  \tag{47}\\
& \left\{p_{a}^{* \lambda}, P_{\mu}\right\} P_{\perp}^{\mu \nu}=0 \tag{48}
\end{align*}
$$

where the projection orthogonal to $P^{\mu}$ is

$$
\begin{equation*}
P_{\nu}^{\mu \nu}=\eta^{\mu \nu}+\hat{P}^{\mu} \hat{P}^{\nu} \tag{49}
\end{equation*}
$$

The relations (47) and (48) state that in the center of mass frame ( $\vec{P}=0$ ) the threevectors $\vec{q}_{a}^{*}$ and $\overrightarrow{\mathrm{p}}_{\mathrm{a}}^{*}$ transform under translations by $\vec{c}$, respectively, as

$$
\begin{equation*}
\overrightarrow{\mathrm{q}}_{\mathrm{a}}^{*} \rightarrow \overrightarrow{\mathrm{q}}_{\mathrm{a}}^{*}+\overrightarrow{\mathrm{c}} \quad \text { and } \quad \overrightarrow{\mathrm{p}}_{\mathrm{a}}^{*} \rightarrow \overrightarrow{\mathrm{p}}_{\mathrm{a}}^{*} \tag{50}
\end{equation*}
$$

No statement is made about the $\mathrm{q}_{\mathrm{a}}^{\mathrm{o}^{*}}$ and $\mathrm{p}_{\mathrm{a}}^{\mathrm{o}^{*}}$ since their transformation properties are fixed in terms of those of the $\overrightarrow{\mathrm{q}}_{\mathrm{a}}^{*}$ and $\overrightarrow{\mathrm{p}}_{\mathrm{a}}^{*}$ by the constraints.

In order to satisfy the requirements (47) and (48) it is both necessary and sufficient that the fixations are restricted by

$$
\begin{equation*}
\left\{x_{\alpha}, P_{\mu}\right\}=\theta_{\alpha} \hat{\mathrm{P}}_{\mu} \tag{51}
\end{equation*}
$$

a condition which is indeed satisfied by (9) as well as (10). This requirement ensures that $P_{\mu}^{*}$ points into the same direction as $P_{\mu}$, as can be seen from (39).

This completes the proof that the physical variables transform correctly under Poincaré transformations which exclude translations in the direction of $\mathrm{P}^{\mu}$.

The world line conditions specified in some papers are therefore identically satisfied and need not be imposed here as an additional restriction.

## VI. CONCLUSIONS

The manifestly covariant formulation of relativistic constraint dynamics presented in this paper for first class constraints exhibits the following results.

The gauge dynamics within a leaf $\Sigma(N)$ of the foliation of the mass she11 constraint hypersurface $\mathscr{M}(7 \mathrm{~N})$ is shown to be weakly equal to the physical dynamics of the reduced variables as an explicit function of the invariant time parameter $\tau$ introduced by the fixations. This is expressed by Eq. (32). No translations are involved in establishing this relationship.

The fixations are not arbitrary. They must not only be Lorentz covariant but must also satisfy (52) the condition that the $P_{\mu}^{*}$ and $P_{\mu}$ are parallel. In addition, they must yield $\omega_{a}$ that are translation invariant and that are consistent with a cluster decomposition. ${ }^{15}$ The fixations (10) satisfy both of these requirements at least for $N=2$. The fixations (9) satisfy neither. Indeed, in general they do not even give translation invariant $\omega_{a}$ (necessary for translation invariance of $d \gamma * / d \tau): \quad\left\{\chi_{a}, K_{b}\right\}$ involves $\left\{\chi_{a}, \phi_{b}\right\}$ which contains a term

$$
\frac{\partial\left(q_{a} \cdot \hat{P}\right)}{\partial p_{c}} \cdot \frac{\partial \phi_{b}}{\partial q_{c}}=\frac{q_{a}^{\perp}}{\sqrt{-p^{2}}} \cdot \frac{\partial \phi_{b}}{\partial q_{c}}
$$

which is not translation invariant unless the $\phi_{b}$ depend on momenta only. Thus, $\Delta$ is not so invariant and consequently neither are the $\omega_{a}$. The
fixations (9) are good only for position independent interactions.
There is in general no simple relation between the evolution generator $H$ and the translation generators $P_{\mu}$. A relation like (41) is spurious because it is based on the assumption (40) that is not satisfied in general. (The commonly used fixations (9) which do satisfy it are not admissible except for position independent interactions.)

Finally, full Poincaré covariance for $L+T\left(\mathbb{R}^{3}\right)$ is demonstrated for the $q_{a}^{*}$ and L-covariance and $T\left(\mathbb{R}^{3}\right)$ invariance for the $p_{a}^{*}$. Here $\mathbb{R}^{3}$ is the three-dimensional hypersurface in $M_{3+1}$ orthogonal to $P_{\mu}$. Covarlance of the world lines is thus ensured.

I want to thank Sidney Drell and the Theory Group at SLAC for their hospitality during the Spring Quarter 1981.

## REFERENCES

1. R. Arens, Nuovo Cimento 21B, 395 (1974).
2. P. Droz-Vincent, Rep. Math. Phys. 8, 79 (1975).
3. I. T. Todorov, Comm. JINR E2-10125, Dubna (1976).
4. A. Komar, Phys. Rev. D18, 1881, 1887 (1978).
5. A. Kihlberg, R. Marnelius, and N. Mukunda, "Relativistic potential models as systems with constraints and their interpretation", University of Göteborg preprint, April 1980.
6. N. Mukunda and E. C. G. Sudarshan, "Forms of relativistic dynamics with world-lines", University of Texas (Austin) preprint, Aug. 1980.
7. J. N. Goldberg, "Relativistically interacting particles and world1ines", Syracuse University preprint, October 1980.
8. L. Lusanna, "Gauge-fixings, evolution generators and world-line conditions in relativistic classical mechanics with constraints", University of Geneva preprint, April 1981.
9. L. P. Horwitz and F. Rohrlich, "Constraint relativistic quantum dynamics", Syracuse University preprint, March 1981.
10. N. Mukunda, Phys. Scripta 21, 801 (1980).
11. D. G. Currie, T. F. Jordan, and E. C. G. Sudarshan, Rev. Mod. Phys. 35, 350 (1965); H. Leutwyler, Nuovo Cimento 37, 556 (1965).
12. P. G. Bergmann and A. Komar, Phys. Rev. Lett. 4, 432 (1960).
13. A. Komar, Phys. Rev. D18, 3617 (1978).
14. F. Rohrlich, Phys. Rev. D23, 1305 (1981).
15. For a given choice of the $\omega_{a}$ and a trajectory parameter $\lambda$ one has

$$
\frac{d \gamma}{d \lambda} \approx \omega_{a}\left\{\gamma, k_{a}\right\}
$$

so that comparison with (6) leads to the identification $\omega_{a} \approx d \lambda_{a} / d \lambda$. It is for this reason that one assumes $\omega_{a}>0$.


[^0]:    * Work supported by the Department of Energy, contract DE-AC03-76SF00515.
    + On leave of absence from Syracuse University, Syracuse, NY 13210.

