# THE DESCRIPTION OF UNSTABLE PARTICLES <br> IN QUANTUM FIELD THEORY* 

H. A. Weldon<br>Stanford Linear Accelerator Center Stanford University, Stanford, California 94305


#### Abstract

The consequences of a pole on the second sheet of the $S$ matrix are investigated. It is shown that corresponding to each such pole is an eigenstate of the Hamiltonian with a complex energy. These eigenstates lie in a natural cxtension of the physical Hilbert space. Because it is the vector space that is modified and not the Hamiltonian, unstable particle states transform covariantly. They have complex energy and momentum but real integer or half-integer spin. Scattering amplitudes involving unstable particles are expressed as residues of poles in a reduction formula of the LSZ type. An example from potential theory is worked out in detail.


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## I. INTRODUCTION

Of the known elementary particles only five are stable; all others eventually decay into combinations of e, P, $\gamma, \nu_{\mathrm{e}}, \nu_{\mu}$ and their antiparticles. For the stable particles field theory provides a beautiful description based on three principles:
a) Stable particles correspond to eigenstates of energy-momentum in a linear vector space.
b) They transform as irreducible representations of the Poincaré group labeled by mass and spin. ${ }^{1}$
c) Scattering amplitudes are matrix elements of Heisenberg operators taken between these states.

Unfortunately most real particles are not stable. The purpose of this paper is to show that even for unstable particles these same principles apply. The essential difference is that stable particles lie in a Hilbert space and unstable ones do not.

Conventional treatments of unstable particles are all based on perturbation theory in that the unstable states are always eigenstates of an effective Hamiltonian which is chosen to contain the essentials of the spectrum. The difference between the total Hamiltonian and the effective Hamiltonian is a small perturbation that induces decays of the eigenstates of the effective Hamiltonian. This approach was the very basis of theoretical atomic physics. There the effective Hamiltonian includes the Coulomb field of the nucleus but neglects the electronphoton interaction. This perturbation results in the radiative decays summarized by the Balmer formula. Dirac ${ }^{2}$ showed how to calculate the effect of such unstable states on the scattering amplitude

$$
\begin{equation*}
T(E)=H_{1}+H_{1} \frac{1}{E-H_{0}} H_{1}+\ldots \tag{1.1}
\end{equation*}
$$

when $E$ is near an eigenvalue of $\mathrm{H}_{0}$. Later, in nuclear physics, the abundance of resonances required more sophistication. A number of approaches were developed based on the energy dependence of the scattering wave function at large distances. Kapur and Peierls ${ }^{3}$ expand the wave function in terms of states with complex eigenvalues corresponding to resonance energies and widths. Wigner and Fisenbud ${ }^{4}$ expand in different states that led to the reactance matrix and the many level formula for overlapping resonances. (See Ref. 5 for complete reviews.) Feshbach ${ }^{6}$ and also Fonda and Newton ${ }^{7}$ generalize the earlier approaches by abandoning coordinate space in favor of an abstract Hamiltonian and projection operators that distinguish open and closed channels. The resonant part of the perturbation series (1.1) is expressed in terms of a general effective Hamiltonian that is both energy dependent and nonhermitian. Feshbach shows that Kapur-Peierls and Wigner-Fisenbud result from different choices of projection operators. Common to all these approaches is the identification of an unstable state with a vector in the usual Hilbert space of stable states that is an eigenstate of a modified Hamiltonian.

The S-matrix theory of unstable particles naturally makes no statement about Hamiltonians or state vectors. It began with the suggestion by Møeller ${ }^{8}$ that an unstable particle corresponds to a pole on the second Riemann sheet of the analytically continued $S$ matrix. This observation is based on the idea that the scattering wave function outside the range of the potential should have only outgoing waves when the energy is exactly that of the unstable state. The presence of a pole on the second sheet is then taken as a Lorentz invariant characterization of an unstable particle. Peierls ${ }^{9}$ brought this idea into field theory by suggesting that the pole should occur in the one particle propagator. Both the relation of this pole to the experimentally measured mass and the
dependence of the pole position upon the choice of field type was questioned by Stapp, ${ }^{10}$ who formulated a pure S-matrix theory of unstable particles.

This paper employs features of both approaches. The most important step is to use the existence of second sheet poles to construct a state vector for unstable particles. The construction is begun in Section II. It is found that whenever there is a pole on the second sheet of a scattering amplitude there is an eigenstate of the full, hermitian Hamiltonian with a complex energy. The usual prohibition against this is that

$$
\begin{equation*}
\langle E \mid E\rangle E=\langle E| H|E\rangle=E^{*}\langle E \mid E\rangle \tag{1.2}
\end{equation*}
$$

forbids E being complex. However, Section III shows that the unstable states lie not in the usual Hilbert-space but in a natural extension that corresponds to the second sheet of the $S$ matrix. In this larger space unstable particles automatically have zero norm and thus escape the prohibition (1.2). Therefore, in contrast to all earlier approaches, it is the vector space and not the Hamiltonian that is modified. The unstable states so constructed are eigenstates of ( $\mathrm{H}, \overrightarrow{\mathscr{P}}$ ) but do not transform as irreducible representations of the Poincaré group. Therefore in Section IV the Casimir corresponding to spin is expressed in a convenient manner that is used in Section V to construct unstable particle states that do transform irreducibly under real Lorentz transformations. The spin is automatically Lorentz invariant and can take on only real integral or half-integral values. The most interesting irreducible representations are the helicity states, which have the particularly simple one dimensional transformation law

$$
\begin{equation*}
\mathrm{U}(\Lambda)|\mathrm{k}, \mathrm{j}, \sigma\rangle=|\Lambda \mathrm{k}, \mathrm{j}, \sigma\rangle \mathrm{e}^{\mathrm{i} \theta_{\mathrm{w}}{ }^{\sigma}} \tag{1.3}
\end{equation*}
$$

Here, in contrast to the familiar massless case, the helicity $\sigma$ may take on all integrally spaced values from $-j$ to $+j$. The practical reason for considering
irreducible representations is given in Section VI. There it is shown that precisely these representations produce poles in partial wave amplitudes. The Lorentz transformation law (1.3) is, however, not observable because of the way in which the complex energy-momentum must be continued to the real axis. The observed transformation law is shown to be just the same as for stable particles. Next, wave functions are investigated. Because unstable particles correspond to eigenstates of energy-momentum, they possess genuine BetheSalpeter wave functions. These are used to calculate scattering amplitudes, which are expressed in a reduction formula of the LSZ type.

To defray any mystery about extending the Hilbert space to include zero norm states a simple noncelativistic example is worked out in Section VII. The Schrœdinger wave functions are displayed and used in calculations. This last section is essentially self-contained and may profitably be read immediately after Section II.

## II. CONSTRUCTION OF AN INTERPOLATING STATE VECTOR

Suppose that when the $S$ matrix is analytically continued in the total energy $\mathrm{k}^{\mathrm{o}}$ in a clockwise direction around a particular n-particle branch point there is a pole in $\mathrm{k}^{\circ}$. The first step in constructing a state vector for the corresponding unstable particle is to find an eigenstate of momentum $\mid \phi(\overrightarrow{\mathrm{K}})>$ such that the function

$$
\begin{equation*}
\left\langle\phi\left(\overrightarrow{\mathrm{k}}^{\mathrm{r}}\right)\right| \frac{1}{\mathrm{k}^{\mathrm{o}}-\mathrm{H}}|\phi(\overrightarrow{\mathrm{k}})\rangle \tag{2.1}
\end{equation*}
$$

has that same pole in $k^{\circ}$ on the second sheet. It is essential for the later analytic continuations that $|\phi(\vec{k})\rangle$ itself does not depend on $\mathrm{k}^{\circ}$. Therefore the similarity of (2.1) to the familiar expression

$$
\begin{equation*}
T_{f i}=\langle f| V+V \frac{1}{k^{0}-H} V|i\rangle \tag{2.2}
\end{equation*}
$$

is of no use; $|\phi\rangle$ cannot be chosen to be $V|i\rangle$ because $|i\rangle$ in (2.2) depends on $k^{0}$. In other words, (2.2) is an on-shell amplitude.

It is only at this point that field theory enters. Its Green's functions are off-shell analogues of (2.2). In particular, the $S$ matrix of a field theory can have a pole only if the Fourier transform of the corresponding time-ordered product of Heisenberg fields also has the same pole. It follows that the state

$$
\begin{equation*}
\int d^{3} X e^{i \vec{k} \cdot \vec{X}} T\left[\psi_{\alpha_{1}}\left(x_{1}\right) \cdots \psi_{\alpha_{n}}\left(x_{n}\right)\right] \mid 0> \tag{2.3}
\end{equation*}
$$

with

$$
\mathrm{X} \equiv \frac{1}{\mathrm{n}} \sum_{1}^{\mathrm{n}} \mathrm{x}_{\mathrm{i}}
$$

is a suitable choice for $|\phi(\overrightarrow{\mathrm{k}})\rangle$. Appendix A contains a more complete argument. It is convenient to sum the Lorentz indices in (2.3) to form eigenstates of $\vec{J}^{2}$ and $J_{3}$ (see also Appendix A) and to integrate the relative coordinates against
some smearing function to achieve the normalization

$$
\begin{equation*}
<\phi\left(\overrightarrow{\mathrm{k}}^{\prime}\right) \mathrm{j}^{\prime}, \mathrm{j}_{3}^{\prime} \mid \phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}>=\delta^{3}\left(\overrightarrow{\mathrm{k}}^{\prime}-\overrightarrow{\mathrm{k}}\right) \delta_{\mathrm{j}^{\prime} \mathrm{j}} \delta_{\mathrm{j}_{3}^{\prime} \mathrm{j}_{3}} \tag{2.4}
\end{equation*}
$$

It should be emphasized that the $|\phi\rangle$ so constructed is not an eigenstate of the Hamiltonian and does not transform into anything simple under boosts. In fact, a boosted $|\phi\rangle$ is not even an eigenstate of $\vec{J}^{2}$ and $J_{3}$.

Given such a $|\phi\rangle$, define projection operators

$$
\begin{aligned}
& A \equiv \int d^{3} k \sum_{j_{3}}\left|\phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}><\phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}\right| \\
& B \equiv 1-A
\end{aligned}
$$

with the multiplication properties

$$
\mathrm{A}^{2}=\mathrm{A}, \quad \mathrm{~B}^{2}=\mathrm{B}, \quad \mathrm{AB}=\mathrm{BA}=0
$$

Split the full Hamiltonian into diagonal and off-diagonal parts with respect to $\mid \phi>$ by defining

$$
\mathrm{H}=\mathrm{H}^{\prime}+\mathrm{H}^{\prime \prime}
$$

where

$$
\begin{align*}
\mathrm{H}^{\prime} & \equiv \mathrm{AHA}+\mathrm{BHB} \\
\mathrm{H}^{\prime \prime} & \equiv \mathrm{AHB}+\mathrm{BHA} \tag{2.6}
\end{align*}
$$

Define the full resolvent by

$$
\mathrm{R}\left(\mathrm{k}^{\mathrm{o}}\right) \equiv \frac{1}{\mathrm{k}^{\mathrm{o}}-\mathrm{H}}
$$

and a reduced resolvent diagonal with respect to $|\phi\rangle$ by

$$
r\left(k^{0}\right) \equiv \frac{1}{k^{0}-\mathrm{II}^{\prime}}
$$

Let the matrix element (2.1) be expressed as

$$
\begin{equation*}
\left\langle\phi\left(\vec{k}^{\prime}\right) \mathrm{j}^{\prime}, \mathrm{j}_{3}^{\prime}\right| R\left(\mathrm{k}^{\mathrm{o}}\right) \mid \phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}>\equiv \frac{\delta^{3} \overrightarrow{\left.\mathrm{i} \overrightarrow{\mathrm{k}}^{\prime}-\overrightarrow{\mathrm{k}}\right)} \delta_{j^{\prime} \mathrm{j}} \delta_{\mathrm{j}_{3}^{\prime} \mathrm{j}_{3}}^{\mathrm{D}(\mathrm{k})}}{}, \tag{2.7}
\end{equation*}
$$

where $D(k)$ depends on both $k^{\circ}$ and $\vec{k}$ and vanishes whenever the $S$ matrix has a pole. (The $j$ dependence of $D(k)$ is suppressed.)

A state vector for the unstable particle corresponding to the pole in (2.1) can now be built from $|\phi\rangle$. The connection between the full resolvent and the matrix element (2.1) is given by the identity

$$
\begin{equation*}
R\left(k^{0}\right)=\operatorname{Br}\left(k^{0}\right) \mathrm{B}+\left[1+\mathrm{r}\left(\mathrm{k}^{\mathrm{O}}\right) \mathrm{H}^{\prime \prime}\right] \operatorname{AR}\left(\mathrm{k}^{\mathrm{O}}\right) \mathrm{A}\left[\mathrm{II}^{\prime \prime} r\left(\mathrm{k}^{\mathrm{o}}\right)+1\right], \tag{2.8}
\end{equation*}
$$

which is proved in Appendix B. Rewrite this as

$$
\begin{equation*}
R\left(k^{o}\right)=\operatorname{Br}\left(k^{o}\right) B+\int d^{3} k \sum_{j_{3}} \frac{\left|\psi(k) j, j_{3}><\psi(k) j, j_{3}\right|}{D(k)} \tag{2.9}
\end{equation*}
$$

where

$$
\begin{equation*}
\left|\psi(k) j, j_{3}\right\rangle \equiv\left[1+r\left(k^{\circ}\right) H^{\prime \prime}\right]\left|\phi(\bar{k}) j_{,} j_{3}\right\rangle \tag{2.10}
\end{equation*}
$$

This may also be written as

$$
\begin{equation*}
\left.\left|\psi(k) j, j_{3}\right\rangle=\frac{1}{k^{\mathrm{O}}-\mathrm{H}} \right\rvert\, \phi(\vec{k}) j, j_{3}>D(k) . \tag{2.11}
\end{equation*}
$$

(From now on the spin indices j and $\mathrm{j}_{3}$ will be suppressed. They will be displayed again in Section $V$ and subsequently:) Equation (2.9) suggests that $|\psi\rangle$ interpolates between eigenstates of H . To demonstrate this, use the definition of $|\psi\rangle$ :

$$
\begin{aligned}
\left(\mathrm{k}^{\left.\mathrm{O}-\mathrm{H}^{\prime}\right)|\psi(\mathrm{k})\rangle}\right. & =\left[\mathrm{k}^{\left.\mathrm{o}-\mathrm{H}^{\prime}+\mathrm{H}^{\prime \prime}\right]|\phi(\vec{k})\rangle}\right. \\
\mathrm{H}^{\prime \prime} \mid \psi(\mathrm{k})> & =\left[\mathrm{H}^{\prime \prime}+\mathrm{H}^{\prime \prime} \mathrm{r}\left(\mathrm{k}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime}\right]|\phi(\overrightarrow{\mathrm{k}})\rangle
\end{aligned}
$$

Subtract to get

$$
\left(\mathrm{k}^{\mathrm{O}}-\mathrm{H}\right)|\psi(\mathrm{k})\rangle=\left[\mathrm{k}^{\mathrm{O}}-\mathrm{H}^{\prime}-\mathrm{H}^{\prime \prime} \mathrm{r}\left(\mathrm{k}^{\mathrm{O}}\right) \mathrm{H}^{\prime \prime}\right]|\phi \overrightarrow{(\vec{k})}\rangle .
$$

The right hand side of this is orthogonal to B and is therefore just $\mid \phi(\overrightarrow{\mathrm{k}})>$ multiplied by a function of $k$. This function is just $D(k)$ because

$$
\begin{equation*}
\left\langle\phi\left(\vec{k}^{\prime}\right)\right| \mathrm{k}^{\mathrm{o}}-\mathrm{H}^{\prime}-\mathrm{H}^{\prime \prime} \mathrm{r}\left(\mathrm{k}^{\mathrm{O}}\right) \mathrm{H}^{\prime \prime}|\phi(\overrightarrow{\mathrm{k}})\rangle=\delta^{3}(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}) \mathrm{D}(\mathrm{k}) \tag{2.12}
\end{equation*}
$$

as proved in Appendix B. Hence

$$
\begin{equation*}
\left(\mathrm{k}^{\mathrm{o}}-\mathrm{H}\right)|\psi(\mathrm{k})\rangle=\mid \phi(\overrightarrow{\mathrm{k}})>\mathrm{D}(\mathrm{k}) . \tag{2.13}
\end{equation*}
$$

This equation contains the principle result: If the $S$ matrix has a pole at some $\widetilde{\mathrm{k}}^{\mathrm{o}}$ then there is an eigenstate of the Hamiltonian with that energy. The state vector $|\psi\rangle$ therefore interpolates between all eigenstates of $H$ with the same quantum numbers as $|\phi\rangle$. When the pole of $S$ is on the physical Riemann sheet, $|\psi\rangle$ lies in the physical Hilbert space. When the pole is reached by continuing in $\mathrm{k}^{\circ}$ to another sheet, the state $\mid \psi(\mathrm{k})>$ must also be "continued" and this is the topic of Section III.

Before proceeding, however, it is a useful comparison to recall the usual applications of the same formulas. Equation (2.12) is essentially Dirac's method ${ }^{2}$ of summing (1.1) near an eigenvalue of $\mathrm{H}_{0}$. Without translation invariance there is no $\overrightarrow{\mathrm{k}}$ dependence. As the energy approaches the real axis from above

$$
\lim _{k^{\mathrm{o}} \rightarrow 0^{+}} \mathrm{D}\left(\mathrm{k}^{\mathrm{o}}\right)=\mathrm{k}^{\mathrm{o}}-\mathrm{E}_{\phi}-\Delta\left(\mathrm{k}^{\mathrm{O}}\right)+\frac{\mathrm{i}}{2} \gamma\left(\mathrm{k}^{\mathrm{o}}\right)
$$

where the unperturbed energy, level shift, and width are given by

$$
\begin{aligned}
& \mathrm{E}_{\phi}=\langle\phi| \mathrm{H}|\phi\rangle \\
& \Delta\left(\mathrm{k}^{\mathrm{O}}\right)=\langle\phi| \mathrm{H}^{\prime \prime} \frac{\mathscr{P}}{\mathrm{k}^{\mathrm{O}}-\mathrm{H}^{\prime}} \mathrm{H}^{\prime \prime}|\phi\rangle \\
& \gamma\left(\mathrm{k}^{\mathrm{O}}\right)=\langle\phi| \mathrm{H}^{\prime \prime} 2 \pi \delta\left(\mathrm{k}^{\mathrm{O}}-\mathrm{H}^{\prime}\right) \mathrm{H}^{\prime \prime}|\phi\rangle .
\end{aligned}
$$

In the Dirac approach $H^{\prime}$ is chosen as the unperturbed Hamiltonian and $|\phi\rangle$ is required to be an eigenstate of $\mathrm{H}^{\prime}$. Feshbach, ${ }^{6}$ on the other hand, let the projection operator (i.e., $\mid \phi>$ ) determine $H^{\prime}$ as done here in (2.6). His effective Ilamiltonian is the energy dependent, nonhermitian operator

$$
\mathrm{H}_{\mathrm{eff}} \equiv \mathrm{H}^{\prime}+\mathrm{H}^{\prime \prime} \mathrm{r}\left(\mathrm{k}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime}
$$

occurring in (2.12).
In all such approaches the final results are dependent upon the choice of the effective Hamiltonian. Write (2.7) with spin indices suppressed:

$$
\begin{equation*}
\frac{\delta^{3}\left(\vec{k}^{\prime}-\overrightarrow{\mathrm{k}}\right)}{\mathrm{D}(\mathrm{k})}=\left\langle\phi\left(\overrightarrow{\mathrm{k}}^{\prime}\right)\right| \frac{1}{\mathrm{k}^{0}-\mathrm{H}}|\phi(\overrightarrow{\mathrm{k}})\rangle \tag{2.14}
\end{equation*}
$$

Clearly $D(k)$ is a functional of $|\phi\rangle$ and $|\phi\rangle^{+}$. Choosing an effective Hamiltonian is equivalent to choosing a particular $|\phi\rangle$. The variation of $D(k)$ when $|\phi\rangle^{+}$is varied independently of $|\phi\rangle$ is

$$
\left.\frac{\delta \mathrm{D}(\mathrm{k})}{\delta|\phi\rangle^{+}}=-\frac{1}{\mathrm{k}^{\mathrm{O}}-\mathrm{H}} \right\rvert\, \phi(\overrightarrow{\mathrm{k}})>\mathrm{D}(\mathrm{k})^{2}
$$

which by (2.11) is just

$$
\begin{equation*}
\frac{\delta \mathrm{D}(\mathrm{k})}{\delta|\phi\rangle^{+}}=-|\psi(\mathrm{k})\rangle \mathrm{D}(\mathrm{k}) \tag{2.15}
\end{equation*}
$$

This variation vanishes only when $D(k)$ vanishes. Thus if a $|\phi\rangle$ has been chosen that produces a pole in (2.14) then the location of that pole is independent of the choice of $|\phi\rangle$ because of (2.15). It is essential to continue $k^{0}$ to the pole. All effective Hamiltonian treatments focus on the real value of $\mathrm{k}^{\circ}$ that produces a bump in (2.14). The location of such a maximum is usually called the resonance energy. However (2.15) shows that this energy clearly depends on the choice made for $|\phi\rangle$ and depends on it even though all orders of the perturbation are summed.

Note that it is the zeros of $D(k)$ that are significant for stable particles too. Equations (2.10) and (2.12) are then the usual formulas of Wigner-Brillouin perturbation theory for the eigenvectors and eigenvalues of $\mathrm{H}^{11}$ (The exact formulas do not depend on $H^{\prime}$ being small, of course.)

## III. EXTENSION OF THE HILBERT SPACE

To apply (2.13) to unstable particles an extension of the physical Hilbert space $h$ must be found that corresponds to different sheets of $\mathrm{D}(\mathrm{k})$. Always the appropriate sheet of $D(k)$ is determined by the resonance and not by the particular scattering process (i.e., $|\phi\rangle$ ) under consideration. For example, to reach the $\rho^{0}$ pole in the $\mathrm{e}^{+} \mathrm{e}^{-}$scattering amplitude requires going around the very same branch points as for the $\pi^{+} \pi^{-}$amplitude (see Fig. 1). Regardless of the amplitude, the pole is always reached by going around the heaviest decay threshold.
A. Inner Products in $\nless \uparrow$

For complex $\mathrm{k}^{0}$ the state

$$
\begin{equation*}
|\psi(\mathrm{k})\rangle=\left[1+\mathrm{r}\left(\mathrm{k}^{\circ}\right) \mathrm{H}^{\prime \prime}\right]|\phi(\overrightarrow{\mathrm{k}})\rangle \tag{3.1}
\end{equation*}
$$

is an ordinary state in the physical Hilbert space $\neq$ as long as the cuts along the real $\mathrm{k}^{\mathrm{o}}$ axis are avoided. A new state vector may always be added to $h$ by specifying the inner product of the new state with every vector of $h$. The means of specifying that inner product here will be analytic continuation and the enlarged space will be called $h \uparrow$. In particular, define the continuation of $\mid \psi(\mathrm{k})>$ as a new state whose inner product with an arbitrary state |f> is the analytic continuation in $\mathrm{k}^{\circ}$ of $\langle\mathrm{f} \mid \psi(\mathrm{k})\rangle$. The inner product to be continued is then

$$
\langle\mathrm{f} \mid \psi(\mathrm{k})\rangle=\langle\mathrm{f}| 1+\mathrm{r}\left(\mathrm{k}^{0}\right) \mathrm{H}^{\prime \prime}|\phi(\overrightarrow{\mathrm{k}})\rangle .
$$

Because of (2.11) this is just

$$
\begin{equation*}
\langle\mathrm{f} \mid \psi(\mathrm{k})\rangle=\langle\mathrm{f}| \frac{1}{\mathrm{k}^{\mathrm{O}}-\mathrm{H}}|\phi(\overrightarrow{\mathrm{k}})\rangle \mathrm{D}(\mathrm{k}) . \tag{3.2}
\end{equation*}
$$

The new states of special interest are those that correspond to a second-sheet pole in the S matrix. Let the pole occur at $\tilde{\mathrm{k}}^{\mathrm{o}}$ on the sheet reached by clockwise continuation around a particular $n$-particle branch point and let $D_{n}(k)$ be the
continuation of $D(k)$ around this branch point. The existence of a pole means that

$$
\begin{equation*}
D_{n}(\tilde{k})=0 \tag{3.3}
\end{equation*}
$$

The corresponding state is $\left|\psi_{\mathrm{n}}(\widetilde{\mathrm{K}})\right\rangle$. For any real-energy eigenstate. $|\mathrm{E}\rangle$ of $h$, (3.2) gives

$$
\langle\mathrm{E} \mid \psi(\mathrm{k})\rangle=\langle\mathrm{E} \mid \phi(\overrightarrow{\mathrm{k}})\rangle \frac{\mathrm{D}(\mathrm{k})}{\mathrm{k}^{\mathrm{O}}-\mathrm{E}}
$$

Analytically continuing to $\widetilde{\mathrm{k}}^{\mathrm{o}}$ yields

$$
\begin{equation*}
<\mathrm{E}\left|\psi_{\mathrm{n}}(\tilde{\mathrm{k}})\right\rangle=0 . \tag{3.4}
\end{equation*}
$$

The unstable particle is therefore orthogonal to all real-energy eigenstates.
At first, this result seems paradoxical because $|\phi\rangle$ itself is in $\nless$ and is therefore just a superposition of real-energy eigenstates. But $|\phi\rangle$ is certainly not orthogonal to the unstable state because

$$
\begin{equation*}
\left\langle\phi\left(\overrightarrow{\mathrm{k}}^{\mathrm{t}}\right) \mid \psi(\mathrm{k})\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}}^{\prime}-\overrightarrow{\mathrm{k}}\right), \tag{3.5}
\end{equation*}
$$

independently of $\mathrm{k}^{\mathrm{o}}$. The explanation lies in the difference between a continuous and a discrete superposition and is discussed further in Appendix C.

The inner product of two new states is the analytic continuation in two variables of their inner product in $h$. Thus the analogue of (3.2) is

$$
\left\langle\psi(\mathrm{k}) \mid \psi\left(\mathrm{k}^{\prime}\right)\right\rangle=\mathrm{D}(\mathrm{k})^{*}\langle\phi(\overrightarrow{\mathrm{k}})| \mathrm{R}\left(\mathrm{k}^{\mathrm{o}}\right)^{+} \mathrm{R}\left(\mathrm{k}^{\mathrm{O}_{\mathrm{r}}}\right)\left|\phi\left(\overrightarrow{\mathrm{k}}^{\prime}\right)\right\rangle \mathrm{D}\left(\mathrm{k}^{\mathrm{r}}\right) .
$$

Using

$$
\begin{equation*}
R\left(k^{0}\right)^{+} R\left(k^{O_{r}}\right)=\frac{R\left(k^{O_{i}}\right)-R\left(k^{O^{+}}\right)^{+}}{k^{O^{*}}-k^{O_{t}}} \tag{3.6}
\end{equation*}
$$

and the definition of $\mathrm{D}(\mathrm{k})$, this may be written as

$$
\begin{equation*}
\left\langle\psi(\mathrm{k}) \mid \psi\left(\mathrm{k}^{\prime}\right)\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}^{\prime}\right) \frac{\mathrm{D}(\mathrm{k})^{*}-\mathrm{D}\left(\mathrm{k}^{\prime}\right)}{\mathrm{k}^{\mathrm{O}^{*}}-\mathrm{k}^{\mathrm{o}^{\prime}}} . \tag{3.7}
\end{equation*}
$$

(Of course, the same result may be obtained directly from (3.1) by using an identity like (3.6) for the reduced resolvent $\mathrm{r}\left(\mathrm{k}^{\mathrm{o}}\right)$ and the alternate expression (2.12) for $D(k)$.) The inner product of an unstable state with itself results from continuing both $\mathrm{k}^{\circ}$ and $\mathrm{k}^{0}$ in (3.7) clockwise around the same branch point. Because

$$
\begin{aligned}
& \mathrm{D}_{\mathrm{n}}(\widetilde{\mathrm{k}})^{*}=\mathrm{D}_{\mathrm{n}}(\widetilde{\mathrm{k}})=0 \\
& \widetilde{\mathrm{k}}^{\mathrm{o}}-\widetilde{\mathrm{k}}^{\mathrm{o}} \neq 0
\end{aligned}
$$

it follows that

$$
\begin{equation*}
\left\langle\psi_{\mathrm{n}}(\widetilde{\mathrm{k}}) \mid \psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right\rangle=0 . \tag{3.8}
\end{equation*}
$$

Obviously (3.7) also implies that two different unstable particles (i.e., with different energies) are orthogonal. The vanishing of (3.8) is a direct consequence of $H$ being hermitian in (3.6). It will now be shown that since II is hermitian in $\boldsymbol{\hbar}$ it is also hermitian in $\boldsymbol{\hbar} \uparrow$ so that the argument (1.2) for a zero norm state was correct.

All the poles of $S$ discussed so far lie in the lower half $\mathrm{k}^{0}$-plane. However, it is well known ${ }^{12}$ that hermitian analyticity requires $S$ to have poles in complex conjugate pairs. From the definition of $D(k)$ in (2.14) it is clearly hermitian analytic for $\overrightarrow{\mathrm{k}}$ real:

$$
\begin{equation*}
\mathrm{D}\left(\mathrm{k}^{\mathrm{O}}, \overrightarrow{\mathrm{k}}\right)^{*}=\mathrm{D}\left(\mathrm{k}^{\mathrm{o}}, \overrightarrow{\mathrm{k}}\right) \tag{3.9}
\end{equation*}
$$

If (3.9) is continued in $\mathrm{k}^{\circ}$ clockwise around an n -particle branch point and into the lower half-plane then $\mathrm{k}^{\mathrm{O}}$ goes counterclockwise around that branch point and into the upper half-plane. See Fig. 2. Since $D_{n}(k)$ denotes the clockwise continuation of $D(k)$, let $D_{\bar{n}}(k)$ denote the counterclockwise continuation of $D(K)$. The continuation of (3.9) is then

$$
\begin{equation*}
\mathrm{D}_{\mathrm{n}}\left(\mathrm{k}^{\mathrm{o}}, \overrightarrow{\mathrm{k}}\right)^{*}=\mathrm{D}_{\overline{\mathrm{n}}}\left(\mathrm{k}^{\mathrm{o}}, \overrightarrow{\mathrm{k}}\right) \tag{3.10}
\end{equation*}
$$

Hence if $D_{n}$ vanishes at $\widetilde{k}^{o}$ the $D_{\bar{n}}$ vanishes at $\widetilde{k}^{0 *}$. Thus unstable particles come in complex conjugate pairs. This again reflects the fact that $H$ is hermitian even in the larger space $h \uparrow$. Denote the state resulting from counterclockwise continuation into the upper half plane with a subscript $\overline{\mathrm{n}}$. Thus

$$
\begin{align*}
& \mathrm{H}\left|\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right\rangle=\widetilde{\mathrm{k}}^{\mathrm{O}}\left|\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right\rangle \\
& \mathrm{H}\left|\psi_{\overline{\mathrm{n}}}\left(\widetilde{\mathrm{k}}^{*}\right)\right\rangle=\widetilde{\mathrm{k}}^{\mathrm{o} *} \mid \psi_{\left.\left.\overline{\mathrm{n}}^{( } \widetilde{\mathrm{k}}^{*}\right)\right\rangle} \tag{3.11a}
\end{align*}
$$

The adjoint states are defined by

$$
\left\langle\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right| \equiv\left|\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right\rangle^{+} .
$$

Thus

$$
\begin{align*}
& \left\langle\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right| \mathrm{H}=\left\langle\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right| \widetilde{\mathrm{k}}^{0 *}  \tag{3.11b}\\
& \left\langle\psi_{\overline{\mathrm{n}}^{-}}\left(\widetilde{\mathrm{K}}^{*}\right)\right| \mathrm{H}=\left\langle\psi_{\overline{\mathrm{n}}^{\left(\widetilde{\mathrm{k}}^{*}\right)} \mid \widetilde{\mathrm{k}}^{\mathrm{o}} .} .\right.
\end{align*}
$$

It is not surprising that

$$
\left.\left.\left\langle\psi_{\overline{\mathrm{n}}}\left(\widetilde{\mathrm{k}}^{*}\right)\right| \psi_{\left.\overline{\mathrm{n}}^{( }\right)} \widetilde{\mathrm{k}}^{*}\right)\right\rangle=0,
$$

just like (3.8). A more interesting calculation is the inner product between the two conjugate partners. This requires continuing both the $\mathrm{k}^{{ }^{\mathrm{O}}}$ in (3.7) clockwise from the upper half-plane around the n-particle branch point to $\widetilde{\mathrm{k}}^{0}$ and the $\mathrm{k}^{\circ}$ in (3.7) counter clockwise from the lower half-plane around the same threshold to the value $\widetilde{\mathrm{k}}^{\mathrm{o}}$. Performing the $\mathrm{k}^{0}$ continuation first gives

$$
\left\langle\psi_{\tilde{\mathrm{n}}^{( }}\left(\tilde{\mathrm{k}}^{*}\right) \mid \psi\left(\mathrm{k}^{\prime}\right)\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}^{\prime}\right) \frac{-\mathrm{D}\left(\mathrm{k}^{\mathrm{I}}\right)}{\tilde{\mathrm{k}}^{0}-\mathrm{k}^{0^{\mathrm{t}}}} .
$$

Because $D\left(k^{1}\right)$ vanishes at $\tilde{k}^{0}$, the continuation in $k^{0^{1}}$ then gives

$$
\left\langle\psi_{\overline{\mathrm{n}}}\left(\tilde{\mathrm{k}}^{*}\right) \mid \psi_{\mathrm{n}}\left(\tilde{\mathrm{k}}^{\prime}\right)\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}^{\prime}\right) \mathrm{N}(\tilde{\mathrm{k}})
$$

where

$$
\begin{equation*}
\left.N(\widetilde{\mathrm{k}}) \equiv \frac{\mathrm{d} \mathrm{D}_{\mathrm{n}}(\mathrm{k})}{\mathrm{dk}{ }^{0}}\right|_{\mathrm{k}=\widetilde{\mathrm{k}}} . \tag{3.12}
\end{equation*}
$$

If instead, the continuation in $\mathrm{k}^{\mathrm{O}}$ is done before that in $\mathrm{k}^{\mathrm{O}}$ the result is

Because of hermitian analyticity this is the same as (3.12). By assumption the pole in the $S$ matrix is of first order so that the zero of $D(k)$ is simple. Thus $\mathrm{N}(\tilde{\mathrm{K}}$ ) is a finite but nonvanishing constant. It is useful to absorb this factor by defining

$$
\begin{align*}
& |\Psi(\mathrm{k})\rangle \equiv\left|\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right\rangle \frac{1}{\sqrt{\mathrm{~N}(\widetilde{\mathrm{k}})}} \\
& |\Psi(\mathrm{k} *)\rangle \equiv\left|\psi_{\mathrm{n}}\left(\widetilde{\mathrm{k}}^{*}\right)\right\rangle \frac{1}{\sqrt{\mathrm{~N}(\widetilde{\mathrm{k}})^{*}}} \tag{3.13}
\end{align*}
$$

so that

$$
\begin{equation*}
\left\langle\Psi\left(\mathrm{k}^{*}\right) \mid \Psi\left(\mathrm{k}^{\prime}\right)\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}^{\prime}\right) . \tag{3.14}
\end{equation*}
$$

(The momentum k in $|\Psi(\mathrm{k})\rangle$ will always be on the mass-shell so there is no need to call it $\widetilde{\mathrm{K}}$.)

Because the unstable particle states have zero norm, it is necessary to find a new definition of norm that is positive-definite. This is necessary so that the equality of two vectors

$$
\begin{equation*}
\left|F_{1}\right\rangle=\left|F_{2}\right\rangle \tag{3.15}
\end{equation*}
$$

will have a precise meaning. In a Hilbert space the meaning of (3.15) is that

$$
\begin{equation*}
\langle\Delta F \mid \Delta F\rangle=0 \tag{3,16}
\end{equation*}
$$

where

$$
|\Delta F\rangle \equiv\left|F_{1}\right\rangle-\left|F_{2}\right\rangle .
$$

Obviously (3.16) is not good enough for $h \uparrow$ because it will still allow $\mid F_{1}>$ and $\mid F_{2}>$ to differ by an arbitrary number of unstable states. In Appendix C a positive definite norm operator $\Omega$ is constructed. The precise meaning of (3.13) is then that

$$
\begin{equation*}
\langle\Delta F| \Omega|\Delta F\rangle=0 . \tag{3.17}
\end{equation*}
$$

Goldberger and Watson ${ }^{13}$ have emphasized that the S matrix may have higher order poles on the second sheet. (The decay law then is the familiar $\mathrm{e}^{-\Gamma \mathrm{t}}$ multiplied by a polynomial in t.) In such a case $N=0$ in (3.12). Appendix D discusses the additional states that are then present in $\not \hbar \uparrow$.
B. A Lee Model Example

A simple example of an unstable particle occurs in the Lee model. ${ }^{14}$ The Hamiltonian for a static N and V is

$$
\begin{align*}
\mathrm{H}= & \mathrm{m}_{\mathrm{V}} \psi_{\mathrm{V}}^{\dagger} \psi_{\mathrm{V}}+\mathrm{m}_{\mathrm{N}^{\psi}} \psi^{\dagger} \mathrm{N}^{\psi} \mathrm{N}+\int \mathrm{d}^{3} \mathrm{p} \omega \mathrm{a}^{\dagger}(\overrightarrow{\mathrm{p}}) \mathrm{a}(\overrightarrow{\mathrm{p}}) \\
& +\left[\frac{\mathrm{g}_{0}}{\sqrt{4 \pi}} \psi_{\mathrm{V}}^{\dagger} \psi_{\mathrm{N}} \int \mathrm{~d}^{3} \mathrm{p} \frac{\mathrm{f}(\omega)}{\sqrt{(2 \omega)}} \mathrm{a}(\overrightarrow{\mathrm{p}})+\text { h.c. }\right] \tag{3.18}
\end{align*}
$$

where $\omega=\sqrt{\mathrm{m}_{\theta}^{2}+\overrightarrow{\mathrm{p}}^{2}}$ and $\mathrm{f}(\omega)$ is a real cutoff function. The interaction produces a dressed V state that is a simple combination of a bare V and a bare $\mathrm{N}+\theta$ :

$$
\begin{equation*}
|\mathrm{V}(\mathrm{E})\rangle=\left[\psi_{\mathrm{V}}^{\dagger}+\frac{\mathrm{g}_{0}}{\sqrt{4 \pi}} \psi_{\mathrm{N}}^{\dagger} \int \frac{\mathrm{d}^{3} \mathrm{p}}{\sqrt{2 \omega}} \frac{\mathrm{f}(\omega)}{\left(\mathrm{E}-\mathrm{m}_{\mathrm{N}}-\omega\right)} \mathrm{a}^{\dagger}(\overrightarrow{\mathrm{p}})\right]|0\rangle \tag{3.19}
\end{equation*}
$$

This state is an eigenstate of H for those values of E that cause

$$
\begin{equation*}
D(E)=E-m_{V}-\frac{g_{0}^{2}}{4 \pi} \int \frac{d^{3} p}{2 \omega} \frac{|f(\omega)|^{2}}{E-m_{N}-\omega} \tag{3.20}
\end{equation*}
$$

to vanish.

The formula for $|V\rangle$ in (3.19) is just a special case of that for $\psi\rangle$ in (3.1) in which $|\phi\rangle=\psi_{V}^{\dagger}|0\rangle, H^{\prime}$ is the free part of (3.18), and $H^{\prime \prime}$ is the interacting term of (3.18). The expression for $D(E)$ is a special case of (2.12). The zeros of $D(E)$ can only lie on the real axis below the branch point at $m_{N}+m_{\theta}$ or in the complex plane on another sheet. The s-wave $\mathrm{N}+\theta \rightarrow \mathrm{N}+\theta$ scattering amplitude is ${ }^{15}$

$$
\mathrm{e}^{2 \mathrm{i} \delta(\mathrm{E})}=\frac{\mathrm{D}_{\mathrm{II}}(\mathrm{E})}{\mathrm{D}_{\mathrm{I}}^{(\mathrm{E})}}
$$

where I and II indicate the physical and unphysical sheets of D. The second sheet poles of the scattering amplitude are therefore the zeros of $D$ on the second sheet. These poles come in complex conjugate pairs and have zero norm:

$$
\langle\mathrm{V}(\mathrm{E}) \mid \mathrm{V}(\mathrm{E})\rangle=\left\langle\mathrm{V}\left(\mathrm{E}^{*}\right) \mid \mathrm{V}\left(\mathrm{E}^{*}\right)\right\rangle=0 .
$$

The inner product of the two is the analytic continuation of

$$
\begin{equation*}
<\mathrm{V}\left(\mathrm{E}^{*}\right)|\mathrm{V}(\mathrm{E})\rangle-1+\frac{\mathrm{g}_{0}^{2}}{4 \pi}, \cdot \frac{\mathrm{~d}^{3} \mathrm{p}}{2 \omega} \frac{|\mathrm{f}(\omega)|^{2}}{\left(\mathrm{E}-\mathrm{m}_{\mathrm{N}}-\omega\right)^{2}} \tag{3.21}
\end{equation*}
$$

as obtained from (3.19). Clearly

$$
\left\langle\mathrm{V}\left(\mathrm{E}^{*}\right) \mid \mathrm{V}(\mathrm{E})\right\rangle=\frac{\mathrm{dD}(\mathrm{E})}{\mathrm{dE}}
$$

as shown generally in (3.12). There is no real benefit in choosing a particular cutoff function $f(\omega)$ and then calculating these integrals explicitly. A more illustrative example is worked out in detail in Section VII .

It should be noted that this version of the Lee model is nothing fancy despite the zero norm states. In particular, it is not the indefinite metric quantization of Källén and Pauli. ${ }^{15}$ They showed that the Lee model violates unitarity in the point source limit (i.e., $f(\omega) \rightarrow 1$ ). To preserve unitarity it is necessary to make $\mathrm{g}_{0}$ imaginary and to let the bare V state have negative norm. Such a
modification changes the sign of $m_{V}$ in (3.18) and (3.20) but, more importantly, makes $\mathrm{g}_{0}^{2}$ negative in (3.20) and (3.21). This allows $\mathrm{D}(\mathrm{E})$ to have a second order zero on the real axis ${ }^{16}$ or zeros at complex $E$ that lie on the physical sheet. ${ }^{17}$ These theories are quite different from the simple proposal made here of taking a well defined cutoff theory, quantized in a conventional Hilbert space with positivedefinite metric, and investigating the consequences of a pole on the second sheet of the scattering amplitude.

Another contrast can be made to the work of Glaser and Källén. ${ }^{18}$ They also treat the cutoff theory with conventional quantization. In order to discuss resonances they do not continue to the zeros of $D$ on the second sheet however, but instead define a new D by replacing the integral in (3.20) with its principal value. The new $D$ has zeros on the real axis above $m_{N}+m_{\theta}$ that give approximate eigenvalues and approximate eigenstates of H . Of course, their answers depend on this prescription.

## IV. POINCARÉ GROUP PRELIMINARIES

Once the eigenstates of energy and momentum are constructed, all the transformation properties can be derived by imitating Wigner. ${ }^{1}$ The present section introduces the covariance notions that will be needed.

## A. Complex Momentum

The unstable particle state $|\Psi(\mathrm{k})\rangle$ was constructed with a real momentum $\overrightarrow{\mathrm{k}}$ and a complex energy $\mathrm{k}^{\circ}$ such that

$$
\begin{align*}
& \mathscr{P}^{\mu}|\Psi(\mathrm{k})\rangle=\mathrm{k}^{\mu}|\Psi(\mathrm{k})\rangle  \tag{4.1}\\
& \left\langle\Psi\left(\mathrm{k}^{*}\right) \mid \Psi\left(\mathrm{k}^{\prime}\right)\right\rangle=\delta^{3}(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}})
\end{align*}
$$

Lorentz transformations are generated by six operators $M^{\mu \nu}$ satisfying

$$
\left[\mathrm{M}^{\mu \nu}, \mathrm{M}^{\alpha \beta}\right]=\mathrm{i}\left(\mathrm{~g}^{\mu \alpha} \mathrm{M}^{\nu \beta}-\mathrm{g}^{\mu \beta_{\mathrm{M}^{\nu \alpha}}}-\mathrm{g}^{\nu \alpha} \mathrm{M}^{\mu \beta}+\mathrm{g}^{\nu \beta} \mathrm{M}^{\mu \alpha}\right)
$$

Corresponding to each Lorentz transformation matrix $\Lambda_{\nu}^{\mu}$ is a linear operator

$$
\mathrm{U}[\Lambda]=\exp \left\{i \lambda_{\mu \nu} M^{\mu \nu}\right\}
$$

The commutation relations of the Poincaré group generators alone imply

$$
\mathrm{U}^{-1}[\Lambda] \mathscr{P}^{\mu} \mathrm{U}[\Lambda]=\Lambda_{\nu}^{\mu} \mathscr{P}^{\nu} .
$$

From this and (4.1) it follows that states with complex three-momentum $\vec{k}^{\prime}$ are automatically generated:

$$
\begin{equation*}
\mathscr{P}^{\mu}\{\mathrm{U}[\Lambda] \mid \Psi(\mathrm{k})>\}=\mathrm{k}^{\prime}\{\mathrm{U}[\Lambda] \mid \Psi(\mathrm{k})>\} \tag{4.2}
\end{equation*}
$$

where $\mathrm{k}^{\mu}=\Lambda^{\mu}{ }_{\nu} \mathrm{k}^{\nu}$.
Complex momenta suggest wave functions which grow exponentially in some spatial direction. In the soluble example of Section VII this is exactly the case. Such exponential growth is, in fact, characteristic of second sheet singularities. It is no problem because inner products are always calculated by analytically
continuing in momentum space. (Heuristically this corresponds to integration along a complex path in position space.) To analytically continue in momentum space requires continuing the $\delta$-function in (4.1). This is discussed in (5.37) and Ref. 24.

At present, it is necessary to introduce some covariant notation. Because of (4.2) the momentum of an unstable particle has the general form

$$
\begin{equation*}
\mathrm{k}^{\mu}=\mathrm{p}^{\mu}+\mathrm{iq}^{\mu}, \tag{4.3}
\end{equation*}
$$

where $p$ and $q$ are real four-vectors. Since $k^{2}$ is Lorentz invariant, so are its real and imaginary parts. Thus put

$$
\begin{equation*}
\mathrm{k}^{2}=\mathrm{M}^{2}-\mathrm{iM} \mathrm{\Gamma} \tag{4.4}
\end{equation*}
$$

where $M$ and $\Gamma$ are real constants that characterize the unstable particle. In terms of p and q this means

$$
\begin{align*}
& M^{2}=p^{2}-q^{2} \\
& M \Gamma=-2 p \cdot q \tag{4.5}
\end{align*}
$$

It is also useful to define the complex number

$$
\mathscr{M} \equiv \frac{\mathrm{M}}{\sqrt{2}}\left(1+\sqrt{1+\left(\frac{\Gamma}{\mathrm{M}}\right)^{2}}\right)^{1 / 2}-\mathrm{i} \frac{\Gamma}{\sqrt{2}}\left(1+\sqrt{1+\left(\frac{\Gamma}{\mathrm{M}}\right)^{2}}\right)^{-1 / 2},(4.6)
$$

in terms of which the mass shell condition (4.4) is

$$
\mathrm{k}^{2}=\mathscr{M}^{2}
$$

Note that $\Gamma$ may be positive or negative because of hermitian analyticity.
The analytic continuation of Section III yields poles with $\overrightarrow{\mathrm{k}}$ real, i.e.,

$$
\begin{align*}
& \mathrm{p}=\left(\operatorname{Rek}^{\mathrm{O}}, \overrightarrow{\mathrm{k}}\right)  \tag{4.7}\\
& \mathrm{q}=\left(\operatorname{Im~k}^{\mathrm{O}}, 0\right)
\end{align*}
$$

Obviously $q^{2}>0$. Because of $(4.5) \mathrm{p}^{2}>0$. Any state whose momentum is related to (4.7) by a real Lorentz transformation must have the same values of $p^{2}$ and $q^{2}$. Conversely, any such state may be transformed by a real Lorentz transformation into one with the standard momentum

$$
\begin{align*}
& \overline{\mathrm{p}} \equiv\left(\frac{\mathrm{p} \cdot \mathrm{q}}{\sqrt{q^{2}}}, 0,0, \sqrt{\frac{(p \cdot q)^{2}-\mathrm{p}^{2} q^{2}}{q^{2}}}\right) \\
& \overline{\mathrm{q}} \equiv\left(\sqrt{\mathrm{q}^{2}}, 0,0,0\right) . \tag{4.8}
\end{align*}
$$

This is just (4.7) with $\overrightarrow{\mathrm{k}}$ rotated into the z -axis.

## B. The Spin Casimir

In order to find unstable particle states that transform irreducibly under the Poincaré group it is necessary to diagonalize the Casimir's $\mathscr{P}^{2}$ and $W^{2}$ where

$$
\begin{equation*}
\mathrm{W}_{\alpha} \equiv \frac{1}{2} \epsilon_{\alpha \beta \mu \nu} \mathscr{P}^{\beta_{\mathrm{M}}}{ }^{\mu \nu} \tag{4.9}
\end{equation*}
$$

in the notation of Gasiorowicz. ${ }^{19}$ This is easily done for a state with real momentum by transforming to the rest frame. This section will show that this technique is the only way to diagonalize $\mathrm{W}^{2}$ even for states with complex momentum.

The application of $W_{\alpha}$ to the eigenstates of energy-momentum constructed in Section III replaces $\mathscr{P}^{\beta}$ by its eigenvalue:

$$
\mathrm{W}_{\alpha}|\Psi(\mathrm{k})\rangle=\mathrm{W}_{\alpha}(\mathrm{k}) \mid \Psi(\mathrm{k})>,
$$

with

$$
\begin{equation*}
\mathrm{w}_{\alpha}(\mathrm{k}) \equiv \frac{1}{2} \epsilon_{\alpha \beta \mu \nu} \mathrm{k}^{\beta_{\mathrm{M}}}{ }^{\mu \nu} \tag{4.10}
\end{equation*}
$$

Now evaluate (4.10) when k has the special form

$$
\begin{align*}
& \overline{\mathrm{k}}^{\mu}=(\alpha, 0,0, \beta)  \tag{4.11}\\
& \alpha^{2}-\beta^{2}=\mathscr{M}^{2}
\end{align*}
$$

introduced in (4.8). The result is

$$
\begin{align*}
& \mathrm{W}_{0}(\overline{\mathrm{k}})=-\beta \mathrm{J}_{3} \\
& \mathrm{~W}_{1}(\mathrm{k})=\alpha \mathrm{J}_{1}-\beta \mathrm{K}_{2} \\
& \mathrm{~W}_{2}(\overline{\mathrm{k}})=\alpha \mathrm{J}_{2}+\beta \mathrm{K}_{1}  \tag{4.12}\\
& \mathrm{~W}_{3}(\overline{\mathrm{k}})=\alpha \mathrm{J}_{3},
\end{align*}
$$

where

$$
\begin{align*}
J_{i} & \equiv-\frac{1}{2} \epsilon_{i j k} M^{j k} \\
K_{i} & \equiv M_{i 0} \tag{4.13}
\end{align*}
$$

are the generators of rotations and boosts, respectively. ${ }^{19}$ The Casimir is then

$$
\begin{equation*}
[\mathrm{W}(\overline{\mathrm{~K}})]^{2}=-\left[\alpha \mathrm{J}_{1}-\beta \mathrm{K}_{2}\right]^{2}-\left[\alpha \mathrm{J}_{2}+\beta \mathrm{K}_{1}\right]^{2}-\mathscr{M}^{2}\left[\mathrm{~J}_{3}\right]^{2} \tag{4.14}
\end{equation*}
$$

This suggests defining

$$
\begin{align*}
& \mathrm{S}_{1}(\overline{\mathrm{k}})=\frac{1}{\mathscr{M}}\left(\alpha \mathrm{~J}_{1}-\beta \mathrm{K}_{2}\right) \\
& \mathrm{S}_{2}(\overline{\mathrm{k}})=\frac{1}{\mathscr{M}}\left(\alpha \mathrm{~J}_{2}+\beta \mathrm{K}_{1}\right)  \tag{4.15}\\
& \mathrm{S}_{3}(\overline{\mathrm{k}})=\mathrm{J}_{3}
\end{align*}
$$

in order that the Casimir be

$$
\begin{equation*}
[\mathrm{W}(\overline{\mathrm{k}})]^{2}=-\mathscr{M}^{2}[\overline{\mathrm{~s}}(\overline{\mathrm{k}})]^{2} \tag{4.16}
\end{equation*}
$$

The operators $S_{a}(\bar{k})$ in (4.15) are the particular combination of rotations and boosts that leave the vector K invariant, i.e., they are the generators of the little group. (In the special case of stable massless particles the appropriate combination provided by (4.14) is obviously $J_{1}-\mathrm{K}_{2}$ and $\mathrm{J}_{2}+\mathrm{K}_{1}$ rather than (4.15). It is well known that these two combinations plus their closure, $\mathrm{J}_{3}$, generate the little group of massless particles. ${ }^{20}$ )

To find the eigenvalues of $\mathrm{W}^{2}$, observe that (4.15) gives the familiar commutation relations

$$
\begin{equation*}
\left[\mathrm{S}_{\mathrm{a}}(\overline{\mathrm{k}}), \mathrm{S}_{\mathrm{b}}(\overline{\mathrm{k}})\right]=\mathrm{i} \epsilon_{\mathrm{abc}} \mathrm{~S}_{\left.\mathrm{c}^{(\mathrm{k}}\right)} \tag{4.17}
\end{equation*}
$$

Obviously both $\mathrm{S}_{1}$ and $\mathrm{S}_{2}$ in (4.15) are nonhermitian because $\alpha, \beta$, and $\mathscr{M}$ are complex. However, because $\operatorname{SU}(2)$ is compact all solutions to the commutation relations (4.17) are equivalent to hermitian operators, ${ }^{21}$ i.e., there exists a hermitian operator $\vec{V}$ and a transformation $T$ such that

$$
\begin{equation*}
\overrightarrow{\mathrm{S}}(\overline{\mathrm{k}})=\mathrm{T} \overrightarrow{\mathrm{~V}} \mathrm{~T}^{-1} \tag{4.18}
\end{equation*}
$$

Hence the eigenvalues of $[\overline{\mathrm{S}}(\overline{\mathrm{k}})]^{2}$ are, as usual, real nonnegative integers or halfintegers and label irreducible representations of the Poincaré group because of (4.16). That these integers or half-integers actually are related to rotations will not be shown until Section V.

To diagonalize the spin Casimir when k is arbitrary requires constructing a general $\overrightarrow{\mathrm{S}}(\mathrm{k})$. It is useful for this and later sections to define complex Lorentz transformation matrices $L^{c}$ by

$$
\begin{equation*}
\mathrm{L}^{\mathrm{c}}\left(\mathrm{k}^{\mathbf{\prime}}, \mathrm{k}\right)_{\beta}^{\alpha} \mathrm{k}^{\beta}=\mathrm{k}^{{ }^{\alpha}}, \tag{4.19}
\end{equation*}
$$

where k and $\mathrm{k}^{\prime}$ are arbitrary complex momenta with the same mass. (Note that (4.19) does not completely specify $L^{c}$.) Now consider the transformation

$$
\begin{equation*}
\mathrm{L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})^{\alpha}{ }_{\mu} \mathscr{M}^{\mu}=\mathrm{k}^{\alpha}, \tag{4.20}
\end{equation*}
$$

where $\mathscr{M}$ denotes the momentum vector

$$
\begin{equation*}
\mathscr{M}^{\mu}=(\mathscr{M}, 0,0,0) \tag{4.21}
\end{equation*}
$$

Use the columns of this matrix to define three complex vectors:

$$
\begin{equation*}
\mathrm{e}^{\alpha}(\mathrm{k}, \mathrm{a}) \equiv \mathrm{L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})_{\mathrm{a}}^{\alpha} \quad \mathrm{a}=1,2,3 \tag{4.22}
\end{equation*}
$$

These vectors are orthonormal

$$
\begin{equation*}
\mathrm{e}^{\alpha}(\mathrm{k}, \mathrm{a}) \mathrm{e}_{\alpha}(\mathrm{k}, \mathrm{~b})=-\delta^{\mathrm{ab}} \tag{4.23}
\end{equation*}
$$

and orthogonal to $\mathrm{k}^{\alpha}$ itself. Together the four vectors form a complete basis in that

$$
\begin{equation*}
-\sum_{\mathrm{a}=1}^{3} \mathrm{e}^{\alpha}(\mathrm{k}, \mathrm{a}) \mathrm{e}^{\beta}(\mathrm{k}, \mathrm{a})+\left(\frac{1}{\mathscr{M}}\right)^{2} \mathrm{k}^{\alpha} \mathrm{k}^{\beta}=\mathrm{g}^{\alpha \beta} . \tag{4.24}
\end{equation*}
$$

This relation is useful in calculating the Casimir

$$
\begin{equation*}
[\mathrm{w}(\mathrm{k})]^{2}=\mathrm{w}_{\alpha}(\mathrm{k}) \mathrm{g}^{\alpha \beta} \mathrm{w}_{\beta}(\mathrm{k}) . \tag{4.25}
\end{equation*}
$$

From the definition (4.10)

$$
\mathrm{k}^{\alpha} \mathrm{W}_{\alpha}(\mathrm{k})=0
$$

so that (4.25) reduces to

$$
\begin{equation*}
[\mathrm{W}(\mathrm{k})]^{2}=-\mathscr{M}^{2}[\overrightarrow{\mathrm{~S}}(\mathrm{k})]^{2} \tag{4.26}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathrm{S}_{\mathrm{a}}(\mathrm{k}) \equiv \frac{1}{\mathscr{M}} \mathrm{e}^{\alpha}(\mathrm{k}, \mathrm{a}) \mathrm{W}_{\alpha}(\mathrm{k}) \quad \mathrm{a}=1,2,3 \tag{4.27}
\end{equation*}
$$

A different expression for $\overrightarrow{\mathrm{S}}(\mathrm{k})$ may be obtained by using

$$
\mathrm{e}^{\alpha}(\mathrm{k}, \mathrm{a}) \frac{\mathrm{k}^{\beta}}{\mathscr{M}} \epsilon_{\alpha \beta \mu \nu}=-\epsilon_{\mathrm{abc}} \mathrm{e}_{\mu}(\mathrm{k}, \mathrm{~b}) \mathrm{e}_{\nu}(\mathrm{k}, \mathrm{c})
$$

in (4.27) to get

$$
\begin{equation*}
\mathrm{S}_{\mathrm{a}}(\mathrm{k})=-\frac{1}{2} \epsilon_{\mathrm{abc}} \mathrm{e}_{\mu}(\mathrm{k}, \mathrm{~b}) \mathrm{e}_{\nu}(\mathrm{k}, \mathrm{c}) \mathrm{M}^{\mu \nu} \tag{4.28}
\end{equation*}
$$

From (4.28) the commutation relations

$$
\begin{equation*}
\left[S_{a}(k), S_{b}(k)\right]=i \epsilon_{a b c} S_{c}(k) \tag{4.29}
\end{equation*}
$$

follow easily.
Again because of the compactness of $\operatorname{SU}(2)$ the eigenvalues of $[\overrightarrow{\mathrm{S}}(\mathrm{k})]^{2}$ are real integers or half-integers. ${ }^{21}$ The equivalence transformation from the
nonhermitian $\overrightarrow{\mathrm{S}}(\mathrm{k})$ to three hermitian operators can now be explicitly constructed. The definitions (4.22) and (4.27) give

$$
\begin{equation*}
\mathrm{S}_{\mathrm{a}}(\mathrm{k})=\frac{1}{\mathscr{M}} \mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})_{\mathrm{a}}^{\alpha} \mathrm{w}_{\alpha}(\mathrm{k}) \tag{4.30}
\end{equation*}
$$

The definition (4.10) of $\mathrm{W}_{\alpha}{ }^{(\mathrm{k})}$ guarantees its covariance under a real Lorentz transformation:

$$
\begin{equation*}
\Lambda_{\mu}^{\alpha} W_{\alpha}(\mathrm{k})=\mathrm{U}[\Lambda] \mathrm{W}_{\mu}\left(\Lambda^{-1} \mathrm{k}\right) \mathrm{U}^{-1}[\Lambda] \tag{4.31}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathrm{U}[\Lambda]=\exp \left\{\dot{\mathrm{i} \lambda_{\mu \nu}} \mathrm{M}^{\mu \nu}\right\} \tag{4.32}
\end{equation*}
$$

as always. When the transformation matrix $\Lambda$ is complex the six parameters $\lambda_{\mu \nu}$ are complex, but (4.31) still holds because it depends only on the commutators of the $\mathrm{M}^{\mu \nu}$ with themselves. Thus using (4.31) in (4.30) gives

$$
\begin{equation*}
\mathrm{S}_{\mathrm{a}}(\mathrm{k})=\frac{1}{\mathscr{M}} \mathrm{U}\left[\mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right] \mathrm{W}_{\mathrm{a}}(\mathscr{M}) \mathrm{U}^{-1}\left[\mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right] . \tag{4.33}
\end{equation*}
$$

Because

$$
\mathrm{W}_{\mathrm{a}}(\mathscr{M})=\mathscr{M} \mathrm{J}_{\mathrm{a}},
$$

(4.33) may be rewritten as

$$
\begin{equation*}
\overrightarrow{\mathrm{S}}(\mathrm{k})=\mathrm{U}^{-1}\left[\mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right] \vec{J} \mathrm{U}\left[\mathrm{~L}^{\mathrm{c}}(\mathscr{M}, \mathrm{k})\right] . \tag{4.34}
\end{equation*}
$$

Note that $U$ here is not unitary because the parameters $\lambda_{\mu \nu}$ in (4.32) are complex. This is the promised equivalence relation between some hermitian operator and the nonhermitian $\vec{S}(\mathrm{k})$. It shows, furthermore, that to construct unstable particle states that transform irreducibly under real Lorentz transformations necessarily requires the use of complex Lorentz transformations.

## V. LORENTZ TRANSFORMATIONS OF UNSTABLE PARTICLES

Sections II and III showed that corresponding to every pole in a scattering amplitude there is an eigenstate of energy momentum in a space $\boldsymbol{h} \uparrow$ larger than the physical Hilbert space. This section will show that corresponding to every pole is a state which transforms irreducibly under the Poincare group.
A. $\underline{J}_{3}$ States

Because of (4.26) a state |k> will transform irreducibly under real Lorentz transformations only if

$$
\begin{equation*}
[\overrightarrow{\mathrm{S}}(\mathrm{k})]^{2}|\mathrm{k}\rangle=j(j+1) \mid \mathrm{k}> \tag{5.1}
\end{equation*}
$$

with j a real integer or half-integer. Right away the states constructed in Section III therefore cannot be-irreducible. They are given in (3.21) by the analytic continuation of

$$
\begin{equation*}
\left.\left|\Psi(\mathrm{k}) \mathrm{j}, \mathrm{j}_{3}\right\rangle=\left[1+\frac{1}{\mathrm{k}^{0}-\mathrm{H}^{\prime}} \mathrm{H}^{\prime \prime}\right] \right\rvert\, \phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}>\frac{1}{\sqrt{\mathrm{N(k)}}} . \tag{5.2}
\end{equation*}
$$

Recall that $\left|\phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}\right\rangle$ are the states constructed in Appendix A for an arbitrary real momentum $\overrightarrow{\mathrm{k}}$. They are eigenstates of $\vec{J}^{2}$ and $J_{3}$ because the discrete field indices are summed against appropriate Clebsch-Gordon coefficients. Both $\mathrm{H}^{\prime}$ and $\mathrm{H}^{\prime \prime}$ commute with $\overrightarrow{\mathrm{J}}^{2}$ and $\mathrm{J}_{3}$ since they are just the projections (2.6) of H determined by $\left|\phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}\right\rangle$. Therefore (5.2) and its analytic continuation satisfy

$$
\begin{align*}
& \overrightarrow{\mathrm{J}}^{2}\left|\Psi(\mathrm{k}) \mathrm{j}, \mathrm{j}_{3}>=\mathrm{j}(\mathrm{j}+1)\right| \Psi(\mathrm{k}) \mathrm{j}, \mathrm{j}_{3}> \\
& \mathrm{J}_{3}\left|\Psi(\mathrm{k}) \mathrm{j}, \mathrm{j}_{3}\right\rangle=\mathrm{j}_{3}\left|\Psi(\mathrm{k}) \mathrm{j}, \mathrm{j}_{3}\right\rangle \tag{5.3}
\end{align*}
$$

and do not satisfy (5.1).
The $|\Psi\rangle$ are, however, irreducible representations when $\overrightarrow{\mathrm{k}}=0$ because (5.1) and (5.3) are then identical due to

$$
\begin{equation*}
\left.\mathrm{S}_{\mathrm{a}}(\mathrm{k})\right|_{\overrightarrow{\mathrm{k}}=0}=\mathrm{J}_{\mathrm{a}} \tag{5.4}
\end{equation*}
$$

An irreducible representation with momentum $k$ then automatically results from boosting such a $|\Psi\rangle$ from rest to k . With the rest momentum again denoted by

$$
\begin{equation*}
\mathscr{M}^{\mu}=(\mathscr{M}, 0,0,0) \tag{5.5}
\end{equation*}
$$

the irreducible representation is just

$$
\begin{equation*}
\left|\mathrm{k}, \mathrm{j}_{,} \mathrm{j}_{3}>\equiv \mathrm{U}\left[\mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right]\right| \Psi(\mathscr{M}) \mathrm{j}, \mathrm{j}_{3}>\sqrt{2 \mathscr{M}} . \tag{5.6}
\end{equation*}
$$

(The $\sqrt{2 \mathscr{M}}$ is included to give a covariant normalization in Section V.D.) The proof of irreducibility is that

$$
\begin{aligned}
& \mathscr{P}^{\mu}\left|\mathrm{k}, \mathrm{j}, \mathrm{j}_{3}>=\mathrm{k}^{\mu}\right| \mathrm{k}, \mathrm{j}, \mathrm{j}_{3}> \\
& {[\overrightarrow{\mathrm{s}}(\mathrm{k})]^{2}\left|\mathrm{k}, \mathrm{j}, \mathrm{j}_{3}\right\rangle=\mathrm{j}(\mathrm{j}+1) \mid \mathrm{k}, \mathrm{j}, \mathrm{j}_{3}>}
\end{aligned}
$$

In addition, (5.6) satisfies

$$
\begin{equation*}
S_{3}(k)\left|k, j, j_{3}>=j_{3}\right| k, j, j_{3}> \tag{5.7}
\end{equation*}
$$

The difference between (5.2) and (5.6) has nothing to do with complex momentum; the difference is equally present for stable particles in the physical Hilbert space. Also note that the precise meaning of (5.6) is that its inner product with an arbitrary state |f> is given by the analytic continuation in $\mathrm{k}^{\circ}$ of

$$
\begin{equation*}
\left\langle\mathrm{f} \mid \mathrm{k}, \mathrm{j}, \mathrm{j}_{3}\right\rangle=\langle\mathrm{f}| \mathrm{e}^{\mathrm{i} \lambda_{\mu \nu} \mathrm{M}^{\mu \nu}} \frac{1}{\mathrm{k}^{\mathrm{o}}-\mathrm{H}}\left|\phi(\overrightarrow{0}) \mathrm{j}, \mathrm{j}_{3}\right\rangle \mathrm{D}\left(\mathrm{k}^{\mathrm{o}}\right) \sqrt{2 \mathscr{M}} \tag{5.8}
\end{equation*}
$$

to the value $\mathrm{k}^{\mathrm{o}}=\mathscr{M}$. Here the $\lambda_{\mu \nu}$ are six complex parameters corresponding to a boost from momentum $\mathscr{M}^{\mu}$ to $k^{\mu}$. Equation (5.8) is an obvious parallel to (3.2). From now on Lorentz transformations will be applied at will but the precise meaning will always be in terms of inner products like (5.8).

The behavior of (5.6) under a real Lorentz transformation $\Lambda$ now follows easily:

$$
\begin{equation*}
\mathrm{U}[\Lambda]\left|\mathrm{k}, \mathrm{j}, \mathrm{j}_{3}\right\rangle=\mathrm{U}\left[\mathrm{~L}^{\mathrm{c}}(\Lambda \mathrm{k}, \mathscr{M})\right] \mathrm{U}\left[\mathrm{R}_{\mathrm{W}}^{\mathrm{c}}\right]\left|\mathscr{M}, \mathrm{j}, \mathrm{j}_{3}\right\rangle \tag{5.9}
\end{equation*}
$$

where

$$
\mathrm{R}_{\mathrm{W}}^{\mathrm{c}} \equiv \mathrm{~L}^{\mathrm{c}}(\Lambda \mathrm{k}, \mathscr{M})^{-1} \Lambda \mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})
$$

Clearly $\mathrm{R}_{\mathrm{W}}^{\mathrm{c}}$ is a complex Lorentz transformation that leaves $\mathscr{M}^{\mu}$ invariant. It is therefore just a complex Wigner rotation whose affect on the rest state is

Here $D$ is the usual rotation matrix for spin $j$. The three rotation angles are complex. The complete transformation law is then

$$
\begin{equation*}
U[\Lambda]\left|k, j, j_{3}>=\sum_{j_{3}^{\prime}}\right| \Lambda k, j, j_{3}^{r}>D_{j_{3}^{t} j_{3}}\left(R_{W}^{c}\right) \tag{5.11}
\end{equation*}
$$

This result is quite similar to the usual transformation law for stable particles. It has been suggested, in fact, that the only momentum $\mathrm{k}^{\mu}$ allowable for an unstable particle are those for which $\frac{1}{\mathscr{M}} \mathrm{k}^{\mu}$ is purely real. ${ }^{22}$ Such a restriction makes the Wigner rotation in $(5.11)$ purely real. The explicitly constructed states in Section III, however, show that other momenta do occur and $R_{W}^{c}$ is therefore complex.

## B. Other States

The states constructed in Section III by analytic continuation have $\mathrm{p}^{2}>0$ and $q^{2}>0$. (See Eq. (4.7).) Once complex boosts are allowed states with any momentum k satisfying $\mathrm{k}^{2}=\mathscr{M}^{2}$ can be constructed from these. (The $\mathrm{J}_{3}$ states of the previous subsection, for example, have no restriction on the sign of $\mathrm{p}^{2}$ or $q^{2}$.) Out of this myriad of possible states the physically important ones are those that can be excited in physical scattering processes. All states used in scattering experiments have a total energy momentum that is timelike. Therefore the only unstable particle states than can be excited in physical scattering
processes must have a four-momentum $k$ that becomes timelike when continued back to the real axis. In short, they must have $\mathrm{p}^{2}>0$ to produce resonances in scattering amplitudes. (See Section VI. A for more details.) For a given timelike $p$ the associated $q$ may be either timelike, lightlike, or spacelike. In each case the momentum k is related by a real Lorentz transformation to a standard momentum $\overline{\mathrm{k}}$ defined as follows:

If $p^{2}>0, q^{2}>0$

$$
\begin{aligned}
& \bar{p} \equiv\left(\frac{p \cdot q}{\sqrt{q^{2}}}, 0,0, \sqrt{\frac{(p \cdot q)^{2}-p^{2} q^{2}}{q^{2}}}\right) \\
& \bar{q} \equiv\left(\sqrt{q^{2}}, 0,0, \theta\right)
\end{aligned}
$$

If $p^{2}>0, q^{2}=0$

$$
\begin{align*}
& \bar{p} \equiv\left(\frac{p^{2} A^{2}+(p \cdot q)^{2}}{2 A(p \cdot q)}, 0,0, \frac{p^{2} A^{2}-(p \cdot q)^{2}}{2 A(p \cdot q)}\right)  \tag{5.12~b}\\
& \bar{q} \equiv(A, 0,0, A) \quad ;
\end{align*}
$$

If $\mathrm{p}^{2}>0, \mathrm{q}^{2}<0$

$$
\begin{align*}
& \bar{p} \equiv\left(\sqrt{\frac{(p \cdot q)^{2}-p^{2} q^{2}}{-q^{2}}}, 0,0, \frac{p \cdot q}{\sqrt{-q^{2}}}\right)  \tag{5.12c}\\
& \bar{q} \equiv\left(0,0,0, \sqrt{-q^{2}}\right)
\end{align*}
$$

Obviously (5.12a) is the prototype (4.8) that resulted from analytic continuation in the variable $\mathrm{k}^{\circ}$.

The $J_{3}$ states constructed previously did not take advantage of the fact that only states with $\mathrm{p}^{2}>0$ are excited in physical scattering processes. Their transformation law (5.11) is based only upon the invariance of momentum (5.5) under arbitrary rotations. The little group thus has three generators. On the other
hand, the only real Lorentz transformations that leave any of the $\bar{k}$ in (5.12) invariant are rotations around the z-axis. The little group thus has only one generator and its irreducible representations are one dimensional. It might therefore be possible to construct states which are irreducible representations both of the full Poincaré group and of this one parameter little group. Because the representation of the translations is not unitary there is no guarantee of finding such states. In the next subsection, however, they are successfully constructed and their transformation law (5.21) is derived.

For completeness, states with momenta other than $\mathrm{p}^{2}>0$ are discussed in Appendix E. Beltrametti and Luzzatto ${ }^{23}$ have shown that for any complex momentum $k$ a standard $\bar{k}$ can be chosen. The corresponding little group is always one dimensional but depends on the sign of

$$
\begin{equation*}
\Delta \equiv(p \cdot q)^{2}-p^{2} q^{2} \tag{5.13}
\end{equation*}
$$

The momenta in (5.12) all have $\Delta>0$ and the states which are irreducible representations of both Poincare group and the little group are constructed in the next subsection. In Appendix $E$ states with $\Delta<0$ on which both groups are represented irreducibly are also constructed. For the case $\Delta=0$, however, such representations are not possible.

## C. Helicity States

The unstable particle states that produce resonances in scattering amplitudes all have a momentum $k$ that is related by a real Lorentz transformation to one of the K of (5.12). All three possibilities in (5.12) have the form

$$
\begin{align*}
\overrightarrow{\mathrm{k}} & =(\alpha, 0,0, \beta)  \tag{5.14}\\
\mathscr{M}^{2} & =\alpha^{2}-\beta^{2}
\end{align*}
$$

Define a new state by

$$
\begin{equation*}
\left|\mathrm{k}, \mathrm{j}, \sigma>\equiv \mathrm{U}\left[\mathrm{~L}(\mathrm{k}, \mathrm{k}) \mathrm{L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right]\right| \Psi(\mathcal{M}) \mathrm{j}, \mathrm{j}_{3}>\left.\sqrt{2 \mathscr{M}}\right|_{\mathrm{j}_{3}=\sigma} \tag{5.15}
\end{equation*}
$$

It should be noted that although (5.15) is very different from (5.6) the only notational distinction will be the use of an index $\sigma$ (later related to helicity) rather than a $\mathrm{j}_{3}$. In spite of the two stage boost in (5.15) it is easy to check irreducibility by

$$
\begin{equation*}
[\overrightarrow{\mathrm{S}}(\mathrm{k})]^{2}\left|\mathrm{k}, \mathrm{j}, \sigma>=\mathrm{U}\left[\mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right] \overrightarrow{\mathrm{J}}^{2} \mathrm{U}[\mathrm{X}]\right| \Psi(\mathscr{M}) \mathrm{j}, \mathrm{j}_{3}>\left.\sqrt{2 \mathscr{M}}\right|_{\mathrm{j}_{3}=\sigma} \tag{5.16}
\end{equation*}
$$

where

$$
\mathrm{X} \equiv \mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})^{-1} \mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}}) \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{~K}}, \mathscr{M})
$$

Because X leaves the four-vector $\mathscr{M}$ invariant it must be some complex rotation. Therefore

$$
\overrightarrow{\mathrm{J}}^{2} \mathrm{U}[\mathrm{X}]=\mathrm{U}[\mathrm{X}] \overrightarrow{\mathrm{J}}
$$

so that (5.16) becomes

$$
\begin{equation*}
[\overrightarrow{\mathrm{S}}(\mathrm{k})]^{2}|\mathrm{k}, \mathrm{j}, \sigma\rangle=\mathrm{j}(\mathrm{j}+1)|\mathrm{k}, \mathrm{j}, \sigma\rangle \tag{5.17}
\end{equation*}
$$

The behavior of (5.15) under a real Lorentz transformation $\Lambda$ follows from

$$
\begin{equation*}
\mathrm{U}[\Lambda]|\mathrm{k}, \mathrm{j}, \sigma\rangle=\mathrm{U}\left[\mathrm{~L}(\Lambda \mathrm{k}, \overline{\mathrm{k}}) \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})\right] \mathrm{U}[\overline{\mathrm{R}}] \mid \Psi(\mathscr{M}) \mathrm{j}, \sigma>\sqrt{2 \mathscr{M}} \tag{5.18}
\end{equation*}
$$

where

$$
\begin{equation*}
\overline{\mathrm{R}} \equiv \mathrm{~L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})^{-1}\left\{\mathrm{~L}(\Lambda \mathrm{k}, \overline{\mathrm{k}})^{-1} \Lambda \mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}})\right\} \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M}) \tag{5.19}
\end{equation*}
$$

Again $\overline{\mathrm{R}}$ leaves $\mathscr{M}^{\mu}$ invariant and must therefore be a complex rotation at worst. Closer inspection shows that the quantity in braces in (5.19) is a real Lorentz transformation that leaves k of (5.15) invariant. Therefore it can only be a rotation about the z -axis:

$$
\begin{equation*}
\mathrm{U}\left[\mathrm{~L}(\Lambda \mathrm{k}, \overline{\mathrm{k}})^{-1} \Lambda \mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}})\right]=\mathrm{e}^{\mathrm{i} \theta W^{\mathrm{J}} 3} \tag{5.20}
\end{equation*}
$$

where $\theta_{\mathrm{W}}$ is a real Wigner rotation angle that depends on k and $\Lambda$. Now $L^{c}(\overline{\mathrm{k}}, \mathscr{M})$ is just a boost along $\hat{z}$. Because

$$
\left[\mathrm{J}_{3}, \mathrm{~K}_{3}\right]=0
$$

it follows that $\overline{\mathrm{R}}$ itself is simply a rotation around the z -axis:

$$
\mathrm{U}[\overline{\mathrm{R}}]=\mathrm{e}^{\mathrm{i} \theta_{\mathrm{W}} \mathrm{~J}_{3}}
$$

Therefore (5.18) becomes

$$
\left.\mathrm{U}[\Lambda]\left|\mathrm{k}, \mathrm{j}, \sigma>=\mathrm{U}\left[\mathrm{~L}(\Lambda \mathrm{k}, \overline{\mathrm{k}}) \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})\right] \mathrm{e}^{\mathrm{i} \theta} \mathrm{~W}^{\mathrm{J}} 3\right| \Psi(\mathscr{M}) \mathrm{j}, \sigma\right\rangle \sqrt{2 \mathscr{M}}
$$

Because $|\Psi\rangle$ is an eigenstate of $J_{3}$ this collapses to

$$
\begin{equation*}
\mathrm{U}[\Lambda]|\mathrm{k}, \mathrm{j}, \sigma>=| \Lambda \mathrm{k}, \mathrm{j}, \sigma>\mathrm{e}^{\mathrm{i} \theta_{\mathrm{W}} \mathrm{~W}^{\sigma}} \tag{5.21}
\end{equation*}
$$

This is the analogue of (5.11). The transformation law for $\mathrm{J}_{3}$ states in (5.11) was more complicated just because those states are an awkward linear combination of the helicity states. The precise combination is

$$
\begin{equation*}
\left|\mathrm{k}, \mathrm{j}, \mathrm{j}_{3}>-\sum_{\sigma}\right| \mathrm{k}, \mathrm{j}, \sigma>\mathrm{D}_{\sigma, \mathrm{j}_{3}}{ }^{(\mathrm{Y})} \tag{5.22}
\end{equation*}
$$

where $Y$ is another complex rotation given by

$$
\begin{equation*}
\mathrm{Y} \equiv \mathrm{~L}^{\mathrm{c}}(\mathscr{M}, \overline{\mathrm{k}}) \mathrm{L}(\overline{\mathrm{k}}, \mathrm{k}) \mathrm{L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M}) \tag{5.23}
\end{equation*}
$$

The transformation law (5.21) is exactly the same as for massless particles because in both cases there is no real Lorentz transformation that takes the three-momentum to zero. The only difference is that here $\sigma$ may take on all integrally spaced values from -j to +j rather than just the two extreme values allowed in the massless case. It should be emphasized that the restriction of $j$ to integers or half-integers comes from the Poincaré group itself in Section IV.B. The construction of $\left|\phi, \mathrm{j}^{2} \mathrm{j}_{3}\right\rangle$ and then $\left|\Psi, \mathrm{j}, \mathrm{j}_{3}\right\rangle$ by combining field indices (as discussed in Appendix A) merely capitalized on that result.

Calling (5.15) helicity states is as yet unjustified. The real justification is the demonstration in Section VI. A that these states occur as resonances in helicity amplitudes. At the present this nomenclature may even seem inappropriate. Clearly (5.15) is an eigenstate of $\overrightarrow{\mathscr{P}} \cdot \vec{J}$ when the momentum is $\overline{\mathrm{k}}$ and hence along $\hat{z}$ :

$$
\begin{equation*}
\mathrm{J}_{3}|\overline{\mathrm{k}}, \mathrm{j}, \sigma\rangle=\sigma|\overline{\mathrm{k}}, \mathrm{j}, \sigma\rangle \tag{5.24}
\end{equation*}
$$

For arbitrary complex $\vec{k}$, however, it is not likely that (5.15) will be an eigenstate of $\overrightarrow{\mathscr{P}} \cdot \vec{J}$ as the massless case might suggest. The correct generalization of the helicity operator is obtained from boosting (5.24) to get

$$
\begin{equation*}
\sum(\mathrm{k})|\mathrm{k}, \mathrm{j}, \sigma>=\sigma| \mathrm{k}, \mathrm{j}, \sigma>, \tag{5.25}
\end{equation*}
$$

where

$$
\begin{equation*}
\sum(\mathrm{k}) \equiv \mathrm{U}[\mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}})] \mathrm{J}_{3} \mathrm{U}^{-1}[\mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}})] \tag{5.26}
\end{equation*}
$$

An alternative expression for this operator that will be derived shortly is

$$
\begin{equation*}
\sum(\mathrm{k})=\frac{\epsilon_{\alpha \beta \mu \nu} \mathrm{p}^{\alpha} \mathrm{q}^{\beta} \mathrm{M}^{\mu \nu}}{2 \vee \overline{\left(\mathrm{p} \cdot \mathrm{q}^{2}-\mathrm{p}^{2} \mathrm{q}^{2}\right.}} \tag{5.27}
\end{equation*}
$$

This helicity operator is manifestly a Lorentz scalar:

$$
\begin{equation*}
\mathrm{U}[\Lambda] \sum(\mathrm{k}) \mathrm{U}^{-1}[\Lambda]=\sum(\Lambda \mathrm{k}) \tag{5.28}
\end{equation*}
$$

Furthermore, whenever $\vec{p}$ and $\vec{q}$ are parallel so that $k$ has the special form

$$
\begin{equation*}
\mathrm{k}_{\mathrm{S}} \equiv\left(\mathrm{p}^{\mathrm{o}}+\mathrm{iq}^{\mathrm{o}}, \hat{\mathrm{n}}|\overrightarrow{\mathrm{p}}| \pm \hat{\mathrm{in}}|\overrightarrow{\mathrm{q}}|\right) \tag{5.29}
\end{equation*}
$$

then (5.27) reduces to

$$
\sum\left(\mathrm{k}_{\mathrm{S}}\right)=\hat{\mathrm{n}} \cdot \overrightarrow{\mathrm{~J}} .
$$

This result at present is only a plausibility argument for the name helicity but it will be very important in Section VI.A.

Naturally $\sum(\mathrm{k})$ is just the generator which leaves a general complex vector k invariant. As an example take

$$
\begin{aligned}
& p=(a, 0,0, b) \\
& q=(c, 0, d, 0) .
\end{aligned}
$$

Then p is left invariant by the three generators

$$
J_{3}, \quad b K_{1}+a J_{2}, \quad a J_{1}-b K_{2}
$$

just as in (4.15). On the other hand, $q$ is left invariant by

$$
\mathrm{J}_{2}, \quad \mathrm{cJ}_{3}-\mathrm{dK}_{1}, \quad \mathrm{cJ}_{1}+\mathrm{dK}_{3}
$$

There is, however, only one combination of generators that leaves both $p$ and $q$ invariant. It is given by

$$
\begin{aligned}
& \epsilon_{\alpha \beta \nu} \mathrm{p}^{\alpha} \mathrm{q}^{\beta} \mathrm{M}^{\mu \nu}=\mathrm{bcJ}_{3}-\mathrm{d}\left(\mathrm{bK}_{1}+\mathrm{aJ}_{2}\right) \\
&=\mathrm{b}\left(\mathrm{cJ}_{3}-\mathrm{dK}_{1}\right)-\mathrm{adJ} \\
& 2
\end{aligned} .
$$

This combination of generators is grouped in two different ways to show how it leaves both p and q invariant.

Now to show that (5.27) is correct, explicitly carry out the transformations in (5.26) to get

$$
\begin{equation*}
\sum(\mathrm{k})=\mathrm{x}^{\mu \nu}(\mathrm{k}, \overline{\mathrm{k}}) \mathrm{M}_{\mu \nu} \tag{5.30}
\end{equation*}
$$

where

$$
x^{\mu \nu}(\mathrm{k}, \overline{\mathrm{k}}) \equiv-\frac{1}{2} \mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}})_{\mathrm{a}}^{\mu} \mathrm{L}(\mathrm{k}, \overline{\mathrm{k}})_{\mathrm{b}}^{\nu} \epsilon^{\mathrm{ab} 3} .
$$

Apply a real Lorentz transformation $\Lambda$ :

$$
\Lambda_{\mu}^{\zeta} \Lambda_{\nu}^{\eta} \mathrm{x}^{\mu \nu}(\mathrm{k}, \overline{\mathrm{k}})=-\frac{1}{2}[\mathrm{~L}(\Lambda \mathrm{k}, \overline{\mathrm{k}}) \mathrm{R}]_{\mathrm{a}}^{\zeta}[\mathrm{L}(\Lambda \mathrm{k}, \overline{\mathrm{k}}) \mathrm{R}]_{\mathrm{b}}^{\eta} \epsilon^{\mathrm{ab} 3}
$$

where

$$
\begin{equation*}
R \equiv L^{-1}(\Lambda k, \bar{k}) \Lambda L(k, \bar{k}) \tag{5.31}
\end{equation*}
$$

Obviously $R$ is a real transformation that leaves $\bar{k}$ invariant. Hence it is only a rotation about the $z$-axis. This means

$$
\mathrm{R}_{\mathrm{a}}^{\mathrm{a}^{\prime}} \mathrm{R}_{\mathrm{b}}^{\mathrm{b}^{\prime}} \epsilon^{\mathrm{ab} 3}=\epsilon^{\mathrm{a}^{\prime} \mathrm{b}^{\prime} 3}
$$

Therefore (5.31) is just

$$
\begin{equation*}
\Lambda_{\mu}^{\zeta} \Lambda_{\nu}^{\eta} \mathrm{x}^{\mu \nu}(\mathrm{k}, \overline{\mathrm{k}})=\mathrm{x}^{\xi \eta}(\Lambda \mathrm{k}, \overline{\mathrm{k}}) \tag{5.32}
\end{equation*}
$$

This means that $x^{\mu \nu}$ is a tensor function of k independently of $\overline{\mathrm{k}}$. Furthermore, from the definition (5.30) it also satisfies

$$
\begin{gathered}
x^{\mu \nu}=-x^{\nu \mu} \\
\mathrm{p}_{\mu} \mathrm{x}^{\mu \nu}=\mathrm{q}_{\mu} \mathrm{x}^{\mu \nu}=0 .
\end{gathered}
$$

The only such tensor is

$$
\mathrm{x}^{\mu \nu}(\mathrm{k})=\mathrm{c} \epsilon^{\alpha \beta \mu \nu} \mathrm{p}_{\alpha} \mathrm{q}_{\beta}
$$

The constant c may be evaluated in a particular frame (like $\mathrm{k}=\mathrm{k}$ ) and gives (5.27) as claimed.

Note that this argument fails, as it should, for stable particles. Then the $R$ in (5.31) that leaves a real vector $\bar{p}$ invariant will contain boosts, though in a special combination like (4.15). Therefore (5.32) fails for massive stable particles and their helicity operator cannot be Lorentz invariant.
D. Covariance of Inner Products and Norms

Both the $\mathrm{J}_{3}$ states and the helicity states are related to $|\Psi(\mathscr{M})\rangle$ by complex, nonunitary boosts. The inner product of either with itself therefore might not be covariant. However, because $H$ is hermitian the inner product vanishes and hence is trivially covariant.

The complex conjugate partner of (5.6) is

$$
\begin{equation*}
\left|\mathrm{k}^{*}, \mathrm{j}, \mathrm{j}_{3}>\equiv \mathrm{U}\left[\mathrm{~L}^{\mathrm{c}}\left(\mathrm{k}^{*}, \mathscr{M}^{*}\right)\right]\right| \Psi\left(\mathscr{M}^{*}\right) \mathrm{j}, \mathrm{j}_{3}>\sqrt{2 \mathscr{M}^{*}} \tag{5.33}
\end{equation*}
$$

and similarly for helicity states. Because the Lorentz generators $\mathrm{M}^{\mu \nu}$ are hermitian

$$
\left[\mathrm{e}^{\mathrm{i} \lambda_{\mu \nu}^{*} \mathrm{M}^{\mu \nu}}\right]^{\dagger}=\mathrm{e}^{-\mathrm{i} \lambda_{\mu \nu} \mathrm{M}^{\mu \nu}}
$$

or equivalently

$$
\mathrm{U}[\mathrm{~L} *]^{\dagger}=\mathrm{U}[\mathrm{~L}]^{-1}
$$

This means that the inner product of (5.33) with its partner (5.6) is

$$
\begin{equation*}
<\mathrm{j}_{3}, \mathrm{j}, \mathrm{k}^{*} \mid \mathrm{k}^{\prime}, \mathrm{j}^{\prime}, \mathrm{j}_{3}^{\prime}>=2 \mathrm{k}^{\mathrm{o}} \delta^{3}(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}) \delta_{\mathrm{j}, \mathrm{j}^{\prime}} \delta_{\mathrm{j}_{3}, \mathrm{j}_{3}^{\prime}} \tag{5.34}
\end{equation*}
$$

The covariance of this product may be explicitly verified. From the definition (4. 19)

$$
\mathrm{L}^{\mathrm{c}\left(\mathrm{k}^{\prime *}, \mathrm{k}^{*}\right)=\mathrm{L}^{\mathrm{c}}\left(\mathrm{k}^{\prime}, \mathrm{k}\right)^{*} .}
$$

Consequently the transformation law of (5.33) is

$$
\mathrm{U}[\Lambda]\left|\mathrm{k}^{*}, \mathrm{j}, \mathrm{j}_{3}>=\sum_{\mathrm{j}_{3}^{\mathrm{j}}}\right| \Lambda \mathrm{k}^{*}, \mathrm{j}_{3} \mathrm{j}_{3}^{\mathrm{j}}>\mathrm{D}_{\mathrm{j}_{3}^{\prime}, \mathrm{j}_{3}}\left(\mathrm{R}_{\mathrm{W}}^{\mathrm{c} *}\right)
$$

where $R_{W}^{c *}$ is just the complex conjugate of the rotation occurring in (5.11). Because

$$
\left[\mathrm{D}_{\mathrm{j}_{3}^{\prime}, \mathrm{j}_{3}}\left(\mathrm{R}^{*}\right)\right]^{*}=\mathrm{D}_{\mathrm{j}_{3}, \mathrm{j}_{3}^{\prime}}\left(\mathrm{R}^{-1}\right)
$$

the inner product (5.34) is genuinely covariant.
For helicity states this check is easier. The transformation law for the conjugate state is

$$
\mathrm{U}[\Lambda]\left|\mathrm{k}^{*}, \mathrm{j}, \sigma\right\rangle=\mid \Lambda \mathrm{k}^{*}, \mathrm{j}, \sigma>\mathrm{e}^{\mathrm{i} \theta} \mathrm{~W}^{\sigma},
$$

where $\theta_{\mathrm{W}}$ is the same real angle as in (5.21). The covariance of the inner product just amounts to

$$
\left[e^{\mathrm{i} \theta} \mathrm{~W}^{\sigma}\right]^{*}=e^{-\mathrm{i} \theta} \mathrm{~W}^{\sigma}
$$

The momenta $\overrightarrow{\mathrm{k}}$ and $\overrightarrow{\mathrm{k}^{\prime}}$ in (5.34) are generally complex and the $\delta$-function must therefore be analytically continued. Moeller proposed this long ago. ${ }^{8}$ The continuation was thoroughly developed by Bremermann and Durand. ${ }^{24}$ The usual representation

$$
\delta\left(\mathrm{k}^{\prime}-\mathrm{k}\right)=\frac{-1}{2 \pi \mathrm{i}}\left\{\frac{1}{\left(\mathrm{k}^{\prime}+\mathrm{i} \epsilon\right)-\mathrm{k}}-\frac{1}{\left(\mathrm{k}^{\prime}-\mathrm{i} \epsilon\right)-\mathrm{k}}\right\},
$$

is valid for k and $\mathrm{k}^{\prime}$ real. For complex variables

$$
\begin{equation*}
\int_{c} d k^{\prime} f\left(k^{\prime}\right) \delta\left(k^{\prime}-k\right) \equiv \int_{c_{1}+c_{2}} d k^{\top} f\left(k^{\top}\right) \frac{-1}{2 \pi i}\left(\frac{1}{\overline{k^{\prime}-k}}\right) \tag{5.35}
\end{equation*}
$$

where $c_{1}$ is a contour parallel to c that passes above k and $\mathrm{c}_{2}$ is a contour antiparallel to c that passes below k .

## VI. WAVE FUNCTIONS AND S-MATRIX ELEMENTS

FOR UNSTABLE PARTICLES
The consequences of the transformation laws derived in Section $V$ for experiments carried out on the real axis will now be discussed. The wave functions for unstable particles and their use in calculating scattering amplitudes will be quite analagous to the usual stable particle results.

## A. How the Irreducible Representations are Produced in Scattering

Two points will be emphasized in this section: First, that the irreducible representations of Section V actually do occur as poles in partial wave amplitudes; Second, that the Lorentz invariance of the helicity in (5.21) is not experimentally observable. It will turn out, in fact, that measurements conducted in two different frames will find the $2 \mathrm{j}+1$ helicity components rotated according to the usual law for stable particles with $\mathrm{m} \neq 0$. In short, it is $(6.18)$ that is observed and not (5.21). (It goes without saying that the transformation law (5.11) for $\mathrm{J}_{3}$ states automatically reduces to that for stable particles when the momentum becomes real.)

For definiteness, consider the scattering of two stable particles that produce a resonance in the $s$ channel. Denote the scattering amplitude by

$$
\begin{equation*}
\left.\mathrm{T}(1+2 \rightarrow \mathrm{~A})=<\mathrm{A}|\mathrm{~T}(\mathrm{p})| \overrightarrow{\mathrm{p}}_{1}, \lambda_{1} ; \overrightarrow{\mathrm{p}}_{2}, \lambda_{2}\right\rangle \tag{6.1}
\end{equation*}
$$

where A is some multiparticle final state, $\lambda_{i}$ are the individual helicities, and $\mathrm{p}=\mathrm{p}_{1}+\mathrm{p}_{2}$. This amplitude may be decomposed into partial waves of definite total angular momentum by following the procedure of Jacob and Wick ${ }^{25}$ to construct eigenstates of $\vec{J}^{2}$ in the center-of-mass frame. These states must be further specified as eigenstates either of $J_{3}$ or of helicity. The partial wave amplitude in any frame is then obtained by boosting these two particle states. Denote the
partial wave amplitude with $s$ channel helicity $\sigma$ by

$$
\begin{equation*}
\mathrm{T}^{\mathrm{j}}(1+2 \rightarrow \mathrm{~A})=\langle\mathrm{A}| \mathrm{T}(\mathrm{p})|\mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle \tag{6.2}
\end{equation*}
$$

where

$$
\begin{aligned}
& \left|j_{1}-j_{2}\right| \leq j \leq\left|j_{1}+j_{2}\right| \\
& -j \leq \sigma \leq j
\end{aligned}
$$

The three labels on the two particle state in (6.2) denote the fact that

$$
\begin{align*}
& \mathscr{P}^{\mu}|\mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle=\mathrm{p}^{\mu}|\mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle  \tag{6.3a}\\
& {[\overrightarrow{\mathrm{S}}(\mathrm{p})]^{2}|\mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle=\mathrm{j}(\mathrm{j}+1)|\mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle}  \tag{6.3b}\\
& \hat{\mathrm{p}} \cdot \overrightarrow{\mathrm{~J}}|\mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle=\sigma|\mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle . \tag{6.3c}
\end{align*}
$$

A Lorentz transformation rotates the $2 \mathrm{j}+1$ values of $\sigma$ according to the usual law

$$
\begin{equation*}
\mathrm{U}[\Lambda]|\mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle=\sum_{\sigma^{1}} \mid \Lambda \mathrm{p}, \mathrm{j}, \sigma^{\prime} ; 1+2>\mathrm{D}_{\sigma^{\prime}, \sigma}(\widetilde{\mathrm{R}}) \tag{6.4}
\end{equation*}
$$

where

$$
\begin{aligned}
& \widetilde{R}=L\left(\tilde{p}^{\prime}, W\right)^{-1}\left\{R\left(p p^{\prime}, \tilde{p}^{\prime}\right)^{-1} \Lambda R(p, \tilde{p})\right\} L(\tilde{p}, W) \\
& W=\left(\sqrt{p^{2}}, 0,0,0\right) \\
& \widetilde{p}=\left(p^{0}, 0,0,|\vec{p}|\right) \\
& p^{\prime}=\Lambda p \\
& \tilde{p}^{\prime}=\left(p^{o}, 0,0,\left|\vec{p}^{\prime}\right|\right)
\end{aligned}
$$

It is, of course, obvious that the different helicity components must mix from the observation that the helicity operator in (6.3c) is not Lorentz invariant.

By hypothesis (6.1) has a resonance pole when analytically continued to the second sheet. The amplitude can be related to the Fourier transform of the

Green's function

$$
\langle 0| \mathrm{T}\left[\mathscr{A}(\mathrm{y}) \psi\left(\mathrm{x}_{1}\right) \bar{\psi}\left(\mathrm{x}_{2}\right)\right]|0\rangle
$$

Just as in Appendix A the pole can then come only from the term

$$
\mathrm{G}(\mathrm{p})=\langle 0| \mathscr{A}(\mathrm{y}) \frac{\mathrm{i}(2 \pi)^{3} \delta^{3}(\overrightarrow{\mathrm{p}}-\overrightarrow{\mathscr{P}})}{\mathrm{p}^{0}-\mathrm{H}} \mathrm{~T}[\psi(\mathrm{r}) \bar{\psi}(-\mathrm{r})]|0\rangle
$$

with $\mathrm{r}=\frac{1}{2}\left(\mathrm{x}_{1}-\mathrm{x}_{2}\right)$. Analytic continuation of $\mathrm{G}(\mathrm{p})$ will then yield a pole whose residue may be expressed in terms of state vectors for the unstable particle:

$$
\begin{equation*}
\left.\mathrm{G}(\mathrm{k}) \rightarrow \frac{1}{\mathrm{k}^{2}-\mathcal{M}^{2}} \sum_{\mathrm{j}^{\prime}, \sigma^{\prime}}<0|\mathscr{A}(\mathrm{y})| \mathrm{k}, \mathrm{j}^{\prime}, \sigma^{\prime}\right\rangle\left\langle\mathrm{k}^{*}, \mathrm{j}^{\prime}, \sigma^{\prime}\right| \mathrm{T}[\psi(\mathrm{r}) \bar{\psi}(-\mathrm{r})]|0\rangle \tag{6.5}
\end{equation*}
$$

The unstable particle states in (6.5) have been chosen as helicity states:

$$
\begin{align*}
& \mathscr{P}^{\mu}\left|\mathrm{k}, \mathrm{j}^{\prime}, \sigma^{\prime}>=\mathrm{k}^{\mu}\right| \mathrm{k}, \mathrm{j}^{\prime}, \sigma^{\prime}>  \tag{6.6a}\\
& {\left[{\overline{\mathrm{s}}(\mathrm{k})]^{2}\left|\mathrm{k}, \mathrm{j}^{\prime}, \sigma^{\prime}>=\mathrm{j}^{\prime}\left(\mathrm{j}^{9}+1\right)\right| \mathrm{k}, \mathrm{j}^{\prime}, \sigma^{\prime}>}^{\left.\sum(\mathrm{k})\left|\mathrm{k}, \mathrm{j}^{\prime}, \sigma^{\prime}>=\sigma^{\prime}\right| \mathrm{k}, \mathrm{j}^{\prime}, \sigma^{\prime}\right\rangle}\right.} \tag{6.6b}
\end{align*}
$$

where

$$
\mathrm{k}^{\mu}=\mathrm{p}^{\mu}+\mathrm{iq}^{\mu}
$$

The contribution of (6.5) to the partial wave amplitude is obtained by taking $q \rightarrow 0$ in the residue of (6.5), projecting out the two particle formation amplitude $<\mathrm{p}, \mathrm{j}^{\mathrm{j}}, \sigma^{\dagger} \mid \overrightarrow{\mathrm{p}}_{1}, \lambda_{1} ; \overrightarrow{\mathrm{p}}_{2}, \lambda_{2}>$, and then combining spins to get

$$
\left.\mathrm{T}^{\mathrm{j}}(1+2 \rightarrow \mathrm{~A}) \approx \frac{1}{\mathrm{p}^{2}-\mathscr{M}^{2}} \sum_{\mathrm{j}^{\prime}, \sigma^{\prime}}<\mathrm{A}\left|\mathrm{p}, \mathrm{j}^{\prime}, \sigma^{\mathrm{r}}><\mathrm{p}, \mathrm{j}^{\prime}, \sigma^{\prime}\right| \mathrm{p}, \mathrm{j}, \sigma ; 1+2\right\rangle
$$

Now, the question of interest is which values of $j^{\prime}$ and $\sigma^{\prime}$ in (6.5) actually survive the partial wave projection in (6.7)? Comparing (6.6b) with (6.3b) shows that $\mathrm{j}^{\prime}=\mathrm{j}$ because

$$
\lim _{q \rightarrow 0} \vec{S}(k)=\vec{S}(p)
$$

(See (4.27) for example.) However the helicity operator

$$
\begin{equation*}
\Sigma(\mathrm{k})=\frac{\epsilon_{\alpha \beta \mu \nu} \mathrm{p}^{\alpha} \mathrm{q}^{\beta_{\mathrm{M}} \mu \nu}}{2 \sqrt{(\mathrm{p} \cdot \mathrm{q})^{2}-\mathrm{p}^{2} \mathrm{q}^{2}}} \tag{6.8}
\end{equation*}
$$

has the property that in general

$$
\lim _{q \rightarrow 0} \sum(k) \neq p \cdot \vec{J} .
$$

Only when $\vec{q}$ is taken to zero along the $\vec{p}$ direction does ( 6.6 c ) correspond to (6.3c), for then

$$
\begin{equation*}
\lim _{\overrightarrow{\mathrm{q}} \| \overrightarrow{\mathrm{p}}} \sum(\mathrm{k})=\mathrm{p} \cdot \overrightarrow{\mathrm{~J}} . \tag{6.9}
\end{equation*}
$$

Thus if $q$ is taken to zero along the plane $\vec{q} \| \vec{p}$ then $\sigma^{t}=\sigma$ in (6.7). Appendix $F$ shows furthermore that the only nonvanishing contribution to (6.7) comes from taking the q limit in this manner. Thus

$$
\begin{equation*}
\mathrm{T}^{\mathrm{j}}(1+2 \rightarrow \mathrm{~A}) \approx \frac{\langle\mathrm{A} \mid \mathrm{p}, \mathrm{j}, \sigma\rangle\langle\mathrm{p}, \mathrm{j}, \sigma \mid \mathrm{p}, \mathrm{j}, \sigma ; 1+2\rangle}{\mathrm{p}^{2}-\mathrm{M}^{2}+\mathrm{iM} \mathrm{\Gamma}} \tag{6.10}
\end{equation*}
$$

The discussion leading to (6.10) shows that the helicity states of Section V.C are actually produced in scattering. In spite of this, the Lorentz invariance of the unstable particle helicity is not observable because the limit required in (6.9) is not covariant. To see exactly what the observed transformation law for the helicity states will be, consider how to calculate (6.10) in two different frames. In the first frame let

$$
\begin{equation*}
\mathrm{p}=\left(\mathrm{p}^{0}, 0,0, \mathrm{p}^{3}\right) \tag{6.11a}
\end{equation*}
$$

Then the residue of the pole in (6.5) that contributes to (6.10) must have

$$
\begin{equation*}
q=\left(q^{0}, 0,0, q^{3}\right) \tag{6.11b}
\end{equation*}
$$

With $q$ in this form $q^{0}$ and $q^{3}$ may be taken to zero independently.

If the same experiment is viewed from a frame moving along the x -axis the observed momenta are

$$
\begin{align*}
& p^{\prime}=\left(p^{0} \cosh \lambda, p^{0} \sinh \lambda, 0, p^{3}\right)  \tag{6.12a}\\
& q^{\prime}=\left(q^{0} \cosh \lambda, q^{0} \sinh \lambda, 0, q^{3}\right) . \tag{6.12b}
\end{align*}
$$

But now $\overrightarrow{q^{1}}$ is not parallel to $\overrightarrow{\mathrm{p}}^{\mathbf{1}}$. To calculate the contribution of the unstable particle pole to the partial wave amplitude in this frame requires continuing $q{ }^{1}$ to zero in the plane of $\overrightarrow{p^{\prime}}$. Let $q^{\prime \prime}$ be this continued value of $q^{\prime}$ and let $p^{\prime \prime}=p^{\prime}$ be the unchanged value of the real momenta:

$$
\begin{align*}
& \mathrm{p}^{\prime \prime}=\left(\mathrm{p}^{0} \sinh \lambda, \mathrm{p}^{0} \sinh \lambda, 0, \mathrm{p}^{3}\right)=\mathrm{p}^{\prime}  \tag{6.13a}\\
& \mathrm{q}^{\prime \prime}=\left(\mathrm{q}^{\prime \prime},\left|\overrightarrow{q^{\prime \prime}}\right| \hat{\mathrm{u}}\right), \tag{6.13b}
\end{align*}
$$

where

$$
\hat{\mathrm{u}}=\frac{\left(\mathrm{p}^{0} \sinh \lambda, 0, \mathrm{p}^{3}\right)}{\left(\mathrm{p}^{0} \sinh \lambda\right)^{2}+\left(\mathrm{p}^{3}\right)^{2}}
$$

For (6.13) to correspond to the same value of mass and width as (6.11) it is necessary that

$$
\begin{equation*}
\left(p^{\prime \prime}+i q^{\prime \prime}\right)^{2}=(p+i q)^{2} . \tag{6.14}
\end{equation*}
$$

If (6.14) has solutions of the form (6.13) then $\mathrm{k}^{\prime \prime}$ and k can still be related by a Lorentz transformation that is complex. Solving (6.14) gives

$$
\begin{align*}
& q^{\prime \prime}=\frac{1}{p^{2}}\left\{p^{\prime o}(p \cdot q)-\left|\overrightarrow{p^{\prime}}\right| \sqrt{(p \cdot q)^{2}-p^{2} q^{2}}\right\}  \tag{6.15}\\
& \left|\vec{q}^{\prime \prime}\right|=\frac{1}{p^{2}}\left\{\left|\overrightarrow{p^{\prime}}\right|(p \cdot q)-p^{\prime} \quad \sqrt{(p \cdot q)^{2}-p^{2} q^{2}}\right\} .
\end{align*}
$$

Thus there is a complex Lorentz transformation $\Lambda^{c}$ for which

$$
p^{\prime \prime}+i q^{\prime \prime}=\Lambda^{c}(p+i q)
$$

Because $\mathrm{k}^{\prime \prime}$ is the appropriate momentum in the moving frame to produce a pole in the partial wave amplitude the observed transformation law will be determined by this complex Lorentz transformation. To find that law requires applying $\Lambda_{c}$ to the $\mid \mathrm{k}, \mathrm{j}, \sigma>$ helicity state in (6.5). The result is directly analagous to (5.18) and (5.19):

$$
\begin{equation*}
\mathrm{U}\left[\Lambda^{\mathrm{c}}\right]\left|\mathrm{k}, \mathrm{j}, \sigma>=\sum_{\sigma^{\mathfrak{p}}}\right| \Lambda^{c} \mathrm{k}, \mathrm{j}, \sigma^{\prime}>\mathrm{D}_{\sigma^{p}, \sigma}\left(\overline{\mathrm{R}}^{\mathrm{c}}\right) \tag{6.16}
\end{equation*}
$$

where

$$
\overline{\mathrm{R}}^{\mathrm{c}} \equiv \mathrm{~L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})^{-1}\left\{\mathrm{~L}\left(\mathrm{k}^{\prime \prime}, \overline{\mathrm{k}}\right)^{-1} \Lambda^{\mathrm{c}} \mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}})\right\} \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})
$$

Recall that in (5.19) $\overline{\mathrm{R}}$ is a real rotation around the z-axis because $\Lambda$ there is real. Here, however, $\Lambda^{c}$ is complex. As a result $\overline{\mathrm{R}}^{\mathrm{c}}$ is neither real nor a z rotation. Consequently all $2 \mathrm{j}+1$ helicity components are rotated in (6.16). The experimentally observed transformation law is the $q^{\prime \prime} \rightarrow 0$ limit of (6.16). Hence the helicity states will be observed to transform covariantly not invariantly. The precise form of this covariance is obtained by taking $q^{\prime \prime} \rightarrow 0$ in (6.16). It will shortly be shown that

$$
\begin{equation*}
\lim _{q^{\prime \prime} \rightarrow 0} \bar{R}^{c}=\widetilde{R} \tag{6.17}
\end{equation*}
$$

where $\widetilde{R}$ is just the usual real rotation (6.4) appropriate to stable helicity states. Thus the observed transformation law following from (6.16) is exactly like (6.4):

$$
\begin{equation*}
\mathrm{U}[\Lambda]|\mathrm{p}, j, \sigma\rangle=\sum_{\sigma^{\prime}} \mid \Lambda \mathrm{p}, j, \sigma^{\prime}>\mathrm{D}_{\sigma^{\prime}, \sigma}(\widetilde{\mathrm{R}}) \tag{6.18}
\end{equation*}
$$

It is, of course, essential that

$$
\begin{equation*}
\lim _{q \rightarrow 0} \bar{R}^{c}=\widetilde{R} \tag{6.19}
\end{equation*}
$$

also. This guarantees that the transformation law is symmetrical between the two observers.

To demonstrate (6.17) it is necessary to specify the transformation $L(\mathbb{k}, k)$ more carefully. Consider, for example, the case $q^{2}>0$. Then $\overline{\mathrm{K}}$ is given by (5.12a). A real Lorentz transformation from k to k may always to accomplished by first boosting $q$ to rest and then rotating into the $z$ axis. Thus

$$
L(\bar{k}, k)=R(\hat{z}, B(\bar{q}, q) p) B(\bar{q}, q)
$$

With the $p$ and $q$ of (6.11) the boost is in the $\hat{z}$ direction so the rotation is just the identity:

$$
\mathrm{L}(\overline{\mathrm{k}}, \mathrm{k})=\mathrm{B}_{\mathrm{z}}(\overline{\mathrm{q}}, \mathrm{q})
$$

On the other hand, for $p^{\prime \prime}$ and $q^{\prime \prime}$ in (6.13)

$$
\mathrm{L}\left(\overline{\mathrm{k}}, \mathrm{k}^{\prime \prime}\right)=\mathrm{R}(\hat{\mathrm{z}}, \hat{\mathrm{u}}) \mathrm{B}_{\hat{\mathrm{u}}}\left(\overline{\mathrm{q}}, \mathrm{q}^{\prime \prime}\right)
$$

Putting this together gives the complex rotation in (6.16) as

$$
\begin{equation*}
\overline{\mathrm{R}}^{\mathrm{c}}=\mathrm{B}_{\mathrm{z}}(\mathscr{M}, \overline{\mathrm{k}})\left\{\mathrm{R}(\hat{\mathrm{z}}, \hat{\mathrm{u}}) \mathrm{B}_{\hat{\mathrm{u}}}\left(\overline{\mathrm{q}}, \mathrm{q}^{\prime \prime}\right) \Lambda^{\mathrm{c}} \mathrm{~B}_{\mathrm{z}}(\mathrm{q}, \overline{\mathrm{q}})\right\} \mathrm{B}_{\mathrm{z}}(\overline{\mathrm{k}}, \mathscr{M}) \tag{6.20}
\end{equation*}
$$

It is convenient to rewrite (6.20) by using

$$
R(\hat{z}, \hat{u}) B_{\hat{u}}\left(\bar{q}, q^{\prime \prime}\right)=B_{z}\left(\bar{q}, q^{\prime \prime}\right) R(\hat{z}, \hat{u})
$$

to get

$$
\begin{equation*}
\overline{\mathrm{R}}^{\mathrm{c}}=\mathrm{B}_{\mathrm{z}}(\mathscr{M}, \overline{\mathrm{k}}) \mathrm{B}_{\mathrm{z}}\left(\overline{\mathrm{q}}, \mathrm{q}^{\prime \prime}\right) \mathrm{R}(\hat{\mathrm{z}}, \hat{\mathrm{u}}) \Lambda^{\mathrm{c}} \mathrm{~B}_{\mathrm{z}}(\mathrm{q}, \overline{\mathrm{q}}) \mathrm{B}_{\mathrm{z}}(\overline{\mathrm{k}}, \mathscr{M}) \tag{6.21}
\end{equation*}
$$

The boosts are explicitly given by

$$
\begin{aligned}
& \mathrm{B}_{\mathrm{z}}(\mathrm{~K}, \mathscr{M})=\mathrm{e}^{\mathrm{i} \mu \mathrm{~K}_{3}} \\
& \mathrm{~B}_{\mathrm{z}}(\mathrm{q}, \overline{\mathrm{q}})=\mathrm{e}^{\mathrm{i} \mu{ }^{\prime} \mathrm{K}_{3}} \\
& \mathrm{~B}_{\mathrm{z}}\left(\mathrm{q}^{\prime \prime}, \overline{\mathrm{q}}\right)=\mathrm{e}^{\mathrm{i} \mu^{\prime \prime} \mathrm{K}_{3}},
\end{aligned}
$$

where

$$
\begin{align*}
& \cosh \mu=\frac{\frac{\mathrm{p} \cdot \mathrm{q}}{\sqrt{q^{2}}}+\mathrm{i} \sqrt{q^{2}}}{\mathscr{M}} \\
& \cosh \mu^{\prime}=\frac{q^{\circ}}{\sqrt{q^{2}}}  \tag{6.22}\\
& \cosh \mu^{\prime \prime}=\frac{q^{\prime \prime}}{\sqrt{q^{2}}} .
\end{align*}
$$

(Note that $\mu^{\prime}$ and $\mu^{\prime \prime}$ are both real.) Thus (6.21) is

$$
\begin{equation*}
\overline{\mathrm{R}}^{\mathrm{c}}=\mathrm{e}^{-\mathrm{i}\left(\mu+\mu^{\prime}\right) \mathrm{K}_{3}} R \hat{R(\underline{z}, \hat{u}) \Lambda^{c} e^{i\left(\mu+\mu^{\prime}\right) K_{3}}} \tag{6.23}
\end{equation*}
$$

By using (6.11), (6.15), and (6.22) one finds that

$$
\begin{align*}
& \cosh \left(\mu+\mu^{\prime \prime}\right)=\frac{\mathrm{p}^{\prime \prime}+\mathrm{i} \mathrm{q}^{\prime \prime}}{\mathscr{M}}  \tag{6.24}\\
& \cosh \left(\mu+\mu^{\prime}\right)=\frac{\mathrm{p}^{\circ}+\mathrm{i} q^{\circ}}{\mathscr{M}}
\end{align*}
$$

Taking $\mathrm{q}^{\prime \prime} \rightarrow 0$ is now easy:

$$
\begin{aligned}
& \lim _{q^{\prime \prime} \rightarrow 0} \cosh \left(\mu+\mu^{\prime \prime}\right)=\frac{p^{\prime \prime}}{M}=\frac{p^{\prime}}{M} \\
& \lim _{q^{\prime \prime} \rightarrow 0} \cosh \left(\mu+\mu^{\prime}\right)=\frac{p^{o}}{M}
\end{aligned}
$$

Hence

$$
\begin{equation*}
\lim _{q^{\prime \prime} \rightarrow 0} \bar{R}^{c}=B\left(M, \tilde{p}^{\prime}\right) R(\hat{z}, \hat{u}) \Lambda B(\tilde{p}, M) \tag{6.25}
\end{equation*}
$$

This is precisely the rotation $\widetilde{R}$ given in (6.4) appropriate to these $p$ and $p^{\prime}$. Hence $\overline{\mathrm{R}}^{\mathrm{C}}$ reduces to $\widetilde{\mathrm{R}}$ as claimed in (6.17). It is also obvious from (6.24) that the same limit $\widetilde{R}$ is obtained if $q \rightarrow 0$ instead of $q^{\prime \prime}$.

## B. Elementary Unstable Particles

The indicated transformation laws apply to any unstable particle. It may or may not be associated with an elementary field occuring in the Lagrangian. In conventional field theories, for example, the $\mu$ has an elementary field but the $\pi$, being composed of quarks, does not. Thus a $\mu$ will be called elementary and a $\pi$ composite though both are unstable.

Since the $\mu$ is, in fact, the only candidate for an unstable particle with an elementary field, it will be treated explicitly here. The generalization to other spins is straightforward. The Feynman propagator for a $\mu$ field is

$$
\left.\mathrm{S}_{\alpha \beta}(\mathrm{k})=-\mathrm{i} \quad d^{4} \mathrm{x} e^{i \mathrm{k} \cdot \mathrm{x}}<0\left|\mathrm{~T}\left[\psi_{\alpha}(\mathrm{x}) \bar{\psi}_{\beta}(\overline{\mathrm{g}})\right]\right| 0\right\rangle .
$$

Performing the Fourier transforms gives

$$
\begin{aligned}
\mathrm{S}_{\alpha \beta}(\mathrm{k})= & \langle 0| \psi_{\alpha}(0) \frac{(2 \pi)^{3} \delta^{3} \overrightarrow{\mathrm{k}}-\overrightarrow{\mathscr{P})}}{\mathrm{k}^{0}-\mathrm{H}+\mathrm{i} \epsilon} \bar{\psi}_{\beta}(0)|0\rangle \\
& \left.-<0\left|\bar{\psi}_{\beta}(0) \frac{(2 \pi)^{3} \delta^{3} \overrightarrow{(\mathrm{k}}-\overrightarrow{\mathscr{P})}}{\mathrm{k}^{0}+\mathrm{H}-\mathrm{i} \epsilon} \psi_{\alpha}(0)\right| 0\right\rangle
\end{aligned}
$$

Continuing in $k^{\circ}$ around the positive energy cuts to the second sheet pole gives

$$
\begin{equation*}
\mathrm{S}_{\alpha \beta}(\mathrm{k}) \underset{\mathrm{k}^{2} \rightarrow \mathscr{M}^{2}}{ } \sum_{\mathrm{j}_{3}}(2 \pi)^{3} \frac{{ }^{\langle 0| \psi} \alpha^{(0)\left|\mathrm{k}, \mathrm{j}_{3}><\mathrm{k}^{*}, \mathrm{j}_{3}\right| \bar{\psi}_{\beta^{(0)}}(0)|0\rangle}}{\mathrm{k}^{2}-\mathcal{M}^{2}} \tag{6.26}
\end{equation*}
$$

In (6.26) a sum over $J_{3}$ states $\left(j_{3}= \pm 1 / 2\right)$ has been used. The sum could just as well have been over helicity states.

The $J_{3}$ wave functions will be discussed first. Just from the definition (5.6) of the $J_{3}$ states the residue must be

$$
\left.<0\left|\psi_{\alpha}(0)\right| \mathrm{k}, \mathrm{j}_{3}\right\rangle=\sqrt{\frac{\sqrt{\mathscr{M}}}{(2 \pi)^{3}}} \mathrm{u}_{\alpha}\left(\mathrm{k}, \mathrm{j}_{3}\right)
$$

where z is some constant and the dimensionless functions $u$ are defined by

$$
\begin{align*}
& \mathrm{u}_{\alpha}(\mathrm{k}, 1 / 2) \equiv \mathrm{S}_{\alpha, 1}\left[\mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right] \\
& \mathrm{u}_{\alpha}(\mathrm{k},-1 / 2) \equiv \mathrm{S}_{\alpha, 2}\left[\mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right] \tag{6.27}
\end{align*}
$$

Here $S[\Lambda]$ denotes the usual $(1 / 2,0) \oplus(0,1 / 2)$ representation of the Lorentz group. ${ }^{20}$ A real Lorentz transformation applied to (6.27) gives

$$
\begin{equation*}
S[\Lambda] u\left(k, j_{3}\right)=\sum_{j_{3}^{\prime}} u\left(\Lambda k, j_{3}^{\prime}\right) S_{j_{3}^{\prime}}, j_{3}\left[R_{W}^{c}\right] \tag{6.28}
\end{equation*}
$$

where $R_{W}^{c}$ is the same complex rotation as (5.9). Define adjoint spinors as

$$
\begin{equation*}
\overline{\mathrm{u}}_{\alpha}\left(\mathrm{k}, \mathrm{j}_{3}\right) \equiv\left(\gamma^{o} \mathrm{u}\left(\mathrm{k}, \mathrm{j}_{3}\right)\right)_{\alpha}^{*} \tag{6.29}
\end{equation*}
$$

The completeness and normalization conditions are

$$
\begin{align*}
& \sum_{j_{3}=-1 / 2}^{1 / 2} u_{\alpha}\left(\mathrm{k}, \mathrm{j}_{3}\right) \bar{u}_{\beta}\left(\mathrm{k}^{*}, \mathrm{j}_{3}\right)=\frac{1}{2 \mathscr{M}}\left(\gamma^{\mu} \mathrm{k}_{\mu}+\mathscr{M}\right)_{\alpha \beta}  \tag{6.30a}\\
& \sum_{\alpha=1}^{4} \overline{\mathrm{u}}\left(\mathrm{k}^{*}, \mathrm{j}_{3}\right) \mathrm{u}_{\alpha}\left(\mathrm{k}, \mathrm{j}_{3}^{\prime}\right)=\delta_{\mathrm{j}_{3}, \mathrm{j}_{3}^{\prime}} \tag{6.30b}
\end{align*}
$$

Using (6.30a) in (6.26) gives

$$
\begin{equation*}
\mathrm{S}_{\alpha \beta}(\mathrm{k}) \underset{\mathrm{k}^{2} \rightarrow \mathscr{M}^{2}}{\longrightarrow} \mathrm{z} \frac{\gamma^{\mu} \mathrm{k}_{\mu}+\mathscr{M}}{\mathrm{k}^{2}-\mathscr{M}^{2}} \tag{6.31}
\end{equation*}
$$

as might be expected.
The location of k and $\mathrm{k}^{*}$ in (6.30) is important. For example, (6.30b) is clearly a reflection of the normalization

$$
<\mathrm{j}_{3}, \mathrm{k}^{*}\left|\mathrm{k}^{\mathrm{r}}, \mathrm{j}_{3}^{\prime}\right\rangle=2 \mathrm{k}^{\mathrm{o}} \delta^{3}\left(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}^{\mathrm{\prime}}\right) \delta_{\mathrm{j}_{3}}, \mathrm{j}_{3}^{\prime}
$$

Note however that

$$
\left\langle\mathrm{j}_{3}, \mathrm{k} \mid \mathrm{k}^{\prime}, \mathrm{j}_{3}^{\prime}\right\rangle=0
$$

has no finite dimensional counterpart:

$$
\begin{equation*}
\sum_{\alpha=1}^{4} \bar{u}_{\alpha}\left(\mathrm{k}, \mathrm{j}_{3}\right) \mathrm{u}_{\alpha}\left(\mathrm{k}, \mathrm{j}_{3}^{\imath}\right) \neq 0 \tag{6.32}
\end{equation*}
$$

In fact, the left hand side of (6.32) is nothing particularly simple: Nor is it needed in calculations.

The spinors may be explicitly displayed by choosing a particular complex Lorentz transformation for (6.27). For example, let this be a pure boost from . $M$ to k . The direction and magnitude of the boost are then given by

$$
\begin{align*}
& \hat{u}=\frac{\vec{p}+\overrightarrow{i q}}{\sqrt{\vec{p}^{2}-\vec{q}^{2}+2 i \vec{p} \cdot \vec{q}}}  \tag{6.33}\\
& \cosh \lambda=\frac{p^{o}+i q^{o}}{\mathscr{N}}
\end{align*}
$$

The $4 \times 4$ spinor representation of this boost is

$$
\begin{align*}
\mathrm{S}\left[\mathrm{~L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right] & =\mathrm{e}^{\frac{\lambda}{2} \hat{\mathrm{u}} \cdot \vec{\alpha}} \\
& =\cosh \frac{\lambda}{2}+\hat{u} \cdot \vec{\alpha} \sinh \frac{\lambda}{2} \tag{6.34}
\end{align*}
$$

where

$$
\vec{\alpha}=\gamma^{0} \vec{\gamma}
$$

This gives

$$
\begin{align*}
& \mathrm{u}_{\alpha}(\mathrm{k}, 1 / 2)=\frac{1}{\sqrt{2 \mathscr{M}\left(\mathrm{k}^{0}+\mathscr{M}\right)}}\left(\gamma^{\mu} \mathrm{k}_{\mu}+\mathscr{M}\right)_{\alpha, 1}  \tag{6.35}\\
& \mathrm{u}_{\alpha}(\mathrm{k},-1 / 2)=\frac{1}{\sqrt{2 \mathscr{M}\left(\mathrm{k}^{0}+\mathscr{M}\right)}}\left(\gamma^{\mu} \mathrm{k}_{\mu}+\mathscr{M}\right)_{\alpha, 2} .
\end{align*}
$$

Obviously as $q^{\mu} \rightarrow 0$ these spinors go smoothly into the usual spinors for real energy momentum.

The helicity wave functions are much more interesting. Let

$$
\left.<0\left|\psi_{\alpha}(0)\right| \mathrm{k}, \sigma\right\rangle=\sqrt{\frac{2 \mathscr{M} \mathrm{z}}{(2 \pi)^{3}}} \mathrm{u}_{\alpha}(\mathrm{k}, \sigma),
$$

where $\sigma= \pm$ to distinguish from $\mathrm{j}_{3}= \pm 1 / 2$. From the definition (5.15) of the helicity states it follows that

$$
\begin{align*}
& \mathrm{u}_{\alpha}(\mathrm{k},+) \equiv \mathrm{S}_{\alpha, 1}\left[\mathrm{~L}\left(\mathrm{k}, \overline{\mathrm{k})} \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})\right]\right. \\
& \mathrm{u}_{\alpha}(\mathrm{k},-) \equiv \mathrm{S}_{\alpha, 2}\left[\mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}}) \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})\right] . \tag{6.36}
\end{align*}
$$

The adjoint spinors are defined just as in (6.29). Completeness and normalization are then the same as (6.30) with $\mathrm{j}_{3}$ replaced by $\sigma$. The novel feature of these spinors is their transformation law. Write (6.36) as

$$
\mathrm{u}(\mathrm{k}, \pm)=\mathrm{S}[\mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}})] \mathrm{u}(\overline{\mathrm{k}}, \pm) .
$$

Then

$$
\begin{equation*}
\mathrm{S}[\Lambda] \mathrm{u}(\mathrm{k}, \pm)=\mathrm{S}[\mathrm{~L}(\Lambda \mathrm{k}, \overline{\mathrm{k}})] \mathrm{S}[\overline{\mathrm{R}}] \mathrm{u}(\overline{\mathrm{k}}, \pm), \tag{6.37}
\end{equation*}
$$

where

$$
\begin{equation*}
\overline{\mathrm{R}}=\mathrm{L}(\Lambda \mathrm{k}, \overline{\mathrm{k}})^{-1} \Lambda \mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}}) . \tag{6.38}
\end{equation*}
$$

Because $\overline{\mathrm{R}}$ is a real transformation that leaves $\overline{\mathrm{K}}$ invariant it can only be a rotation around z just as in Section V.C. Thus

$$
\begin{equation*}
\mathrm{S}[\overline{\mathrm{R}}]=\mathrm{e}^{\mathrm{i} \theta_{\mathrm{W}} \frac{1}{2} \sigma_{3}} \tag{6.39}
\end{equation*}
$$

where $\sigma_{3}$ is the usual Pauli matrix. This gives for (6.37)

$$
\begin{equation*}
\mathrm{S}[\Lambda] \mathrm{u}(\mathrm{k}, \pm)=\mathrm{u}(\Lambda \mathrm{k}, \pm) \mathrm{e}^{ \pm \frac{1}{2} \theta} \mathrm{~W} . \tag{6.40}
\end{equation*}
$$

Note that (6.40) holds even when $\mathrm{q} \rightarrow 0$. This is because $\mathrm{L}(\mathrm{k}, \overline{\mathrm{k}})$ does not reduce to the identity even when $q \rightarrow 0$.

## C. Composite Unstable Particles

By definition a composite unstable particle is one which does not produce a pole in any two-point function. It does, of course, produce poles in Green's functions with more lines. Thus in quark theories all hadrons are composite. All of these except the proton are unstable when weak and electromagnetic interactions are included.

Composite particles that are stable produce a pole in a Green's function when continued in the total energy-momentum onto the real axis below the production threshold. Composite particles that are unstable produce poles when the continuation is onto the second sheet. The residue of the pole is a product of Bethe-Salpeter wave functions like

$$
\begin{equation*}
\chi_{\mathrm{k}}\left(\mathrm{r}_{1}, \ldots, \mathrm{r}_{\mathrm{n}}\right) \equiv\langle 0| \mathrm{T}\left[\psi\left(\mathrm{r}_{1}\right) \ldots \psi\left(\mathrm{r}_{\mathrm{n}}\right)\right]|\mathrm{k}, \mathrm{j}, \sigma\rangle \tag{6.41a}
\end{equation*}
$$

The only difference between (6.41a) and the usual Bethe-Salpeter wave functions is that here k is complex and the state $|\mathrm{k}, \mathrm{j}, \sigma\rangle$ does not lie in the physical Hilbert space. The Lorentz transformation property of (6.41a) follows from (5.21). Denote the adjoint wave function by

$$
\begin{equation*}
\bar{\chi}_{\mathrm{k}}\left(\mathrm{r}_{\mathrm{n}}, \ldots, \mathrm{r}_{1}\right)=\langle\sigma, \mathrm{j}, \mathrm{k}| \mathrm{T}\left[\bar{\psi}\left(\mathrm{r}_{\mathrm{n}}\right) \ldots \bar{\psi}\left(\mathrm{r}_{1}\right)\right]|0\rangle . \tag{6.41b}
\end{equation*}
$$

As always, the complex conjugate of (6.41a) involves antitime ordering and is not simply related to (6.41b). Note, of course, that either of (6.41) may also have momentum $\mathrm{k}^{*}$.

The momentum space wave functions are real Fourier transforms of (6.41) in spite of the fact that $k$ is complex. As Appendix A illustrates, if the Green's function depends on ( $x_{1}, \ldots, x_{n}$ ) and ( $x_{1}^{1}, \ldots, x_{n}^{1}$ ) the pole occurs in the Fourier transform with respect to the average coordinate X given by

$$
\begin{equation*}
x_{i}=X+r_{i}, \quad \sum_{1}^{n} r_{i}=0 \tag{6.42a}
\end{equation*}
$$

Therefore only the relative coordinates $r_{i}$ appear in the residue factors like (6.41). If $k_{i}$ are the momenta conjugate to $x_{i}$ then it is convenient to define

$$
\begin{equation*}
\mathrm{k}_{\mathrm{i}}=\frac{1}{\mathrm{n}} \mathrm{k}+\ell_{\mathrm{i}}, \quad \sum_{1}^{\mathrm{n}} \ell_{\mathrm{i}}=0 \tag{6.42b}
\end{equation*}
$$

The total momentum k is then conjugate to the average position X and the relative momenta $l_{i}$ are conjugate to the relative coordinates $r_{i}$ :

$$
\begin{equation*}
\sum_{1}^{n} k_{i} \cdot x_{i}=k \cdot x+\sum_{1}^{n} \ell_{i} \cdot r_{i} \tag{6.43}
\end{equation*}
$$

Even though the $\mathrm{k}_{\mathrm{i}}$ are complex all the $\ell_{i}$ may be taken as real because of (6.43). The momentum space wave functions are then

$$
\begin{align*}
& x_{k}\left(\ell_{1}, \ldots, \ell_{n}\right)=\int(d r) x_{k}\left(r_{1}, \ldots, r_{n}\right) e^{i \sum \ell_{i} \cdot r_{i}}  \tag{6.44a}\\
& \bar{x}_{k}\left(\ell_{1}, \ldots, \ell_{n}\right)=j(d r) \bar{x}_{k}\left(r_{1}, \ldots, r_{n}\right) e^{-i \sum \ell_{i} \cdot r_{i}}, \tag{6.44b}
\end{align*}
$$

where

$$
\int(d r) \equiv \int d^{4} r_{1} \ldots d^{4} r_{n} \delta^{4}\left(\frac{1}{n} \sum_{1}^{n} r_{i}\right)
$$

It is important that all the relative momenta $\ell_{i}$ are real for (6.44) to make sense.
The Bethe-Salpeter integral equation and normalization condition may be written either in coordinate space or momentum space. They are displayed graphically in Fig. 3. The derivation of these equations and the graphical notation is the same as in Ref. 26. The only difference is that now the normalization is between $\chi_{\mathrm{k}}$ and $\bar{\chi}_{\mathrm{k}^{*}}$. This is because

$$
\begin{equation*}
\mathrm{G}\left(\mathrm{k}^{\mathrm{\prime}}, \mathrm{k}\right) \underset{\mathrm{k}^{2} \rightarrow \mathscr{M}^{2}}{ } \frac{\chi_{\mathrm{k}^{\prime}} \mathrm{i}(2 \pi)^{4} \delta^{4}\left(\mathrm{k}^{\prime}-\mathrm{k}\right) \bar{\chi}_{\mathrm{k}^{*}}}{\mathrm{k}^{2}-\mathscr{M}^{2}} . \tag{6.45}
\end{equation*}
$$

(See (6.5) for example.)

## D. S Matrix Elements

Scattering amplitudes involving unstable particles are residues of poles in the off-shell Green's functions in direct analogy with the stable case. The reduction formula expresses just this relation. Its practical value is that the Green's functions may be expanded in a perturbation series and the poles extracted from the external legs. This has been done for stable particles (elementary or composite) in Ref. 26. The graphical analysis leading to the pole extraction applies equally to unstable particles.

A simple example of the reduction formula for unstable particles is provided by e- $\mu$ elastic scattering. The scattering amplitude is

$$
\begin{equation*}
{ }^{\text {out }}{ }_{<\mathrm{k}_{2}}, \mathrm{p}_{2}\left|\mathrm{p}_{1}, \mathrm{k}_{1}\right\rangle^{\text {in }}=\int \mathrm{d}^{4} \mathrm{x}_{1} \mathrm{~d}^{4} \mathrm{x}_{2} \mathrm{~d}^{4} \mathrm{y}_{1} \mathrm{~d}^{4} \mathrm{y}_{2}\left[\mathrm{U}\left(\mathrm{k}_{2}, \mathrm{y}_{2}\right) \overline{\mathrm{U}}\left(\mathrm{p}_{2}, \mathrm{x}_{2}\right) \mathscr{R} \mathrm{U}\left(\mathrm{p}_{1}, \mathrm{x}_{1}\right) \mathrm{u}\left(\mathrm{k}_{1}, \mathrm{y}_{1}\right)\right] \tag{6.46}
\end{equation*}
$$

where
$\mathscr{R} \equiv\left(\vec{\square}_{\mathrm{y}_{2}}+\mathscr{N}^{*^{2}}\right)\left(\vec{\square}_{\mathrm{x}_{2}}+\mathrm{m}^{2}\right)<0\left|\mathrm{~T}\left[\psi\left(\mathrm{y}_{2}\right) \psi\left(\mathrm{x}_{2}\right) \bar{\psi}\left(\mathrm{x}_{1}\right) \bar{\psi}\left(\mathrm{y}_{1}\right)\right] 10\right\rangle\left(\bar{\square}_{\mathrm{x}_{1}}+\mathrm{m}^{2}\right)\left(\overleftarrow{\square}_{\mathrm{y}_{1}}+\mathscr{M}^{2}\right)$.
Here x and p are the coordinates and momenta of the $\mathrm{e} ; \mathrm{y}$ and k , of the $\mu$. The external wave functions are spinors times plane waves:

$$
\begin{equation*}
\mathrm{U}(\mathrm{p}, \mathrm{x})=\mathrm{u}(\mathrm{p}) \mathrm{e}^{\mathrm{ip} \cdot \mathrm{x}} \tag{6.47}
\end{equation*}
$$

Continuing the right hand side of (6.46) to complex $k_{i}$ such that

$$
\left(\mathrm{k}_{1}\right)^{2}=\mathscr{M}^{2}, \quad\left(\mathrm{k}_{2}\right)^{2}=\mathscr{M}^{*^{2}}
$$

gives the matrix element as the residue of the Green's function poles. Note that energy-momentum conservation reads

$$
\mathrm{p}_{2}+\mathrm{k}_{2}^{*}=\mathrm{p}_{1}+\mathrm{k}_{1}
$$

For stable particles the proof of the reduction formula depends on the existence of asymptotic states ${ }^{27}$ like

$$
\begin{equation*}
|p\rangle^{\text {in }}=\lim _{t \rightarrow-\infty} \int d^{3} y \bar{\psi}(y) i \frac{\vec{\partial}}{\partial t} u(p) e^{i p \cdot y}|0\rangle \tag{6.48}
\end{equation*}
$$

(This is, of course, only true in the weak operator sense.) For unstable particles the energy is complex so that this limit either diverges or vanishes. (This reflects the physical reality.) The absence of an asymptotic state has an important consequence for the connectedness of the scattering amplitudes. Consider e-e scattering instead of $e-\mu$. Then the complex $k_{i}$ in (6.46) become real $\ell_{i}$. The connected amplitude is

$$
\begin{equation*}
\text { out } \left._{<\ell_{2}, p_{2} \mid p_{1}, \ell_{1}>C}^{\text {in }}=i \int d^{4} y \text { out }_{<\ell_{2}}, p_{2}|\bar{\psi}(y)| p_{1}\right\rangle^{\text {in }}\left(\overleftarrow{\square}+\mathrm{m}^{2}\right) \mathrm{U}\left(\ell_{1}, y\right) \tag{6.49}
\end{equation*}
$$

Strictly speaking, the external wave functions should really be square integrable solutions of the Klein-Gordon equation that equal (6.47) only in the plane wave limit. Before this limit is taken the spatial derivatives in (6.49) may be integrated by parts to get

Because the wave function satisfies the Klein-Gordon equation this is

$$
\begin{align*}
\text { out }_{<\ell_{2}, p_{2} \mid p_{1}, \ell_{1}>C}^{\text {in }} & =-\int \mathrm{d}^{4} \mathrm{y} \frac{\partial}{\partial \mathrm{t}}\left\{\text { out }_{\left.\left\langle\ell_{2}, p_{2}\right| \bar{\psi}(\mathrm{y})\left|p_{1}\right\rangle^{\text {in }} \mathrm{i} \frac{\vec{\partial}}{\partial \mathrm{t}} \mathrm{U}\left(\ell_{1}, \mathrm{y}\right)\right\}}\right. \\
& =\left(\lim _{\mathrm{t}=-\infty}-\lim _{\mathrm{t}=+\infty}\right) \int \mathrm{d}^{3} \mathrm{y}{ }^{\text {out }}\left\langle\ell_{2}, \mathrm{p}_{2}\right| \bar{\psi}(\mathrm{y})\left|\mathrm{p}_{1}\right\rangle^{\text {in }} \mathrm{i} \frac{\vec{\partial}}{\partial \mathrm{t}} \mathrm{U}\left(\ell_{1}, \mathrm{y}\right) . \tag{6.51}
\end{align*}
$$

Because of the asymptotic condition ${ }^{27} \bar{\psi}(\mathrm{y})$ creates an in state at $\mathrm{t}=-\infty$ and creates an out state at $t=+\infty$ :

$$
\begin{equation*}
\text { out }_{<l_{2}}, \mathrm{p}_{2} \mid \mathrm{p}_{1}, \ell_{1}>\mathrm{C}{ }^{\text {in }}={ }^{\text {out }}\left\langle l_{2}, \mathrm{p}_{2} \mid \mathrm{p}_{1}, \ell_{1}\right\rangle^{\text {in }}-\left\langle\ell_{2} \mid l_{1}\right\rangle\left\langle p_{2} \mid p_{1}\right\rangle . \tag{6.52}
\end{equation*}
$$

This shows that stable particles may propagate freely or interact as Fig. 4 illustrates.

It is easy to see that unstable particles do not give this connectedness structure. If $\vec{k}$ is complex, the integration by parts in (6.50) is not possible and the derivation fails. If, however, $\overrightarrow{\mathrm{k}}$ is actually real then (6.50) is valid but because $\mathrm{k}^{\circ}$ is complex the limits in (6.51) either diverge or vanish. Thus (6.52) does hold for unstable particles. The connectedness structure for the S-matrix elements is then that given in Fig. 5 where the wavy line denotes the unstable particle.

For unstable particles the proof of the reduction formula must naturally be modified. The method of Appendix A is again the key. The Fourier transform of the three $\theta$-functions in $(6.46)$ produce three resolvents. These may be analytically continued to the poles whose residues are then the unstable state vectors. The right hand side of $(6.46)$ thus becomes

$$
\mathrm{i} \int \mathrm{~d}^{4} \mathrm{y}{ }^{\text {out }}\left\langle\mathrm{k}_{2}, \mathrm{p}_{2}\right| \bar{\psi}(\mathrm{y}) \mid \mathrm{p}_{1}>\text { in }\left[\overleftarrow{\square}+\mathscr{N}^{2}\right] \mathrm{U}\left(\mathrm{k}_{1}, \mathrm{y}\right)
$$

analogously to (6.49). But now there is no fourth $\theta$-function to produce a resolvent. Instead, crossing and translational invariance give

$$
\begin{align*}
\mathrm{i} \int \mathrm{~d}^{4} \mathrm{y} & \text { out }_{<\mathrm{k}_{2}, \mathrm{p}_{2},} \overline{\mathrm{p}}_{1}\left|\mathrm{e}^{-\mathrm{i} \mathscr{P} \cdot \mathrm{y}} \bar{\psi}(0)\right| 0>\left[\overleftarrow{\square}^{2}+\mathscr{N}^{2}\right] \mathrm{U}\left(\mathrm{k}_{1}, \mathrm{y}\right) \\
& =\mathrm{i}^{\text {out }}<\mathrm{k}_{2}, \mathrm{p}_{2}, \overline{\mathrm{p}}_{1}\left|(2 \pi)^{4} \delta^{4}\left(\mathrm{k}_{1}-\mathscr{P}\right) \vec{\psi}(0)\right| 0>\left(-\mathrm{k}_{1}^{2}+\mathscr{M}^{2} \mathrm{U}\left(\mathrm{k}_{1}\right)\right. \tag{6.53}
\end{align*}
$$

To go to $\left(\mathrm{k}_{1}\right)^{2}=\mathscr{M}^{2}$ requires analytically continuing the $\delta$ function. Suppose $\overrightarrow{\mathrm{k}}_{1}$ is kept real and $\mathrm{k}_{1}^{\mathrm{O}}$ is taken complex. Then as discussed in Section V. C,

Bremermann and Durand ${ }^{24}$ have shown

$$
2 \pi \mathrm{i} \delta\left(\mathrm{k}^{\mathrm{O}}-\mathrm{H}\right)= \begin{cases}\frac{1}{\mathrm{k}^{\mathrm{O}}-\mathrm{H}} & \operatorname{Im}\left(\mathrm{k}^{\mathrm{O}}-\mathrm{H}\right)>0  \tag{6.54}\\ \frac{1}{\mathrm{k}^{\mathrm{O}}-\mathrm{H}} & \operatorname{Im}\left(\mathrm{k}^{\mathrm{O}}-\mathrm{H}\right)<0\end{cases}
$$

The continuation appropriate to (6.53) starts with $\mathrm{k}_{1}^{0}$ in the upper half plane. Thus (6.53) becomes

$$
\begin{equation*}
\left.{ }^{\text {out }}<\mathrm{k}_{2}, \mathrm{p}_{2}, \overline{\mathrm{p}}_{1} \mid(2 \pi)^{3} \delta^{3} \overrightarrow{\mathrm{k}}_{1}-\overrightarrow{\mathscr{P}}\right) \left.\frac{1}{\mathrm{k}_{1}^{0}-\mathrm{H}} \bar{\psi}(0) \right\rvert\, 0>\left(\mathrm{k}_{1}^{2}-\mathscr{M}^{2}\right) \mathrm{U}\left(\mathrm{k}_{1}\right) \tag{6.55}
\end{equation*}
$$

when $\operatorname{Im} \mathrm{k}_{1}^{\mathrm{O}}>0$. Now the resolvent can be continued clockwise onto the second sheet. Because of (2.9) it has a pole whose residue is a product of unstable state vectors. The continuation of (6.55) to the pole is then

Using crossing again shows this is

$$
{ }^{\text {out }}\left\langle\mathrm{k}_{2}, \mathrm{p}_{2} \mid \mathrm{p}_{1}, \mathrm{k}_{1}\right\rangle{ }^{\text {in }} .
$$

This demonstrates the reduction formula (6.46) as claimed. (Note that for stable particles the continuation is to a real energy below the production threshold and the same argument is rather weak because of the double contribution from (6.54). Already knowing the correct answer (6.52) shows that the $\pm i \epsilon$ actually corresponds to the creation of both an in state and an out state.) If the unstable particles are composite rather than elementary essentially the same proof applies.

## VII. A SOLUBLE EXAMPLE

To show how simple it is to describe unstable particles as eigenstates of the Hamiltonian, the familiar problem of elastic scattering from a threedimensional square well potential will now be examined. This problem is nonrelativistic, but otherwise contains all the features of unstable particle states. The exact $S$ matrix is given in (7.7). The wave functions that correspond to the second sheet poles are displayed in (7.28). These are exponentially growing but nevertheless have zero norm (7.32) when properly calculated by analytic continuation. This method of calculating with unstable particle wave functions is put to an exact test subsection E .
A. The S Matrix

The Schroedinger equation for a constant potential of depth $\mathrm{V}>0$ and range $R$ is

$$
\begin{equation*}
\left[-\frac{1}{2 \mathrm{~m}} \nabla^{2}-\mathrm{V} \theta(\mathrm{R}-\mathrm{r})\right] \Phi(\overrightarrow{\mathrm{r}})=\mathrm{E} \Phi(\overrightarrow{\mathrm{r}}) \tag{7.1}
\end{equation*}
$$

The $\ell=0$ solution is

$$
\Phi_{E}(r)= \begin{cases}\frac{C(E)}{r} \sin K r & r \leq R  \tag{7.2}\\ -\frac{D(E)}{r}\left[e^{-i k r}-\eta(E) e^{i k r}\right] & r \geq R,\end{cases}
$$

where

$$
\begin{equation*}
\mathrm{k}=\sqrt{2 \mathrm{mE}}, \quad \mathrm{~K}=\sqrt{2 \mathrm{~m}(\mathrm{E}+\mathrm{V})} . \tag{7.3}
\end{equation*}
$$

The three quantities $\mathrm{C}, \mathrm{D}$, and $\eta$ are determined by the continuity of $\Phi$ and its derivative at $r=R$ and by the overall normalization.

The scattering amplitude is determined by the phase shifts:

$$
\begin{equation*}
\left\langle\overrightarrow{\mathrm{p}}_{2}\right| \mathrm{S}\left|\overrightarrow{\mathrm{p}}_{1}\right\rangle=(2 \pi)^{3} \frac{\delta\left(\mathrm{p}_{2}-\mathrm{p}_{1}\right)}{4 \pi \mathrm{p}_{2} \mathrm{p}_{1}} \sum_{\ell=0}^{\infty}(2 \ell+1) \mathrm{e}^{2 \mathrm{i} \delta}{ }^{\infty} \mathrm{P}_{\ell}(\cos \theta) \tag{7.4}
\end{equation*}
$$

where plane waves are normalized as

$$
\begin{equation*}
\left\langle\overrightarrow{\mathrm{p}}_{2} \mid \overrightarrow{\mathrm{p}}_{1}\right\rangle=(2 \pi)^{3} \delta^{3}\left(\overrightarrow{\mathrm{p}}_{2}-\overrightarrow{\mathrm{p}}_{1}\right) . \tag{7.5}
\end{equation*}
$$

The phase shifts may be calculated from the asymptotic behavior of the wave function because

$$
\Phi(\vec{r}) \underset{r \rightarrow \infty}{\longrightarrow} e^{i \vec{k} \cdot \vec{r}}+f(\theta) \frac{e^{i k r}}{r}
$$

where

$$
\begin{equation*}
\mathrm{f}(\theta)=\frac{1}{2 \mathrm{ik}} \sum_{\ell=0}^{\infty}(2 \ell+1)\left(\mathrm{e}^{2 \mathrm{i} \delta_{\ell}}-1\right) \mathrm{P}_{\ell}(\cos \theta) \tag{7.6}
\end{equation*}
$$

Comparing the asymptotic form of (7.2) with (7.6) gives

$$
\eta(\mathrm{F})=\mathrm{e}^{2 \mathrm{i} \delta_{0}(\mathrm{E})}
$$

Thus $\eta$ determines the scattering amplitude. To find $\eta$ explicitly, apply the boundary conditions to (7.2) to get

$$
\begin{equation*}
\eta_{\mathrm{I}}(\mathrm{E})=\frac{\mathrm{K} \cot \mathrm{KR}+\mathrm{ik}}{\mathrm{~K} \cot \mathrm{KR}-\mathrm{ik}} \mathrm{e}^{-2 \mathrm{ikR}} \tag{7.7a}
\end{equation*}
$$

This is the exact $S$ matrix. It has a cut in $E$ whose location depends on the definition chosen for the square root in (7.3). It is conventional to take this cut along the positive real axis. Thus

$$
0<\operatorname{Arg} \sqrt{\mathrm{E}} \leq \pi,
$$

i.e.,

$$
\begin{align*}
& \operatorname{Im} k \geq 0  \tag{7.8}\\
& \operatorname{Im} K \geq 0 .
\end{align*}
$$

The analytic continuation of $\eta_{\mathrm{I}}(\mathrm{E})$ onto the second sheet is then

$$
\begin{equation*}
\eta_{I I}(\mathrm{E})=\frac{\mathrm{K} \cot \mathrm{KR}-\mathrm{ik}}{\mathrm{~K} \cot \mathrm{KR}+\mathrm{ik}} \mathrm{e}^{2 \mathrm{ikR}} \tag{7.7b}
\end{equation*}
$$

where k and K are still the principle square roots given in (7.8).

The two-sheeted S matrix $\eta(E)$ explicitly satisfies extended unitarity and hermitian analyticity. Physical unitarity is the statement

$$
\left|\eta_{\mathrm{I}}(\mathrm{E})\right|^{2}-1 \quad \text { (E positive real) }
$$

The extension of this to all values of E is

$$
\begin{equation*}
\left.\eta_{\mathrm{I}}(\mathrm{E}) \eta_{\mathrm{II}}(\mathrm{E})=1 \quad \text { ( } \mathrm{E} \text { complex }\right) . \tag{7.9}
\end{equation*}
$$

Hermitian analyticity follows from the principle square root definition in (7.8) which satisfies

$$
(\sqrt{\mathrm{E}})^{*}=-\sqrt{\mathrm{E}^{*}}
$$

Consequently

$$
\begin{gather*}
\eta_{\mathrm{I}}(\mathrm{E})^{*}=\eta_{\mathrm{I}}\left(\mathrm{E}^{*}\right)  \tag{7.10}\\
\eta_{\mathrm{II}}(\mathrm{E})^{*}=\eta_{\mathrm{II}} \mathrm{I}^{\left(\mathrm{E}^{*}\right)}
\end{gather*}
$$

It is well known that resonance bumps occur along the real axis at energies for which

$$
\cot K_{B} R=0
$$

This means

$$
\mathrm{K}_{\mathrm{B}} \mathrm{R}=(\mathrm{n}+1 / 2) \pi
$$

or equivalently

$$
\begin{equation*}
\mathrm{E}_{\mathrm{B}}=\frac{(\mathrm{n}+1 / 2)^{2} \pi^{2}}{2 \mathrm{mR}^{2}}-\mathrm{V} \tag{7.11a}
\end{equation*}
$$

Precisely at this energy

$$
\eta\left(\mathrm{E}_{\mathrm{B}}\right)=-\mathrm{e}^{-2 i \mathrm{~K}_{\mathrm{B}} \mathrm{R}}
$$

The width of the resonance is obtained by writing

$$
\eta_{\mathrm{I}}(\mathrm{E})=\frac{\mathrm{f}(\mathrm{E})-\mathrm{E}}{\mathrm{f}(\mathrm{E})+\mathrm{E}} \mathrm{e}^{-2 i k R}
$$

where

$$
\begin{equation*}
f(E) \equiv i \frac{k K}{2 m} \cot K R \tag{7.12}
\end{equation*}
$$

When $E$ is near $E_{B}$

$$
\mathrm{f}(\mathrm{E}) \approx\left(\mathrm{E}-\mathrm{E}_{\mathrm{B}}\right) \mathrm{f}^{\prime}\left(\mathrm{E}_{\mathrm{B}}\right)+\mathscr{O}\left(\mathrm{E}-\mathrm{E}_{\mathrm{B}}\right)^{2}
$$

Therefore

$$
\eta(E) \approx \frac{E-E_{B}-\frac{i}{2} \gamma_{B}}{E-E_{B}+\frac{i}{2} \gamma_{B}} e^{-2 i k_{B} R}
$$

where

$$
-\frac{i}{2} \gamma_{\mathrm{B}} \equiv \frac{\mathrm{E}_{\mathrm{B}}}{\mathrm{f}^{\prime}\left(\mathrm{E}_{\mathrm{B}}\right)}
$$

Explicitly calculating this gives

$$
\begin{equation*}
\gamma_{\mathrm{B}}=\frac{2 \sqrt{(\mathrm{n}+1 / 2)^{2} \pi^{2}-2 \mathrm{mVR}^{2}}}{\mathrm{mR}^{2}} \tag{7.11b}
\end{equation*}
$$

The resonances will be distinct only if the width is much smaller than the energy spacing. This requires that

$$
\begin{equation*}
\frac{2 \sqrt{(n+1 / 2)^{2} \pi^{2}-2 m V R^{2}}}{(n+1) \pi^{2}} \ll 1 \tag{7.13}
\end{equation*}
$$

## B. Location of the Poles

The discussion of these resonance bumps is quite standard and is given in many textbooks. The emphasis of the present work, however, is not on these bumps but rather on second sheet poles. In this example, as always, there is no guarantee that real axis bumps are a manifestation of second sheet poles. This section will show that, in fact, there are second sheet poles. These locations will turn out to agree with the energies and widths in (7.11).

Let the poles occur at a complex energy $\tilde{E}$ on the second sheet. Put

$$
\begin{align*}
\alpha & =\mathrm{R} \sqrt{2 \mathrm{~m}(\widetilde{\mathrm{E}}+\mathrm{V})} \\
\beta & =\mathrm{R} \sqrt{2 \mathrm{~m} \widetilde{\mathrm{E}}} \tag{7.14}
\end{align*}
$$

where

$$
\operatorname{Im} \alpha>0, \quad \operatorname{Im} \beta>0 .
$$

Then the $S$ matrix (7.7b) has a second sheet pole only if

$$
\begin{align*}
\alpha \cot \alpha & =-\mathrm{i} \beta \\
\alpha^{2}-\beta^{2} & =2 \mathrm{mVR}^{2} . \tag{7.15}
\end{align*}
$$

Now eliminate $\beta$ to get

$$
\begin{equation*}
\frac{\alpha}{\sin \alpha}= \pm \sqrt{2 \mathrm{mVR}^{2}} . \tag{7.16}
\end{equation*}
$$

The real part of the left hand side must vanish and the imaginary part must equal the right hand side. Let

$$
\begin{align*}
& \alpha=\alpha_{1}+\mathrm{i} \alpha_{2} \\
& \beta=\beta_{1}+\mathrm{i} \beta_{2} . \tag{7.17}
\end{align*}
$$

Then (7.16) gives

$$
\begin{gather*}
\frac{\tanh \alpha_{2}}{\alpha_{2}}=\frac{\tan \alpha_{1}}{\alpha_{1}}  \tag{7.18a}\\
\alpha_{2}= \pm \cos \alpha_{1} \sqrt{\left(\frac{\alpha_{1}}{\sin \alpha_{1}}\right)^{2}-2 \mathrm{mVR}^{2}} . \tag{7.18b}
\end{gather*}
$$

The existence of solutions to these equations is proven graphically in Fig. 6. There the solutions of (7.18a) and of (7.18b) are both plotted. The intersections of the plots give the $\alpha_{1}$ and $\alpha_{2}$ of the corresponding unstable particle pole. Figure 6 is plotted for a weak potential. As $V$ is increased the widths become smaller and $\alpha_{1}$ approaches ( $\mathrm{n}+1 / 2$ ) $\pi$. If V is increased so much that $2 \mathrm{mVR}{ }^{2}>\alpha_{1}$
the unstable particles become stable. Then $\alpha_{2}=0$ and (7.18b) determines $\alpha_{1}$. As $\mathrm{V} \rightarrow \infty$, the stable particle energies approach the usual value $\alpha_{1}=n \pi$.

Once $\alpha_{1}$ and $\alpha_{2}$ are found, the corresponding values of $\beta_{1}$ and $\beta_{2}$ given by (7.15) are

$$
\begin{align*}
& \beta_{1}=\alpha_{2} \tan \alpha_{1}\left[\frac{\csc ^{2} \alpha_{1}-\operatorname{csch}^{2} \alpha_{2}}{\cot ^{2} \alpha_{1}+\operatorname{coth}^{2} \alpha_{2}}\right] \\
& \beta_{2}=\alpha_{1} \cot \alpha_{1}\left[\frac{\csc ^{2} \alpha_{1}+\operatorname{csch}^{2} \alpha_{2}}{\cot ^{2} \alpha_{1}+\operatorname{coth}^{2} \alpha_{2}}\right] . \tag{7.19}
\end{align*}
$$

Figure 6 shows that the only allowed values of positive $\alpha_{1}$ lie in the strips

$$
\mathrm{n} \pi<\alpha_{1}<(\mathrm{n}+1 / 2) \pi
$$

Because of (7.19) this means that all the poles have $\beta_{2}>0$ and hence all lie on the sccond shect as claimed. (Of course, $\alpha_{2}$ must also be positive but its sign is lost in going from (7.15) to (7.16).)

Let the real and imaginary parts of the energy at the pole be given by

$$
\begin{equation*}
\widetilde{\mathrm{E}}=\mathrm{E}_{\mathrm{P}}+\frac{\dot{i}}{2} \gamma_{\mathrm{P}} \tag{7.20}
\end{equation*}
$$

Then

$$
\begin{align*}
& \mathrm{E}_{\mathrm{P}}=\frac{\alpha_{1}^{2}-\alpha_{2}^{2}}{2 \mathrm{mR}^{2}}-\mathrm{V}=\frac{\beta_{1}^{2}-\beta_{2}^{2}}{2 \mathrm{mR}^{2}}  \tag{7.21}\\
& \gamma_{\mathrm{P}}=\frac{2 \alpha_{1} \alpha_{2}}{\mathrm{mR}^{2}}=\frac{2 \beta_{1} \beta_{2}}{\mathrm{mR}^{2}} \tag{7.22}
\end{align*}
$$

because of the definitions (7.14). Equations (7.18) have the property that if $\alpha_{1}+\mathrm{i} \alpha_{2}$ is a solution then so is $-\alpha_{1}+\mathrm{i} \alpha_{2}$. Therefore from (7.19) both $\beta_{1}+\mathrm{i} \beta_{2}$ and $-\beta_{1}+\mathrm{i} \beta_{2}$ are solutions. Thus in (7.22) there are automatically poles at $\mathrm{E}_{\mathrm{P}}+\frac{\mathbf{i}}{2} \gamma_{\mathrm{P}}$ and $\mathrm{E}_{\mathrm{P}}-\frac{\mathrm{i}}{2} \gamma_{\mathrm{P}}$ as claimed.

The existence of the second sheet poles is thus proved but it would be nice to determine their locations more accuratcly. Fortunatcly, approximate solutions to the coupled transcendental equations (7.18) may be obtained for $\alpha_{1}$ large and $\alpha_{2}$ small. In this limit Fig. 6 shows that $\alpha_{1}$ is a little less than ( $\mathrm{n}+1 / 2$ ) $\pi$ for n a large positive integer. Therefore let

$$
\begin{equation*}
\alpha_{1}=(\mathrm{n}+1 / 2) \pi-\frac{\lambda}{(\mathrm{n}+1 / 2) \pi} \tag{7.23}
\end{equation*}
$$

and determine $\lambda$. Equation (7.18b) gives

$$
\begin{equation*}
\alpha_{2}=\frac{\lambda \sqrt{b^{2}+\lambda(\lambda-2)}}{\mathrm{a}}\left[1+\mathscr{O}\left(\frac{\lambda^{2}}{\mathrm{a}^{2}}\right)\right] \tag{7.24}
\end{equation*}
$$

where

$$
\begin{aligned}
& a \equiv(\mathrm{n}+1 / 2) \pi \\
& \mathrm{b} \equiv \sqrt{(\mathrm{n}+1 / 2)^{2} \pi^{2}-2 \mathrm{mVR}^{2}} .
\end{aligned}
$$

For $\alpha_{2}$ to be small it is necessary that

$$
\mathrm{b} \ll \mathrm{a}
$$

This is in accordance with (7.13). Substituting (7.23) and (7.24) into (7.18a) determines $\lambda$ via the equation

$$
\frac{1}{\lambda}+\frac{1}{a^{2}}\left(1-\frac{\lambda}{3}\right)=1-\frac{1}{3} \frac{\lambda^{2}\left(b^{2}+\lambda(\lambda-2)\right)}{a^{2}}+\mathscr{O}\left(\frac{1}{a^{4}}\right) .
$$

Because a is large, $\lambda$ must be near 1. In fact

$$
\lambda=1+\frac{b^{2}+1}{3 a^{2}} .
$$

The solution is then

$$
\alpha_{1}=a-\frac{1}{a}-\frac{b^{2}+1}{3 a^{3}}+\ldots
$$

$$
\begin{equation*}
\alpha_{2}=\frac{\sqrt{\mathrm{b}^{2}-1}}{\mathrm{a}}\left[1+\frac{\mathrm{b}^{2}+1}{3 \mathrm{a}^{2}}+\ldots\right] \tag{7.25a}
\end{equation*}
$$

The corresponding values of $\beta$ are

$$
\begin{align*}
& \beta_{1}=\sqrt{\mathrm{b}^{2}-1}\left[1-\frac{1}{2 \mathrm{a}^{2}}+\ldots\right]  \tag{7.25b}\\
& \beta_{2}=1+\frac{2 \mathrm{~b}^{2}-1}{6 \mathrm{a}^{2}}+\ldots
\end{align*}
$$

These solutions give for the real energy (7.21)

$$
E_{P}=\frac{1}{2 m R^{2}}\left[b^{2}-2+\frac{4-5 b^{2}}{3 a^{2}}+\cdots\right]
$$

This is in good agreement with the energy at which the bump on the real axis occurs which according to (7.11a) is

$$
\mathrm{E}_{\mathrm{B}}=\frac{\mathrm{b}^{2}}{2 \mathrm{mR}^{2}}
$$

The width which follows from (7.22) is

$$
\begin{equation*}
\gamma_{P}=\frac{2 \sqrt{b^{2}-1}}{m R^{2}}\left[1+\frac{b^{2}-2}{3 a^{2}}+\ldots\right] \tag{7.26a}
\end{equation*}
$$

whereas the width of the real axis bump is

$$
\begin{equation*}
\gamma_{\mathrm{B}}=\frac{2 \mathrm{~b}}{\mathrm{mR}^{2}} \tag{7.26b}
\end{equation*}
$$

according to (7.11b).

## C. Schrœdinger Wave Functions

The wave functions for an unstable state is the analytic continuation in $\widetilde{E}$ of $\Phi(\mathrm{E}, \mathrm{r})$ to the point E on the second sheet. At this value of energy, $\eta=\infty$ and $\mathrm{D}=0$ but the product $\mathrm{D}_{\eta}$ is finite. To calculate norms and inner products with such
states it is useful to introduce the wave function

$$
\psi(E, r)= \begin{cases}\frac{C(E)}{r} \sin K r & r \leq R  \tag{7.27}\\ \frac{D(E) \eta(E)}{r} e^{i k r} & r \geq R\end{cases}
$$

where C, D, and $\eta$ are the same functions of $E$ as in (7.2). Note that whereas $\Phi$ is an energy eigenstate for any energy $E$ because it satisfies the boundary conditions for any E, $\psi$ certainly is not. However $\psi$ is an energy eigenstate at energy $\widetilde{E}$ on the second sheet because it coincides with $\Phi$ there:

$$
\psi_{\Pi}(\widetilde{\mathrm{E}}, \mathrm{r})=\Phi_{\mathrm{II}}(\widetilde{\mathrm{E}}, \mathrm{r})
$$

On the second sheet $\psi$ is given by

$$
\psi_{\Pi I}(\mathrm{E}, \mathrm{r})= \begin{cases}\frac{-\mathrm{C}_{\mathrm{II}}(\mathrm{E})}{\mathrm{r}} \sin \mathrm{Kr} & \mathrm{r} \leq \mathrm{R}  \tag{7.28}\\ i \frac{\mathrm{D}_{\mathrm{II}}(\mathrm{E}) \eta_{\Pi}(\mathrm{E})}{\mathrm{r}} \mathrm{e}^{-\mathrm{ikr}} & \mathrm{r} \geq \mathrm{R}\end{cases}
$$

where

$$
\operatorname{Im} \mathrm{k} \geq 0, \quad \operatorname{Im} K \geq 0
$$

as always. The continuity of $\psi_{\mathrm{II}}$ and its derivative at $\widetilde{\mathrm{E}}$ imply

$$
\begin{align*}
-\mathrm{C} \sin \alpha & =(\mathrm{D} \eta) \mathrm{e}^{-\mathrm{i} \beta}  \tag{7.29}\\
\mathrm{C} \cos \alpha & =\mathrm{i} \frac{\beta}{\alpha}(\mathrm{D} \eta) \mathrm{e}^{-\mathrm{i} \beta}
\end{align*}
$$

where $\alpha$ and $\beta$ are given by (7.14) and $C, D, \eta$ are the values at $\widetilde{\mathrm{E}}$ on the second sheet. (Note that the boundary conditions (7.29) contain the pole constraints (7.15) on $\alpha$ and $\beta$.)

The norm of $\psi_{\text {II }}$ obviously cannot be obtained by integrating the absolute square of the wave function (7.28) because it diverges exponentially in $r$. This divergence in results from continuing to the second sheet; $\psi$ decreases
exponentially on the first sheet as (7.27) shows. The norm of $\psi_{\mathrm{II}}$ is obtained by calculating the norm on the first sheet and then analytically continuing the answer to the second sheet. (See Section III.) The first sheet norm is simply

$$
\begin{align*}
\langle\psi(\mathrm{E}) \mid \psi(\mathrm{E})\rangle= & \int \mathrm{d}^{3}{ }_{\mathrm{r}}|\psi(\mathrm{E}, \mathrm{r})|^{2} \\
= & 2 \pi|\mathrm{C}(\mathrm{E})|^{2}\left[\frac{\sin \left(\mathrm{~K}^{*}-\mathrm{K}\right) \mathrm{R}}{\mathrm{~K}^{*}-\mathrm{K}}-\frac{\sin \left(\mathrm{K}^{*}+\mathrm{K}\right) R}{\mathrm{~K}^{*}+\mathrm{K}}\right] \\
& +4 \pi|\mathrm{D}(\mathrm{E}) \eta(\mathrm{E})|^{2} \frac{\mathrm{e}^{i\left(\mathrm{k}-\mathrm{k}^{*}\right) \mathrm{R}}}{\mathrm{i}\left(\mathrm{k}-\mathrm{k}^{*}\right)} \tag{7.30}
\end{align*}
$$

Analytically continuing (7.30) to the second sheet gives

$$
\begin{gather*}
\left\langle\psi_{\mathrm{II}}(\widetilde{\mathrm{E}}) \mid \psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle=2 \pi \mathrm{R}|\mathrm{C}|^{2}\left[\frac{\sinh 2 \alpha_{2}}{2 \alpha_{2}}-\frac{\sin 2 \alpha_{1}}{2 \alpha_{1}}\right] \\
-2 \pi \mathrm{R}|\mathrm{D} \eta|^{2} \frac{\mathrm{e}^{2 \beta_{2}}}{\beta_{2}} \tag{7.31}
\end{gather*}
$$

Because of the two boundary conditions (7.29) this whole quantity vanishes. To see this, multiply the first boundary condition by the complex conjugate of the second and vice versa to get

$$
\begin{aligned}
& |C|^{2} \sin \alpha \cos \alpha^{*}=i \frac{\beta^{*}}{\alpha^{*}}|\mathrm{D} \eta|^{2} \mathrm{e}^{2 \beta_{2}} \\
& |\mathrm{C}|^{2} \sin \alpha^{*} \cos \alpha=-i \frac{\beta}{\alpha}|\mathrm{D} \eta|^{2} \mathrm{e}^{2 \beta_{2}} .
\end{aligned}
$$

Add and subtract these:

$$
\begin{aligned}
& |C|^{2} \sin 2 \alpha_{1}-i\left(\frac{\beta^{*}}{\alpha^{*}}-\frac{\beta}{\alpha}\right)|\mathrm{D} \eta|^{2} \mathrm{e}^{2 \beta_{2}} \\
& |\mathrm{C}|^{2} \sinh 2 \alpha_{2}=\left(\frac{\beta^{*}}{\alpha^{*}}+\frac{\beta}{\alpha}\right)|\mathrm{D} \eta|^{2} \mathrm{e}^{2 \beta_{2}} .
\end{aligned}
$$

Substitute this into (7.31) to get

$$
\begin{aligned}
\left\langle\psi_{\mathrm{II}}\left(\mathrm{E}_{\mathrm{p}}\right) \mid \psi_{\mathrm{II}}\left(\mathrm{E}_{\mathrm{p}}\right)\right\rangle & =2 \pi \mathrm{R}|\mathrm{D} \eta|^{2} \mathrm{e}^{2 \beta_{2}}\left[\frac{1}{2 \alpha_{2}}\left(\frac{\beta^{*}}{\alpha^{*}}+\frac{\beta}{\alpha}\right)-\frac{i}{2 \alpha_{1}}\left(\frac{\beta^{*}}{\alpha^{*}}-\frac{\beta}{\alpha}\right)-\frac{1}{\beta_{2}}\right] \\
& =2 \pi \mathrm{R}|\mathrm{D} \eta|^{2} \mathrm{e}^{2 \beta_{2}}\left[\frac{\beta_{1}}{\alpha_{1} \alpha_{2}}-\frac{1}{\beta_{2}}\right]
\end{aligned}
$$

Now from the definition of $\alpha$ and $\beta$

$$
\alpha^{2}-\beta^{2}=2 \mathrm{mVR}
$$

Separating real and imaginary parts gives

$$
\alpha_{1} \alpha_{2}-\beta_{1} \beta_{2}=0
$$

Thus

$$
\begin{equation*}
\left\langle\dot{\psi}_{\mathrm{II}}(\widetilde{\mathrm{E}}) \mid \psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle=0 . \tag{7.32}
\end{equation*}
$$

Note that the vanishing of this norm depends essentially on the reality of $V$, i.e., on the hermiticity of the Hamiltonian.

As indicated in Section III the effective norm of an unstable state is the inner product with the state of energy $\widetilde{\mathbb{E}} *$. This will be explicitly calculated now and used in several later subsections. It was shown subsequent to (7.22) that such poles do exist. Because of time reversal invariance

$$
\Phi_{E^{*}}(\vec{r})=\Phi_{E}(\vec{r})^{*}
$$

Therefore (7.27) satisfies

$$
\begin{equation*}
\psi\left(\mathrm{E}^{*}, \mathrm{r}\right)=\psi(\mathrm{E}, r)^{*} \tag{7.33}
\end{equation*}
$$

This means that evaluating

$$
\left\langle\psi_{\mathrm{II}}\left(\widetilde{\mathrm{E}}^{*}\right) \mid \psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle
$$

first requires calculating

$$
\begin{aligned}
\left\langle\psi\left(\mathrm{E}^{*}\right) \mid \psi(\mathrm{E})\right\rangle & =\int \mathrm{d}^{3} \mathrm{x} \psi\left(\mathrm{E}^{*}, \mathrm{r}\right)^{*} \psi(\mathrm{E}, \mathrm{r}) \\
& =\int \mathrm{d}^{3} \mathrm{x}[\psi(\mathrm{E}, \mathrm{r})]^{2}
\end{aligned}
$$

and then continuing the result to E . Performing the integral gives

$$
\begin{aligned}
\left\langle\psi\left(\mathrm{E}^{*}\right) \mid \psi(\mathrm{E})\right\rangle= & 2 \pi[\mathrm{C}(\mathrm{E})]^{2}\left[\mathrm{R}-\frac{\sin 2 \mathrm{KR}}{2 \mathrm{~K}}\right] \\
& -4 \pi[\mathrm{D}(\mathrm{E}) \eta(\mathrm{E})]^{2} \frac{\mathrm{e}^{2 \mathrm{ikR}}}{2 \mathrm{ik}}
\end{aligned}
$$

The continuation of this to the second sheet is

$$
\begin{equation*}
\left\langle\psi_{\Pi}\left(\widetilde{\mathrm{E}}^{*}\right) \mid \psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle=2 \pi R \mathrm{C}^{2}\left[1-\frac{\sin 2 \alpha}{2 \alpha}\right]+2 \pi \mathrm{R}(\mathrm{D} \eta)^{2} \frac{\mathrm{e}^{-2 \mathrm{i} \beta}}{\dot{\mathrm{i} \beta}} \tag{7.34}
\end{equation*}
$$

The boundary conditions (7.29) may be manipulated into another useful form by multiplying them together and by adding the squares:

$$
\begin{aligned}
& -\mathrm{C}^{2} \sin \alpha \cos \alpha=\mathrm{i} \frac{\beta}{\alpha}(\mathrm{D} \eta)^{2} \mathrm{e}^{-2 \mathrm{i} \beta} \\
& \mathrm{C}^{2}=\left(1-\frac{\beta^{2}}{\alpha^{2}}\right)(\mathrm{D} \eta)^{2} \mathrm{e}^{-2 \mathrm{i} \beta}
\end{aligned}
$$

Using these in (7.34) gives

$$
\left\langle\psi_{\mathrm{II}}\left(\widetilde{\mathrm{E}}^{*}\right)\right| \psi_{\mathrm{II}}(\widetilde{\mathrm{E}})>=2 \pi \mathrm{R}(\mathrm{D} \eta)^{2} \mathrm{e}^{-2 \mathrm{i} \beta}\left(1-\frac{\beta^{2}}{\alpha^{2}}\right)\left(1+\frac{1}{\mathrm{i} \beta}\right) .
$$

For subsequent purposes it is useful to define normalized states

$$
\begin{align*}
& \left|\Psi_{I I}(\widetilde{\mathrm{E}})\right\rangle=\left\lvert\, \psi_{\mathrm{II}}(\widetilde{\mathrm{E}})>\frac{1}{\sqrt{\mathrm{~N}}}\right. \\
& \left|\Psi_{\mathrm{II}}\left(\widetilde{\mathrm{E}}^{*}\right)\right\rangle=\left|\psi_{\mathrm{II}}\left(\widetilde{\mathrm{E}}^{*}\right)\right\rangle \frac{1}{\sqrt{\mathrm{~N}^{*}}} \tag{7.35}
\end{align*}
$$

where

$$
\begin{align*}
\mathrm{N} & =2 \pi \mathrm{R}(\mathrm{D} \eta)^{2} \mathrm{e}^{-2 \mathrm{i} \beta}\left(1-\frac{\beta^{2}}{\alpha^{2}}\right)\left(1+\frac{1}{\mathrm{i} \beta}\right) \\
& =2 \pi \mathrm{RC}^{2}\left(1+\frac{1}{\mathrm{i} \beta}\right) . \tag{7.36}
\end{align*}
$$

## D. Calculation of $\gamma_{F}$

One simple application of these curious wave functions is to calculate the decay width $\gamma_{\mathrm{F}}$ given by Fermi's Golden Rule and see how it compares with $\gamma_{\mathrm{B}}$ and $\gamma_{\mathrm{P}}$. This requires evaluating the matrix element

$$
\begin{equation*}
\left.<\overrightarrow{\mathrm{k}}^{*}|\mathrm{~V}| \Psi \mathrm{II}(\widetilde{\mathrm{E}})\right\rangle \tag{7.37}
\end{equation*}
$$

where

$$
\overrightarrow{\mathrm{k}}^{2}=2 \mathrm{~m} \widetilde{\mathrm{E}}
$$

To calculate the decay amplitude use the first sheet wave functions (7.27) to get

$$
\begin{aligned}
\left\langle\overrightarrow{\mathrm{k}}^{*}\right| V|\psi(\mathrm{E})\rangle & =-\int_{\mathrm{d}}{ }^{3} \mathrm{x} \mathrm{e}^{-\mathrm{i} \overrightarrow{\mathrm{k}} \cdot \overrightarrow{\mathrm{x}}} \mathrm{~V} \theta(\mathrm{R}-\mathrm{r}) \frac{\mathrm{C}(\mathrm{E})}{\mathrm{r}} \sin \mathrm{Kr} \\
& =-\dot{C}(\mathrm{E}) \frac{2 \pi V}{\mathrm{k}}\left[\frac{\sin (\mathrm{~K}-\mathrm{k}) \mathrm{R}}{\mathrm{~K}-\mathrm{k}}-\frac{\sin (\mathrm{K}+\mathrm{k}) \mathrm{R}}{\mathrm{k}+\mathrm{k}}\right]
\end{aligned}
$$

Continuing this to the second sheet and using the definitions of $\alpha$ and $\beta$ gives

$$
\left\langle\overrightarrow{\mathrm{k}}^{*}\right| \mathrm{V}\left|\psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle=\mathrm{C} \frac{2 \pi \mathrm{VR}^{2}}{\beta}\left[\frac{\sin (\alpha-\beta)}{\alpha-\beta}-\frac{\sin (\alpha+\beta)}{\alpha+\beta}\right]
$$

Using (7.15) simplifies this to

$$
\left.<\overrightarrow{\mathrm{k}}^{*}|\mathrm{~V}| \psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle=\frac{2 \pi \mathrm{C}}{\mathrm{~m}}\left[\sin \alpha \cos \beta-\frac{\alpha}{\beta} \cos \alpha \sin \beta\right]
$$

The boundary condition (7.29) can be used to eliminate $\alpha$ and get

$$
\left\langle\overrightarrow{\mathrm{k}}^{*}\right| \mathrm{V}\left|\psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle=\frac{-2 \pi}{\mathrm{~m}} \mathrm{D} m
$$

The normalized amplitude (7.37) is obtained by dividing this by N in (7.36):

$$
\begin{equation*}
\left\langle\overrightarrow{\mathrm{k}}^{*}\right| V\left|\Psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle=-\frac{2 \pi}{\mathrm{~m}} \frac{\mathrm{e}^{\mathrm{i} \beta}}{\sqrt{2 \pi \mathrm{R}\left(1-\frac{\beta^{2}}{\alpha^{2}}\right)\left[1+\frac{1}{\mathrm{i} \beta}\right]}} \tag{7.38}
\end{equation*}
$$

Denote the continuation of this amplitude to real energy by T :

$$
\begin{align*}
T & =\lim _{\widetilde{\mathrm{E}} \rightarrow \text { real }}\left[\left\langle\overrightarrow{\mathrm{k}}^{*}\right| V \mid \mathrm{II}^{(\widetilde{\mathrm{E}})\rangle]}\right.  \tag{7.39}\\
& =-\frac{2 \pi}{\mathrm{~m}} \frac{\mathrm{e}^{\mathrm{i} \beta_{1}}}{\sqrt{2 \pi \mathrm{R}\left[1-\left(\frac{\beta_{1}}{\alpha_{1}}\right)^{2}\right]\left[1+\frac{1}{\mathrm{i} \beta_{1}}\right]}}
\end{align*}
$$

The real part of $\widetilde{E}$ is defined as $E_{p}$ in (7.20). The decay rate is then

$$
\gamma_{\mathrm{F}}=\int \frac{\mathrm{d}^{3} \mathrm{k}}{(2 \pi)^{3}} 2 \pi \delta\left(\frac{\mathrm{k}^{2}}{2 \mathrm{~m}}-\mathrm{E}_{\mathrm{P}}\right)|\mathrm{T}|^{2}
$$

The phase space integration gives

$$
\begin{aligned}
\gamma_{\mathrm{F}} & -\frac{\mathrm{m} \sqrt{2 \mathrm{mE}} \mathrm{P}}{\pi}|\mathrm{~T}|^{2} \\
& =\frac{\mathrm{m} \beta_{1}}{\pi \mathrm{R}}|\mathrm{~T}|^{2}
\end{aligned}
$$

By (7.39) this is

$$
\gamma_{\mathrm{F}}=\left(\frac{2 \beta_{1}}{\mathrm{mR}^{2}} \frac{1}{\left[1-\left(\frac{\beta_{1}}{\alpha_{1}}\right)^{2}\right] \sqrt{1+\left(\frac{1}{\beta_{1}}\right)^{2}}}\right.
$$

Using the explicit values of $\alpha_{1}$ and $\beta_{1}$ in (7.25) gives

$$
\begin{equation*}
\gamma_{F}=\frac{2 \sqrt{b^{2}-1}}{m R^{2}}\left[1+\frac{b^{2}-1}{a^{2}}+\frac{1}{2\left(b^{2}-1\right)}+\cdots\right] \tag{7.40}
\end{equation*}
$$

This value of the width is to be compared with (7.26) for $\gamma_{B}$ and $\gamma_{P}$.

## E. An Exact Test

The idea of introducing a wave function for an unstable particle may be checked exactly. The general expression for the S matrix

$$
\begin{equation*}
\left\langle\overrightarrow{\mathrm{k}}_{2}\right| \mathrm{S}\left|\overrightarrow{\mathrm{k}}_{1}\right\rangle=\left\langle\overrightarrow{\mathrm{k}}_{2} \mid \overrightarrow{\mathrm{k}}_{1}\right\rangle-2 \pi \mathrm{i} \hat{o}\left(\mathrm{E}_{2}-\mathrm{E}_{1}\right)\left\langle\overrightarrow{\mathrm{k}}_{2}\right| \mathrm{V}+\mathrm{V} \frac{1}{\mathrm{E}-\mathrm{H}} \mathrm{~V}\left|\overrightarrow{\mathrm{k}}_{1}\right\rangle \tag{7.41}
\end{equation*}
$$

must have a pole on the second sheet at $\widetilde{\mathrm{E}}$. According to Section II this pole is produced by the resolvent in (7.41) and has a residue determined by the wave function of the unstable state. The scattering amplitude at the pole is therefore predicted to be

$$
\left\langle\overrightarrow{\mathrm{k}}_{2}\right| \mathrm{S} \left\lvert\, \overrightarrow{\mathrm{k}}_{1}>\underset{\mathrm{E} \rightarrow \widetilde{\mathrm{E}}}{\rightarrow-2 \pi \mathrm{i} \delta\left(\mathrm{E}_{2}-\mathrm{E}_{1}\right)} \frac{\mathrm{C}}{\mathrm{E}-\widetilde{\mathrm{E}}}\right.
$$

where $\mathrm{E}=\mathrm{E}_{1}=\mathrm{E}_{2}$ and the residue is

$$
\mathrm{C} \equiv\left\langle\overrightarrow{\mathrm{k}}^{*}\right| \mathrm{V}\left|\Psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle\left\langle\Psi_{\mathrm{II}}\left(\widetilde{\mathrm{E}}^{*}\right)\right| \mathrm{V}|\overrightarrow{\mathrm{k}}\rangle
$$

Because of the time reversal property (7.33)

$$
\left\langle\Psi_{\mathrm{II}}\left(\widetilde{\mathrm{E}}^{*}\right)\right| \mathrm{V}|\overrightarrow{\mathrm{k}}\rangle=\left\langle\overrightarrow{\mathrm{k}}^{*}\right| \mathrm{V}\left|\Psi_{\mathrm{II}}(\widetilde{\mathrm{E}})\right\rangle
$$

This is the matrix element already calculated in (7.38). Hence the residue C is predicted to be

$$
\begin{equation*}
\mathrm{C}=\frac{2 \pi}{\mathrm{~m}^{2} \mathrm{R}} \frac{\mathrm{e}^{2 \mathrm{i} \beta}}{\left(1-\frac{\beta^{2 \backslash}}{\alpha^{2}}\left(1+\frac{1}{\mathrm{i} \beta}\right)\right.} \tag{7.42}
\end{equation*}
$$

The real test of this calculation is the knowledge of the exact $S$ matrix for $\ell=0$. If the calculations that have been done with the exponentially diverging wave functions for unstable states are valid then (7.42) must be an exact result.

According to (7.4) for s -wave scattering

$$
\begin{align*}
\left\langle\overrightarrow{\mathrm{p}}_{2}\right| \mathrm{S}\left|\overrightarrow{\mathrm{p}}_{1}\right\rangle & =(2 \pi) \frac{3}{3} \frac{\delta\left(\mathrm{p}_{2}-\mathrm{p}_{1}\right)}{4 \pi \mathrm{p}_{2} \mathrm{p}_{1}} \eta(\mathrm{E}) \\
& =-2 \pi \mathrm{i} \delta\left(\mathrm{E}_{2}-\mathrm{E}_{1}\right)\left[\frac{\mathrm{i} \pi}{\mathrm{mp}} \eta(\mathrm{E})\right] \tag{7.43}
\end{align*}
$$

On the second sheet

$$
\eta_{I I}(E)=\frac{f(E)+E}{f(E)-E} e^{2 i k R}
$$

where $f(E)$ is given by (7.12). At the pole

$$
\eta_{I I}(\mathrm{E}) \underset{\mathrm{E} \rightarrow \widetilde{\mathrm{E}}}{\longrightarrow} \frac{2 \widetilde{\mathrm{E}}}{(\mathrm{E}-\widetilde{\mathrm{E}})\left[\mathrm{f}^{\mathfrak{\imath}}(\widetilde{\mathrm{E}})-1\right]} \mathrm{e}^{2 \mathbf{i} \beta}
$$

This gives for (7.43)

$$
\left\langle\overrightarrow{\mathrm{p}}_{2}\right| S\left|\overrightarrow{\mathrm{p}}_{1}\right\rangle \underset{\mathrm{E} \rightarrow \widetilde{\mathrm{E}}^{-}}{ }-2 \pi \mathrm{i} \delta\left(\mathrm{E}_{2}-\mathrm{E}_{1}\right) \frac{\mathrm{C}^{\prime}}{\mathrm{E}-\widetilde{\mathrm{E}}}
$$

with

$$
\begin{equation*}
\mathrm{C}^{\prime} \equiv \frac{-\mathrm{i} \pi}{m \widetilde{\mathrm{k}}} \frac{2 \widetilde{\mathrm{E}}}{\left[\mathrm{f}^{\prime}(\widetilde{\mathrm{E}})-1\right]} \mathrm{e}^{2 \mathrm{i} \beta} \tag{7.44}
\end{equation*}
$$

Calculating the required derivative gives

$$
\mathrm{f}^{\prime}(\widetilde{\mathrm{E}})-1=\frac{-\mathrm{i} \beta}{2}\left(1-\frac{\beta^{2}}{\alpha^{2}}\right)\left(1+\frac{1}{\mathrm{i} \beta}\right)
$$

Putting this into (7.44) gives

$$
\begin{equation*}
\mathrm{C}^{\prime}=\frac{2 \pi}{\mathrm{~m}^{2} \mathrm{R}} \frac{\mathrm{e}^{2 \mathrm{i} \beta}}{\left(1-\frac{\beta^{2}}{\alpha^{2}}\right)\left(1+\frac{1}{\mathrm{i} \beta}\right)} \tag{7.45}
\end{equation*}
$$

Hence $\mathrm{C}^{\prime}=\mathrm{C}$ and the wave function calculation is exact.

## APPENDIX A

EXISTENCE OF $\mid \phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}>$
The resonance pole will occur in many scattering amplitudes. To demonstrate the existence of the state $|\phi\rangle$ used in Section II consider, in particular, an elastic scattering amplitude in which the pole occurs in total center-of-mass energy, i.e., in s. This scattering amplitude is obtained from the Fourier transform of a Green's function with an equal number, $n$, of incoming and outgoing lines. Denote this Green's function by

$$
\begin{equation*}
\mathscr{G}\left(\mathrm{x}^{\prime}, \mathrm{x}\right)=\langle 0| \mathrm{T}\left[\mathscr{A}^{\dagger}\left(\mathrm{x}^{\prime}\right) \mathscr{A}(\mathrm{x})\right]|0\rangle \tag{A.1}
\end{equation*}
$$

where

$$
\begin{align*}
& x=\left(x_{1}, x_{2}, \ldots x_{n}\right) \\
& \mathscr{A}(\mathrm{x})=\mathrm{T}\left[\psi_{\alpha_{1}}\left(\mathrm{x}_{1}\right) \psi_{\alpha_{2}}\left(\mathrm{x}_{2}\right) \ldots \psi_{\alpha_{\mathrm{n}}}\left(\mathrm{x}_{\mathrm{n}}\right)\right] \tag{A.2}
\end{align*}
$$

and the $\psi(\mathrm{x})$ are renormalized Heizenberg fields. Let $\mathrm{k}^{\mu}$ be the total energymomentum four vector of the corresponding elastic scattering amplitude (e.g., $s=k^{2}$ ). Introduce average and relative coordinates by

$$
x_{i}=X+r_{i} \quad, \quad \sum_{1}^{n} r_{i}=0
$$

Then X and k are conjugate variables.
By hypothesis, analytic continuation of the scattering amplitude in the variable $k^{\circ}$ leads to a pole on the second sheet independently of the other momentum variables. The pole must therefore occur in the analytic continuation of

$$
\begin{equation*}
G\left(k^{\prime}, k\right) \equiv \int d^{2} X^{\prime} d^{4} X e^{i k^{\prime} \cdot X^{\prime}} \mathscr{G}\left(x^{\prime}, x\right) e^{-i k \cdot X} \tag{A.3}
\end{equation*}
$$

independently of the $2 \mathrm{n}-2$ relative coordinates r and $\mathrm{r}^{\prime}$. Among the ( 2 n ) : time orderings in (A.1), isolate the $(n!)^{2}$ terms for which $\left(x_{i}\right)^{0}>\left(x_{j}\right)^{0}$, for all i and $j$,
by separating

$$
\left.\mathscr{G}\left(\mathrm{x}^{\prime}, \mathrm{x}\right)=<0\left|\mathscr{A}^{\dagger}\left(\mathrm{x}^{\prime}\right) \theta\left(\mathrm{t}_{\min }^{\prime}-\mathrm{t} \max \right) \mathscr{A}(\mathrm{x})\right| 0\right\rangle+\Delta \mathscr{G}\left(\mathrm{x}^{\top}, \mathrm{x}\right)
$$

When there are only two fields in $\mathscr{A}$,

$$
\begin{aligned}
& t_{\max }=\frac{1}{2}\left(t_{1}+t_{2}\right)+\frac{1}{2}\left|t_{1}-t_{2}\right| \\
& t_{\min }=\frac{1}{2}\left(t_{1}+t_{2}\right)-\frac{1}{2}\left|t_{1}-t_{2}\right| .
\end{aligned}
$$

Generally for n fields

$$
\begin{aligned}
& \mathrm{t}_{\max }=\mathrm{x}^{\mathrm{o}}+\alpha\left(\mathrm{r}_{1}{ }^{\mathrm{o}}, \mathrm{r}_{2}^{\mathrm{o}}, \ldots \mathrm{r}_{\mathrm{n}}{ }^{\mathrm{o}}\right) \\
& { }^{\mathrm{t}} \min =\mathrm{X}^{\mathrm{o}}+\beta\left(\mathrm{r}_{1}^{\mathrm{o}}, \mathrm{r}_{2}^{\mathrm{o}}, \ldots \mathrm{r}_{\mathrm{n}}{ }^{\mathrm{o}}\right) .
\end{aligned}
$$

This separation and the fact that

$$
\begin{equation*}
\mathscr{A}(\mathrm{X}, \mathrm{r})=\mathrm{e}^{i \mathrm{X}^{\mathrm{O}} \mathrm{H}} \mathscr{A}(\overrightarrow{\mathrm{X}}, \mathrm{r}) \mathrm{e}^{-\mathrm{i} \mathrm{X}^{\mathrm{O}} \mathrm{H}} \tag{A.4}
\end{equation*}
$$

allows the Fourier transform (A.3) to be performed. The result is

$$
\begin{align*}
\mathrm{G}\left(\mathrm{k}^{\mathrm{\prime}}, \mathrm{k}\right)=\left\langle\phi\left(\overrightarrow{\mathrm{k}^{\mathrm{r}}}\right)\right. & \left.\mathrm{e}^{-\mathrm{i} \beta\left(\mathrm{k}^{\circ}-\mathrm{H}\right)} \frac{2 \pi \mathrm{i} \delta\left(\mathrm{k}^{\mathrm{o}}-\mathrm{k}^{\mathrm{o}}\right)}{\mathrm{k}^{\mathrm{o}}-\mathrm{H}} \mathrm{e}^{\mathrm{i} \alpha\left(\mathrm{k}^{\mathrm{O}}-\mathrm{H}\right)} \right\rvert\, \phi(\overrightarrow{\mathrm{k}})> \\
& +\Delta \mathrm{G}\left(\mathrm{k}^{\mathrm{t}}, \mathrm{k}\right) \tag{A.5}
\end{align*}
$$

where

$$
\begin{gather*}
\operatorname{Im} \mathrm{k}^{\mathrm{o}}>0 \\
\alpha=\alpha\left(\mathrm{r}_{\mathrm{i}}^{\mathrm{o}}\right) \\
\beta=\beta\left(\mathrm{r}_{\mathrm{i}}^{\prime}\right) \\
\left|\phi(\overrightarrow{\mathrm{k}})>\equiv \int \mathrm{d}^{3} \mathrm{X}^{\mathrm{o}} \mathrm{e}^{\mathrm{i} \overrightarrow{\mathrm{k}} \cdot \overrightarrow{\mathrm{X}}} \mathscr{A}(\overrightarrow{\mathrm{X}}, \mathrm{r})\right| 0> \tag{A.6}
\end{gather*}
$$

Because the resonance pole in the scattering amplitude is on the second sheet, it can only come from terms in (A.5) that have a branch point in $\mathrm{k}^{0}$. The displayed term has branch points along the positive real $\mathrm{k}^{\circ}$ axis coming from thresholds in $H$. Of the terms in $\Delta G,(n!)^{2}$ of them have cuts only along the
negative real $k^{0}$ axis coming from $\left(k^{\circ}+H\right)^{-1}$. None of the other terms in $\Delta G$ have branch points at all because they come from finite ranges of integration of the variable $\mathrm{X}^{{ }^{\circ}}-\mathrm{X}^{\circ}$. Thus the unstable particle pole can only come from the analytic continuation of the first term of (A.5). Since this term has a pole then so does the analytic continuation of

$$
\begin{equation*}
\left\langle\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\right| \frac{1}{\mathrm{k}^{\mathrm{O}}-\mathrm{H}}|\phi(\overrightarrow{\mathrm{k}})\rangle \tag{A.7}
\end{equation*}
$$

Because the spatial dependence of the Heisenberg fields is

$$
\mathscr{A}(\overrightarrow{\mathrm{X}}, \mathrm{r})=\mathrm{e}^{-\mathrm{i} \overrightarrow{\mathrm{X}} \cdot \overrightarrow{\mathscr{P}}} \mathscr{A}(0, \mathrm{r}) \mathrm{e}^{\mathrm{i} \overrightarrow{\mathrm{X}} \cdot \overrightarrow{\mathscr{P}}}
$$

where $\overrightarrow{\mathscr{F}}$ is the full momentum operator. This means that $|\phi\rangle$ in (A.6) is an eigenstate of momentum:

$$
\begin{equation*}
\overrightarrow{\mathscr{P}}|\phi(\overrightarrow{\mathrm{k}})\rangle=\overrightarrow{\mathrm{k}}|\phi(\overrightarrow{\mathrm{k}})\rangle . \tag{A.8}
\end{equation*}
$$

As yet $|\phi\rangle$ has $n$ discrete field indices ( $\alpha_{1}, \alpha_{2}, \ldots \alpha_{n}$ ) that are suppressed in (A.6). These indices may be coupled in pairs with ordinary Clebsch-Gordon coefficients leading to an n-fold Clebsch-Gordon coefficient that satisfies
where $\mathscr{D}^{l}$ are the rotation appropriate to the field representation (usually $(\ell, 0) \oplus(0, \ell))$. The final form for $|\phi\rangle$ is then

$$
\begin{align*}
& \left|\phi(\vec{k}) j, j_{3}\right\rangle-\sum_{\alpha} \int\left(d^{4} x\right)^{n} \delta\left(X^{o}\right) e^{i \vec{k} \cdot x}\left\{T\left[\psi_{\alpha_{1}}\left(x_{1}\right) \ldots \psi_{\alpha_{n}}\left(x_{n}\right)\right]|0\rangle\right. \\
& \left.\left(\begin{array}{llll}
\ell_{1} & & \ell_{n} & j \\
\alpha_{1} & \cdots & \alpha_{n} & ; \\
j_{3}
\end{array}\right) f\left(r_{1}, \ldots r_{n}\right)\right\} . \tag{A.9}
\end{align*}
$$

Here $f$ is a convenient smearing function of Lorentz invariants $r^{2}$ and $r_{i} \cdot r_{j}$. If $|\phi\rangle$ in (A.6) gives a pole in (A. 7) then at least one of the ( $\mathrm{j}, \mathrm{j}_{3}$ ) combinations in
(A.9) must also because of the completeness of the Clebsch-Gordon coefficients. Under rotations (A.9) transforms as

$$
\mathrm{U}[\mathrm{R}]\left|\phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}>=\sum_{\mathrm{j}_{3}^{\prime}}\right| \phi(\mathrm{Rk}) \mathrm{j}_{3} \mathrm{j}_{3}^{\mathrm{j}}>\mathrm{D}_{\mathrm{j}_{3}^{\prime}, \mathrm{j}_{3}}{ }^{(\mathrm{R})} .
$$

Also,

$$
\begin{aligned}
& \overrightarrow{\mathrm{J}}^{2}\left|\phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}>=\mathrm{j}(\mathrm{j}+1)\right| \phi(\overrightarrow{\mathrm{k}}) \mathrm{j}_{,} \mathrm{j}_{3}> \\
& \mathrm{J}_{3}\left|\phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}>=\mathrm{j}_{3}\right| \phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}>
\end{aligned}
$$

The function f may be chosen to guarantee the normalization

$$
<\mathrm{j}_{3}^{\mathrm{t}}, \mathrm{j}^{\prime} \phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\left|\phi(\overrightarrow{\mathrm{k}}) \mathrm{j}, \mathrm{j}_{3}\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}^{\prime}}-\overrightarrow{\mathrm{k}}\right) \delta_{\mathrm{j}^{\prime}, \mathrm{j}^{2}} \delta_{3}, \mathrm{j}_{3}
$$

## APPENDIX B

## RESOLVENT IDENTITIES

The identities (2.8) and (2.12) follow very easily from the definitions

$$
\begin{aligned}
& R\left(k^{0}\right)=\frac{1}{k^{0}-H} \\
& r\left(k^{0}\right)=\frac{1}{k^{0}-H^{\prime}}
\end{aligned}
$$

The full resolvent satisfies the three integral equations

$$
\begin{aligned}
& R\left(k^{0}\right)=r\left(k^{0}\right)+R\left(k^{0}\right) H^{\prime \prime} r\left(k^{0}\right) \\
& R\left(k^{0}\right)=r\left(k^{0}\right)+r\left(k^{0}\right) H^{\prime \prime} R\left(k^{0}\right) \\
& R\left(k^{0}\right)=r\left(k^{0}\right)+r\left(k^{0}\right) H^{\prime \prime} r\left(k^{0}\right)+r\left(k^{0}\right) H^{\prime \prime} R\left(k^{0}\right) H^{\prime \prime} r\left(k^{0}\right) .
\end{aligned}
$$

Use the fact that $\mathrm{r}\left(\mathrm{k}^{\mathrm{o}}\right)$ commutes with A and B to calculate the projections

$$
\begin{aligned}
& \operatorname{AR}\left(k^{0}\right) B=\left[\operatorname{AR}\left(k^{0}\right) A\right] H^{\prime \prime} r\left(k^{0}\right) \\
& \operatorname{BR}\left(k^{0}\right) A=r\left(k^{0}\right) H^{\prime \prime}\left[\operatorname{AR}\left(k^{0}\right) A\right] \\
& \operatorname{BR}\left(k^{0}\right) B=\operatorname{Br}\left(k^{0}\right) B+r\left(k^{0}\right) H^{\prime \prime}\left[\operatorname{AR}\left(k^{0}\right) A\right] H^{\prime \prime} r\left(k^{o}\right) .
\end{aligned}
$$

Sum these three projections together with $A R\left(k^{0}\right) A$ itself to get

$$
\begin{equation*}
\mathrm{R}\left(\mathrm{k}^{\mathrm{o}}\right)=\mathrm{Br}\left(\mathrm{k}^{\mathrm{O}}\right) \mathrm{B}+\left[1+\mathrm{r}\left(\mathrm{k}^{\mathrm{O}}\right) \mathrm{H}^{\prime}\right] \operatorname{AR}\left(\mathrm{k}^{\mathrm{O}}\right) \mathrm{A}\left[\mathrm{H}^{\prime \prime} \mathrm{r}\left(\mathrm{k}^{\mathrm{O}}\right)+1\right] \tag{B.1}
\end{equation*}
$$

as claimed in (2.8).
Now to simplify $\operatorname{AR}\left(\mathrm{k}^{\mathrm{O}}\right)$ A use the integral equation

$$
R\left(k^{\circ}\right)=r\left(k^{0}\right)+r\left(k^{0}\right) H^{\prime \prime} r\left(k^{\circ}\right)+r\left(k^{0}\right) H^{\prime \prime} r\left(k^{O}\right) H^{\prime \prime} R\left(k^{\circ}\right) .
$$

The diagonal projection of this is

$$
\begin{equation*}
A R\left(k^{0}\right) A=\frac{A}{k^{0}-E_{\phi}}+\frac{1}{k^{0}-E_{\phi}} H^{\prime \prime} r\left(k^{0}\right) H^{\prime \prime}\left[A R\left(k^{0}\right) A\right] \tag{B.2}
\end{equation*}
$$

where $\mathrm{E}_{\phi}$ is the expectation value

$$
\left.\left\langle\phi\left(\overrightarrow{\mathrm{k}}_{1}\right)\right| \mathrm{H}^{\prime}\left|\phi\left(\overrightarrow{\mathrm{k}}_{2}\right)\right\rangle=\mathrm{E}_{\phi} \delta^{3} \overrightarrow{(\vec{k}}_{1}-\overrightarrow{\mathrm{k}}_{2}\right) .
$$

Rewrite (B. 2) as

$$
\left[\mathrm{k}^{\mathrm{o}}-\mathrm{E}_{\phi}-\mathrm{H}^{\prime \prime} \mathrm{r}\left(\mathrm{k}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime}\right] \mathrm{AR}\left(\mathrm{k}^{\mathrm{o}}\right) \mathrm{A}=\mathrm{A} .
$$

The expectation value of this is

$$
\begin{equation*}
\int \mathrm{d}^{3} \mathrm{k}_{2}<\phi\left(\overrightarrow{\mathrm{k}}_{1}\right)\left|\mathrm{k}^{\mathrm{o}}-\mathrm{H}^{\prime}-\mathrm{H}^{\prime \prime} \mathrm{r}\left(\mathrm{k}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime}\right| \phi\left(\overrightarrow{\mathrm{k}}_{2}\right)><\left.\phi\left(\overrightarrow{\mathrm{k}}_{2}\right)\right|_{\mathrm{k}^{\mathrm{O}}-\mathrm{H}}\left|\phi\left(\overrightarrow{\mathrm{k}}_{3}\right)\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}}_{1}-\overrightarrow{\mathrm{k}}_{3}\right) \tag{B.3}
\end{equation*}
$$

The definition of $D(k)$ in (2.7) is

$$
<\phi\left(\dot{\overrightarrow{\mathrm{k}}}_{2}\right)\left|\frac{1}{\mathrm{k}^{\mathrm{o}}-\mathrm{H}}\right| \phi\left(\overrightarrow{\mathrm{k}}_{3}\right)>\equiv \frac{\delta^{3}\left(\overrightarrow{\mathrm{k}}_{2}-\overrightarrow{\mathrm{k}}_{3}\right)}{\mathrm{D}\left(\mathrm{k}_{2}\right)} .
$$

Therefore (B. 3) shows that an equivalent formula for $D(k)$ is

$$
\left.<\phi\left(\overrightarrow{\mathrm{k}}_{1}\right)\left|\mathrm{k}^{\mathrm{o}}-\mathrm{H}^{\prime}-\mathrm{H}^{\prime \prime} \mathrm{r}\left(\mathrm{k}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime}\right| \phi\left(\overrightarrow{\mathrm{k}}_{2}\right)\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}}_{1}-\overrightarrow{\mathrm{k}}_{2}\right) \mathrm{D}(\mathrm{k}),
$$

as claimed in (2.12).

## APPENDIX C

## THE TOPOLOGY OF $\not \subset \uparrow$

Because the unstable particles have zero norm a new, positive-definite norm operator must be found to give meaning to (3.15). This norm should be positive-definite both in the unstable particle sector of $h \uparrow$ as well as in the physical Hilbert space $h \subset \not \subset$. The difficulty is that certain states like $\mid \phi>$ are in the Hilbert space $h$ and yet are not orthogonal to the unstable states. In particular,

$$
\begin{equation*}
<\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\left|\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}^{\prime}}-\overrightarrow{\mathrm{k}}\right) \tag{C.1}
\end{equation*}
$$

Essentially the problem is that the three states $\left|\psi_{n}\right\rangle,\left|\psi_{\bar{n}}\right\rangle$, and $|\phi\rangle$ are linearly independent but not orthogonal. To construct an orthogonal basis is straightforward. First let

$$
\begin{equation*}
\mathrm{I}_{1 \mathrm{u}}={ }_{\mathrm{d}}{ }^{3} \mathrm{k}\left\{\left|\Psi(\mathrm{k})><\Psi\left(\mathrm{k}^{*}\right)\right|+\left|\Psi\left(\mathrm{k}^{*}\right)><\Psi(\mathrm{k})\right|\right\} \tag{C.2}
\end{equation*}
$$

This is the identity operator in the one unstable particle sector:

$$
\begin{align*}
& \mathrm{I}_{1 \mathrm{u}}|\Psi(\mathrm{k})\rangle=|\Psi(\mathrm{k})\rangle  \tag{C.3a}\\
& \mathrm{I}_{1 \mathrm{u}}\left|\Psi\left(\mathrm{~K}^{*}\right)\right\rangle=\left|\Psi\left(\mathrm{k}^{*}\right)\right\rangle \tag{C.3b}
\end{align*}
$$

The projection of $|\phi\rangle$ along the unstable sector is

$$
\mathrm{I}_{1 \mathrm{u}}|\phi(\overrightarrow{\mathrm{k}})\rangle=|\Psi(\mathrm{k})\rangle \frac{1}{\sqrt{\mathrm{~N}}}+\left|\Psi\left(\mathrm{K}^{*}\right)\right\rangle \frac{1}{\sqrt{\mathrm{~N}^{*}}}
$$

with $N$ taken from (3.12). The part of $|\phi\rangle$ orthogonal to the unstable particle sector is just

$$
\begin{equation*}
\mathrm{I}_{1 \mathrm{~s}}|\phi(\overrightarrow{\mathrm{k}})\rangle=|\phi(\overrightarrow{\mathrm{k}})\rangle-|\Psi(\mathrm{k})\rangle \frac{1}{\sqrt{\mathrm{~N}}}+\left|\Psi\left(\mathrm{k}^{*}\right)\right\rangle \frac{1}{\sqrt{\mathrm{~N}^{*}}} \tag{C.3c}
\end{equation*}
$$

An orthogonal basis is therefore provided by the three states (C. 3a), (C.3b), and (C. 3c).

The identity operator (C.2) for the unstable particle sector may be extended in a hermitian manner to include many unstable particles or mixtures of stable particles and at least one unstable particle. With the extension $I_{u}$ so defined let

$$
I_{s}|F\rangle \equiv \begin{cases}|F\rangle-I_{u}|F\rangle & \text { if }|F\rangle \epsilon h  \tag{C.4}\\ 0 & \text { if }|F\rangle \notin h\end{cases}
$$

Thus $I_{S}$ projects out what may be called the very stable part of the Hilbert space h. The identity operator in the full space $h \uparrow$ is then

$$
\begin{equation*}
I=I_{S}+I_{u} \tag{C.5}
\end{equation*}
$$

where $I_{s} I_{u}=I_{u} I_{s}=0$. The important point is that the first term of (C.5) is $I_{s}$ and not the usual identity in h. Note, too, that (C.5) assumes that the eigenstates of $H \operatorname{span} \not \hbar$.

Before constructing a positive definite norm operator (see (C.13)), it is worthwhile to investigate in more detail why the inner product (C.1) does not vanish. To do this, express $|\phi\rangle$ as a superposition of energy eigenstates. By assumption the asymptotic scattering states, either in or out, span $h$. Thus put

$$
\begin{equation*}
|\phi(\overrightarrow{\mathrm{k}})\rangle=\int \mathrm{dE}|\mathrm{E}, \overrightarrow{\mathrm{k}}\rangle^{\text {in }} \varphi(\mathrm{E}) \tag{C.6}
\end{equation*}
$$

The state $|\mathrm{E}, \overrightarrow{\mathrm{k}}\rangle^{\text {in }}$ is a direct product of asymptotic stable particle states (e, P, $\gamma, \nu_{e}, \nu_{\mu}$ ) with total energy $E$ and total momentum $\vec{k}$. (The relative energies and momenta are suppressed.) The state $|\phi\rangle$ was specially chosen in Section II and Appendix A so that a pole would occur in the matrix elements

$$
\frac{\delta^{3}\left(\overrightarrow{\mathrm{k}^{\mathrm{t}}}-\overrightarrow{\mathrm{k}}\right)}{\mathrm{D}(\mathrm{k})}=\left\langle\phi\left(\overrightarrow{\mathrm{k}^{\mathrm{l}}}\right)\right| \frac{1}{\mathrm{k}^{0}-\mathrm{H}}|\phi(\overrightarrow{\mathrm{k}})\rangle
$$

Because of (C.6)

$$
\begin{equation*}
\frac{1}{D(k)}=\int d E \frac{\varphi(E)^{*} \varphi(E)}{k^{o}-E} \tag{C.7}
\end{equation*}
$$

Analytically continuing the $\mathrm{k}^{\mathrm{o}}$ dependence of (C.7) clockwise into the lower half plane merely distorts the dE contour ahead of it. A pole can occur in (C. 7) at $\mathrm{k}^{\mathrm{O}}=\widetilde{\mathrm{k}}^{\mathrm{o}}$ only if the contour is trapped against a fixed pole of the integrand at $\mathrm{E}=\widetilde{\mathrm{k}}^{0}$ in the lower half plane. Thus

$$
\begin{equation*}
\varphi^{*}(\mathrm{E}) \varphi(\mathrm{E}) \rightarrow \frac{\mathrm{c}}{\mathrm{E}-\widetilde{\mathrm{k}}^{\mathrm{o}}} \tag{C.8}
\end{equation*}
$$

where c is finite but nonvanishing. The pole in (C.8) cannot be of higher order because the pole in (C. 7) would not be first order then. Furthermore, $\varphi(E)$ cannot have ad hoc poles and cuts because they would produce spurious singularities in (C.6) and hence in the S-matrix. Hence it is the pole in the weighting function (C. 8) that allows certain states in the physical Hilbert space $h$ to have a nonzero overlap with the unstable particle states.

All this is in preparation for discussing the norm. For a general state

$$
\begin{equation*}
|\mathrm{G}\rangle=\int \mathrm{d}^{3} \mathrm{k}\left|\Psi(\mathrm{k})>\mathrm{G}_{\mathrm{A}}(\overrightarrow{\mathrm{k}})+\int \mathrm{d}^{3} \mathrm{k}\right| \Psi\left(\mathrm{k}^{*}\right)>\mathrm{G}_{\mathrm{B}}(\overrightarrow{\mathrm{k}}) \tag{C.9}
\end{equation*}
$$

the inner product norm is

$$
\begin{equation*}
\langle\mathrm{G} \mid \mathrm{G}\rangle=\int \mathrm{d}^{3} \mathrm{k}\left\{\mathrm{G}_{\mathrm{A}}^{*} \mathrm{G}_{\mathrm{B}}+\mathrm{G}_{\mathrm{A}} \mathrm{G}_{\mathrm{B}}^{*}\right\} \tag{C.10}
\end{equation*}
$$

This norm is, of course, real but not necessarily positive. Clearly (C. 10) is just the expectation value of (C. 2):

$$
\langle G \mid G\rangle=\langle G| I_{1 u}|G\rangle .
$$

This suggests defining a diagonal operator

$$
\begin{equation*}
\Omega_{1 \mathrm{u}} \equiv \int \mathrm{~d}^{3} \mathrm{k}\left\{|\Psi(\mathrm{k})\rangle\langle\Psi(\mathrm{k})|+\left|\Psi\left(\mathrm{k}^{*}\right)\right\rangle\left\langle\Psi\left(\mathrm{k}^{*}\right)\right|\right\} \tag{C.11}
\end{equation*}
$$

The expectation value

$$
\begin{equation*}
\langle G| \Omega_{1 u}|G\rangle=\int d^{3} k\left\{\left|G_{A}\right|^{2}+\left|G_{B}\right|^{2}\right\} \tag{C.12}
\end{equation*}
$$

is then positive-definite. Again (C. 11) can be generalized to the entire unstable sector. Thercfore definc the full norm in $\not \subset \uparrow$ by

$$
\begin{gather*}
\Omega=\mathrm{I}_{\mathrm{S}}+\Omega_{\mathrm{u}}  \tag{C.13}\\
\|\mathrm{~F}\|^{2} \equiv\langle\mathrm{~F}| \Omega|\mathrm{F}\rangle \text { for }|\mathrm{F}\rangle \epsilon \not \subset \uparrow
\end{gather*}
$$

To see how (C.13) acts, calculate the norm of $|\phi\rangle$. From (C.11)

$$
\langle\phi(\overrightarrow{\mathrm{k}})| \Omega_{\mathrm{u}}\left|\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\right\rangle=\left(\frac{1}{\mathrm{~N}}+\frac{1}{\mathrm{~N}^{*}}\right) \delta^{3}\left(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}^{\prime}\right) .
$$

This, of course, is not generally positive. To calculate the $I_{S}$ contribution to the norm use

$$
\left.<\phi(\overrightarrow{\mathrm{k}})\left|\mathrm{I}_{\mathrm{S}}\right| \phi\left(\overrightarrow{\mathrm{k}}^{\prime}\right)\right\rangle=\left\{\langle\phi(\overrightarrow{\mathrm{k}})| \mathrm{I}_{\mathrm{S}}\right\}\left\{\mathrm{I}_{\mathrm{S}}|\phi(\overrightarrow{\mathrm{k}})\rangle\right\} .
$$

Substituting (C. 3c) gives

$$
\left.\langle\phi(\overrightarrow{\mathrm{k}})| \mathrm{I}_{\mathrm{s}}\left|\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\right\rangle=\left(1-\frac{1}{\mathrm{~N}}-\frac{1}{\mathrm{~N}^{*}}\right) \delta^{3} \overrightarrow{(\overrightarrow{\mathrm{k}}}-\overrightarrow{\mathrm{k}^{\prime}}\right)
$$

so that

$$
\langle\phi(\overrightarrow{\mathrm{k}})| \Omega\left|\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\right\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}^{\prime}\right) .
$$

This shows that for states in $\not \subset$ the $\Omega$ norm agrees with the usual norm in $\not \approx$ :

$$
\langle\mathrm{F}| \Omega|\mathrm{F}\rangle=\langle\mathrm{F} \mid \mathrm{F}\rangle \quad \text { for }|\mathrm{F}\rangle \epsilon \hbar .
$$

Thus $\|F\|$ is a genuine norm. It satisfies
a) $\quad\|F\| \geq 0 ; \quad\|F\|=0 \quad$ iff $\quad|F\rangle=0$
b) $\quad\|\alpha \mathrm{F}\|=|\alpha| \cdot\|\mathrm{F}\|$
c)

$$
\left\|F_{1}+F_{2}\right\| \leq\left\|F_{1}\right\|+\left\|F_{2}\right\|
$$

A further bonus is that $h \uparrow$ is Cauchy complete in this norm. More precisely, any sequence of states $\mid F_{n}>$ of the form (C.9) that is Cauchy complete, i.e., for which

$$
\left\|\left|F_{n}>-\right| F_{m}>\right\| \rightarrow 0 \quad \text { as } n, m \rightarrow \infty
$$

must necessarily have a limit $\mid F>\epsilon \nmid \uparrow$. This is a trivial consequence of (C.12) and the fact that square integrable functions are Cauchy complete. The more substantive version of completeness is why only states of the form (C. 10) should be considercd. This is just the question of whether the eigenstates of $H$ span $h \uparrow$ previously mentioned.

The definition of the norm operator in (C.13) is not unique. For example, define linear combinations

$$
\begin{aligned}
& \left|\pi_{1}(\mathrm{k})>\equiv\right| \Psi(\mathrm{k})>\alpha+\mid \Psi\left(\mathrm{k}^{*}\right)>\beta^{*} \\
& \left|\pi_{2}(\mathrm{k})>\equiv\right| \Psi(\mathrm{k})>\frac{|\alpha \beta|}{\beta}-\left\lvert\, \Psi\left(\mathrm{k}^{*}\right)>\frac{|\alpha \beta|}{\alpha^{*}}\right.
\end{aligned}
$$

where $\alpha$ and $\beta$ are complex numbers satisfying

$$
\alpha \beta+\alpha^{*} \beta^{*}=1
$$

The $\mid \pi>$ have inner products

$$
\begin{aligned}
& <\pi_{1}\left(\mathrm{k}^{\prime}\right) \mid \pi_{1}(\mathrm{k})>=\delta^{3} \overrightarrow{\left(\mathrm{k}^{\prime}-\overrightarrow{\mathrm{k}}\right)} \\
& <\pi_{2}\left(\mathrm{k}^{\prime}\right) \mid \pi_{2}(\mathrm{k})>=-\delta^{3} \overrightarrow{\left(\mathrm{k}^{\prime}-\overrightarrow{\mathrm{k}}\right)} \\
& <\pi_{1}\left(\mathrm{k}^{\prime}\right) \mid \pi_{2}(\mathrm{k})>=0
\end{aligned}
$$

The identity operator (C.2) written in terms of these new states is

$$
\begin{equation*}
\mathrm{I}_{1 \mathrm{u}}=\int \mathrm{d}^{3} \mathrm{k}\left\{\left|\pi_{1}(\mathrm{k})><\pi_{1}(\mathrm{k})\right|-\left|\pi_{2}(\mathrm{k})><\pi_{2}(\mathrm{k})\right|\right\} \tag{C.14}
\end{equation*}
$$

The norm operator $\Omega$ in (C.11) is diagonal in the $|\Psi\rangle$ basis but not in the $|\pi\rangle$ basis. The minus sign in (C.14) suggests defining a new positive definite norm operator that is diagonal in $\mid \pi>$ :

$$
\Omega^{\prime} \equiv \int \mathrm{d}^{3} \mathrm{k}\left\{\left|\pi_{1}(\mathrm{k})><\pi_{1}(\mathrm{k})\right|+\left|\pi_{2}(\mathrm{k})><\pi_{2}(\mathrm{k})\right|\right\}
$$

Obviously $\mid \pi_{2}>$ here is formally analogous to the timelike photons of the GuptaBleuler quantization of QED and $\Omega^{\prime}$ is analogous to their $\eta$ trick.

Sets which are open in the $\Omega$ norm are also open in the $\Omega^{\prime}$ norm because the two are connected by the transformation

$$
\mathrm{T} \equiv \int \mathrm{~d}^{3} \mathrm{k}\left\{\left|\pi_{1}(\mathrm{k})\right\rangle\langle\Psi(\mathrm{k})|+\left|\pi_{2}(\mathrm{k})\right\rangle\left\langle\Psi\left(\mathrm{k}^{*}\right)\right|\right\}
$$

in that

$$
\Omega=\mathrm{T}^{+} \Omega^{\prime} \mathrm{T}
$$

Because of (C. 2)

$$
\mathrm{T}^{+} \mathrm{T}=\mathrm{I}_{1 \mathrm{u}}
$$

Because of (C. 14)

$$
\mathrm{TT}^{+}=\mathrm{I}_{1 \mathrm{u}}
$$

Thus T is unitary. The topology of $\boldsymbol{\kappa} \uparrow$ is therefore not changed by such a change in the definition of norm. In particular if

$$
\langle\Delta F| \Omega|\Delta F\rangle=0
$$

then

$$
\langle\Delta F| \Omega^{\prime}|\Delta F\rangle=0 .
$$

Therefore

$$
\left|F_{1}\right\rangle=\left|F_{2}\right\rangle
$$

has an invariant meaning.

## APPENDIX D

## MULTIPLE POLES AND MULTIPOLE GHOSTS

The analysis of Section III applies to any first order pole in the scattering amplitude. It is well known that unitarity and the asymptotic condition forbid multiple poles on the physical sheet. ${ }^{28}$ No such prohibition applies to unphysical sheets. ${ }^{13}$ An analysis similar to Section III will show that whenever there are multiple poles the state space $\not \uparrow \uparrow$ necessarily contains some curious new states.

The term multipole ghosts originated in Heisenberg's indefinite metric quantization of the point source Lee model. ${ }^{16}$ In such a quantization the usual prohibition does not apply and, indeed the scattering amplitude has a second order pole on the real axis. (See Section III.B.) Associated with the double pole are two states: an energy eigenstate and a ghost of that state. Nakanishi has discussed the possibility of multiple ghost states generally in an indefinite metric quantization and shown that if there are such states the scattering amplitude has a multiple pole on the physical sheet. ${ }^{29}$

This discussion borrows the name multipole ghosts but differs in two respects. First, the existence of the states is deduced from the existence of the multiple poles (rather than conversely). Second, the quantization is conventional in that the multiple poles are only on unphysical sheets.

Suppose then that the scattering amplitude contains a pole of order $L$ on the second sheet at $\mathrm{k}^{\mathrm{O}}=\mathrm{N}^{0}$. This sheet is reached just as in (3.3) by continuing clockwise around a particular n-particle branch point. At the pole

$$
\mathrm{D}_{\mathrm{n}}(\widetilde{\mathrm{~K}})=0
$$

Furthermore, because the zero is of order $L$

$$
\left.\frac{d D_{n}(k)}{d k^{o}}\right|_{\widetilde{k}}=0
$$

and generally

$$
\frac{\mathrm{d}^{\ell-1} \mathrm{D}_{\mathrm{n}}(\mathrm{k})}{\mathrm{d}\left(\left.\mathrm{k}^{\mathrm{o})^{l-1}}\right|_{\widetilde{k}}=0 \quad \ell=1,2, \ldots \mathrm{~L} . . . . . . . . .\right.}
$$

Using expression (2.12) for $D(k)$, these three equations are

$$
\begin{align*}
& \left.<\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right) \mid \widetilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}^{\prime}-\mathrm{H}^{\prime \prime} \mathrm{r}_{\mathrm{n}} \widetilde{\mathrm{k}}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime} \mid \phi(\overrightarrow{\mathrm{k}})>=0  \tag{D.1a}\\
& <\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\left|1+\mathrm{H}^{\prime \prime}\left[r_{\mathrm{n}}\left(\widetilde{\mathrm{k}}^{\mathrm{o}}\right)\right]^{2} \mathrm{H}^{\prime \prime}\right| \phi(\overrightarrow{\mathrm{k}})>=0  \tag{D.1b}\\
& \left.<\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right) \mid \mathrm{H}^{\prime \prime}\left[r_{\mathrm{n}} \widetilde{\mathrm{k}}^{\mathrm{o}}\right)\right]^{\ell} \mathrm{H}^{\prime \prime} \mid \phi(\overrightarrow{\mathrm{k}})>=0 \quad \ell=3,4, \ldots \mathrm{~L} . \tag{D.1c}
\end{align*}
$$

The $L^{1}$ th derivative of $D(k)$ does not vanish. Let it be

$$
\begin{equation*}
<\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\left|\mathrm{H}^{\prime \prime}\left[\mathrm{r}_{\mathrm{n}}\left(\mathrm{k}^{\mathrm{o}}\right)\right]^{\mathrm{L}+1} \mathrm{H}^{\prime \prime}\right| \phi(\overrightarrow{\mathrm{k}})>\equiv \mathrm{C} \delta\left(\overrightarrow{\mathrm{k}^{\prime}}-\overrightarrow{\mathrm{k}}\right) \tag{D.2}
\end{equation*}
$$

where C is a finite but nonvanishing constant analogous to N in (3.12).
The analytic continuation of (2.13) is still

$$
\left(\mathrm{k}^{\mathrm{o}}-\mathrm{H}\right)\left|\psi_{\mathrm{n}}(\mathrm{k})\right\rangle=\mid \phi(\overrightarrow{\mathrm{k}})>\mathrm{D}_{\mathrm{n}}(\mathrm{k})
$$

Thus

$$
\left(\widetilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}\right)\left|\psi_{\mathrm{n}}(\widetilde{\mathrm{k}})\right\rangle=0
$$

as before. Now, however, the equation

$$
\left(\mathrm{k}^{\mathrm{o}}-\mathrm{H}\right)^{2}\left|\psi_{\mathrm{n}}(\mathrm{k})\right\rangle=\left(\mathrm{k}^{\mathrm{o}}-\mathrm{H}\right) \mid \phi(\overrightarrow{\mathrm{k}})>\mathrm{D}_{\mathrm{n}}(\mathrm{k})
$$

contains new information. Differentiate this with respect to $\mathrm{k}^{\circ}$ and then set $\mathrm{k}^{\mathrm{o}}=\widetilde{\mathrm{k}}^{\mathrm{o}}$ :

$$
\left.\widetilde{\mathrm{k}}^{0}-\mathrm{H}\right)^{2}\left[\left.\frac{\mathrm{~d} \mid \psi_{\mathrm{n}}(\mathrm{k})>}{\mathrm{dk}}{ }^{0}\right|_{\widetilde{\mathrm{k}}}\right]=0
$$

Generally if the equation

$$
\left(\mathrm{k}^{\mathrm{o}}-\mathrm{H}\right)^{\ell}\left|\psi_{\mathrm{n}}(\mathrm{k})\right\rangle=\left(\mathrm{k}^{\mathrm{o}}-\mathrm{H}\right)^{\ell-1} \mid \phi(\overrightarrow{\mathrm{k}})>\mathrm{D}_{\mathrm{n}}(\mathrm{k})
$$

is differentiated $\ell-1$ times the result is

$$
\left.\widetilde{(\mathrm{k}}^{\mathrm{o}}-\mathrm{H}\right)^{\ell}\left[\left.\frac{\mathrm{d}^{\ell-1} \mid \psi_{\mathrm{n}}(\mathrm{k})>}{\mathrm{d}\left(\mathrm{k}^{\mathrm{o}}\right)^{\ell-1}}\right|_{\tilde{\mathrm{k}}}\right]=0 \quad . \quad \ell=1,2,3, \ldots \mathrm{~L} .
$$

These equations may be summarized by

$$
\begin{equation*}
\left.\widetilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}\right)^{\ell} \mid \psi_{\mathrm{n}}^{\ell}(\widetilde{\mathrm{k}})>=0 \tag{D.3a}
\end{equation*}
$$

where

These equations are useless unless $\mid \psi_{n}(k)>$ can actually be differentiated. The explicit form (2.10) shows that it can. Furthermore, because the $\mathrm{k}^{\mathrm{O}}$ dependence of (2.10) is so simple the results are just

$$
\begin{align*}
& \left.\left|\psi_{\mathrm{n}}^{1}(\widetilde{\mathrm{k}})\right\rangle=\left[1+\mathrm{r}_{\mathrm{n}} \widetilde{\mathrm{k}}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime}\right]|\phi(\overrightarrow{\mathrm{k}})\rangle  \tag{D.4a}\\
& \left.\left.\left.\mid \psi_{\mathrm{n}}^{\ell} \widetilde{\mathrm{k}}\right)\right\rangle=\left[\mathrm{r}_{\mathrm{n}} \widetilde{\mathrm{k}}^{\mathrm{o}}\right)\right]^{\ell} \mathrm{H}^{\prime \prime}|\phi(\overrightarrow{\mathrm{k}})\rangle \quad \ell=2,3, \ldots \mathrm{~L} . \tag{D.4b}
\end{align*}
$$

Now, of course,

$$
\begin{equation*}
\left.\left.\left.\widetilde{\mathrm{k}}^{\mathrm{O}}-\mathrm{H}\right) \mid \psi_{\mathrm{n}}^{1} \widetilde{\mathrm{k}}\right)\right\rangle=0 \tag{D.5}
\end{equation*}
$$

as always. Next try to evaluate

$$
\left.\widetilde{\mathrm{k}}^{\mathrm{O}}-\mathrm{H}\right)\left|\psi_{\mathrm{n}}^{2}(\widetilde{\mathrm{k}})\right\rangle
$$

Because

$$
\left(\widetilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}^{\prime}\right) \mathrm{r}_{\mathrm{n}}\left(\widetilde{\mathrm{k}}^{\mathrm{o}}\right)=1
$$

it is clear from (D. 4b) that

$$
\begin{equation*}
\left.\left.\left.\left(\widetilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}^{\prime}\right) \mid \psi_{\mathrm{n}}^{2} \widetilde{\mathrm{k}}\right)\right\rangle=\mathrm{r}_{\mathrm{n}} \widetilde{\mathrm{k}}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime}|\phi(\overrightarrow{\mathrm{k}})\rangle \tag{D.6}
\end{equation*}
$$

Furthermore

$$
\left.\left.H^{\prime \prime} \mid \psi_{n}^{2} \widetilde{\mathrm{k}}\right)\right\rangle=\mathrm{H}^{\prime \prime}\left[\mathrm{r}_{\mathrm{n}}\left(\widetilde{\mathrm{k}}^{\mathrm{o}}\right)\right]^{2} \mathrm{H}^{\prime \prime}|\phi(\overrightarrow{\mathrm{k}})\rangle
$$

From the definitions of $\mathrm{H}^{\prime \prime}$ and the reduced resolvent in Section II the right hand side of this equation is proportional to $|\phi\rangle$. Thus

$$
\begin{equation*}
\mathrm{H}^{\prime \prime}\left|\psi_{\mathrm{n}}^{2}(\widetilde{\mathrm{k}})\right\rangle=\mid \phi(\widetilde{\mathrm{k}})>\lambda \tag{D.7}
\end{equation*}
$$

where $\lambda$ is given by

$$
\left.\left\langle\phi\left(\overrightarrow{\mathrm{k}^{\prime}}\right)\right| \mathrm{H}^{\prime \prime}\left[r_{\mathrm{n}} \widetilde{\mathrm{k}}^{\mathrm{o}}\right)\right]^{2} \mathrm{H}^{\prime \prime}|\phi(\overrightarrow{\mathrm{k}})\rangle=\delta^{3}\left(\overrightarrow{\mathrm{k}^{\prime}}-\overrightarrow{\mathrm{k}}\right) \lambda
$$

Referring to (D. 1b) shows that $\lambda=-1$. Subtracting (D. 7) from (D.6) gives

$$
\left.\left.\tilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}\right)\left|\psi_{\mathrm{n}}^{2} \tilde{(\mathrm{k})}\right\rangle=\left[1+\mathrm{r}_{\mathrm{n}} \widetilde{\mathrm{k}}^{\mathrm{o}}\right) \mathrm{H}^{\prime \prime}\right]|\phi(\widetilde{\mathrm{k}})\rangle
$$

Hence

$$
\begin{equation*}
\left.\left(\tilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}\right)\left|\psi_{\mathrm{n}}^{2} \tilde{(\tilde{\mathrm{k}})\rangle}=\right| \psi_{\mathrm{n}}^{1} \widetilde{(\mathrm{k})}\right\rangle . \tag{D.8}
\end{equation*}
$$

Now do the same thing for the higher order states:

$$
\left.\left.\left.\left.\widetilde{\mathrm{k}}^{o}-\mathrm{H}\right) \mid \psi_{\mathrm{n}}^{\ell} \widetilde{\mathrm{k}}\right)\right\rangle=\left(\widetilde{\mathrm{k}}^{o}-\mathrm{H}^{\prime}\right)\left|\psi_{\mathrm{n}}^{\ell}(\widetilde{\mathrm{k}})>-\mathrm{H}^{\prime \prime}\right| \psi_{\mathrm{n}}^{\ell}(\widetilde{\mathrm{k}})\right\rangle .
$$

For $\ell \geq 3$ the second term vanishes because of (D. 1c). The first term is just

$$
\left.\left.\left.\widetilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}^{\prime}\right) \mid \psi_{\mathrm{n}}^{\ell} \tilde{\mathrm{k}}\right)>=\left[\mathrm{r}_{\mathrm{n}} \widetilde{\mathrm{k}}^{\mathrm{o}}\right)\right]^{\ell-1} \mathrm{H}^{\prime \prime}|\phi(\overline{\mathrm{k}})\rangle \quad \ell=3,4, \ldots \mathrm{~L} .
$$

Therefore the general result is

$$
\begin{equation*}
\left.\tilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}\right)\left|\psi_{\mathrm{n}}^{\ell}(\widetilde{\mathrm{k}})\right\rangle=\left|\psi_{\mathrm{n}}^{\ell-1}(\widetilde{\mathrm{k}})\right\rangle \quad \ell=2,3, \ldots \mathrm{~L} \tag{D.9}
\end{equation*}
$$

even though the $\ell=2$ case required special treatment. Because of (D.9) the fact that

$$
\left.\tilde{\mathrm{k}}^{\mathrm{o}}-\mathrm{H}\right)^{\ell}\left|\psi_{\mathrm{n}}^{\ell}(\tilde{\mathrm{k}})\right\rangle=0
$$

merely reflects the fact that

$$
\left.\left.\left(\tilde{\mathrm{k}}^{0}-\mathrm{H}\right) \mid \psi_{\mathrm{n}}^{1} \widetilde{\mathrm{k}}\right)\right\rangle=0
$$

The states $\left|\psi_{\mathrm{n}}^{\ell}\right\rangle$ are sometimes called multipole ghost states. It is easy to show that

$$
\begin{equation*}
\left\langle\psi_{\mathrm{n}}^{\ell} \widetilde{\mathrm{k}}\right)\left|\psi_{\mathrm{n}}^{\ell}(\widetilde{\mathrm{k}})\right\rangle=0 \tag{D.10}
\end{equation*}
$$

because the Hamiltonian is hermitian as in Section III. Again hermitian analyticity requires that $D(k)$ has a zero of order $L$ when continued counterclockwise to $\mathrm{k}^{\mathrm{o}}=\widetilde{\mathrm{k}}^{\mathrm{o}^{*}}$ in the upper half plane. There are thus L additional states $\left|\psi_{\overline{\mathrm{n}}}{ }^{\ell}\left(\widetilde{\mathrm{k}}^{*}\right)\right\rangle$ analogous to (D.4). The inner product of conjugate partners is

$$
\left.\left.\left\langle\psi_{\overline{\mathrm{n}}}^{\ell}\left(\widetilde{\mathrm{k}}^{*}\right)\right| \psi_{\mathrm{n}}^{\ell^{\prime}} \widetilde{\mathrm{k}}^{\prime}\right)\right\rangle= \begin{cases}\mathrm{C} \delta^{3}\left(\overrightarrow{\mathrm{k}}-\overrightarrow{\mathrm{k}}^{\prime}\right) & \text { if } \ell+\ell^{\prime}=\mathrm{L}+1  \tag{D.11}\\ 0 & \text { otherwise }\end{cases}
$$

The constant C is given in (D. 2).

## APPENDIX E

## OTHER ONE DIMENSIONAL IRREDUCIBIE REPRESENTATIONS

The states constructed in Section III by analytic continuation automatically have $q^{2}>0$. Unfortunately they are not irreducible representations of the Poincaré group. To construct irreducible representations requires introducing complex boosts because of (4.34). Once these boosts are allowed, however, a state with any momentum $\mathrm{k}=\mathrm{p}+\mathrm{iq}$ satisfying $\mathrm{k}^{2}=\mathscr{M}^{2}$ can be generated. Beltrametti and Luzzatto ${ }^{23}$ observed that there are three classes of momentum vectors that are distinguished by whether the plane containing $p$ and $q$ (1) intersects the interior of the light cone, (2) is tangent to it, or (3) is completely spacelike. The quantity $(p: q)^{2}-p^{2} q^{2}$ is correspondingly positive, zero, or negative. A suitable real Lorentz transformation will bring any $k$ into one of the three standard forms as follows:

Type 1: $\Delta>0$

$$
\begin{align*}
\mathrm{Z} & =(\alpha, 0,0, \beta)  \tag{E.1}\\
\mathscr{A}^{2} & =\alpha^{2}-\beta^{2}
\end{align*}
$$

Type 2: $\Delta=0$

$$
\begin{equation*}
\overline{\mathrm{k}}=(\alpha, 0, \mathrm{i} \mathscr{M}, \alpha) \tag{E.2}
\end{equation*}
$$

Type 3: $\Delta<0$

$$
\begin{align*}
\overline{\mathrm{k}} & =(0,0, \alpha, \beta)  \tag{E.3}\\
\mathscr{M}^{2} & =-\alpha^{2}-\beta^{2} .
\end{align*}
$$

The subgroup of the real Lorentz group that leave $\overline{\mathrm{K}}$ invariant is one dimensional in each case. Denote the little group generator by $g_{i}$. Then the three classes have, respectively,

$$
\mathrm{g}_{1}=\mathrm{J}_{3}
$$

$$
\begin{align*}
& \mathrm{g}_{2}=\mathrm{J}_{2}+\mathrm{K}_{1} \\
& \mathrm{~g}_{3}=\mathrm{K}_{1} \tag{E.4}
\end{align*}
$$

(Of course for any of the momentum classes the subgroup of the complex Lorentz group that leaves a particular K invariant has not one but three generators, viz. the three $S_{i}(\bar{k})$ given by (4.27). In each case, however, only one of these three generators is proportional to a hermitian generator $\mathrm{g}_{\mathrm{i}}$.)

Regardless of the type, an irreducible representation of the full Lorentz group may be defined as in (5.15) by

$$
\left|\mathrm{k}, \mathrm{j}, \sigma>\equiv \mathrm{U}\left[\mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}}) \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})\right]\right| \Psi(\mathscr{M}) \mathrm{j}, \mathrm{j}_{3}>\left.\right|_{\mathrm{j}_{3}=\sigma} \sqrt{2 \mathscr{M}}
$$

where the meaning of $\sigma$ may differ from one class to another. For Type 1 momenta (E.5) actually yielded the states of Section V.C on which the one dimensional little group was represented irreducibly. It will now be shown that states of Type 3 can also have their one dimensional little group represented irreducibly but states of Type 2 do not have such a representation. To see this, apply a real Lorentz transformation $\Lambda$ to (E.5) to get

$$
\begin{equation*}
\mathrm{U}[\Lambda]|\mathrm{k}, \mathrm{j}, \sigma\rangle=\mathrm{U}\left[\mathrm{~L}(\Lambda \mathrm{k}, \mathrm{k}) \mathrm{L}^{\mathrm{c}}(\mathrm{k}, \mathscr{M})\right] \mathrm{U}\left[\mathrm{Z}_{\mathrm{i}}\right]|\mathscr{M}, \mathrm{j}, \sigma\rangle \tag{E.6}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathrm{Z}_{\mathrm{i}} \equiv \mathrm{~L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})^{-1}\left\{\mathrm{~L}(\Lambda \mathrm{k}, \overline{\mathrm{k}})^{-1} \Lambda \mathrm{~L}(\mathrm{k}, \overline{\mathrm{k}})\right\} \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M}) \tag{E.7}
\end{equation*}
$$

The quantity in braces in each case is a real Lorentz transformation that leaves k invariant. It is therefore generated by the corresponding $\mathrm{g}_{\mathrm{i}}$ so that

$$
\begin{equation*}
\mathrm{U}\left[\mathrm{~L}(\Lambda \mathrm{k}, \overline{\mathrm{k}})^{-1} \Lambda \mathrm{~L}(\mathrm{k}, \overline{\mathrm{k})}]=\mathrm{e}^{\mathrm{i} \lambda \mathrm{~W}^{\mathrm{g}} \mathrm{i}}\right. \tag{E.8}
\end{equation*}
$$

where $\lambda_{\mathrm{W}}$ is a real Wigner parameter. (F or Type $1, \lambda_{\mathrm{W}}$ is the real Wigner rotation angle around the z axis, for Type $3, \lambda_{\mathrm{W}}$ gives the real boost velocity along
the x axis, and for Type $2, \lambda_{\mathrm{W}}$ is a mixture of rotation angle around y and boost along x.)

States with momenta of Type 1 are the helicity states discussed in Section V.C. Consider now states of Type 3. Choose the complex boost in (E.5) to be

$$
\begin{equation*}
\mathrm{U}\left[\mathrm{~L}^{\mathrm{c}}(\overline{\mathrm{~K}}, \mathscr{M})\right]=\mathrm{e}^{\mathrm{i} \theta \mathrm{~J}} 1 \mathrm{e}^{-\frac{\pi}{2} \mathrm{~K}_{2}} \tag{E.9}
\end{equation*}
$$

(The velocity boost in the y direction is purely imaginary.) The corresponding Lorentz matrix is

$$
L^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})_{\nu}^{\mu}=\left\{\left.\begin{array}{cccc}
0 & 0 & \mathrm{i} & 0 \\
0 & 1 & 0 & 0 \\
\mathrm{i} \cos \theta & 0 & 0 & -\sin \theta \\
i \sin \theta & 0 & 0 & \cos \theta
\end{array} \right\rvert\,\right.
$$

Obviously

$$
\mathrm{L}^{\left.\mathrm{c}(\overline{\mathrm{k}}, \mathscr{M})_{\nu}^{\mu} \mathscr{M}^{\nu}=\begin{array}{c}
i \\
\vdots \\
\vdots \\
\mathrm{i} \mathscr{M} \cos \theta \\
\mathrm{i} \mathscr{H} \cdot \sin \theta
\end{array} \right\rvert\,}
$$

in accordance with the K of (E.3). To calculate Z use (E.8) and

$$
\left[J_{1}, K_{1}\right]=0
$$

to get

$$
\mathrm{U}[\mathrm{Z}]=\mathrm{e}^{\frac{\pi}{2} \mathrm{~K}_{2}} \mathrm{e}^{\mathrm{i} \mathrm{\lambda} \mathrm{~W}^{K_{1}}} \mathrm{e}^{-\frac{\pi}{2} \mathrm{~K}_{2}}
$$

Because

$$
\left[\mathrm{K}_{1}, \mathrm{~K}_{2}\right]=-\mathrm{i} \mathrm{~J}_{3}
$$

this becomes

$$
\mathrm{U}[\mathrm{Z}]=\mathrm{e}^{-\lambda} \mathrm{W}^{\mathrm{J}} 3 .
$$

The transformation law (E.6) for Type 3 states is therefore one dimensional as claimed:

$$
\mathrm{U}[\Lambda]|\mathrm{k}, \mathrm{j}, \sigma\rangle=\mid \Lambda \mathrm{k}, \mathrm{j}, \sigma>\mathrm{e}^{-\lambda \mathrm{W}^{\sigma}}
$$

Because $\lambda_{\mathrm{W}}$ and $\sigma$ are real this representation is not unitary. More importantly, it is not even unitary when $k$ is continued to real values.

For states with momentum of Type 2 there is no choice of the complex boost $\mathrm{L}^{\mathrm{C}}(\overline{\mathrm{K}}, \boldsymbol{M})$ that will represent the one dimensional little group irreducibly. Suppose there were such a choice. The Z in (E.7) could only be a rotation around the z-axis if

$$
\begin{equation*}
\mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})^{-1}\left(\mathrm{~J}_{2}+\mathrm{K}_{1}\right) \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{k}}, \mathscr{M})=\zeta \mathrm{J}_{3} \tag{E.15}
\end{equation*}
$$

for some constant $\zeta$. This would then insure that

$$
\mathrm{U}[\mathrm{Z}]=\mathrm{e}^{\mathrm{i} \lambda} \mathrm{~W}^{\xi \mathrm{J}_{3}}
$$

and (E.6) would then represent the little group irreducibly. Unfortunately there is no $L^{c}$ that satisfies (E. 15). To see this write (E. 15) as

$$
\begin{equation*}
\left(\mathrm{J}_{2}+\mathrm{K}_{1}\right) \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{~K}}, \mathscr{M})=\zeta \mathrm{L}^{\mathrm{c}}(\overline{\mathrm{~K}}, \mathscr{M}) \mathrm{J}_{3} . \tag{E.16}
\end{equation*}
$$

The generators may be represented by the $4 \times 4$ matrices

$$
\begin{aligned}
J_{2}+\mathrm{K}_{1} & =\left(\left.\begin{array}{cccc}
0 & -i & 0 & 0 \\
-\mathrm{i} & 0 & 0 & \mathrm{i} \\
0 & 0 & 0 & 0 \\
0 & -i & 0 & 0
\end{array} \right\rvert\,\right. \\
J_{3} & =\left|\begin{array}{cccc}
0 & 0 & 0 & 0 \\
0 & 0 & -i & 0 \\
0 & i & 0 & 0 \\
0 & 0 & 0 & 0
\end{array}\right|
\end{aligned}
$$

Explicit multiplication shows that the only matrices $L^{\mathrm{C}}$ that satisfy ( E .16 ) have all zeros in the second row. Such matrices have zero determinant and thus cannot represent Lorentz transformations. This shows that there is no state like (E.5) which represents the little group of Type 2 momenta irreducibly.

## APPENDIX F

## HOW $q \rightarrow 0$ IN HELICITY AMPLITUDES

It is crucial that $q$ goes to zero in the plane $\vec{p} \| \vec{q}$. Because of this, the helicity states of Section V. C are actually produced in (6.10) and yet their observed transformation law is (6.18). To see how $q \rightarrow 0$ begin with the transformation law

$$
\begin{aligned}
&\left.<\Lambda \mathrm{k}^{*}, \mathrm{j}, \sigma^{\prime}\left|\mathrm{T}\left[\psi_{\alpha}(\Lambda \mathrm{r}) \bar{\psi}_{\beta}(-\Lambda \mathrm{r})\right]\right| 0\right\rangle \\
&=\mathrm{e}^{-\mathrm{i} \theta \mathrm{~W}^{\prime}}\left\langle\mathrm{k}^{*}, \mathrm{j}, \sigma^{\prime}\right| \mathrm{T}\left[\psi_{\alpha^{\prime}}(\mathrm{r}) \bar{\psi}_{\beta^{\prime}}(-\mathrm{r})\right]|0\rangle \mathscr{D}_{\alpha^{\prime}, \alpha^{\prime}}(\Lambda) \mathscr{D}_{\beta^{\prime}, \beta^{\prime}}^{(\Lambda)}
\end{aligned}
$$

appropriate to the residue in (6.5). Let $q \rightarrow 0$ in some unspecified manner. Then project out the two particle irreducible representation (6.4):

$$
\begin{align*}
& <\Lambda p, j, \sigma^{\prime}|\Lambda p, j, \sigma ; 1+2\rangle \\
&  \tag{o}\\
& \left.=\mathrm{e}^{-\mathrm{i} \theta^{\prime} \sigma^{\prime}} \sum_{\sigma^{\prime \prime}}<p, j, \sigma^{\prime} \mid \mathrm{p}, \mathrm{j}, \sigma^{\prime \prime} ; 1+2>\mathscr{D}_{\sigma^{\prime \prime}, \sigma} \widetilde{\mathrm{R}}\right)
\end{align*}
$$

where

$$
\begin{equation*}
\theta^{q} \equiv \lim _{q \rightarrow 0} \theta_{W} \tag{F.2}
\end{equation*}
$$

and $\widetilde{R}$ is given by (6.4).
Equation (F. 1) may be used to investigate how $q \rightarrow 0$. Recall from (6.4) that

$$
\widetilde{\mathrm{R}}=\mathrm{L}\left(\mathrm{~W}, \tilde{\mathrm{p}}^{\prime}\right) \mathrm{R}\left(\tilde{\mathrm{p}}^{\prime}, p^{\prime}\right) \Lambda R(\mathrm{p}, \tilde{\mathrm{p}}) \mathrm{L}(\tilde{\mathrm{p}}, \mathrm{~W})
$$

Choose

$$
\begin{aligned}
& p=\left(p^{o}, 0,0,|\vec{p}|\right) \\
& \Lambda=e^{i \theta J} 3
\end{aligned}
$$

Then $p=\tilde{p}^{\prime}=p^{\prime}=\tilde{p}^{\prime}$ so that

$$
\widetilde{\mathrm{R}}=\mathrm{e}^{\mathrm{i} \theta J_{3}}
$$

This greatly simplifics (F.1) because now

$$
\mathrm{D}_{\sigma^{\prime \prime}, \sigma}(\widetilde{R})=\delta_{\sigma^{\prime \prime}, \sigma} \mathrm{e}^{\mathrm{i} \theta \sigma}
$$

and (F.1) reduces to

$$
\begin{equation*}
<p, j, \sigma^{\prime}\left|p, j, \sigma ; 1+2>=\mathrm{e}^{-\mathrm{i} \theta^{\prime} \sigma^{\prime}}<\mathrm{p}, \mathrm{j}, \sigma^{\prime}\right| \mathrm{p}, \mathrm{j}, \sigma ; 1+2>\mathrm{e}^{\mathrm{i} \theta \sigma} \tag{F.3}
\end{equation*}
$$

The content of (F.3) is that the amplitude vanishes unless

$$
\theta^{\prime} \sigma^{\prime}=\theta \sigma
$$

The angle $\theta^{\prime}$ will depend on whether or not $\overrightarrow{\mathrm{q}} \| \overrightarrow{\mathrm{p}}$. In fact, it will be shown that

$$
\theta^{\prime}= \begin{cases}\theta & \text { if } \overrightarrow{\mathrm{q}} \| \overrightarrow{\mathrm{p}}  \tag{F.4}\\ 0 & \text { otherwise }\end{cases}
$$

This has two consequences: If $\vec{q} \| \vec{p}$ then the amplitude vanishes unless $\sigma^{\prime}=\sigma$; If $\vec{q}$ and $\vec{p}$ are not parallel, the amplitude vanishes regardless of the value of $\sigma^{\prime}$.

To demonstrate (F.4) take

$$
\mathrm{q}=\left(\mathrm{q}^{\circ},|\overrightarrow{\mathrm{q}}| \sin \eta, 0,|\overrightarrow{\mathrm{q}}| \cos \eta\right)
$$

Since $\Lambda$ is a z rotation

$$
\mathbf{q}^{1}=\left(q^{\circ},|\vec{q}| \sin \eta \cos \theta,|\vec{q}| \sin \eta \sin \theta,|\vec{q}| \cos \eta\right) .
$$

Thus $\eta$ and $\theta$ are the polar and azimuthal angles of $q^{\prime}$ and $\vec{q} \| \vec{p}$ corresponds to $\eta=0$. To calculate $\theta^{\prime}$ requires first calculating $\theta_{\mathrm{W}}$ from

$$
\begin{equation*}
e^{i \theta} W^{J_{3}}=L\left(\bar{k}, k^{\prime}\right) e^{i \theta J_{3}} L(k, \bar{k}) \tag{F.5}
\end{equation*}
$$

A specific form for the real transformation $L(k, \bar{k})$ is necessary. Consider the case $q^{2}>0$ as in (6.20). Then

$$
\bar{p}=\left(\frac{p \cdot q}{\sqrt{q^{2}}}, 0,0, \sqrt{\frac{(p \cdot q)^{2}-p^{2} q^{2}}{q^{2}}}\right)
$$

$$
\left.\overline{\mathrm{q}}=\left(\sqrt{\mathrm{q}^{2}}, 0,0,0\right)\right)
$$

from (5.12a). A real transformation $L(\bar{k}, k)$ from $k$ to $\bar{k}$ may always be accomplished by boosting $q$ to the rest form $\bar{q}$ and then rotating into the $z$ axis:

$$
\mathrm{L}(\overline{\mathrm{k}}, \mathrm{k})=\mathrm{R}(\hat{\mathrm{z}}, \mathrm{~B}(\overline{\mathrm{q}}, \mathrm{q}) \mathrm{p}) \mathrm{B}(\overline{\mathrm{q}}, \mathrm{q})
$$

Thus

$$
\begin{equation*}
\mathrm{e}^{\mathrm{i} \theta_{\mathrm{W}} \mathrm{~J}_{3}}=\mathrm{R}\left(\hat{\mathrm{z}}, \mathrm{~B}\left(\overline{\mathrm{q}}, \mathrm{q}^{\mathrm{q}}\right) \mathrm{p}\right) \mathrm{B}\left(\overline{\mathrm{q}}, \mathrm{q}^{\prime}\right) \mathrm{e}^{\mathrm{i} \theta \mathrm{~J}_{3}} \mathrm{~B}(\mathrm{q}, \overline{\mathrm{q}}) \mathrm{R}(\mathrm{~B}(\overline{\mathrm{q}}, \mathrm{q}) \mathrm{p}, \hat{\mathrm{z}}) \tag{F.6}
\end{equation*}
$$

where $p^{\prime}=p$ has been used in the argument of the leftmost rotation. The two boosts are given by

$$
\begin{aligned}
& B(q, \bar{q})=e^{i \mu \hat{n} \cdot \vec{K}} \\
& B\left(q^{\prime}, \bar{q}\right)=e^{i \mu \hat{n^{\prime}} \cdot \overrightarrow{\mathrm{K}}}
\end{aligned}
$$

where

$$
\begin{aligned}
& \hat{\mathrm{n}}=(\sin \eta, 0, \cos \eta) \\
& \hat{\mathrm{n}}^{\prime}=(\sin \eta \cos \theta, \sin \eta \sin \theta, \cos \eta) \\
& \tanh \mu=\frac{|\vec{q}|}{\mathrm{q}^{0}} .
\end{aligned}
$$

Because $\hat{n}^{\prime}$ is just $\hat{n}$ rotated by $\theta$ around $\hat{z}$,

$$
\mathrm{B}\left(\overline{\mathrm{q}}, \mathrm{q}^{\prime}\right) \mathrm{e}^{\mathrm{i} \theta \mathrm{~J}_{3}}=\mathrm{e}^{\mathrm{i} \theta \mathrm{~J}_{3}} \mathrm{~B}(\overline{\mathrm{q}}, \mathrm{q})
$$

Hence (F.6) is just

To calculate $\theta_{\mathrm{W}}$ these rotations must be explicitly displayed. The rightmost rotation above rotates a vector with momentum along $\hat{z}$ until it is parallel
to

$$
B(\bar{q}, q) p=\left(\begin{array}{c}
p^{o} \cosh \mu-|\vec{p}| n_{3} \sinh \mu  \tag{F.8}\\
-p^{o} n_{1} \sinh \mu+|\vec{p}| n_{1} n_{3}(\cosh \mu-1) \\
0 \\
-p^{o} n_{3} \sinh \mu+|\vec{p}|\left[1+n_{3} n_{3}(\cosh \mu-1)\right]
\end{array}\right)
$$

where the vector components are written ( $t, x, y, z$ ) vertically. Thus the rotation is around the $y$ axis and is given by

$$
R(\mathrm{~B}(\mathrm{q}, \overline{\mathrm{q}}) \mathrm{p}, \hat{\mathrm{z}})=\left(\begin{array}{ccc}
\cos \omega & 0 & \sin \omega  \tag{F.9}\\
0 & 1 & 0 \\
-\sin \omega & 0 & \cos \omega
\end{array}\right)
$$

where

$$
\begin{equation*}
\tan \omega=\frac{-\mathrm{p}^{\mathrm{o}} \mathrm{n}_{1} \sinh \mu+|\overrightarrow{\mathrm{p}}| \mathrm{n}_{1} \mathrm{n}_{3}(\cosh \mu-1)}{-\mathrm{p}^{\mathrm{o}} \mathrm{n}_{3} \sinh \mu+|\overrightarrow{\mathrm{p}}|\left[1+\mathrm{n}_{3} \mathrm{n}_{3}(\cosh \mu-1)\right]} \tag{F.10}
\end{equation*}
$$

This gives for (F.7)

$$
\mathrm{e}^{\mathrm{i} \theta} \mathrm{~W}^{\mathrm{J}_{3}}=\mathrm{R}\left(\hat{\mathrm{z}}, \mathrm{~B}\left(\overline{\mathrm{q}}, \mathrm{q}^{\prime}\right) \mathrm{p}\right)\left(\begin{array}{ccc}
\cos \theta & \sin \theta & 0  \tag{F.11}\\
-\sin \theta & \cos \theta & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{ccc}
\cos \omega & 0 & \sin \omega \\
0 & 1 & 0 \\
-\sin \omega & 0 & \cos \omega
\end{array}\right)
$$

The remaining rotation to be calculated rotates a vector with momentum along $\hat{z}$ until it is parallel to

$$
B\left(\bar{q}, q^{\prime}\right) p=\left(\begin{array}{l}
p^{o} \cosh \mu-|\vec{p}| n_{3}^{\prime} \sinh \mu  \tag{F.12}\\
-p^{o} n_{1}^{\prime} \sinh \mu+|\vec{p}| n_{1}^{\prime} n_{3}^{\prime}(\cosh \mu-1) \\
-p^{o} n_{2}^{\prime} \sinh \mu+|\vec{p}| n_{2}^{\prime} n_{3}^{\prime}(\cosh \mu-1) \\
-p^{o} n_{3}^{\prime} \sinh \mu+|\vec{p}|\left[1+n_{3}^{\prime} n_{3}^{\prime}(\cosh \mu-1)\right]
\end{array}\right) .
$$

Comparing $\hat{\mathrm{n}}^{\boldsymbol{\gamma}}$ with $\hat{\mathrm{n}}$ shows that this vector is just (F.8) rotated around the z axis by $\theta$. Thus

$$
R\left(\mathrm{~B}\left(\mathrm{q}^{\prime}, \bar{q}\right) \mathrm{p}, \hat{\mathrm{z}}\right)-\left(\begin{array}{ccc}
\cos \theta & \sin \theta & 0 \\
-\sin \theta & \cos \theta & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{ccc}
\cos \omega & 0 & \sin \omega \\
0 & 1 & 0 \\
-\sin \omega & 0 & \cos \omega
\end{array}\right) .(\mathrm{F} \cdot 13)
$$

This rather simplifies (F.11), which becomes

$$
\mathrm{e}^{\mathrm{i} \theta} \mathrm{~W}^{\mathrm{J}} 3=\left(\begin{array}{ccc}
1 & 0 & 0  \tag{F.14}\\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

This means that $\theta_{\mathrm{W}}=0$ and hence $\theta^{\prime}=0$ as claimed for $\overrightarrow{\mathrm{q}}$ not parallel to $\overrightarrow{\mathrm{p}}$.
It appears, in fact, that $\theta_{\mathrm{W}}=0$ regardless of the direction of $\overrightarrow{\mathrm{q}}$. This is not correct. Suppose that $\vec{q} \| \vec{p}$. Then $\hat{n}$ is along the $z$ axis. Thus the vector in (F.8) has only $t$ and $z$ components so that

$$
R(B(\bar{q}, q) p, \hat{z})=\left(\begin{array}{lll}
1 & 0 & 0  \tag{F.15a}\\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

In other words $\omega=0$. The vector (F.12) in this case also has only t and z components so that

$$
R\left(B\left(\bar{q}, q^{\prime}\right) p, \hat{z}\right)=\left(\begin{array}{lll}
1 & 0 & 0  \tag{F.15b}\\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

and hence (F.7) becomes

$$
\mathrm{e}^{\mathrm{i} \theta} \mathrm{~W}^{J} 3=\left(\begin{array}{ccc}
\cos \theta & \sin \theta & 0  \tag{F.16}\\
-\sin \theta & \cos \theta & 0 \\
0 & 0 & 1
\end{array}\right)
$$

as claimed in (F.4). Generally, of course, the rotation that takes the vector $(0,0,1)$ into itself is not just the identity. It can be any fixed rotation around the
z axis. This fixed rotation must appear in both (F.15a) and (F.15b). It then cancels in ( F .16 ) so that $\theta_{\mathrm{W}}=\theta$ regardless. This is the reason that ( F .15 b ) is not just the $\omega=0$ limit of (F.13).

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## FIGURE CAPTIONS

1. Example showing that the continuation to the unstable particle pole is always around the same branch point (viz., the heaviest decay mode) independently of the scattering process considered.
2. The path followed in continuing hermitian analyticity away from the physical sheet to obtain

$$
\mathrm{D}_{\mathrm{n}}(\mathrm{k})^{*}=\mathrm{D}_{\mathrm{n}^{( }}\left(\mathrm{k}^{*}\right)
$$

3. Bethe-Salpeter equation and normalization condition for a composite unstable particle with complex momentum $\mathrm{k}^{\mu}$.
4. The usual connectedness structure that results from the LSZ asymptotic condition.
5. The connectedness structure that results when one particle (denoted by the wavy line) is unstable.
6. The values of $\alpha_{1}$ and $\alpha_{2}$ that solve (7.18) and correspond to second sheet poles in the $S$ matrix (7.7). The broken line is a plot of (7.18a); the solid line, of (7.18b).


Fig. 1


Fig. 2

$$
\begin{gathered}
{\left[\left(\frac{\square}{\vdots}\right)^{-1}-[Z]=0\right.} \\
\left.K^{*} \sqrt{\square} \frac{\partial}{\partial K_{\mu}}\left[\left(\frac{\square}{\vdots}\right)^{-1}-\boxed{Z}\right]=\right\rangle_{K}=n!\frac{2 K^{\mu}}{i}
\end{gathered}
$$

Fig. 3


Fig. 4

$$
\begin{aligned}
& \frac{m(S) n}{}=\underline{m(C) n} \\
& \overline{m(S) m}=\overline{m(C) m}+\overline{m(C) m}+\overline{m(C) m}
\end{aligned}
$$

Fig. 5


Fig. 6


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