# CURRENT COMMUTATORS AND ELECTRON SCATTERING AT HIGH MOMENTUM TRANSFER* 

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#### Abstract

Sum rules for electron proton total cross sections are deduced from the vanishing of various equal time commutators; it is shown how these cross sections determine the (spin-averaged) proton expectation value of all equal time commutators of components of the electric current and time derivatives thereof.


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[^0]Electron scattering on a proton at rest is described by the cross section (electron mass $=0$ )

$$
\begin{equation*}
\frac{d^{2} \sigma^{e p}}{d q^{2} d \nu}=\frac{\alpha^{2}}{M^{2} E^{2}} \frac{1}{q^{2}}\left[\left(2 M^{2} E E^{\prime}-\frac{q^{2}}{2} M^{2}\right) A_{1}\left(q^{2}, \nu\right)+\nu^{2} A_{1}\left(q^{2}, \nu\right)-q^{2} A_{2}\left(q^{2}, \nu\right)\right] \tag{1}
\end{equation*}
$$

where

$$
\alpha=\mathrm{e}^{2} / 4 \pi=(137)^{-1}
$$

$E\left(E^{\prime}\right)=$ initial (final) electron energy

$$
\begin{aligned}
& \mathrm{q}_{\mu}=\mathrm{e}_{\mu}-\mathrm{e}_{\mu}^{\prime}=\text { momentum imparted to proton } \\
& \nu=\left(\mathrm{E}-\mathrm{E}^{\prime}\right) \mathrm{M}=-\mathrm{q} \cdot \mathrm{p}(\mathrm{p}=\text { proton momentum }),
\end{aligned}
$$

and the $A_{i}$ are the absorbtive parts of the forward, off-shell Compton amplitudes $F_{i}$ defined by

$$
\begin{align*}
\mathrm{T}_{\mu \nu}(\mathrm{q}, \mathrm{p}) & =\mathrm{i} \int \mathrm{~d}_{4} \mathrm{x} \mathrm{e}^{-\mathrm{i} \underline{q} \cdot \underline{x}}\langle\mathrm{p}| \mathrm{T}\left(\mathrm{j}_{\mu}(\mathrm{x}) \mathrm{j}_{\nu}(0)\right)|\mathrm{p}\rangle_{\mathrm{c}}  \tag{2a}\\
& =\left[\mathrm{q}^{2} \mathrm{p}_{\mu} \mathrm{p}_{\nu}+\nu\left(\mathrm{q}_{\mu} \mathrm{p}_{\nu}+\mathrm{q}_{\nu} \mathrm{p}_{\mu}\right)+\nu^{2} \delta_{\mu \nu}\right] \mathrm{F}_{1}+\left(\mathrm{q}_{\mu} \mathrm{q}_{\nu}-\mathrm{q}^{2} \delta_{\mu \nu}\right) \mathrm{F}_{2} \tag{2b}
\end{align*}
$$

where $\mathrm{j}_{\mu}$ is the electric current, and a spin average is implicit. The subscript " $c$ " indicates the covariant time ordered product; thus $\mathrm{T}_{\mu \nu}$ differs from the ordinary time ordered product by a polynomial in $q$ if the equal-time commutator of $j_{0}$ and $j_{i}$ has a connected matrix element.

Bjorken ${ }^{1}$ has pointed out that for large $q_{0}$ and fixed $\underline{q}$ the coefficient of $\mathrm{q}_{0}^{-\ell-1}(\ell \geq 0)$ in an expansion of $\mathrm{T}_{\mu \nu}$ gives the matrix element of the equal-time
commutator of the electric current and its $\ell^{\text {th }}$ time derivative; in particular

$$
\begin{align*}
& \mathrm{T}_{\mu \nu} \xrightarrow{q_{0} \rightarrow \infty} \text { polynomial in } q_{0} \\
& \quad-\sum_{\ell=0}^{\infty} \frac{(-i)^{\ell}}{q_{0}^{\ell+1}} \int d \underline{x} e^{-i \underline{q} \cdot \underline{x}}\langle p|\left[\partial_{0}^{\ell} j_{\mu}(\underline{x}, 0), j_{\nu}(0)\right]|\mathrm{p}\rangle \tag{3}
\end{align*}
$$

Thus, all of these commutators are determined by the coefficients $C_{i}^{\ell}\left(\underline{q} \cdot \underline{p}, \underline{q}^{2}, p_{0}\right)$ occurring in

$$
\begin{equation*}
\mathrm{F}_{\mathrm{i}} \xrightarrow{\mathrm{q}_{0} \rightarrow \infty} \text { polynomial }+\sum_{\ell=0}^{\infty} \mathrm{C}_{\mathrm{i}}^{\ell} \mathrm{q}_{0}^{-l-1} \tag{4}
\end{equation*}
$$

We wish to show how the $C_{i}^{l}$ can be constructed from the $A_{i}$, and thus from the electron proton scattering data.

We assume that each of the $\mathrm{F}_{\mathrm{i}}\left(\mathrm{q}^{2}, \nu\right)$ satisfy the D. G.S. representation ${ }^{2,3}$

$$
\begin{equation*}
F_{i}=\sum_{m=0}^{M_{i}} F_{i}^{m}=\frac{1}{\pi} \sum_{m=0}^{M_{i}} \int_{0}^{\infty} d \sigma \int_{-1}^{1} d \beta \frac{\nu^{m_{h}}{ }_{i}^{m}(\sigma, \beta)}{q^{2}+2 \beta \nu+\sigma} \tag{5}
\end{equation*}
$$

where by crossing symmetry

$$
\begin{equation*}
\mathrm{h}_{\mathrm{i}}^{\mathrm{m}}(\sigma,-\beta)=(-1)^{m} \mathrm{~h}_{\mathrm{i}}^{\mathrm{m}}(\sigma, \beta) . \tag{6}
\end{equation*}
$$

Expanding this form of $\mathrm{F}_{\mathrm{i}}$ as in Eq. (4), we obtain after some combinatorics

$$
\begin{equation*}
G_{i}^{\ell}=\sum_{n, s, t}(-q \cdot p)^{\ell-1-2 s+2 n}\left(q^{2}\right)^{s-2 n-t}\left(p_{0}\right)^{2 s-\ell t+1} \frac{s!(2 n)!}{(2 s-\ell+1)!(\ell-1-2 s+2 n)!(s-2 n-t)!} K_{i}^{n t} \tag{7a}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathrm{K}_{\mathrm{i}}^{\mathrm{nt}}=\sum_{\cdot m=0}^{\mathrm{M}_{\mathrm{i}}} \frac{1}{(2 \mathrm{n}-\mathrm{m})!(t+\mathrm{m})!} \int_{0}^{\infty} \mathrm{d} \sigma \int_{-1}^{1} \mathrm{~d} \beta(2 \beta)^{2 \mathrm{n}-\mathrm{m}} \sigma^{\mathrm{t}+\mathrm{m}} \mathrm{~h}_{\mathrm{i}}^{\mathrm{m}}(\sigma, \beta), \tag{7b}
\end{equation*}
$$

and where the sum over $n, s$ and $t$ is restricted by $n \geq 0, s \geq 0$ and the requirement that the arguments of all the factorials are non-negative. Thus, for example, $\mathrm{t} \geq$ maximum ( $-\mathrm{m},-2 \mathrm{n}$ ).

The absorbtive parts $A_{i}$ of the $F_{i}$ can be read off from Eq. (5).

$$
\begin{equation*}
A_{i}=\sum_{m=0}^{M_{i}} A_{i}^{m}=\sum_{m=0}^{M_{i}} \int d \sigma d \beta \nu^{m} h_{i}^{m} \delta\left(q^{2}+2 \beta \nu+\sigma\right) \tag{8}
\end{equation*}
$$

and hence for $\mathrm{n}=1,2,3, \ldots$

$$
\begin{equation*}
\int_{0}^{\infty} \frac{d \nu}{\nu^{2 n+1}} A_{i}=\sum_{\substack{m=0 \\ 2 n-m \geq 0}}^{M_{i}}(-1)^{m} \int_{0}^{\infty} d \sigma \int_{-1}^{0} d \beta \frac{(2 \beta)^{2 n-m}}{\left(q^{2}+\sigma\right)^{2 n+1-m}} h_{i}^{m} \tag{9}
\end{equation*}
$$

As indicated, the sum on $m$ includes only terms for which $2 n-m \geq 0 .{ }^{4}$ We will show in a moment that the R.H.S. of Eq. (9) can be expanded for large $q^{2}$ in terms of the $\mathrm{K}_{\mathrm{i}}^{\mathrm{nt}}$ occurring in Eq. (7). However, to complete this connection, we first must extend Eq. (9) to $\mathrm{n}=0$.

Assuming Regge asymptotic behavior of the $A_{i}$ for large $\nu$, we guess that ${ }^{5}$

$$
\begin{align*}
& \mathrm{A}_{1}\left(\mathrm{q}^{2}, \nu\right) \xrightarrow{\nu \rightarrow \infty} 0  \tag{10a}\\
& \mathrm{~A}_{2}\left(\mathrm{q}^{2}, \nu\right) \xrightarrow{\nu \rightarrow \infty} \mathrm{f}_{\mathrm{p}}\left(\mathrm{q}^{2}\right) \nu+\sum_{\alpha} \mathrm{f}_{\alpha}\left(\mathrm{q}^{2}\right) \nu^{\alpha}, \tag{10b}
\end{align*}
$$

where, in addition to the Pomeranchon, the sum on $\alpha$ includes whatever other trajectories contribute with $0<\alpha<1$; for example, the $A_{2}$ and $f_{0}$. Because of Eq. (10a) the integral on the L.H.S. of Eq. (9) exists for $A_{1}$ when $n=0,{ }^{6}$ and Eq. (9) holds for this case. However, for $A_{2}$ the L.H.S. of Eq. (9) must be modified for $\mathrm{n}=0$.

As suggested by the form of Eq. (5), we assume that the Regge limit in Eq. (10b) arises only from the terms with $m \geq 1$ in Eq. (8). That is, we assume that (10b) is satisfied with $A_{2}$ replaced by $A_{2}-A_{2}^{0}$. Since $F_{2}-F_{2}^{0}$ vanishes at $\nu=0$, a subtracted dispersion relation for this difference reads

$$
\begin{equation*}
\mathrm{F}_{2}-\mathrm{F}_{2}^{0}=\frac{2 \nu^{2}}{\pi} \int_{0}^{\infty} \frac{\mathrm{d} \nu^{\prime}}{\nu^{\prime}\left(\nu^{\prime}-\nu^{2}\right.}\left(\mathrm{A}_{2}\left(\mathrm{q}^{2}, \nu^{\prime}\right)-\mathrm{A}_{2}^{0}\left(\mathrm{q}^{2}, \nu^{\prime}\right)\right) \tag{11}
\end{equation*}
$$

and if we add and subtract the R.H.S. of (10b) to the integrand of this expression, and in the first casc do the integral explicitly, we obtain

$$
\left.\begin{array}{rl}
\mathrm{F}_{2}-\mathrm{F}_{2}^{0}=\mathrm{if}  \tag{12}\\
\mathrm{p}
\end{array}\right)-\sum_{\alpha} \frac{\left(1+\mathrm{e}^{-\mathrm{i} \pi \alpha}\right)}{\sin \pi \alpha} \mathrm{f}_{\alpha} \nu^{\alpha} .
$$

Let us now assume that the behavior of $\mathrm{F}_{2}$ for large $\nu$ is given entirely by the Regge terms, namely, that there is no part of $\mathrm{F}_{2}$ constant in $\nu$ in this limit. Then, since both $F_{2}^{0}$ and the bracket in the integrand of Eq. (12) vanish as $\nu \rightarrow \infty$, it follows that ${ }^{7}$

$$
\begin{equation*}
\int_{0}^{\infty} \frac{\mathrm{d} \nu}{\nu}\left[\mathrm{~A}_{2}-\mathrm{A}_{2}^{0}-\mathrm{f}_{\mathrm{p}} \nu-\sum_{\alpha} \mathrm{f}_{\alpha} \nu^{\alpha}\right]=0 \tag{13}
\end{equation*}
$$

which from (8) gives

$$
\begin{equation*}
\int_{0}^{\infty} \frac{\mathrm{d} \nu}{\nu}\left[\mathrm{~A}_{2}-\mathrm{f}_{\mathrm{p}} \nu-\sum_{\alpha} \mathrm{f}_{\alpha} \nu^{\alpha}\right]=\int_{0}^{\infty} \mathrm{d} \sigma \int_{-1}^{0} \mathrm{~d} \beta \frac{\mathrm{~h}_{2}^{0}(\sigma, \beta)}{\mathrm{q}^{2}+\sigma} \tag{14}
\end{equation*}
$$

The R. H.S. of this expression is the same as the R.H.S. of (9) with $n=0$ and $i=2$. Thus Eq. (9) can be extended to $n=0$ : with no change in form for $i=1$, and with (9) replaced by (14) for $\mathrm{i}=2$. Summarizing this result, and expanding the R.H.S. of (9) in a power series in $\left(q^{2}\right)^{-1}$, we obtain for $n=0,1,2, \ldots$

$$
\begin{equation*}
\int_{0}^{\infty} \frac{\mathrm{d} \nu}{\nu^{2 \mathrm{n}+1}}\left[\mathrm{~A}_{\mathrm{i}}-\delta_{\mathrm{i} 2} \delta_{\mathrm{n} 0}\left(\mathrm{f}_{\mathrm{p}} \nu+\sum_{\alpha} \mathrm{f}_{\alpha} \nu^{\alpha}\right)\right]=\frac{1}{2} \sum_{\mathrm{t}}(-1)^{\mathrm{t}} \frac{(2 \mathrm{n}+\mathrm{t})!}{\left(\mathrm{q}^{2}\right)^{2 \mathrm{n}+1+\mathrm{t}}} K_{\mathrm{i}}^{\mathrm{nt}} \tag{15}
\end{equation*}
$$

with $K_{i}^{n t}$ given in (7b).
The connection between the commutators in Eq. (3) and the integrals for $n \geq 0$ on the L.H.S. of (15) can be read off from Eqs. (2, 3, 4, 7) and (15). In principle, the integrals in Eq. (15), and thus the $K_{i}$ nt , can be determined from the electron scattering data, and from these results the $C_{i}^{l}$ and the matrix elements of all commutators in Eq. (3) can be constructed.

Rather than pursue the connection between commutators and the integrals in (15) in more detail, let us present some restrictions on the electron scattering cross sections which follow from the vanishing of various equal-time commutators. These restrictions can be derived straight forwardly from our previous results. Actually there are many more restrictions (involving higher values of n in (15)) than the ones we list; however, all of these others are implied from the relations that we write explicitly because of the conditions

$$
\begin{gather*}
0 \leq \mathrm{A}_{1}  \tag{16a}\\
-\mathrm{M}^{2} \mathrm{~A}_{1} \leq \mathrm{A}_{2} \leq \nu^{2} / \mathrm{q}^{2} \mathrm{~A}_{1} \tag{16b}
\end{gather*}
$$

$$
-5-
$$

These inequalities follow from the definitions in Eq. (2), or equivalently, from the requirement of non-negative cross sections for both transverse and longitudinally polarized photons. ${ }^{8}$ Since $2 \nu>q^{2}$ when $A_{i} \neq 0$ in the integrand of (15), these inequalities are useful in allowing us to conclude that if

$$
\begin{equation*}
\left(q^{2}\right)^{p} \int \frac{\mathrm{~d} \nu}{\nu^{2 \mathrm{n}+1}} \mathrm{~A}_{1} \xrightarrow{q^{2} \rightarrow \infty} 0 \tag{17a}
\end{equation*}
$$

then

$$
\begin{equation*}
\left(q^{2}\right)^{p+2} \int \frac{d \nu}{\nu^{2 n+3}} A_{1} \xrightarrow{q^{2} \rightarrow \infty} 0 \tag{17b}
\end{equation*}
$$

and

$$
\begin{equation*}
\left(\mathrm{q}^{2}\right)^{\mathrm{p}+1} \int \frac{\mathrm{~d} \nu}{\nu^{2 \mathrm{n}+3}} \mathrm{~A}_{2} \xrightarrow{\mathrm{q}^{2} \longrightarrow \infty} 0 \tag{17c}
\end{equation*}
$$

The restrictions listed below can be extended to larger values of $n$ by Eqs. (17). They are similar in form, and also in their origin, to relations discussed recently by Bander and Bjorken. ${ }^{9}$
(i) if $\langle p|\left[j_{0}(x), j_{0}(0)\right]|p\rangle=0$, then for $i=1,2$

$$
\begin{equation*}
\mathrm{q}^{2} \int \frac{\mathrm{~d} \nu}{\nu^{3}} \mathrm{~A}_{\mathrm{i}}\left(\mathrm{q}^{2}, \nu\right) \xrightarrow{\mathrm{q}^{2} \longrightarrow \infty} 0 \tag{18}
\end{equation*}
$$

(ii) if $\langle p|\left[j_{0}(x), j_{i}(0)\right]|p\rangle=0$, then for $i=1,2$

$$
\begin{equation*}
\mathrm{q}^{2} \int_{0}^{\infty} \frac{\mathrm{d} \nu}{\nu}\left[\mathrm{~A}_{\mathrm{i}}-\delta_{\mathrm{i} 2}\left(\mathrm{f}_{\mathrm{p}} \nu+\sum_{\alpha} \mathrm{f}_{\alpha} \nu^{\alpha}\right)\right] \xrightarrow{\mathrm{q}^{2} \rightarrow \infty} 0 \tag{19}
\end{equation*}
$$

(iii) if $\langle\mathrm{p}|\left[\mathrm{j}_{\mathrm{i}}(\mathrm{x}), \mathrm{j}_{\mathrm{k}}(0)\right]|\mathrm{p}\rangle=0$, then there are no further restrictions if (i) and (ii) hold.
(iv) if $\langle\mathrm{p}|\left[\partial_{0} \mathrm{j}_{\mathrm{i}}(\mathrm{X}), \mathrm{j}_{\mathrm{k}}(0)\right]|\mathrm{p}\rangle=0$, as well as (i) and (ii), then for $\mathrm{i}=1,2$

$$
\begin{equation*}
\left(q^{2}\right)^{2} \int_{0}^{\infty} \frac{\mathrm{d} \nu}{\nu}\left[A_{\mathrm{i}}-\delta_{\mathrm{i} 2}\left(\mathrm{f}_{\mathrm{p}} \nu+\sum_{\alpha} \mathrm{f}_{\alpha} \nu^{\alpha}\right)\right] \xrightarrow{\mathrm{q}^{2} \rightarrow \infty} 0 \tag{20}
\end{equation*}
$$

The conditions (18) - (20) for $A_{1}$ can be expressed simply in terms of high energy, high momentum transfer sum rules for the electron scattering cross section in Eq. (1). The amplitude $\mathrm{A}_{2}$ makes its presence felt for backward scattering angles ${ }^{10}$ and is relatively more difficult to extract from the data. For $A_{1}$

$$
\begin{equation*}
\int \frac{\mathrm{d} \nu}{\nu} \frac{\mathrm{~d}^{2} \sigma^{\mathrm{ep}}}{\mathrm{dq}^{2} \mathrm{~d} \nu} \stackrel{\mathrm{E} \rightarrow \infty}{ } \frac{2 \alpha^{2}}{\mathrm{q}^{2}} \int_{0}^{\infty} \frac{\mathrm{d} \nu}{\nu} \mathrm{~A}_{1}+0(1 / \mathrm{E}) \tag{21}
\end{equation*}
$$

Thus, for example, if the commutators (i) - (iv) are all satisfied, as has been suggested, ${ }^{11}$ then it follows from (20) and (21) that

$$
\begin{equation*}
\lim _{E \rightarrow \infty}\left(q^{2}\right)^{3} \int \frac{d \nu}{\nu} \frac{d^{2} \sigma^{e p}}{d q^{2} d \nu} \xrightarrow{q^{2} \infty} 0 \tag{22}
\end{equation*}
$$

If only (i) - (iii) are valid, we would expect the R.H.S. of (22) to be a (non-zero) constant. In fact, for the Sugawara model, ${ }^{12}$ this constant can (almost) be calculated exactly. ${ }^{13}$

To give some perspective to these results, let us compare them to the inequalities derived previously by Bjorken. ${ }^{10,14}$ The first inequality, derived from isospin manipulations on Adler's sum rule, reads (for all $q^{2}$ )

$$
\begin{equation*}
\mathrm{q}^{2} \int_{0}^{\infty} \mathrm{d} \nu \mathrm{~A}_{1 \text { (isoscalar) }} \geq \frac{\pi}{2} \tag{23}
\end{equation*}
$$

The second inequality, based upon quark commutation rules for the space components of the isospin current, is

$$
q^{2} \int_{0}^{\infty} d \nu\left(A_{1}-\frac{q^{2}}{\nu^{2}} A_{2}\right) \text { isoscalar }^{q^{2} \rightarrow \infty} \geq \frac{\pi}{2}
$$

It is clear that we have been optimistic in assuming the existence of all moments of the $h_{i}^{m}$ in Eq. (7b). It is a likely possibility that for $t$ greater than some value the $\mathrm{K}_{\mathrm{i}}^{\mathrm{nt}}$ do not exist; that is, essentially, that for high order derivatives, the commutators in Eq. (3) are not defined. An interesting possibility ${ }^{16}$ is that the integrals in Eq. (15) $\sim \exp \left[-\sqrt{q^{2}}\right]$. The $K_{i}^{n t}$ determined from scattering data according to Eq. (15) would then all be zero, and the connection given in (7) between the commutators and the $\mathrm{K}_{\mathrm{i}}^{\mathrm{nt}}$ would break down. ${ }^{16}$ Except for this kind of occurrence, the conditions listed above as (i) - (iv) are reversible; that is, if théy are satisfied, then the corresponding commutators vanish.

Finally, we remark that radiative corrections and/or multiple photon exchange would tend to decrease the significance of our results.

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3. There also could be a polynomial in $q^{2}$ in the numerator of Eq. (5), but it could be eliminated in favor of polynomials in $\nu$ and in $q^{2}+2 \beta \nu+\sigma$. The latter contribute neither to the commutators nor to the absorbtive parts and hence are ignored.
4. To show this: a) write a twice subtracted dispersion relation for $\mathrm{F}_{\mathrm{i}}^{\mathrm{m}}(\mathrm{m}>2 \mathrm{n} \geq 2)$ noting that both subtraction constants vanish, since $\mathrm{F}_{\mathrm{i}}^{\mathrm{m}} \rightarrow \nu \nu^{\mathrm{m}}$ for small $\nu$; b) expand the result for small $\nu$ and set the coefficient of $\nu^{2 n}(2 n<m)$ equal to zero.
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6. For $A_{1}^{m}, n=0$ and $m=1$, the argument of footnote 4 restricting $2 n-m \geq 0$ is applicable if an unsubtracted dispersion relation is used.
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