# INEQUALITY FOR <br> BACKWARD EIECTRON-AND MUON-NUCLEON SCATTERING AT HIGH MOMENTUM TRANSFER * 

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#### Abstract

From a sum rule for backward $\nu-p$ scattering, valid only in the limit of large four-momentum transfer $q^{2}$, we obtain an inequality for backward e-p inelastic scattering which depends upon the commutator of space components of isospin currents. Given chiral $U(6) \times U(6)$ current algebra, the total backward scattering at fixed large $q^{2}$. is predicted to be at least as great as that from a point Dirac particle with charge $\pm e / 2$.


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[^0]Recently, from Adler's sum rule for neutrino processes ${ }^{1}$

$$
\begin{equation*}
\lim _{\mathrm{E} \rightarrow \infty}\left[\frac{\mathrm{~d} \sigma(\bar{\nu} \mathrm{p})}{\mathrm{dq}^{2}}-\frac{\mathrm{d} \sigma(\nu \mathrm{p})}{\mathrm{dq}^{2}}\right]=\frac{\mathrm{G}^{2}}{\pi}\left(\cos ^{2} \theta_{\mathrm{c}}+2 \sin ^{2} \theta_{\mathrm{c}}\right) \tag{1}
\end{equation*}
$$

we have derived ${ }^{2}$ an inequality for electron and muon-nucleon scattering by isospin manipulation

$$
\begin{equation*}
\lim _{\mathrm{E} \rightarrow \infty}\left[\frac{\mathrm{~d} \sigma_{\mathrm{ep}}}{\mathrm{dq}^{2}}+\frac{\mathrm{d} \sigma_{\mathrm{en}}}{\mathrm{dq}^{2}}\right] \geq \frac{2 \pi \alpha^{2}}{\mathrm{q}^{4}} \tag{2}
\end{equation*}
$$

This inequality is of some interest inasmuch as it predicts a large amount of inelastic scatiering at high momentum transfer $q_{i}^{2}$, something which can be experimentaily tested. The magnitude is comparable to that resulting from scattering off poimi charges; this result can be traced back to the assumption of locality of the isospin current.

However, electron-nucleon scattering is described by two form factors, and the sum rule, Eq. (2) involves only one of them, the "charge" form factor whicin contributes to forward scattering. There arises the question of whether there is any such relation for the other form factor which describes backward scattering. The purpose of this paper is to provide a partial answer ior large $q^{2}$. We write

$$
\begin{equation*}
\frac{\pi \mathrm{d} \sigma_{\mathrm{ep}}}{\mathrm{EE}^{\prime} \mathrm{d} \Omega \mathrm{dE}^{\prime}}=\frac{\mathrm{d} \sigma_{\mathrm{ep}}}{\mathrm{dq}^{2} \mathrm{dE}^{\prime}}=\frac{\mathrm{E}^{\prime}}{\mathrm{E}}\left\{\cos ^{2} \frac{\theta}{2} \sigma_{1 \mathrm{p}}\left(\mathrm{q}^{2}, \nu\right)+\sin ^{2} \frac{\theta}{2} \sigma_{2 p}\left(\mathrm{q}^{2}, \nu\right)\right\} \tag{3}
\end{equation*}
$$

Here $E$ and $E$ ' are incident and final lepton energy and $\theta$ the scattering argie; $\psi_{i}^{z}=-4 E E^{r} \sin ^{2} \frac{\theta}{2}$ and $\nu=E-E^{\prime}$, the laboratory energy of the virtual photon. All hadron states of appropriate momentum have been summed over in writing (2q. (3).

The old inequality is ${ }^{2}$

$$
\begin{equation*}
\int_{o}^{\infty} \mathrm{d} \nu\left\{\sigma_{1 \mathrm{n}}\left(\mathrm{q}^{2}, \nu\right)+\sigma_{1 \mathrm{p}}\left(\mathrm{q}^{2}, \nu\right)\right\} \geq \frac{2 \pi \alpha^{2}}{\mathrm{q}^{4}} \tag{4}
\end{equation*}
$$

The new inequality is (as $\left|q^{2}\right| \rightarrow \infty$ only)

$$
\begin{equation*}
\int_{0}^{\infty} \frac{d \nu}{\nu}\left\{\sigma_{2 n}\left(q^{2}, \nu\right)+\sigma_{2 p}\left(q^{2}, \nu\right)\right\} \geq \frac{4 \pi \alpha^{2}}{\left|q^{6}\right|} \int \mathrm{d}^{3} \mathrm{xe}^{\mathrm{iq} \cdot \mathrm{x}}\langle\mathrm{p}|\left[\mathrm{j}_{\mathrm{x}}^{+}(\mathrm{x}), \mathrm{j}_{\mathrm{x}}^{-}(0)\right]|\mathrm{p}\rangle \tag{5}
\end{equation*}
$$

and $\mathrm{j}_{\mathrm{x}}^{+}$is the plus component of isovector current, normalized such that the commutator in Eq. (5) is unity for the $\mathrm{U}(6) \times \mathrm{U}(6)$ algebra. ${ }^{3}$ Corresponding to Adler's old neutrino sum rule ${ }^{1}$ [the $\beta$-sum rule] for $\sigma_{1}$

$$
\begin{align*}
\int_{0}^{\infty} \mathrm{d} \nu\left[\sigma_{I}^{\nu} \mathrm{p}\left(\mathrm{q}^{2}, \nu\right)-\sigma_{1}^{\nu \mathrm{p}}\left(\mathrm{q}^{2}, \nu\right)\right] & =\frac{\mathrm{G}^{2}}{\pi} \int_{\mathrm{d}^{3} \mathrm{x}}^{3}\langle\mathrm{P}|\left[\mathrm{J}_{0}^{+}\left(\mathrm{x}_{n}\right), \mathrm{J}_{0}^{-}(0)\right]|\mathrm{P}\rangle \mathrm{e}^{\mathrm{i} q} \cdot \mathrm{x} \\
& =\frac{\mathrm{G}^{2}}{\pi}\left(\cos ^{2} \theta_{\mathrm{c}}+2 \sin ^{2} \theta_{\mathrm{c}}\right) \tag{6}
\end{align*}
$$

we also find (as $\left|q^{2}\right| \rightarrow \infty$ only)

$$
\begin{aligned}
\left.\frac{\left|q^{2}\right|}{2} \int_{0}^{\infty} \frac{d \nu}{\nu}\left[\sigma_{2}^{(\nu} p\right)\left(q^{2}, \nu\right)-\sigma_{2}^{(\nu p)}\left(q^{2}, \nu\right)\right] & =\frac{\mathrm{G}^{2}}{\pi} \int \mathrm{~d}^{3} \mathrm{x}\langle\mathrm{P}|\left[\mathrm{J}_{\mathrm{x}}^{+}(\mathrm{x}), \mathrm{J}_{\mathrm{x}}^{-}(0)\right]|\mathrm{P}\rangle \mathrm{e}^{\mathrm{iq} \cdot \mathrm{x}} \mathrm{~m} \\
& =\left\{\begin{array}{c}
\frac{\mathrm{G}^{2}}{\pi}\left(\cos ^{2} \theta_{\mathrm{c}}+2 \sin ^{2} \theta_{\mathrm{c}}\right) \mathrm{U}(6) \mathrm{xU}(6) \text { algebra } \\
0 \quad \text { Spin } O \text { constituents }
\end{array}\right\}
\end{aligned}
$$

$\mathrm{J}_{\mu}^{ \pm}(\mathrm{x})$ is now the full Cabbibo current $(\mathrm{V}-\mathrm{A}, \Delta \mathrm{S}=0,1)$. Although similar, Eq. (7) is not Adler's $\alpha$-sum rule, ${ }^{1}$ which lacks the convergence factor $\mathrm{q}^{2} / \nu^{2}$.

As might be expected, the result depends upon the structure of the commutator of space components of isovector currents. With the chiral $U(6) \times U(6)$ algebra ${ }^{3}$ the commutator on the right-hand side of Eq. (5) is unity, and we expect relatively large scattering. However, one can imagine models in which the isospin current is carried by spinless objects; in this case
the commutator vanishes and there is no lower bound to the backward scattering cross sections.

We start, as with the derivation of the forward-scattering inequality, Eq. (4), with the amplitude $M_{\mu \nu}$ for scattering an isovector current $j_{\mu}^{+}(x)$ from a proton in the forward direction. ${ }^{4}$ (See Fig. 1).

$$
\begin{align*}
M_{\mu \nu}(q, P)= & {\left[q^{2} P_{\mu} P_{\nu}-\left(q_{\mu} P_{\nu}+q_{\nu} P_{\mu}\right) q \cdot P+\left(q \cdot P^{2} g_{\mu \nu}\right] F_{1}\left(q^{2}, q \cdot P\right)\right.} \\
& +\left[q_{\mu} q_{\nu}-g_{\mu \nu} q^{2}\right] F_{2}\left(q^{2}, q \cdot P\right)+\frac{\left[q_{\mu} P_{\nu}+q_{\nu} P_{\mu}-g_{\mu \nu} q \cdot P\right]}{q^{2}}  \tag{8}\\
& +\left[\begin{array}{l}
\text { Polynomial } \\
\text { in } q \text { and } p
\end{array}\right]
\end{align*}
$$

We include Born terms ${ }^{5}$ in the definition of $F_{1}$ and $F_{2} . M_{\mu \nu}$ is defined (up to normalization factors) such that when lepton pairs are att ached it is a piece of the S-matrix. It is not necessarily the time-ordered product of currents. Notice

$$
\begin{equation*}
\mathrm{q}_{\mu} \mathrm{M}^{\mu \nu}=\mathrm{P}^{\nu}+[\text { Polynomial in } \mathrm{q} \text { and } \mathrm{P}] \tag{9}
\end{equation*}
$$

and

$$
\begin{equation*}
P_{\mu}=P_{o} \int\langle P|\left[j_{o}^{+}(x), j_{\mu}^{-}(0)\right]|P\rangle \quad d^{3} x \tag{10}
\end{equation*}
$$

The neutrino- (and antineutrino- proton scattering cross section is proportional to $\operatorname{Im} F_{1}$ and $\operatorname{Im} F_{2}$. The backward-scattering cross sections $\sigma_{2}$ are proportional to the coefficient of $g_{\mu \nu}$

$$
\begin{equation*}
\sigma_{2} \alpha \operatorname{Im}\left\{(\mathrm{q} \cdot \mathrm{P})^{2} \mathrm{~F}_{1}-\mathrm{q}^{2} \mathrm{~F}_{2}\right\} \tag{11}
\end{equation*}
$$

Adler's sum rule is obtained by demanding, as is suggested by Regge theory, ${ }^{6}$ asymptotic behavior for the coefficient of $q_{\mu} P_{\nu}$ less strong than constant.

Thus

$$
\begin{equation*}
\frac{1}{\pi} \int \mathrm{~d} \nu^{\prime} \operatorname{Im} \mathrm{F}_{1}\left(\mathrm{q}^{2}, \nu^{\prime}\right)=\frac{-1}{\mathrm{q}^{2}} \tag{12}
\end{equation*}
$$

Regge behavior also suggests ${ }^{6,7}$ that $F_{2}$ needs one subtraction. We shall assume this is the case:

$$
\begin{equation*}
\mathrm{F}_{2}\left(\mathrm{q}^{2}, \nu\right)=\mathrm{F}_{2}\left(\mathrm{q}^{2}, 0\right)+\frac{\nu}{\pi} \int \frac{\mathrm{d} \nu^{\prime} \operatorname{Im~} \mathrm{F}_{2}\left(\mathrm{q}^{2}, \nu^{\prime}\right)}{\nu^{\prime}\left(\nu^{\prime}-\nu\right)} \tag{13}
\end{equation*}
$$

We now study $\mathrm{M}_{\mu \nu}$ as $\mathrm{q}_{\mathrm{o}} \rightarrow \mathrm{i} \infty$, $\mathrm{q}_{m=1}$ fixed. As in Ref. 4, the coefficient of $1 / q_{0}$ is an equal-time commutator. In the limit $q_{0} \longrightarrow i \infty$,

$$
\begin{align*}
F_{1}\left(q^{2}, \nu\right)= & \frac{1}{\pi} \int^{\infty} \frac{\mathrm{d} \nu^{\prime} \operatorname{Im} F_{1}\left(q^{2}, \nu\right)}{\nu^{\prime}-\nu} \rightarrow \frac{1}{\pi} \int^{\infty} \frac{\mathrm{d} \nu^{\prime}}{\nu^{\prime}} \operatorname{Im} F_{1}\left(q^{2}, \nu^{\prime}\right)  \tag{14}\\
& \left|\nu^{\prime}\right| \geq \frac{\left|q^{2}\right|}{2 m}
\end{align*}
$$

The most reasonable estimate is that

$$
\begin{equation*}
\int \frac{\mathrm{d} \nu^{\prime}}{\nu^{\prime}} \operatorname{Im} \mathrm{F}_{1}\left(\mathrm{q}^{2}, \nu^{\prime}\right) \sim \frac{\text { const }}{\mathrm{q}^{2}} \int \mathrm{~d} \nu^{\prime} \operatorname{Im} \mathrm{F}_{1} \sim \frac{\text { const }}{\mathrm{q}^{4}} \tag{15}
\end{equation*}
$$

which would be rigorously true if $\operatorname{Im} F_{1}$ does not change sign. We assume that u.ere are no delicate cancellations here and we may use Eq. (15). With this estimate the terms involving $F_{1}$ are of order $1 / q_{0}^{2}$ in the limit. Writing $\eta_{\mu}=(1,0,0,0)$, we find, barring pathological cancellations,

$$
\begin{align*}
& \mathrm{M}_{\mu \nu}(\mathrm{q}, \mathrm{P}) \rightarrow[\text { Polynomial }]+\left[\eta_{\mu} \mathrm{q}_{\nu}+\eta_{\nu} \mathrm{q}_{\mu}-2 \eta \cdot \mathrm{q}_{\mu} \eta_{\nu}+\left(\eta_{\mu} \eta_{\nu}-\mathrm{g}_{\mu \nu}\right) \mathrm{q}^{2}\right] \mathrm{F}_{2}\left(\mathrm{q}^{2}, 0\right) \\
& \mathrm{q}_{\mathrm{o}} \rightarrow \mathrm{i} \infty \\
& \quad+\left(\eta_{\mu} \eta_{\nu}-\mathrm{g}_{\mu \nu}\right) \frac{\mathrm{q}_{\mathrm{o}}^{3} \mathrm{P}_{\mathrm{o}}}{\pi} \int_{-\infty}^{\infty} \frac{\mathrm{d} \nu^{\nu}}{\nu^{\prime}} \operatorname{Im} \mathrm{F}_{2}\left(\mathrm{q}^{2}, \nu^{\prime}\right)+\frac{\left[\eta_{\mu} \mathrm{P}_{\nu}+\eta_{\nu} \mathrm{P}_{\mu}-\mathrm{g}_{\mu \nu} \eta \cdot \mathrm{P}\right]}{\mathrm{q}_{\mathrm{o}}}  \tag{16}\\
& \quad+\left[\text { terms more convergent as } \mathrm{q}_{\mathrm{o}} \rightarrow \infty\right]
\end{align*}
$$

[The axial part can be treated in a similar way] • On the ot her hand, the term $0\left(\frac{1}{q_{o}}\right)$ is

$$
M_{\mu \nu} \rightarrow P_{o} \int \frac{\langle p|\left[j_{\mu}^{+}(x), j_{\nu}^{-}(0)\right]|\mathrm{P}\rangle \mathrm{d}^{3} \mathrm{x} \mathrm{e}^{\mathrm{i} q^{\cdot} \times \mathrm{x}}}{\mathrm{q}_{\mathrm{o}}}+\left[\begin{array}{l}
\text { terms with different }  \tag{17}\\
\text { asymptotic behavior }
\end{array}\right]
$$

Thus the term multiplying $\mathrm{F}_{2}\left(\mathrm{q}^{2}, 0\right)$ contributes to any operator Schwinger terms involving $\left[j_{o}^{+}, j_{i}^{-}\right]$. A deviation of the commutator of space components of the currents from the chiral algebra prediction is measured by $\operatorname{Im} \mathrm{F}_{2}$. Indeed

$$
\begin{align*}
\int\langle p|\left[j_{i}^{+}(x), j_{i}^{-}(0)\right]|P\rangle d^{3} x e^{i q \cdot x} & =\left[1+\frac{q^{4}}{\pi} \int \frac{d \nu^{\prime}}{\nu^{\prime}} \operatorname{Im} F_{2}\left(q^{2}, \nu^{\prime}\right)\right]  \tag{18}\\
& =\frac{q^{2}}{\pi} \int_{-\infty}^{\infty} \frac{d \nu^{\prime}}{\nu^{\prime}}\left[-\nu^{\prime}{ }^{2} \operatorname{Im} F_{1}\left(q^{2}, \nu^{\prime}\right)+q^{2} \operatorname{Im} F_{2}\left(q^{2}, \nu^{\prime}\right)\right]
\end{align*}
$$

where we have used Eq. (12). The quantity in brackets is proportional to the vector piece of $\sigma_{2}^{\bar{\nu} \mathrm{p}}\left(\mathrm{q}^{2}, \nu\right)$ or $\sigma_{2}^{\nu \mathrm{p}}$ as defined in Eqs. (3) and (11). After a routine struggle with normalization factors (most simply done by considering free fields) one arrives at the sum rule Eq. (7). The same isospin manipulations ${ }^{2}$ as used in obtaining Eq. (4) from Eq. (6) are sufficient to get Eq. (5) from Eq. (7).

It is tempting to assume the result Eq. (7) to be generally valid for all $q^{2}$. However, consideration of the limit as $\mathrm{q}^{2} \longrightarrow 0$ gives

$$
\begin{equation*}
\left(\mu_{p}-\mu_{n}\right)^{2}=1 \tag{19}
\end{equation*}
$$

in considerable disagreement with experiment.
The following physical picture of the result Eq. (5) suggests itself: If the "elementary constituents" (if any) of the nucleon, which couple to isospin, were spinless, there would be very little backward scattering at
large $q^{2}$, because backward scattering demands helicity flip. If the constituents have spin $1 / 2$ the scattering should be incoherent and proportional to the sum of squares of the magnetic moments of the constituents. ${ }^{8}$

Experimental verification of the inequality Eq. (5) may be difficult because of the problems of radiative corrections.

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4. We follow the derivation outlined in J. D. Bjorken, Phys. Rev. 148, 1467 (1966); see also "Symmetry Principles at High Energy" (1967 Coral Gables Conference), Freeman, 1967.
5. The apparent pole at $q^{2}=0$ in Eq. (8) is cancelled by another pole in the Born terms. We set the nucleon mass equal to unity.
6. This assumption is, we believe, the least trustworthy in the derivation of that result. This is because non-Regge behavior has been shown to exist in the coefficient of $P_{\mu} P_{\nu}$ (J. Bronzan, I. Gerstein, B. Lee, and F. Low, Phys. Rev. Letters 18, 32 (1967), V. Singh, Phys.Rev. Letters 18, 36 (1967)). If similar asymptotic behavior, corresponding to a fixed pole or Kronecker $\delta$ at $J=1$, also occurs in the coefficient of $q_{\mu} P_{\nu}$, we lose the sum rule Eq. (6) as well as the once-subtracted dispersion relation, Eq. (13). Arguments that this does not happen in simple models have been given by the above authors; there are other arguments by deAlfaro, Fubini, Furlan, and Rosetti (to be published).
7. H. Harari, Phys. Rev. Letters 17, 1303 (1966).
8. This picture is similar to that discussed for forward scattering by K. Gottfried (Phys. Rev. Letters, to be published). There also exist sum rules of this kind in nuclear physics; for a review see deForest and Walecka, Advances in Physics, 15, 1 (1966).

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Kinematics for forward scattering of a current from a nucleon.


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