

The Hierarchy Problem : It's Solutions (?) and Their Collider Signatures



- Successes of the Standard Model
- What IS the Hierarchy Problem?
 - ⇒ Related Problems $\left\{ \begin{array}{l} \text{Fine-tuning} \\ \text{EWSB} \\ \vdots \end{array} \right.$
- Possible 'Solutions' ...
- Signatures at future colliders ...
LHC and LC ...
- Summary / Conclusions

The Standard Model of Particle Physics

Symmetry: ^(GCD) $SU(3)_C \times \underbrace{SU(2)_L \times U(1)_Y}_{EWK}$ } Spontaneously broken

Particles: spin-1/2 “matter” fermions

- 3 Generations of Quarks ($N_C = 3$)

$$\begin{pmatrix} u \\ d \end{pmatrix} \quad \begin{pmatrix} c \\ s \end{pmatrix} \quad \begin{pmatrix} t \\ b \end{pmatrix}$$

- 3 Generations of Leptons ($N_C = 1$)

$$\begin{pmatrix} \nu_e \\ e \end{pmatrix} \quad \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix} \quad \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}$$

spin-1 gauge bosons – “force carriers”

- 8 gluons (g) mediate strong interactions
- Photon (γ) mediates electromagnetism
- W^\pm and Z mediate the weak interactions

→ [Higgs boson ?]

This structure is experimentally confirmed!

? → GRAVITY ??

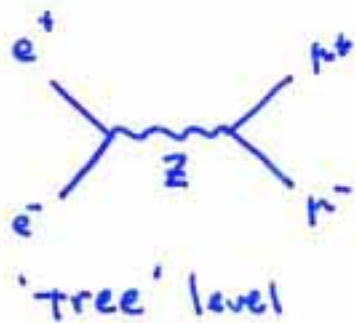
The SM generally agrees w/ precision data at
 the 0.1% level or better...[†]

$$\left\{ \begin{array}{l} M_{\nu}/m \approx 5 \text{ ppm} \\ m_{\nu 1 \text{ loop}} \approx 44 \text{ ppm} \end{array} \right\}$$

→ full 1-loop calculations completed

→ almost all 2-loop + many important 3-loop diagrams calculated...

(QED : $(g-2)_e$ has been done to 4-loops !!
 but the SM is much more complicated!)

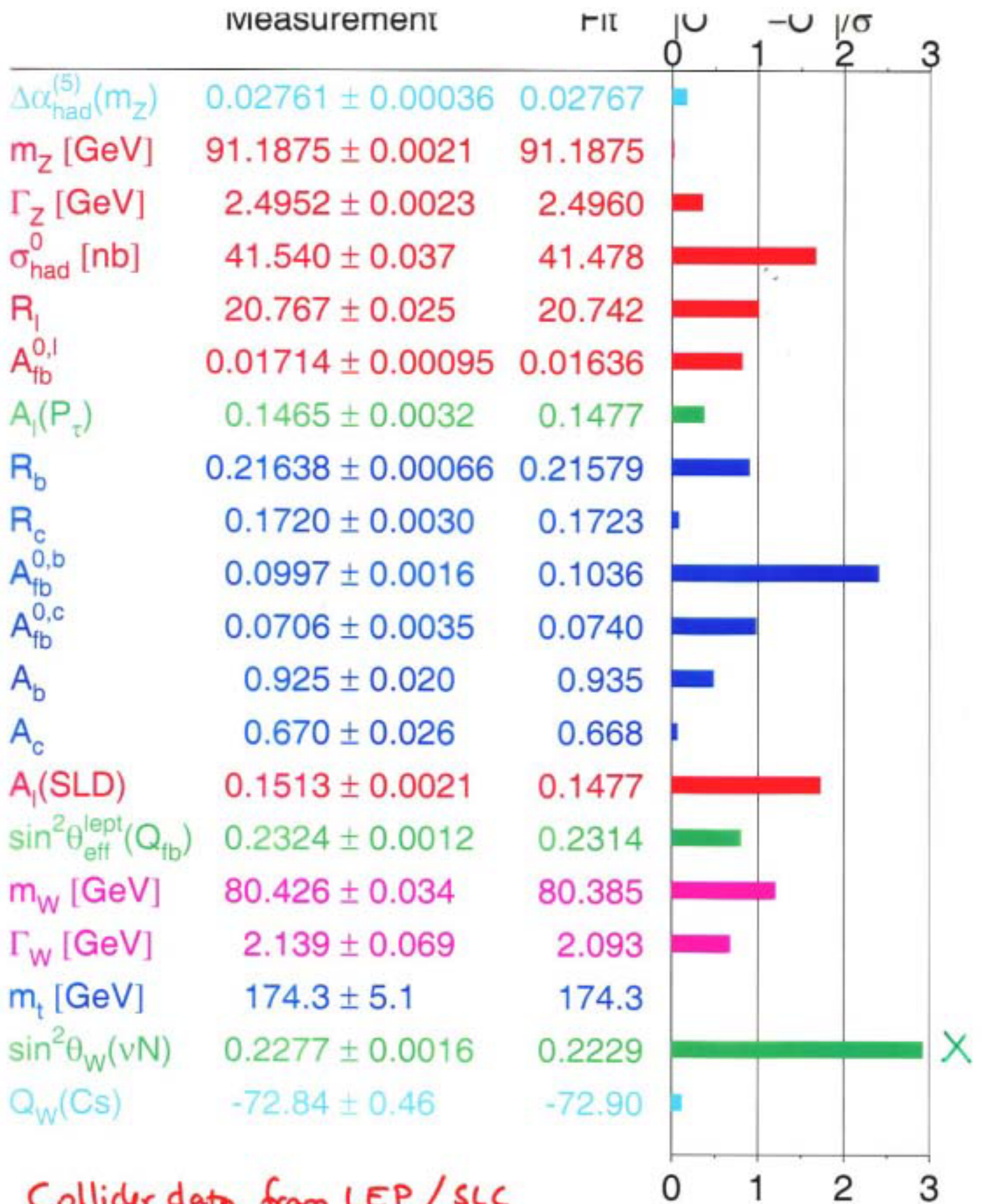


$$O(\alpha)$$



• However many issues still remain.. as do many unanswered questions....

† + plans to go another factor of 10 in future collider experiments



Collider data from LEP / SLC
+ Tevatron

⊕ DIS + APV

"Why is gravity so much weaker than EM?"

In SM language this is rephrased ...

Why is $M_{\text{Planck}} \approx \sqrt{G_N^{-1}} \gg M_{\text{Weak}} \approx \sqrt{G_F^{-1}} ?$
 $\sim 10^{19} \text{ GeV}$ $\sim 1 \text{ TeV}$

→ the (gauge) Hierarchy problem ($10^{15}!$)

- Another (related?) problem:

$\left[\Lambda_{\text{cosmological constant}} \right]^{1/4} \sim M_{\text{Pl}} \leftrightarrow$ the only scale in GR!!

instead $\sim 10^{-3} \text{ eV}!$ $(10^{120}!)^{1/4}$

Note the curiosity: $M_{\text{Pl}} (m_W/M_{\text{Pl}})^2 \sim 10^{-3} \text{ eV}!$

⇒ Is this telling us something??

- The HP is related to other issues in the SM many 'solutions' deal with these.

Hierarchy Problem

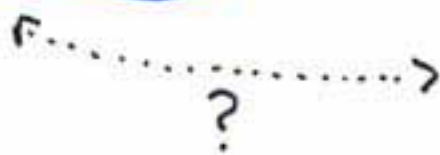
Sine-tuning problem

origin of EW Symmetry breaking

i.e., what breaks the symmetry at $\sim 1\text{TeV}$?

(if unbroken, no HP!)

Cosmological constant problem



loop correction in SM w/ cutoff Λ

large loop momenta

$$\sim \int^{\Lambda} \frac{d^4k}{k^4} \dots$$

momentum flowing in a loop Feynman graph

gauges: $\frac{\delta m_W^2}{m_W^2} \sim \frac{\alpha}{4\pi} \ln \frac{\Lambda}{m_W}$



fermions: $\frac{\delta m_e}{m_e} \sim \frac{\alpha}{4\pi} \ln \Lambda/m_e$



.. are 'log-sensitive' to cut-off. Even if $\Lambda \sim M_{\text{pl}}$ these corrections are not too large since $\frac{\alpha}{4\pi} \ll 1$

• Logs occur because of gauge/chiral symmetries

• What about the Higgs scalar[†] in the SM?

No protection!

[†] naive method of EWSB



$$m_H^2 = m_{H_0}^2 + \frac{3G_F}{4\sqrt{2}\pi^2} \Lambda^2 \left\{ m_H^2 + 2m_W^2 + m_Z^2 - 4m_t^2 + \dots \right\}^\dagger$$

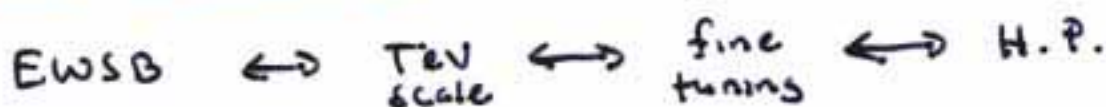
\uparrow QUADRATICALLY sensitive to Λ !!

N.B.!

We expect $m_H \sim 125 \text{ GeV}$ or so in SM [Fig 1]

If $\Lambda \sim M_{Pl}$, $m_{H_0}^2 + C \{ \}$ have to cancel by 30 significant figures!

Fine-tuning problem : Clearly



If $m_H \sim M_{Pl}$ no problem!

'Solutions' clearly link all 3 problems

\dagger and $m_H^2 + m_Z^2 + 2m_W^2 - 4m_t^2 \dots \neq 0!$

not allowed by data (this is fine-tuning too!)

⊕ want to remove these divergences at higher order as well

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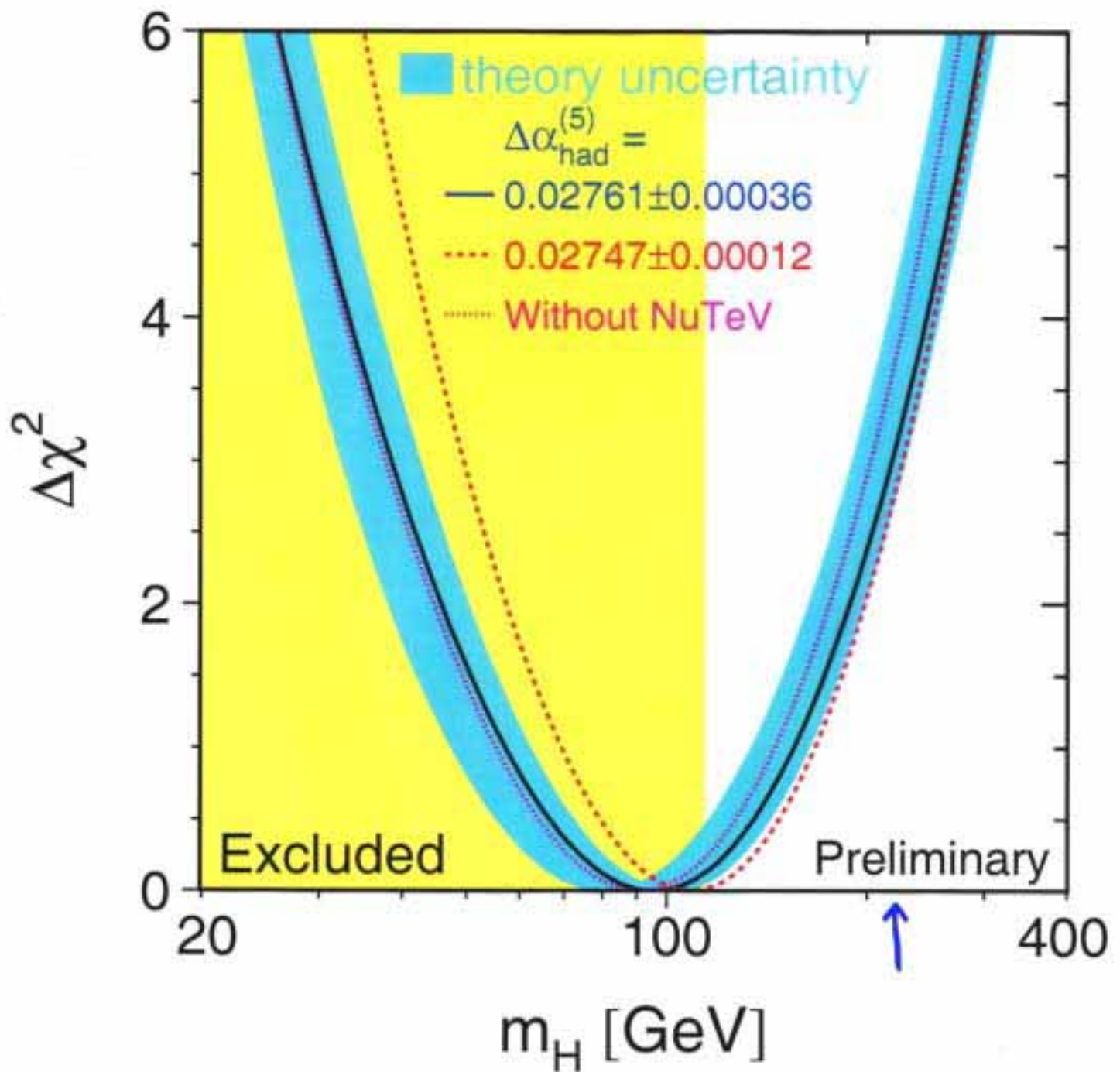
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The SM ⊕ existing data places
a strong upper bound on $m_H \lesssim 219$ GeV.

Direct searches → $m_H \gtrsim 114$ GeV.



How do we get

$$\frac{3G_F}{4\sqrt{2}\pi^2} \Lambda^2 \{ \dots \} \sim (100-200 \text{ GeV})^2 \quad ?$$

"Solutions"

perturbative weakly-coupled dynamics {
Supersymmetry (SUSY) [80][†]
"Little Higgs" (LH) [02]

Idea: add a symmetry { and new particles associated with it } to make $\Lambda \rightarrow \log \Lambda^2$

Strong dynamics: { technicolor [79]

Idea: replace Higgs by condensates - not favored by EW data - won't discuss here

avoid the problem {
 M_{Pl} is really \sim few TeV [99]
no Higgs at all: warped ED w/ mixed BC [03]
(rather complicated)

† most 'familiar'

† They couple exactly the same as do SM fields - only differ by spin- $\frac{1}{2}$ + mass

New particles in SUSY Theory †, ‡

- Spin $\frac{1}{2}$ quarks \Rightarrow spin 0 squarks
- Spin $\frac{1}{2}$ leptons \Rightarrow spin 0 sleptons
- Spin 1 gauge bosons \Rightarrow spin $\frac{1}{2}$ gauginos
- Spin 0 Higgs \Rightarrow spin $\frac{1}{2}$ Higgsino

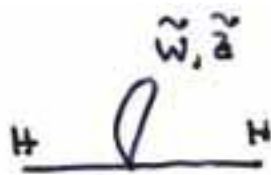
} neutralinos
charginos

Experimentalists dream....many particles to search for!

What mass scale?

Supersymmetry is broken....no scalar with mass of electron

† recall we 'doubled' the # of particles before thanks to Dirac
- anti particles !!



subleading term

$$\Delta m_H^2|_{\text{top}} \sim \alpha \left\{ \Lambda^2 (m_t^2 - m_{\tilde{t}}^2) + N (m_{\tilde{t}}^2 - m_t^2) \ln \frac{\Lambda}{M_W} + \dots \right\}$$

(same couplings) \leftrightarrow Stop loop $\left\{ \begin{array}{l} \text{top loop} \\ \text{recall fermion sign} \end{array} \right.$ Kinematic mass in loop

cancel!

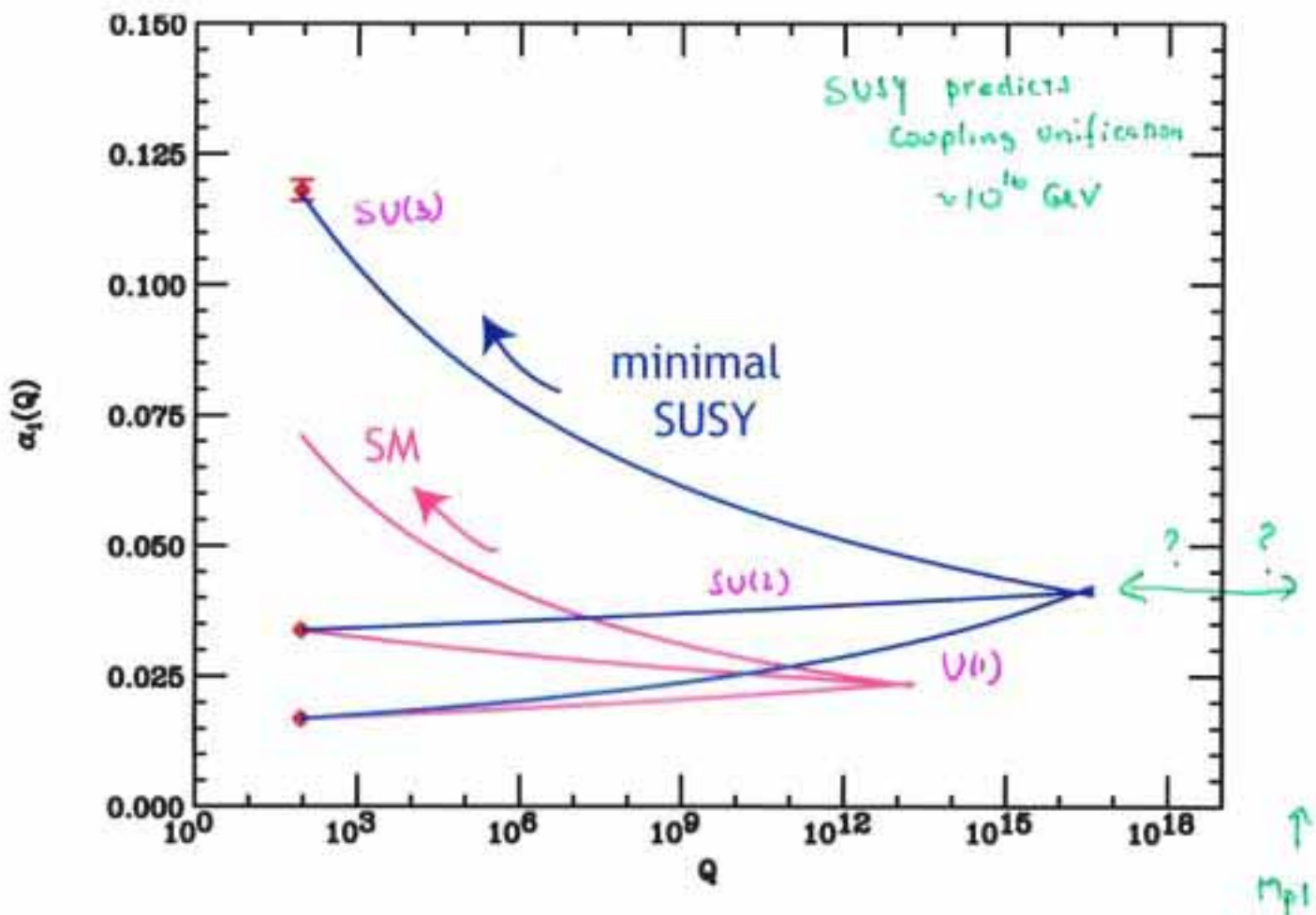
$$\Delta m_H^2 \sim \underbrace{|m_{\text{susy}}^2 - m_{\text{SM}}^2|}_{\text{can't be too big}} \ln \Lambda / M_W \approx \text{TeV or so}$$

now
In fact:
no $\propto \Lambda^2$ terms
at all

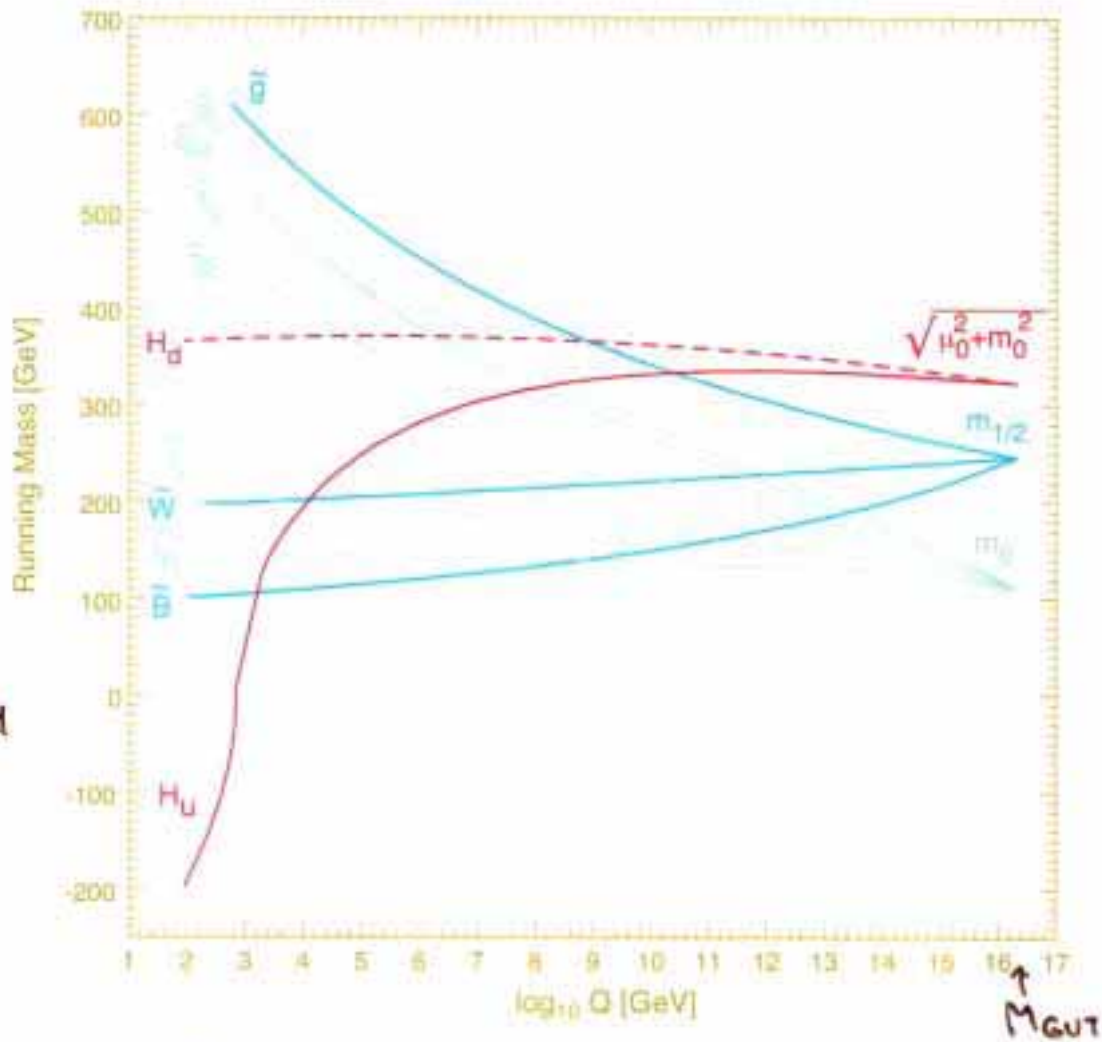
If $m_{\text{susy}} = m_{\text{SM}}$ $\Delta m_H^2 = 0!$ but $m_{\text{susy}} \neq m_{\text{SM}}!$

SUSY is a broken symmetry - how??
(many ideas...)

- Nice things about SUSY;
 - allows for coupling unification $M \sim 10^{16}$ GeV
 - 'radiative' breaking of SM via RGE
and big $m_{\text{top}} \approx 175$ GeV



For heavy top, SSB may follow naturally in SUSY



Higgs
mass driven
negative
by RGE

(sign of M^2 indicated)

Kane, et al. (hep-ph/9312272, *Phys. Rev. D* **49**, 6173 (1994))

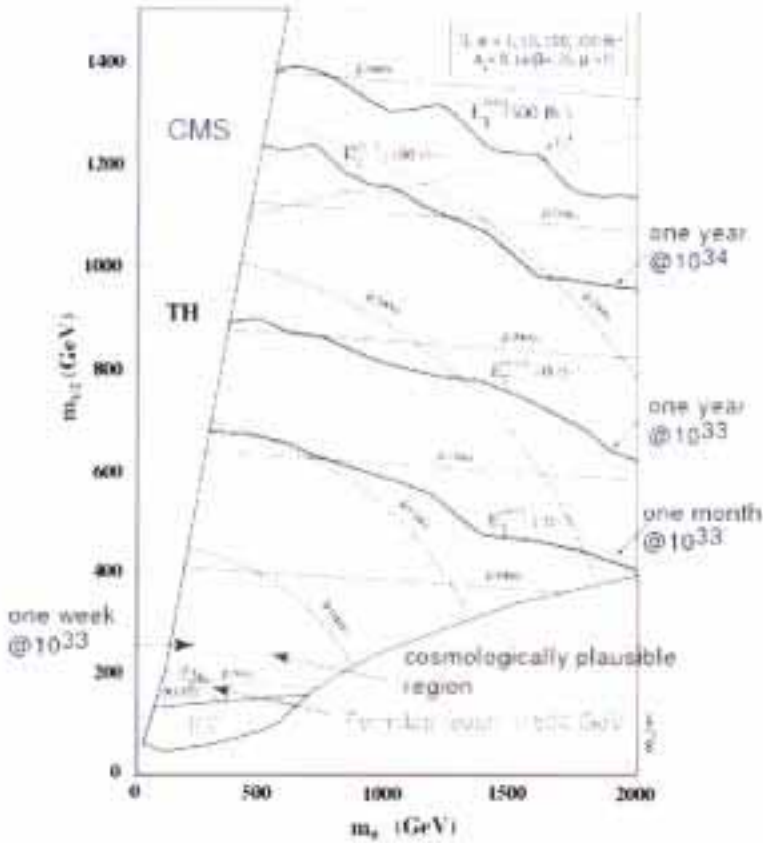
(cont)

- Perhaps a better fit to Precision data (?)
- Lightest SUSY particle (χ_1^0) is a dark matter candidate
- Should be found at LHC / LC (linear ^{e⁺e⁻} collider)
 - discover it at LHC...
 - measure masses / spins / couplings at LC
 - + verify its SUSY!
- link to Superstrings

SUSY 'stabilizes' the hierarchy but does not explain why it occurs... nor why it is a broken symmetry

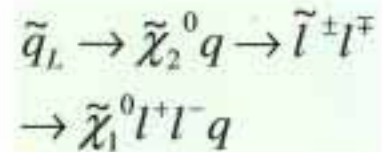
If no susy signs at LHC - theory probably dead!

•LHC/Tevatron will find SUSY



Catania, CMS

- Discovery of many SUSY particles is straightforward
- Untangling spectrum is difficult
 - ⇒ all particles produced together
- SUSY mass differences from cascade decays; eg



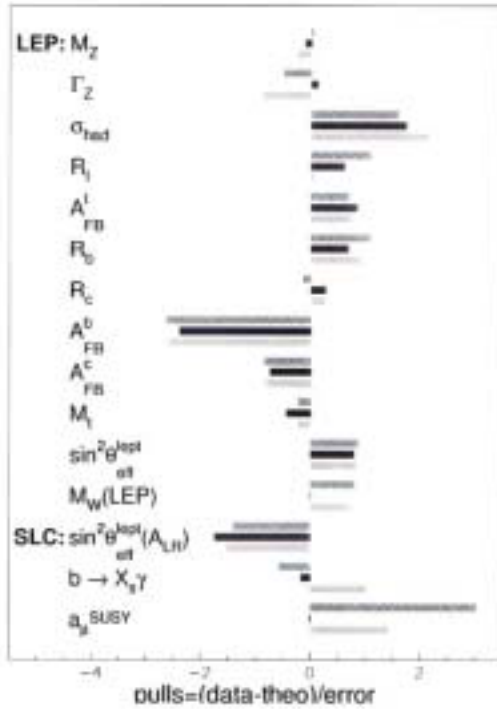
- $M_{\tilde{\chi}^0}$ limits extraction of other masses

Precision measurements can't tell you source of new physics

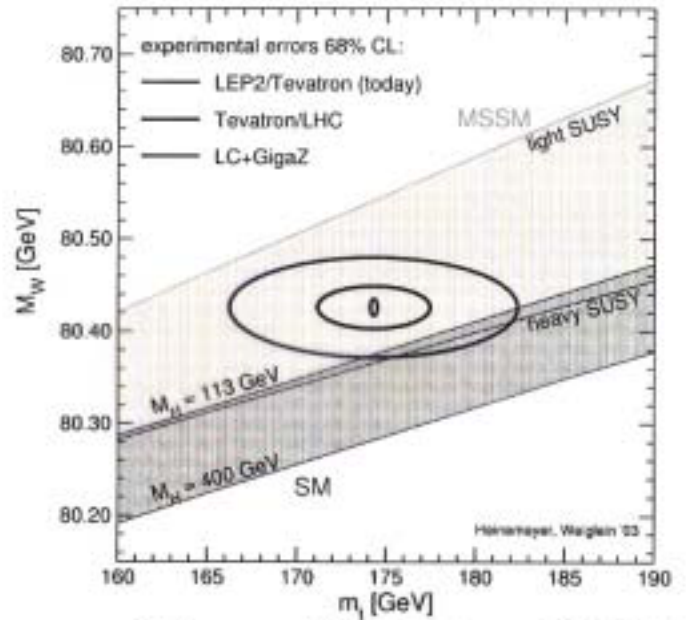
- Example: try to fit precision data to MSSM

— SM: $\chi^2/\text{d.o.f} = 27.2/16$
 — MSSM: $\chi^2/\text{d.o.f} = 16.4/12$
 — CMSSM: $\chi^2/\text{d.o.f} = 23.2/16$

**MSSM slightly better fit (17% prob)
 vs SM (5% prob)**
MSSM prefers "light" SUSY



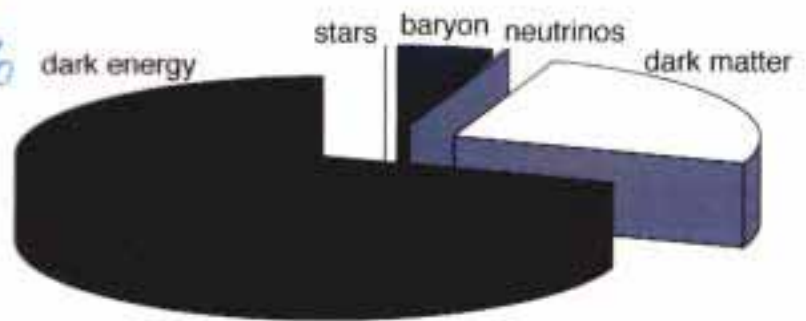
deBoer & Sanders, hep-ph/0307049



Heinemeyer & Weiglein, hep-ph/0307177

What is the universe made of?

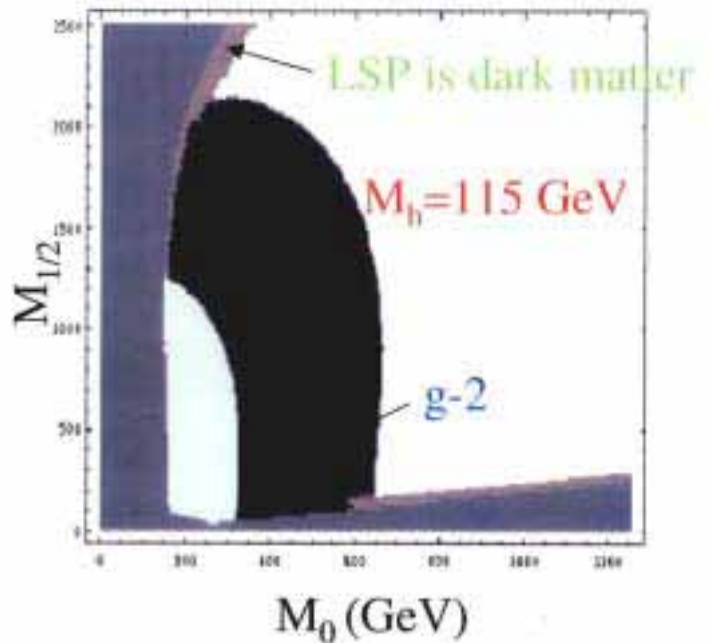
- Stars and galaxies are only 0.1%
- Neutrinos are ~0.1–10%
- Electrons and protons are ~5%
- Dark Matter ~25%
- Dark Energy ~70%



H. Murayama

Supersymmetry provides understanding of dark matter?

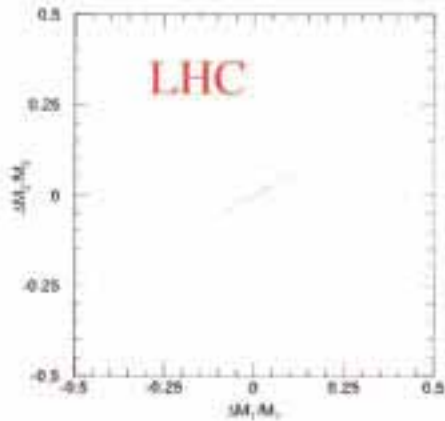
- Lightest SUSY particle (LSP) could be dark matter candidate!
- LSP is weakly interacting, neutral, and stable
- LSP in range of LC/LHC
- LC can determine LSP mass; check dark matter predictions



Drees, hep-ph/0210142

LHC & LC improves SUSY mass resolution

- LSP mass constrained at LHC at 10% level



Bachacou, Hinchliffe, Paige, hep-ph/9907518

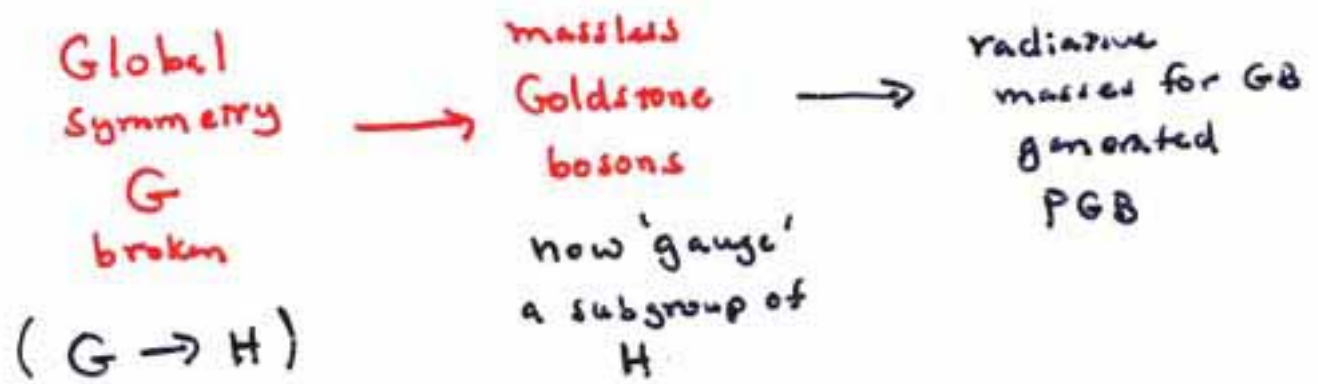
⇒ LC input improves accuracy significantly

- Take LSP mass as input from LC

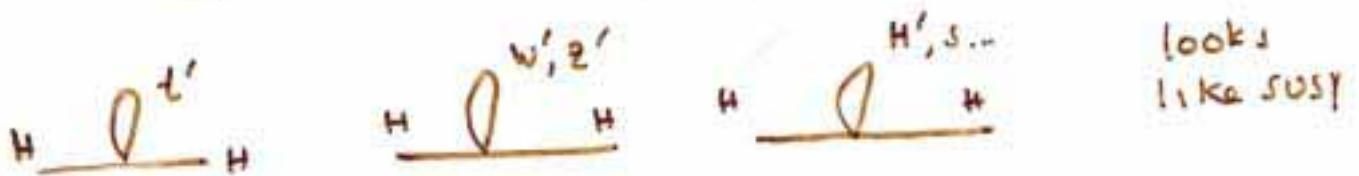
(GeV)	LHC	LHC+ LC(.2%)	LHC+ LC (1%)
$\Delta m(\tilde{z}_1^0)$	9.2	.2	1
$\Delta m(\tilde{l}_R)$	9.2	.5	1
$\Delta m(\tilde{\chi}_2^0)$	9.0	.3	1
$\Delta m(\tilde{b}_1)$	23	17	17
$\Delta m(\tilde{q}_L)$	15	5	5

Weiglein, LHC/LC Study

Little Higgs : less ambitious \rightarrow "Higgs is a pseudo goldstone boson of a non-linearly realized broken symmetry." - whoa!



Enlarge Symmetry of Higgs field \rightarrow extra matter couplings



But these guys have the same spin as sm ones!
w/ opposite sign couplings [careful group theory]

- New particles at $\sim 4\pi M_{\text{weak}} \sim 1-3 \text{ TeV}$
- New 'scale' at $\sim (4\pi)^2 M_{\text{weak}} \sim 10-20 \text{ TeV} = \tilde{\Lambda}$

$$\Delta m_H^2 \sim \alpha \left\{ 0 \cdot \tilde{\Lambda}^2 + N \ln \tilde{\Lambda}/m_W \right\} + N' \alpha^2 \tilde{\Lambda}^2$$

$\tilde{\Lambda}^2$ terms pushed to 2-loops \therefore small if $\tilde{\Lambda} \gtrsim 10-20 \text{ TeV}$

Delays H.P. until after $\tilde{\Lambda}$... new theory is required.. - not a real 'solution'

↳ "Ultraviolet Completion"

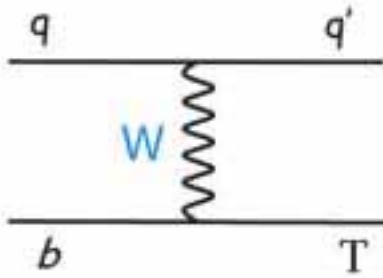
→ Why is $\tilde{\Lambda}$ ($\sim 10-20\text{TeV}$) present + not M_{pl} ?

"The SM is not applicable above $\tilde{\Lambda}$." Hmmm...

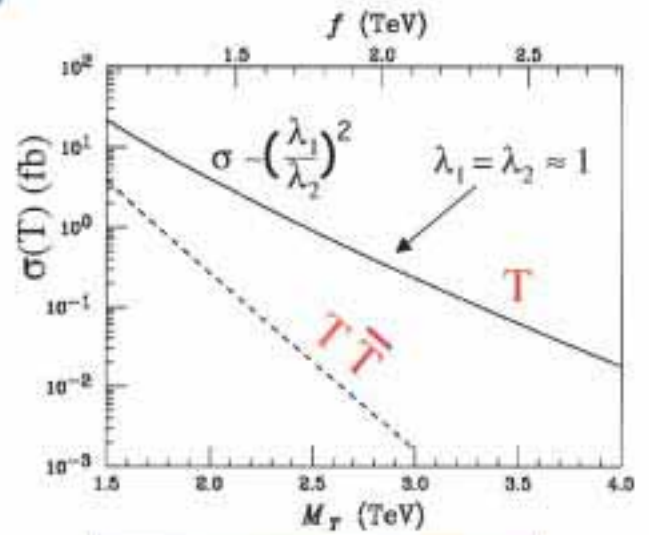
⇒ To test the scenario, find t', W', Z' etc
at $\sim \text{TeV}$ w/ right couplings.

This can be done at LHC/LC....

- Production mechanism = Wb fusion



BR	{	$T \rightarrow bW$	50 %
		$T \rightarrow tZ$	25 %
		$T \rightarrow th$	25 %



λ_1, λ_2 Yukawa couplings

$$\frac{\Gamma}{M} \approx \frac{1}{16\pi} \approx 2 \%$$

$$q\bar{q} \rightarrow Z_H \rightarrow e^+e^-$$

$\mu^+\mu^-$ not used due to invariant mass resolution

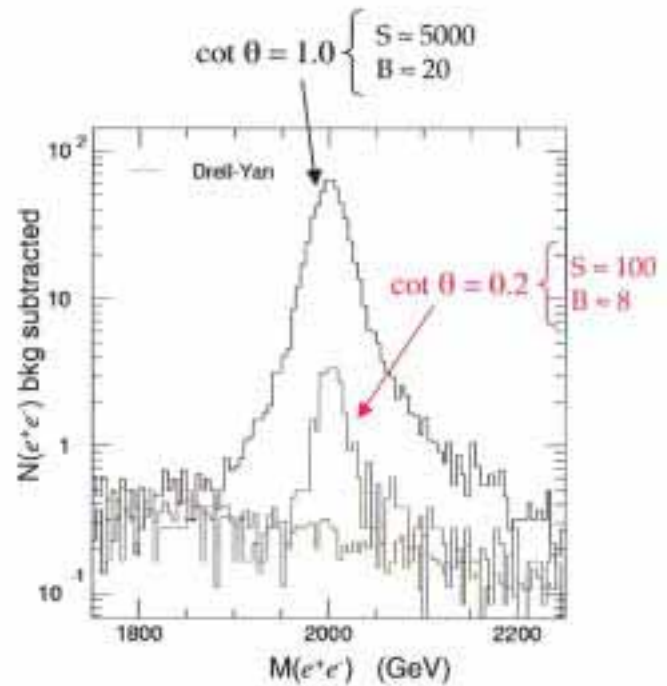
Selection cuts:

- 2 isolated electrons with $p_T > 20$ GeV and $|\eta| < 2.5$
- minimum invariant mass equal to 800 GeV

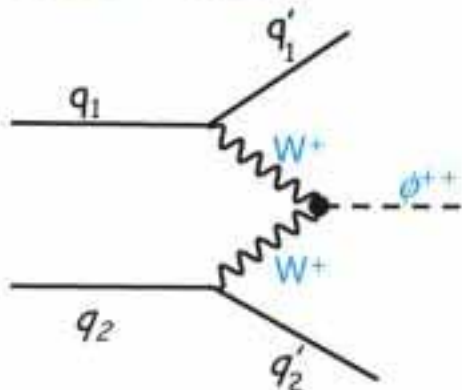
Background:

Drell-Yan ($q\bar{q} \rightarrow Z/\gamma \rightarrow e^+e^-$)
other bkg's are under evaluation

$$M(Z_H) = 2 \text{ TeV} \quad \mathcal{L} = 3 \cdot 10^5 \text{ pb}^{-1}$$



- Production: VBF mechanism (vector boson fusion)

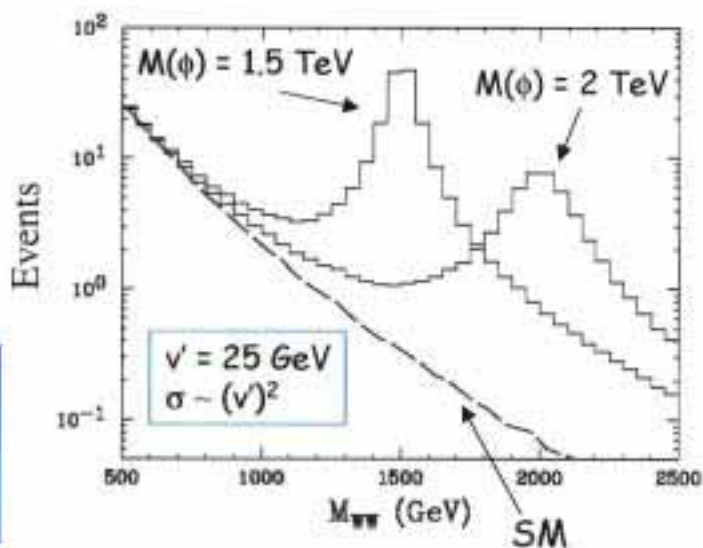


$$\mathcal{L} \approx 3 \cdot 10^5 \text{ pb}^{-1}$$

$$W^+W^+ \rightarrow \phi^{++} \rightarrow W^+W^+$$

Decays $\left\{ \begin{array}{l} \phi^{++} \rightarrow W^+W^+ \\ \phi^{++} \rightarrow \tau^+\tau^+, e^+e^+, \dots \end{array} \right.$

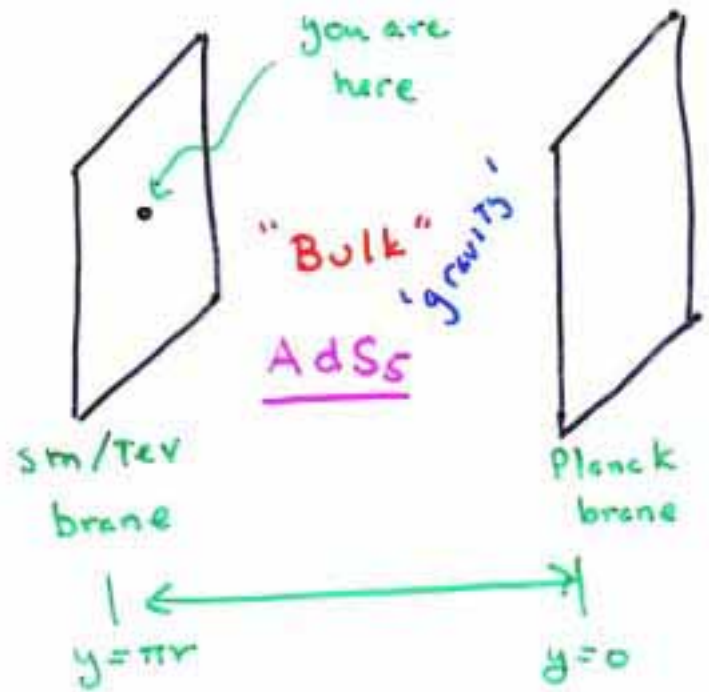
might be suppressed



To 'avoid' the HP we need Extra Dimensions!

Randall-Sundrum Set-up

- 2, 4-d surfaces 'branes' in 1 extra finite dimension



- The space between is curved a la GR.

- anti deSitter space (const. neg. curvature)

$$R_{\mu\nu} = -20k^2 g_{\mu\nu}$$

(Ricci curvature)

$$ds^2 = e^{-2ky} \delta_{\mu\nu} dx^\mu dx^\nu - dy^2$$

warp factor (!)

$$k \sim M_{Pl}$$

→ Any mass^(M) you write down is naturally $\sim M_{Pl}$

BUT if $kr \sim 10$ we'd observe $m \sim e^{-kr} M$ which is $\sim \text{TeV}$. [Example]

→ No large numbers!

$$\Lambda \rightarrow \text{TeV}$$

[We can get $kr \sim 10$ naturally]

Hierarchy is an artifact of the warping of space

How do we test this??

Kaluza-Klein Excitations : gravity is 'trapped'

in the extra dim 'box' \Rightarrow momentum quantization

$$P^2 = \underbrace{P_\mu P^\mu}_{m^2} + \underbrace{P_5 P^5}_{\substack{\text{some quantized } m_n^2 \\ \text{set of values}}}$$

- If space were flat
- Here we have AdS5

$$P_5 = n/r \text{ from } \underline{Q.M}$$

$$P_5 = x_n k e^{-\pi k r_c}$$

$\hookrightarrow J_1(x_n) = 0$
a Bessel function!

graviton KK's..

$$m_n^{KK} \sim \text{few hundred GeV} - \text{few TeV}$$

couple to matter as $\sim \frac{1}{\Lambda_\pi} T_{\mu\nu} h_{(n)}^{\mu\nu}$

w/ $\Lambda_\pi = M_{pl} e^{-k r_c} \sim \text{few TeV}$

These are massive, Spin-2 resonances you can see at a collider...

$\Lambda \rightarrow \Lambda_\pi$

- Spin -2
- Branching fractions } is it a graviton KK?

Imagine the Higgs field on the TeV brane..

$$S = \int d^4x dy \sqrt{-g} \left\{ g^{\mu\nu} \partial_\mu H^\dagger \partial_\nu H - \lambda (H^2 - v_0^2)^2 \right\} \delta(y - \pi r)$$

$$\left\{ \left[e^{-2ky} \right]^4 \right\}^{1/2} \rightarrow e^{2ky} \delta^{\mu\nu}$$

Higgs vev $\sim M_{pl}$

$$S = \int d^4x \left\{ e^{-2kr_2\pi} \partial_\mu H^\dagger \partial^\mu H - e^{-4kr_2\pi} \lambda (H^2 - v_0^2)^2 \right\}$$

$$\text{let } H \rightarrow e^{kr_2\pi} \hat{H}$$

$$S = \int d^4x \left\{ \partial_\mu \hat{H}^\dagger \partial^\mu \hat{H} - \lambda \left(\hat{H}^2 - \underbrace{v_0^2 e^{-2kr_2\pi}}_{\text{vev is TeV!}} \right)^2 \right\}$$

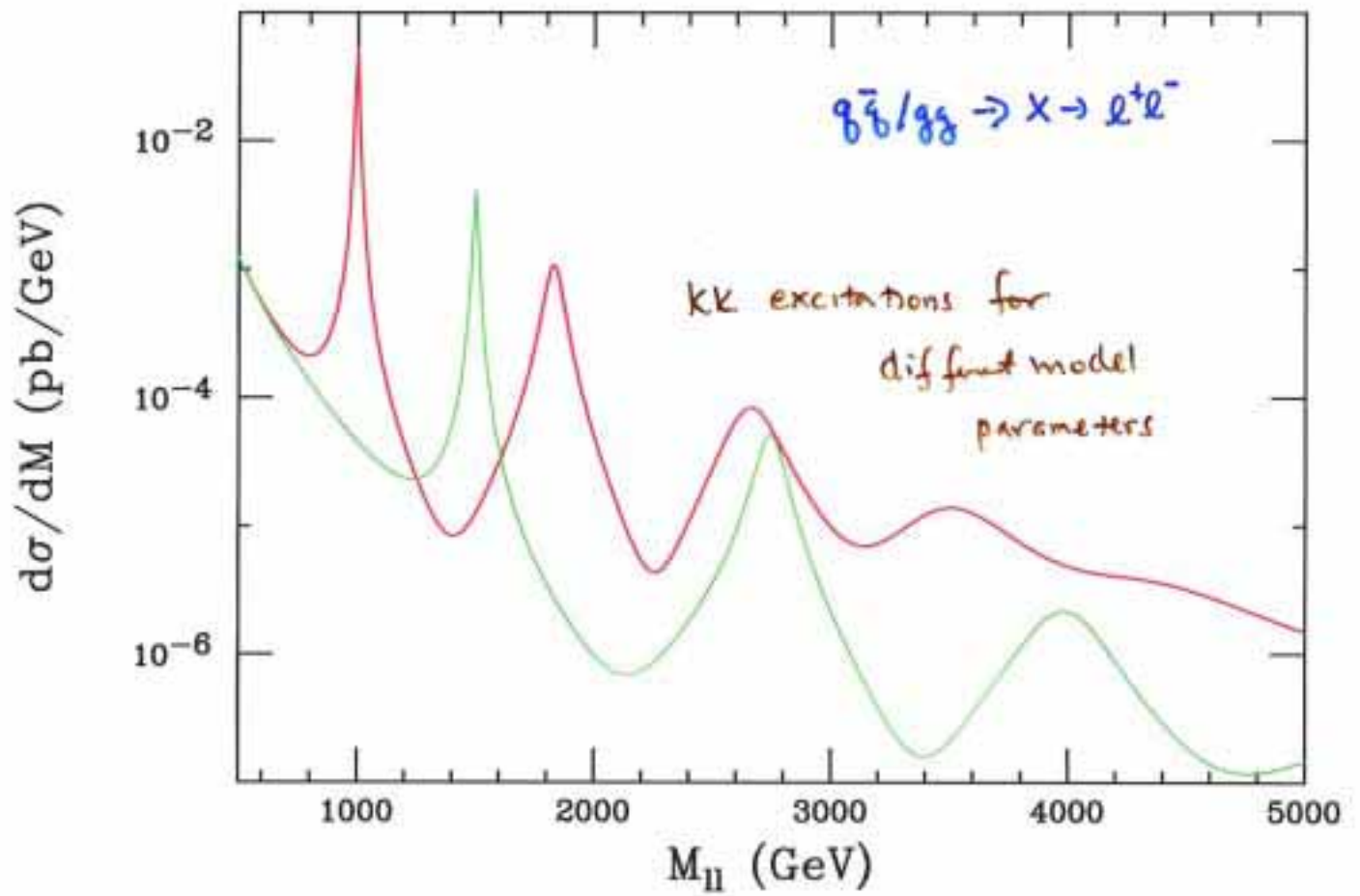
"canonically" normalized field..

- the one we'd just write down...

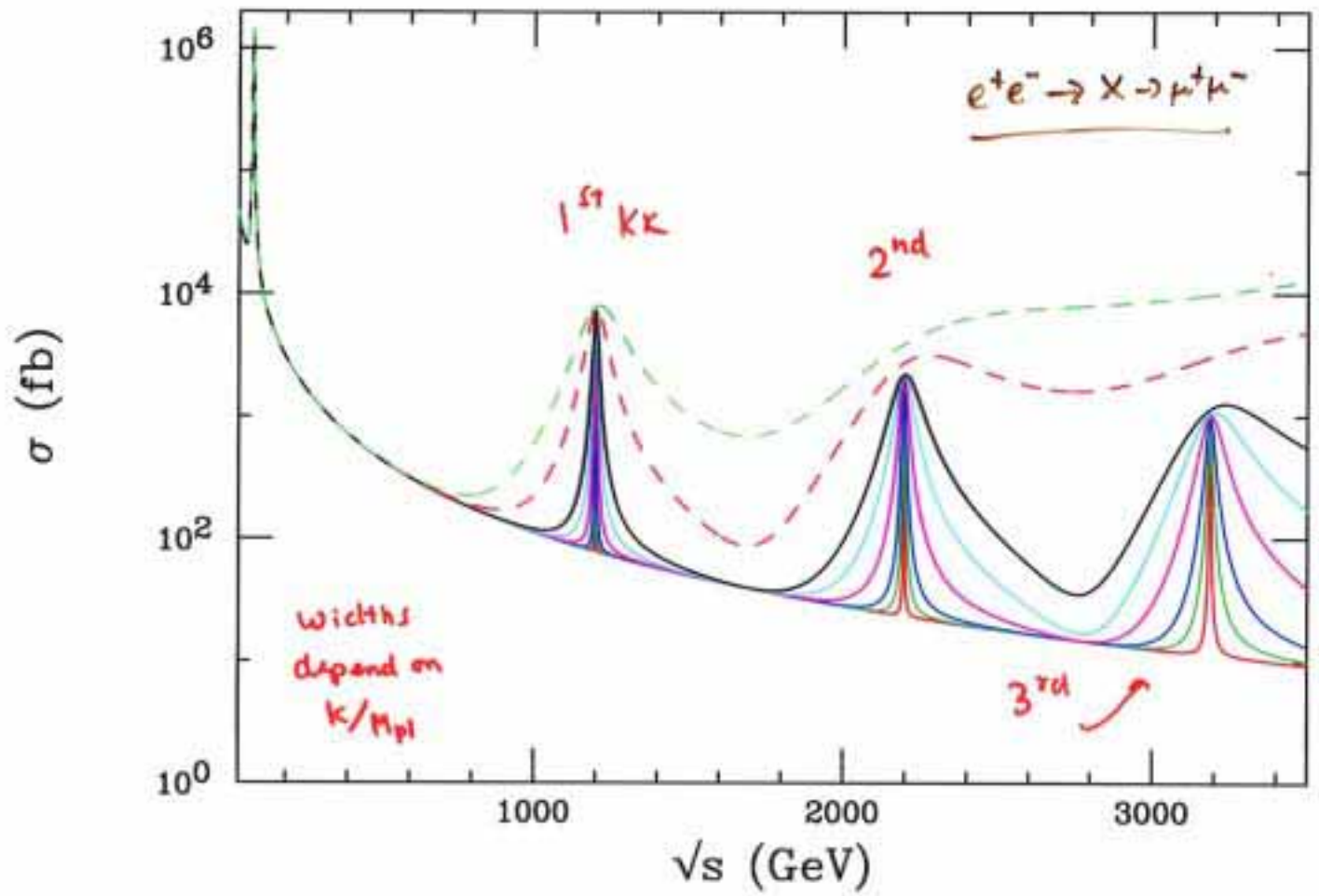
↑
mass scale associated w/
Higgs are TeV

Amazingly simple!

graviton resonance production at LHC

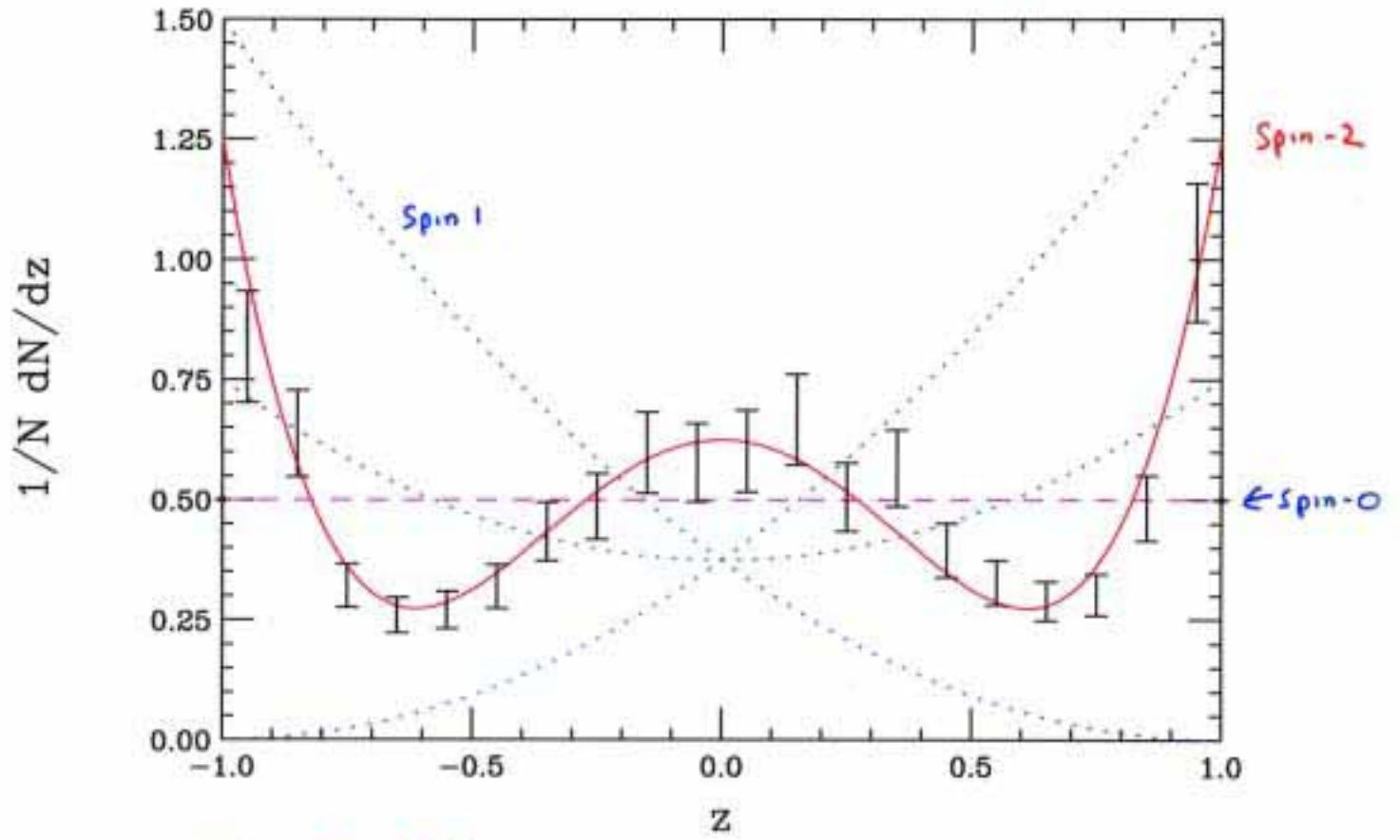


gravitons at TeV scale LC



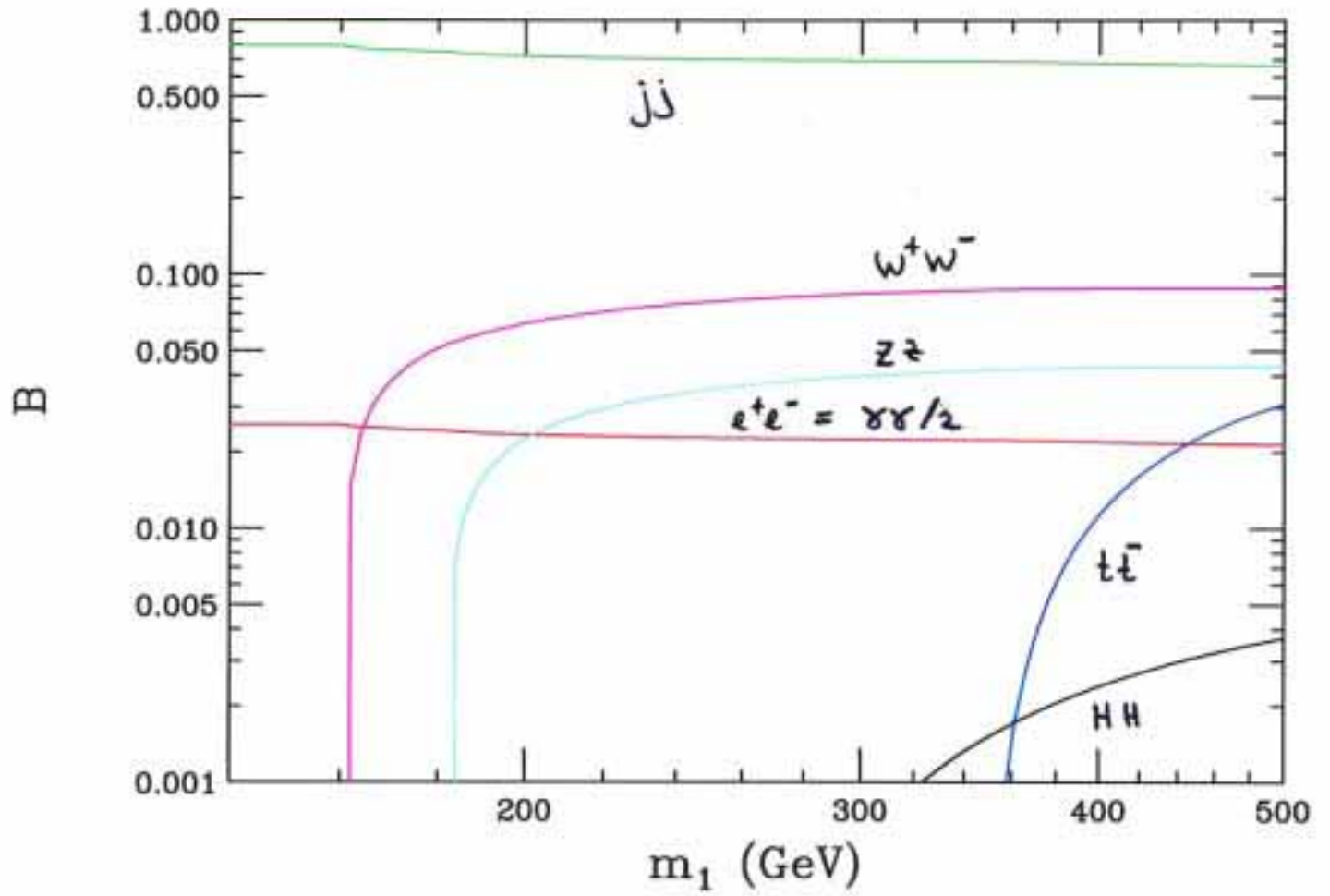
angular distributions determine

Valence Spin



Spin 2 ??

Decay branching fractions for gravitinos



No Higgs at All

- a massless 4-d field has 2 d.o.f A_μ
- " " " 3 d.o.f $\{A_\mu, A_5\}$

can A_5 be the Higgs?

↑
in 4d, an extra scalar

- It would be nice since (i) no Higgs 'by-hand' and (ii) gauge invariance protects A_5 mass

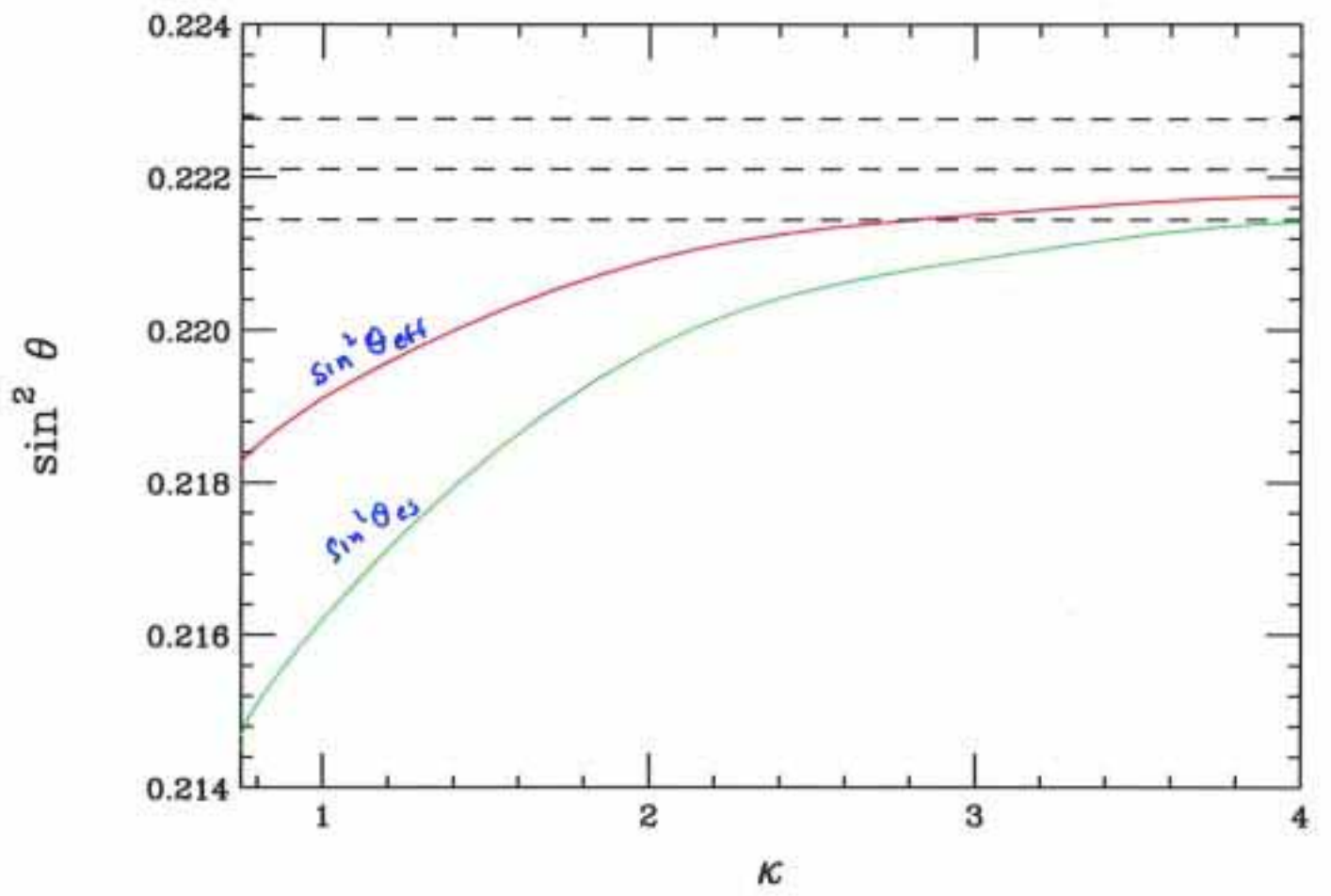
⇒ Not trivially obvious this can be made to work

e.g., in $SU(2)$ A_5 is a triplet but we know H is a doublet to get m_W/m_Z correct...

This is a hot research area at present BUT it appears that (at least approximate) a model NOW EXISTS with all the right features ...

Csaki, Grujean
Pilo, Terning, Nomura

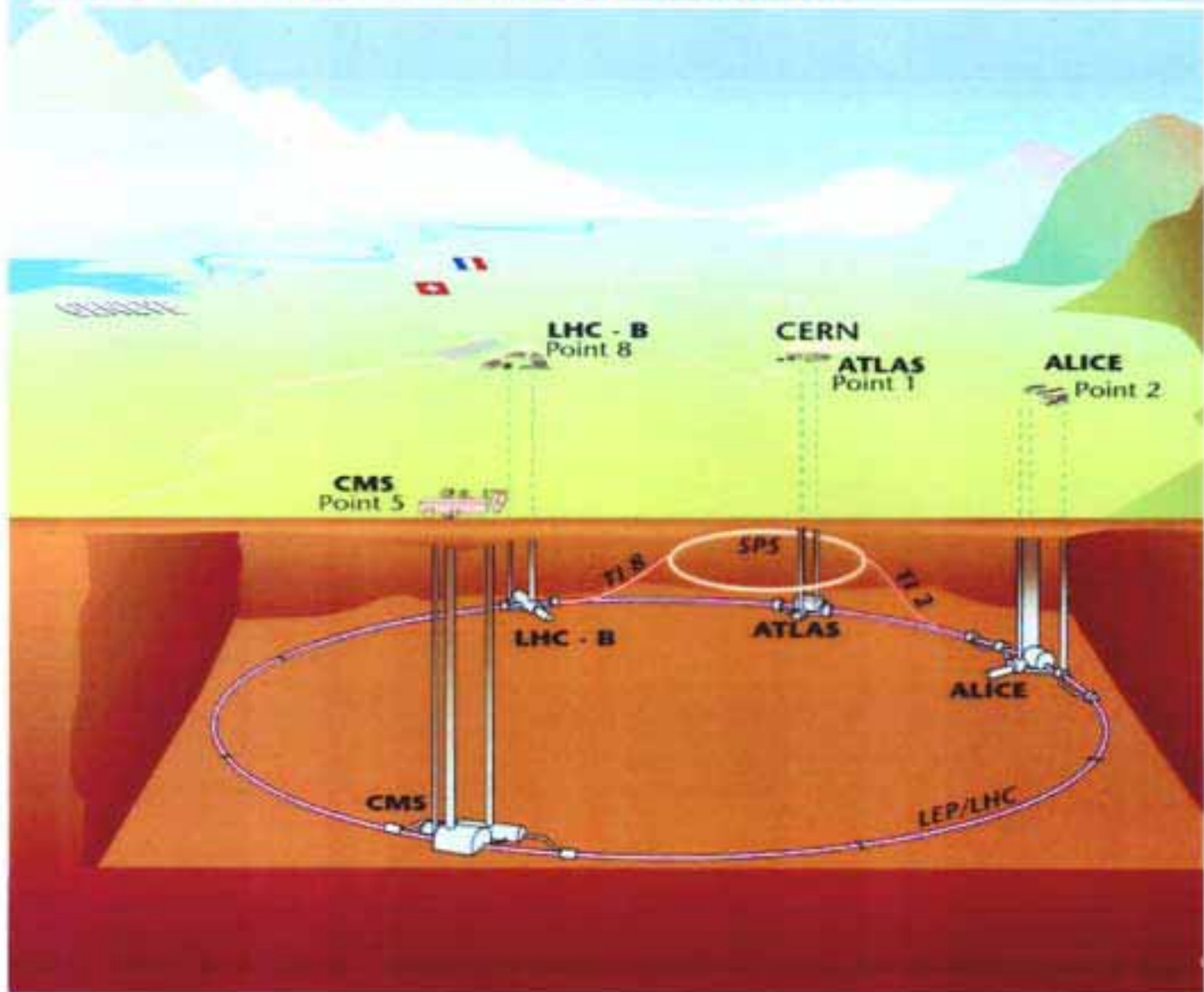
- We are exploring this + extensions of this model now to see how well it works.... Davoudiasl, Hewett, Lillie + TGR

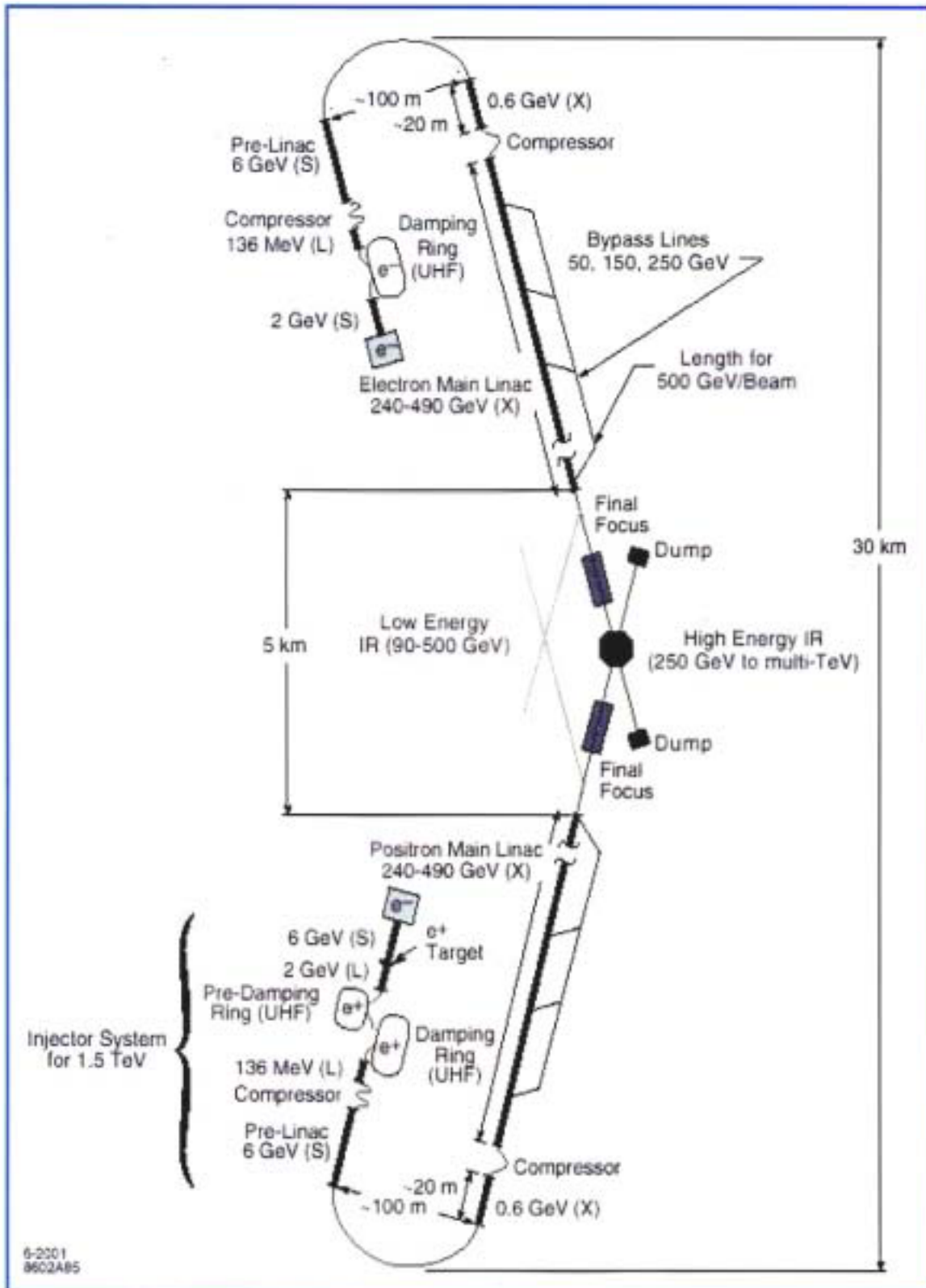


Summary + Conclusions

- The HP is difficult, but some partial sol'n's have been proposed
- The HP is intimately linked to EWSB and fine-tuning problems
 - maybe also to the cosmological constant problem as well!
- Each sol'n leads to testable predictions at the TeV we're only now beginning to probe
- The LHC + LC will provide critical info. needed to attack this problem further.
- While we wait, there's lots of work to do!

Overall view of the LHC experiments.

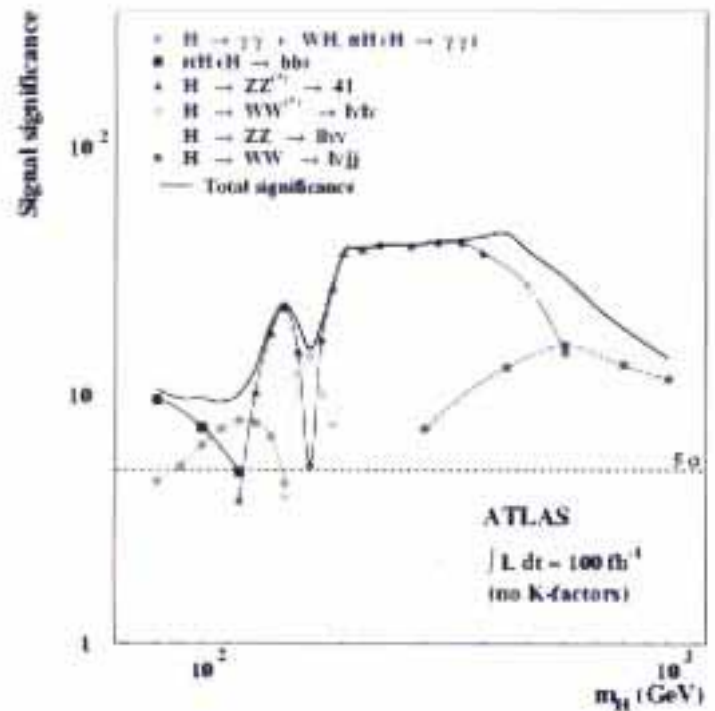




NLC Layout

CERN Large Hadron Collider (LHC)

- pp interactions at $\sqrt{s} = 14$ TeV
- LHC will discover Higgs boson if it exists
- Sensitive to M_{H^0} from 100-1000 GeV
- Higgs signal in just a few channels
- Physics circa 2008



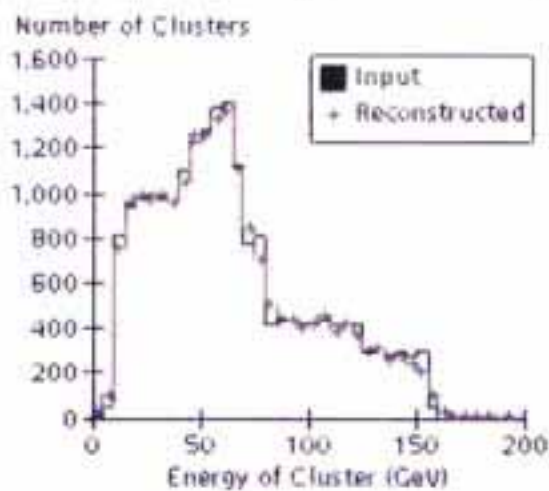
ATLAS TDR

LC mass measurements from endpoints

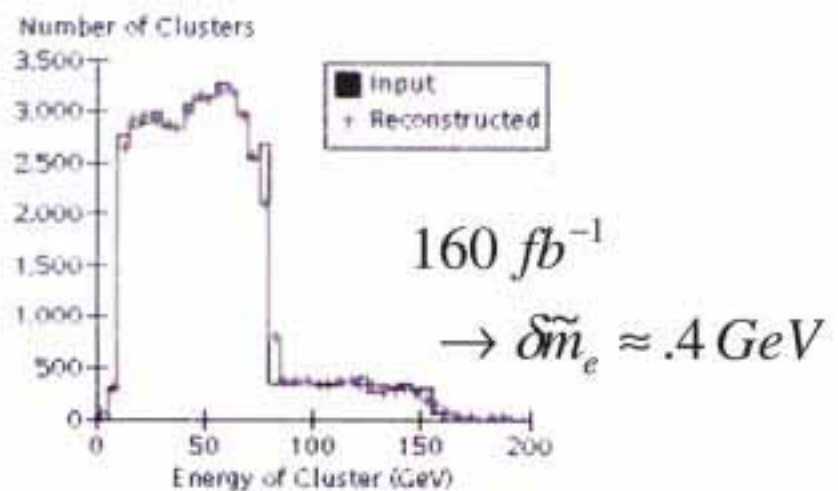
$$e^+e^- \rightarrow \tilde{e}^+\tilde{e}^- \rightarrow e^+e^-\tilde{\chi}_1^0\tilde{\chi}_1^0$$

$$m_{\tilde{l}}^2 = \frac{sE_{\max}E_{\min}}{(E_{\max} + E_{\min})^2}, \quad 1 - \frac{m_{\tilde{\chi}_1^0}^2}{m_{\tilde{l}}^2}$$

Selectron Calorimetry (Left Beam Pol.)

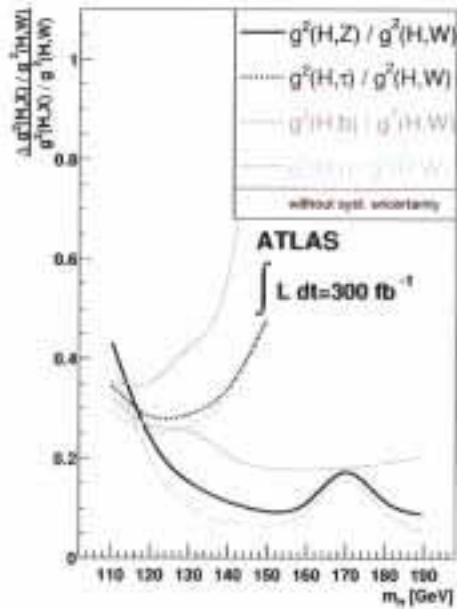


Selectron Calorimetry (Right Beam Pol.)



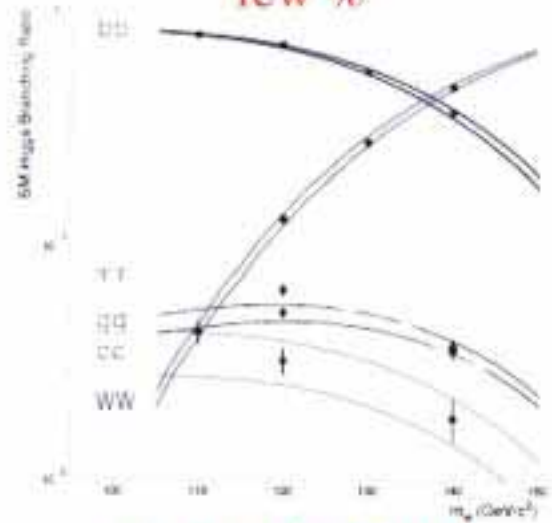
Once we find the Higgs, we need to measure its couplings

Ratios of coupling constants measured quite precisely at LHC



Duhrssen, ATL-PHYS-2003-030

LC measures couplings to a few %



e^+e^- LC at $\sqrt{s}=350$ GeV

$L=500 \text{ fb}^{-1}$, $M_{H^\pm}=120$ GeV

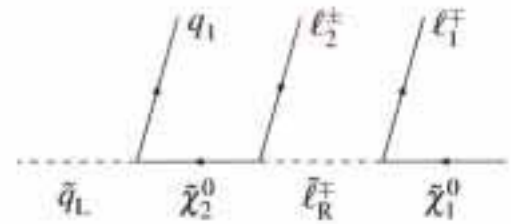
Battaglia & Desch, hep-ph/0101165

Linear Collider is the place!

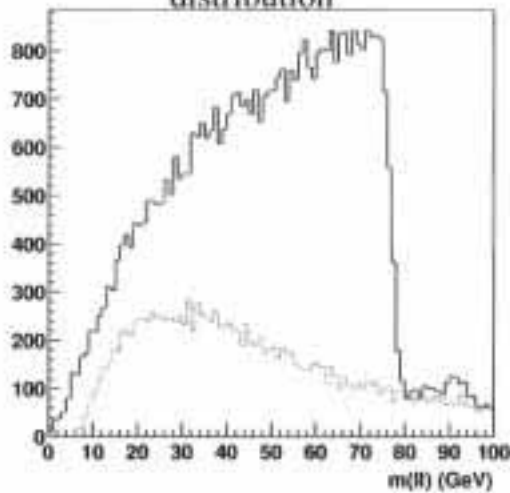
Measuring SUSY masses at hadron colliders

- Complicated decay chains
- Main tool: dilepton edge from $\tilde{\chi}_2^0 \rightarrow l^+ l^- \tilde{\chi}_1^0$
- Sbottom/squark and gluino reconstruction
- Proportional to mass differences: strong mass correlations

$$\tilde{\chi}_2^0 \rightarrow l^+ l^- \tilde{\chi}_1^0$$



Same flavor-opposite sign lepton distribution



Invariant qll mass

