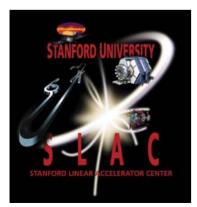
Angular Distribution of Gamma-rays from Up-to-date p-p Interactions

GLAST Science Lunch Talk Aug 3, 2006

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□ Recap

- p-p interaction model
- parameterization of gamma-ray spectra
- □ Angular distribution of gamma-rays simulations
- □ Parameterization of angular distribution
- □ Final words, ideas, input!

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p-p Interaction Model

- Kamae et al. (2005): previous models of inelastic proton-proton interaction lacked several key features:
 - diffraction dissociation process
 - violation of Feynman scaling
 - logarithmically rising inelastic p-p cross section
- □ Up-to-date model created
 - missing features included

□ Simulations show

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- 30%-80% more π^0 produced
- gamma-ray spectrum harder than incident proton spectrum
- Kamae et al. (2005) claimed to explain 50% of the GeV-excess in EGRET data



Inelastic Proton-Proton Interaction

□ Inelastic proton-proton interaction model:

- non-diffractive interaction including violation of Feynman scaling
 - □ high energy (Tp > 52.6 GeV): simulated with Pythia 6.2
 - □ low energy (Tp ≤ 52.6 GeV): parametric model of pion cross sections by Blattnig et al. (2000)
- diffraction dissociation process

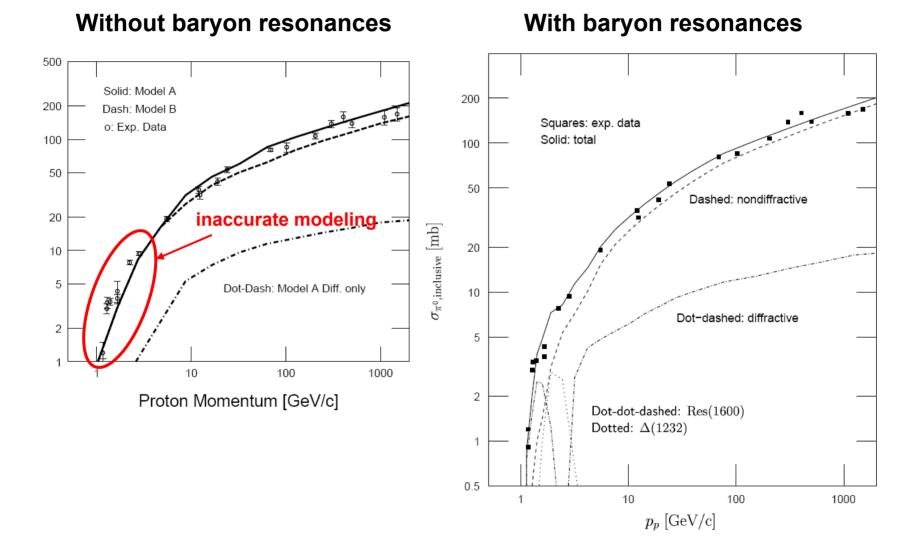
□ Above not accurate enough near pion production threshold

- add modeling of two baryon resonances
- Δ(1232) and Res(1600)
- □ MC simulations for above processes
- Calculated inclusive cross sections for the stable secondary particles
 - $-\gamma$ -rays, electrons, positrons and four neutrino species

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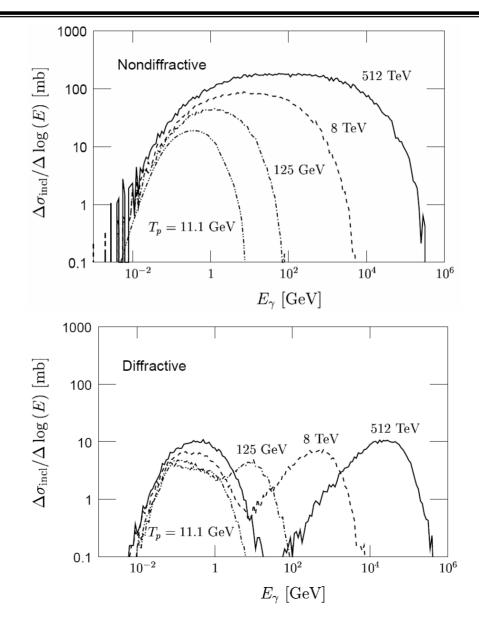


π^0 Multiplicity





Inclusive Gamma-ray Cross Sections



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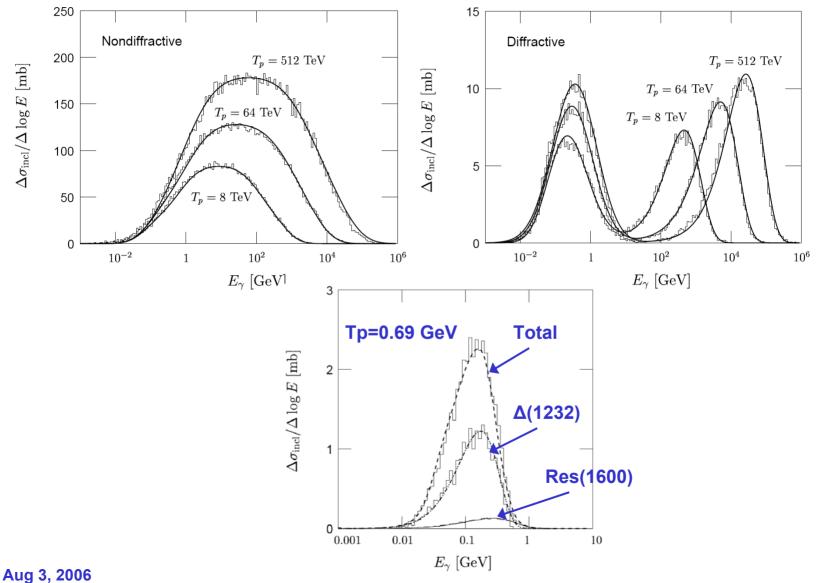
Parameterization of Incl. Cross Section

- Wanted: the inclusive cross section as a function of proton kinetic energy and photon energy
- □ For each of the four components:
 - fit to common function forms (functions of log(E_{sec}), one form for each component
 - \rightarrow set of parameters for each Tp
 - then fit these parameters as functions of Tp
- Parameterization has been completed for all stable secondary particles
 - $-\gamma$ -rays, electrons, positrons and four neutrinos





Comparison: MC vs Parametric model

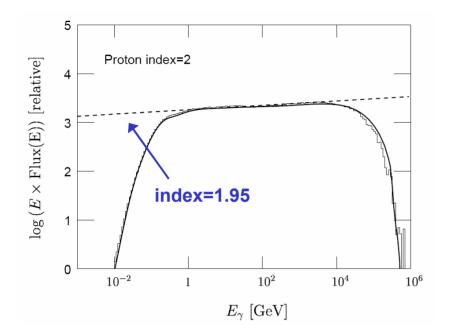


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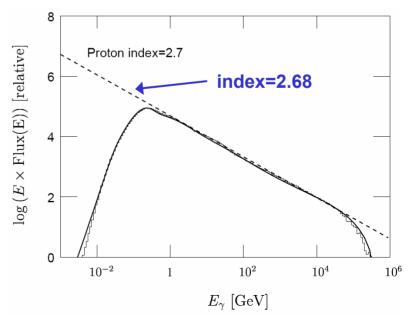


Gamma-ray Spectrum



Histograms: Monte Carlo simulations Solid lines: calculations with parametric model Dashed lines: asymptotic power-law

(no absolute normalization!)









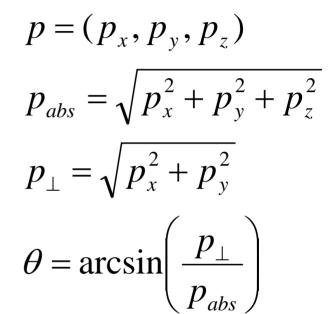
Angular Distribution - Simulations

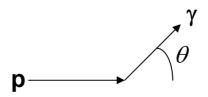
□ Assumptions:

- stationary target experiment
- projectile proton along z-axis

For each event

- calculate p_T (GeV/c) and θ (arcmin)
- bin in 2D histograms
 pixel size: Δp_T=10MeV/c, Δθ=2 arcmin



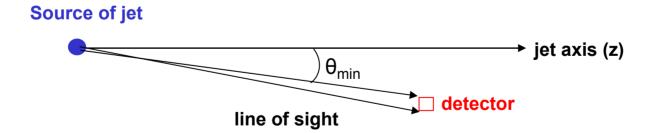




Pencil Proton Beam

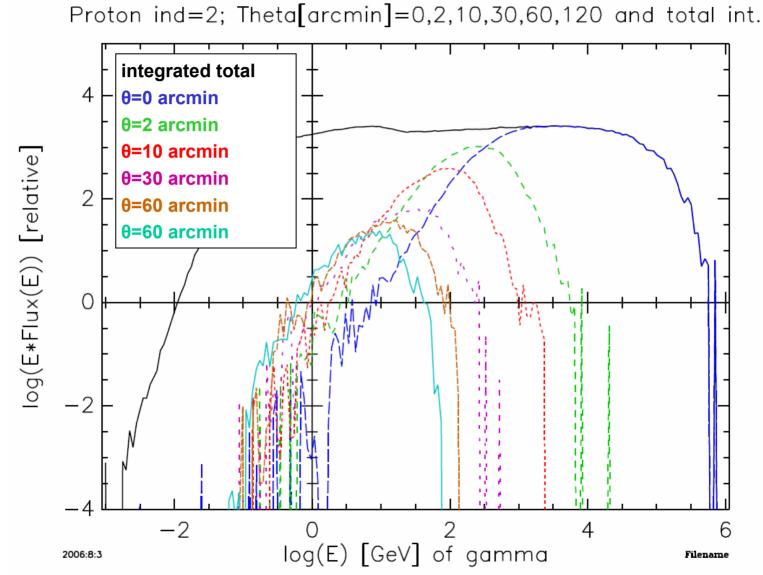
 \Box Jet axis offset an angle θ from the line of sight

- \Box How does the spectrum change with θ ?
- □ Difference in angular distribution between energy bands?





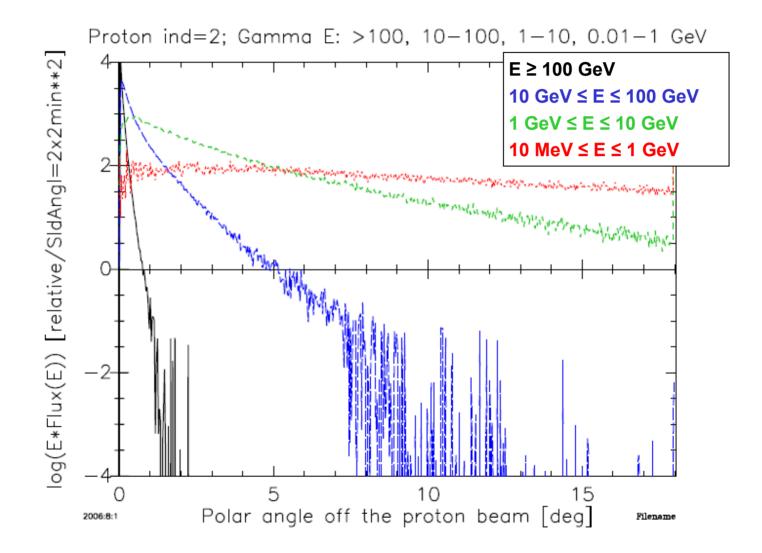
Pencil Proton Beam (cont.)



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Pencil Proton Beam (cont.)



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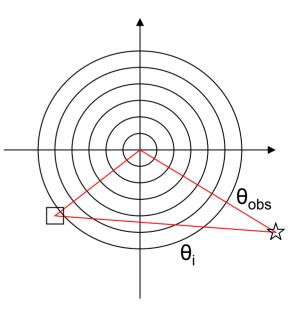


Gaussian Proton Beam

- Beam has a Gaussian (2D) profile
 - FWHM 2 deg $\rightarrow \sigma \approx$ 1 deg

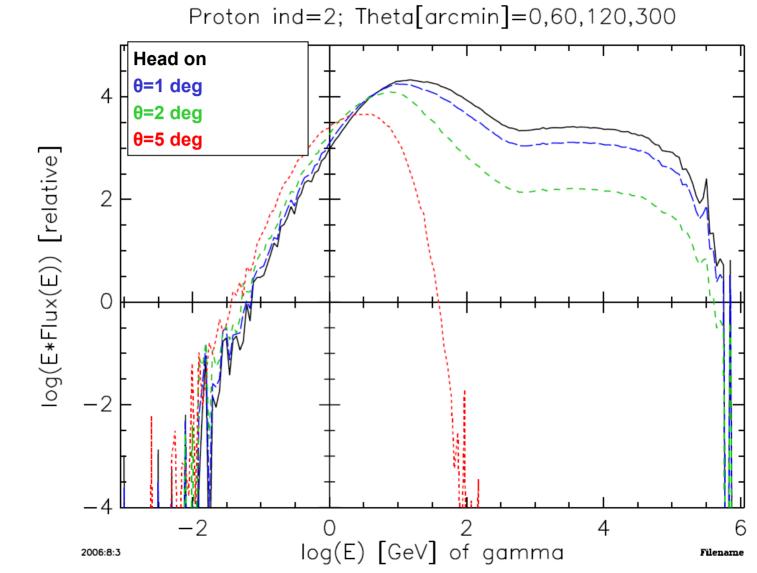
□ Convolve spectrum using beam profile

- bin the profile in 5 arcmin bins
- calculate relative angle to each bin
- weight with profile of the beam
- sum contributions from all bins

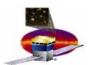




Gaussian Proton Beam (cont.)



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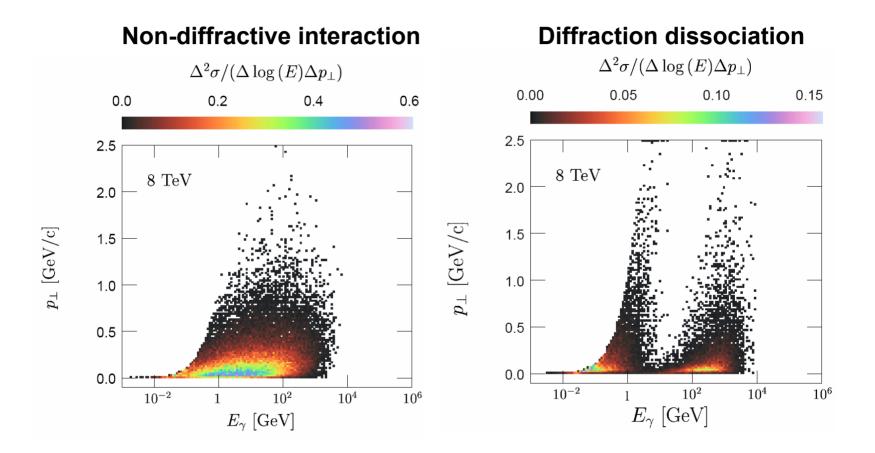


Parameterization of Angular Distribution

- □ Same idea as for total inclusive cross section
- □ Work with transverse momentum not angle
- □ Only gamma-rays worth while
- □ For each component
 - parameterize p_T distribution as functions of E_{γ} and p_T
 - first, for every Tp, fit p_T for each bin of $E_{\gamma} \rightarrow$ set of parameters
 - fit parameters as 2D functions of E_{γ} and Tp
 - work in phase space: $\rho = N/p_T$

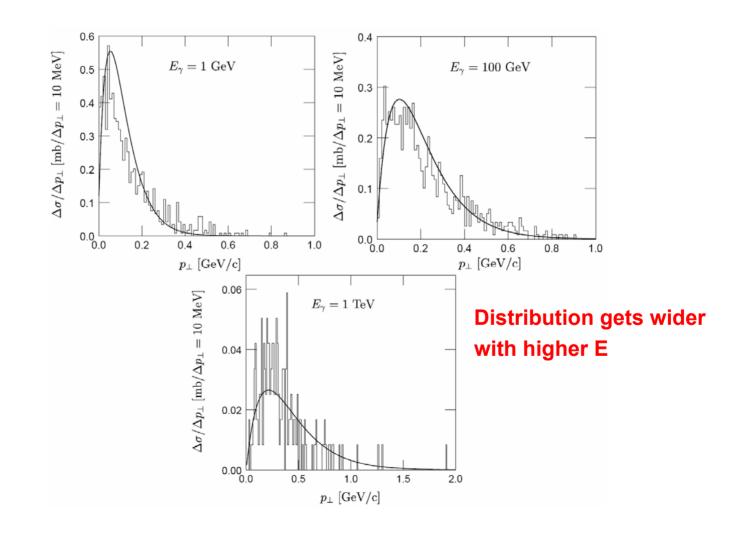


Gamma-ray pT Distributions



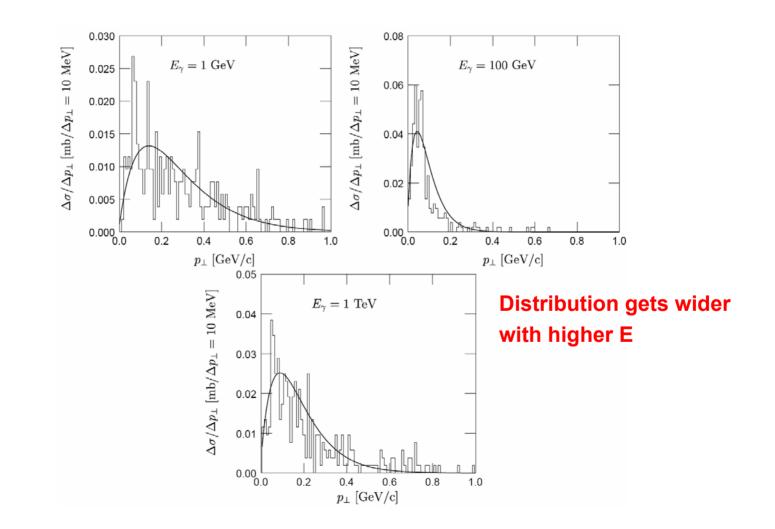


Fits: Non-diffractive Interaction





Fits: Diffraction Dissociation

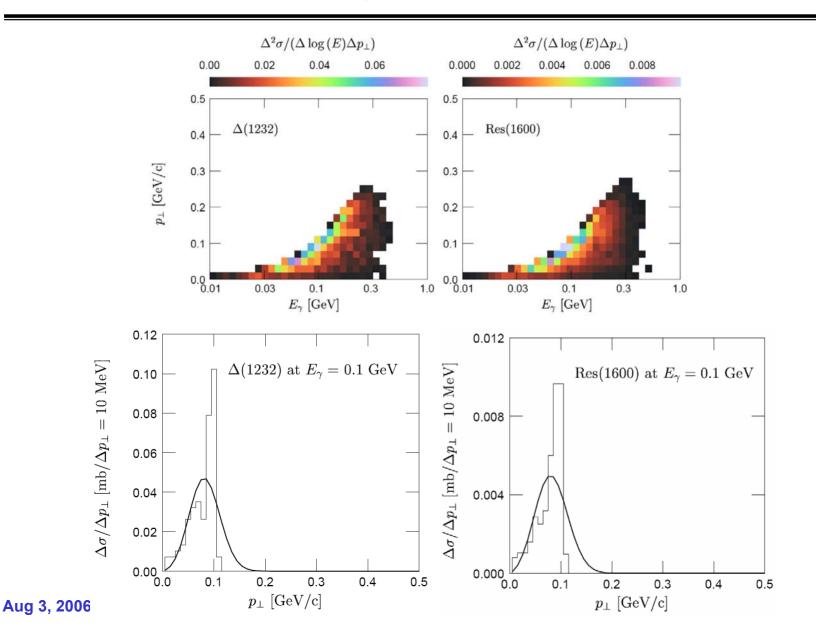


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Fits: Baryon Resonances



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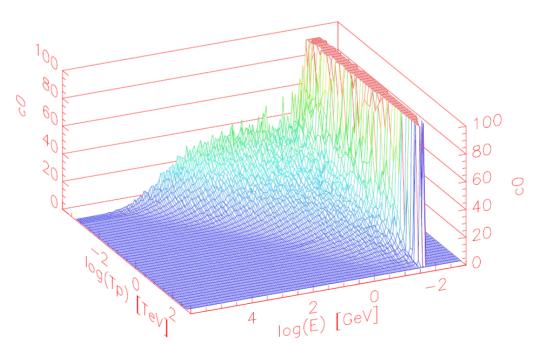


Final Step – Fit Parameters

- Parameters depend on both E and Tp
 - non-diff and diffraction: 2 parameters
 - baryon resonances: 3 parameters
- □ Find a way to parameterize these
- Artifacts from binning must be taken care of

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Final words

- □ Up-to-date p-p inelastic model important
- □ Parameterization of secondary particle spectra very successful
 - paper to be published in ApJ
- Gamma-rays from anisotropic relativistic outflows might be detectable even if jet is viewed off-axis

□ Parameterization of angular distribution for gamma-rays started

- looks promising
- paper is being prepared
- □ The two parameterizations gives us a great tool for modeling anisotropic relativistic outflows from SNRs, AGNs and GRBs