

# Flavor Physics in the LHC Era:

## The (Supersymmetric) Case for a SuperB Factory

Christopher Kolda  
University of Notre Dame

SuperB III Workshop  
14 June 2006

# Greetings from...

# SUSY06

newport beach, california  
june 12-17, 2006

**14TH INTERNATIONAL CONFERENCE ON  
SUPERSYMMETRY AND THE UNIFICATION  
OF FUNDAMENTAL INTERACTIONS**

**HTTP://SUSY06.PHYSICS.UCI.EDU  
EMAIL: SUSY06@UCI.EDU**



## **advisory committee**

Igor Anisimov (CERN), Jonathan Berger (Johns Hopkins), Vernon Berger (Wiscronin), Lars Bergstrom (Stockholm), Menelaos Cerna (Fermilab), Nirjan Das (Pennsylvania), Athanasios Dedes (Durham), Keith Denno (Arizona), Heide Dreier (Bonn), Jonathan Ellis (CERN), Gian Giudice (CERN), Rohini Godbole (IG Bangalore), Katsu Hagihara (KEK), Gordon Kane (Michigan), Dmitri Kazakov (Dubna), Jim E. Kim (Seoul), Young-Kee Kim (Chicago), Paul Langacker (Pennsylvania), Andrei Linde (Stanford), Joseph Lykken (Fermilab), Roberto Machado (Maryland), Hilarion Muncyama (UC Berkeley), Fran Nath (Northeastern), Ann Nelson (Washington), Michael Parikh (SLAC), Graham Ross (Oxford), Bernard Sadoulet (UC Berkeley), John Schwarz (Caltech), Edmund Willen (IAS Princeton), Takahiro Yanagida (Tokyo), Peter Zander (DESY)

## **program committee**

Katell Bala (Columbia State), Laura Bauda (Florida), Jeffrey Barger (UC Santa Barbara), Kiwoon Choi (KAIST), John Conway (UC Davis), Sacha Davidson (Lyon), Keith Denno (Arizona), Nicolaas Gaiotto (CERN), Concha Guadalupe-Garcia (Story Brook), Christophe Guehen (Sackler), JoAnna Hewell (SLAC), Junji Hasebe (Tokyo), Clifford Johnson (USC), Manoj Kaplinghat (UC Irvine), Alexander Kusenko (UCLA), Greg Landsberg (Brown), Dorey Marotta (Kansas), Konstantin Melnikov (Florida), Takao Mori (Tohoku), Mitsuaki Nagai (KEK), Yasuhiro Nomura (UC Berkeley), Maxim Pospelov (Camel), Avishai Rajaraman (UC Irvine), Yael Shadmi (Technion), Maria Spiropulu (CERN), Shigeru Yafuku (Chiba)

## **organizing committee**

**uc irvine** Myron Bandet, Steven Barwick, Jose Ruiz Cembranos, Jonathan L. Feng (Chair), Herbert Haber, Manoj Kaplinghat, William Mazon, Avishai Rajaraman, Hank Sobel, Virginia Trimble

sponsors:



University of California, Irvine



US Department of Energy



National Science Foundation

A common sentiment:

The Standard Model works **unreasonably** well!

Over past 20 years, (almost) all pieces verified:

- LEP checked weak gauge sector at 1–2 loops
- Tevatron has completed fermion spectrum
- Tevatron + HERA + LEP + ... have verified strong interactions/QCD
- B-factories have verified CKM picture of flavor-mixing and CP violation
- Neutrino masses found -- AS EXPECTED
- **Only EW symmetry breaking mechanism remains undiscovered.**

If Standard Model works so well, why do we KNOW it must be wrong?

The most serious problems arise from astrophysics. Standard Model does not produce:

- dark matter
- enough baryons
- dark energy

N.B. Not on my list: that neutrinos have mass.

“Post-Wilsonian” **minimal** Standard Model **MUST** have Majorana  $\nu$ -masses UNLESS we extend its symmetries.

(Also missing: strong CP problem.)

The most serious problems:

- dark matter
- enough baryons
- dark energy

- Astrophysics requires a new weakly-interacting (at most), massive particle with  $\tau > 10^{10}$  yr.

- For  $\Omega_{\text{DM}} \approx 0.3$  today, want  $M_{\text{DM}} \sim \sigma^{-1/2} \sim M_{\text{weak}}!$

➡ So physics of dark matter may be tied to EW scale!

## The most serious problems:

- dark matter
- enough baryons
- dark energy

- Sakharov conditions require B-violation, CP-violation, and baryon production out of equilibrium.

- All exist at EW phase transition, but require light Higgs,  $m_h < 40$  GeV.

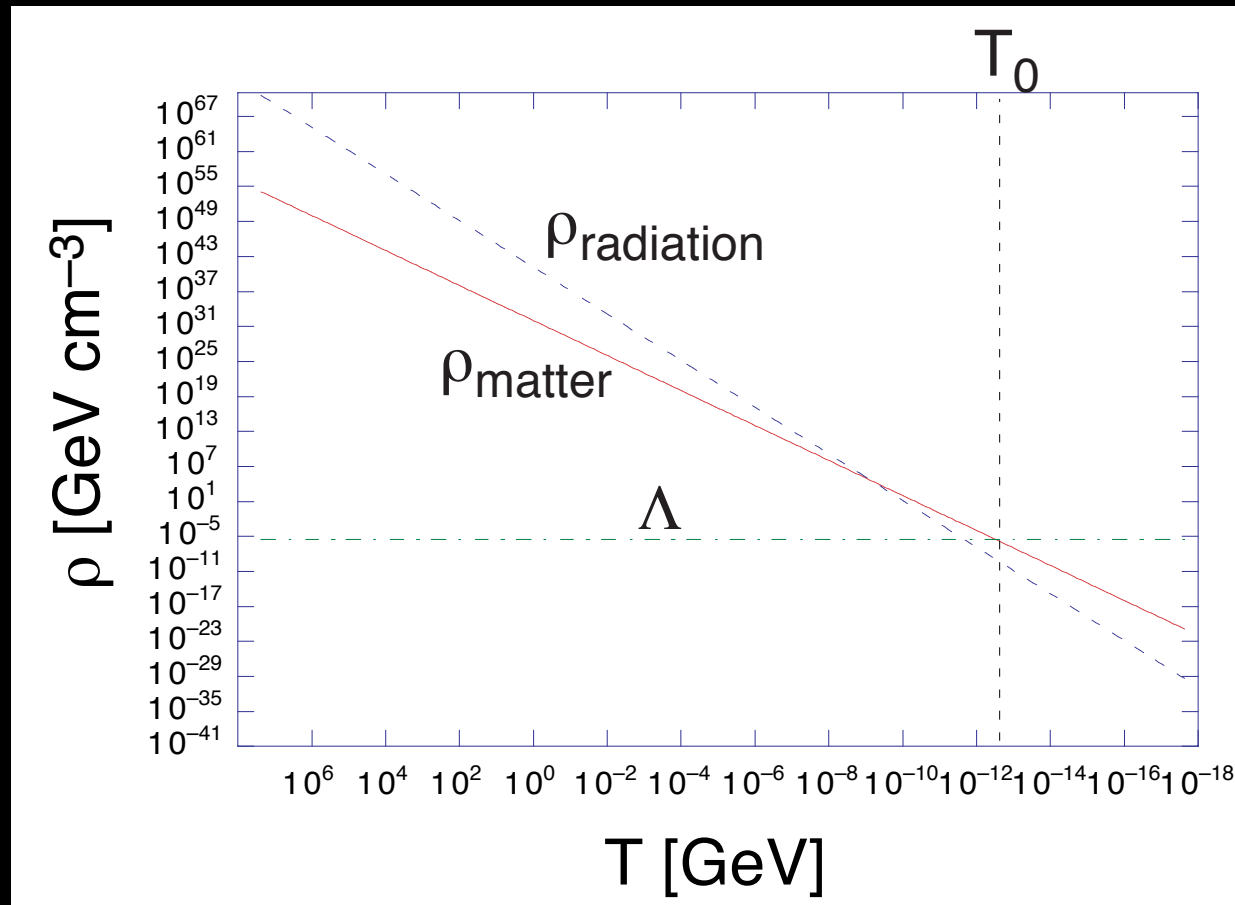
➔ The CKM explanation of CPV can't be whole story.

## The most serious problems:

- dark matter
  - enough baryons
  - dark energy
- 
- SM unable to make sensible prediction for cosmological constant, or any energy density  $\rho \sim (10^{-3} \text{ GeV})^4$ .
  - Question may or may not be tied to quantum gravity / M-theory.
  - Evidence of a triple-coincidence problem, which is solved by tying dark energy to weak scale.  
(Arkani-Hamed, Hall, Murayama, Kolda)

## The most serious problems:

- dark matter
- enough baryons
- dark energy



$$\rho_{\Lambda} \sim \frac{m_W^8}{m_{Pl}^4}$$

## “Major” theoretical problems:

- The SM doesn't include gravity (at the quantum level).
- The SM cannot be embedded into a more fundamental, high-energy theory thanks to quadratic divergences  $\Rightarrow$  **hierarchy problem**.
- Source of weak scale (What sets scale in Higgs potential?)

Proposed solutions to first include: string theory, quantum loop gravity, not much else.

$\Rightarrow$  Implies new physics at  $10^{17-18}$  GeV.

Proposed solutions to other two: SUSY, extra dimensions, technicolor, etc.

$\Rightarrow$  Implies new physics at  $10^{2-3}$  GeV.

## “Annoying” theoretical problems:

- Why 3 separate gauge groups?
- Why such strange quantum numbers?
- Why 3 generations?
- Why the large fermion mass hierarchies?
- Why is MNS matrix so unlike CKM?

Three classes of answers to all these problems:

### 1. **Grand unification conjecture:**

Physics at very high scales answers all these in a (nearly?) unique way. Goal of physics is to find this ultimate T.O.E.

## “Annoying” theoretical problems:

- Why 3 separate gauge groups?
- Why such strange quantum numbers?
- Why 3 generations?
- Why the large fermion mass hierarchies?
- Why is MNS matrix so unlike CKM?

Three classes of answers to all these problems:

1. Grand unification conjecture:

2. **Intermediate flavor dynamics hope:**

Perhaps physics of flavor explained by physics at intermediate scales,  $10^{6-16}$  GeV. If low enough, expect to see evidence in precision flavor studies.

## “Annoying” theoretical problems:

- Why 3 separate gauge groups?
- Why such strange quantum numbers?
- Why 3 generations?
- Why the large fermion mass hierarchies?
- Why is MNS matrix so unlike CKM?

Three classes of answers to all these problems:

1. Grand unification conjecture:
2. Intermediate flavor dynamics hope:
3. **Landscape surrender:**

$\exists$  billions of possible universes. Cosmology chose this one by chance. Gauge groups, q-numbers, masses, etc are pure “anarchy”.

Most problems have a preferred scale for solution:

- Solve hierarchy problem/dark matter problems at  $10^{2-3}$  GeV.
- Explain small neutrino masses at  $10^{14}$  GeV.
- Solve unification/quantum gravity problems at  $10^{16-18}$  GeV.

Of these, hierarchy/dark matter within our reach.

Most problems have a preferred scale for solution:

- Solve hierarchy problem/dark matter problems at  $10^{2-3}$  GeV.
- Explain small neutrino masses at  $10^{14}$  GeV.
- Solve unification/quantum gravity problems at  $10^{16-18}$  GeV.

Of these, hierarchy/dark matter within our reach.

But there is no preferred scale for new flavor physics, except that:

$$M_{\text{flavor}} > 10^6 \text{ GeV (or } 10^7 \text{ GeV if CPV)}$$

Worse, flavor dynamics anywhere near  $10^6$  GeV would probably corrupt hierarchy solution.

E.g. In mSUGRA, flavor dynamics below  $M_{\text{GUT}}$  can be ruled out in almost all cases.

So arguments for SuperB factories based on

- Deciphering the 3 generation riddle
- Determining dynamics leading to CKM structure
- Discovering ultimate source of CPV

will go unheeded by many physicists.

So arguments for SuperB factories based on

- Deciphering the 3 generation riddle
- Determining dynamics leading to CKM structure
- Discovering ultimate source of CPV

will go unheeded by many physicists.

Theorists' view: We are used to “decoupling” flavor physics from new physics that solves hierarchy problem. We generally don't expect non-trivial flavor dynamics anywhere near weak scale.

Flavor measurements are useful constraints on new physics, but are unlikely to teach us much about new physics once it is found. (E.g. LHC Theory Initiative)

So arguments for SuperB factories based on

- Deciphering the 3 generation riddle
- Determining dynamics leading to CKM structure
- Discovering ultimate source of CPV

will go unheeded by many physicists.

Theorists' view: We are used to “decoupling” flavor physics from new physics that solves hierarchy problem. We generally don't expect non-trivial flavor dynamics anywhere near weak scale.

Flavor measurements are useful constraints on new physics, but are unlikely to teach us much about new physics once it is found. (E.g. LHC Theory Initiative)

WRONG

Europe seems to have the right idea:



# Flavour in the era of the LHC

a Workshop on the interplay of flavour and collider physics

First meeting:

## CERN, November 7-10 2005

<http://mlm.home.cern.ch/mlm/FlavLHC.html>



- BSM signatures in B/K/D physics, and their complementarity with the high-pT LHC discovery potential
- Flavour phenomena in the decays of SUSY particles
- Squark/slepton spectroscopy and family structure
- Flavour aspects of non-SUSY BSM physics
- Flavour physics in the lepton sector
- $g-2$  and EDMs as BSM probes
- Flavour experiments for the next decade

**Local Organizing Committee**

A. Ceccucci (CERN, Geneva)  
D. Denegri (Saclay, Gif sur Yvette)  
J. Ellis (CERN, Geneva)  
R. Fleischer (CERN, Geneva)  
G. Giudice (CERN, Geneva)  
T. Hurth (CERN, Geneva)  
M. Mangano (CERN, Geneva)  
T. Nakada (EPFL, Lausanne)  
G. Polesello (INFN, Pavia)  
M. Smizanska (Lancaster Univ)

**International Advisory Committee**

A. Ali (DESY, Hamburg)  
A. Buras (TUM, Munich)  
P. Cooper (FNAL, Batavia)  
P. Franzini (LNF, Frascati)  
M. Giorgi (Universita' di Pisa)  
K. Hagiwara (KEK, Tsukuba)  
S. Jin (IHEP, Beijing)  
L. Littenberg (BNL, Brookhaven)  
G. Martinelli (La Sapienza, Roma)  
A. Masiero (Universita' di Padova)  
H. Murayama (UC and LBNL, Berkeley)  
A. Sanda (Nagoya University)  
Y. Semertzidis (BNL, Brookhaven)  
S. Stone (Syracuse University)  
M. Yamauchi (KEK, Tsukuba)  
P. Zerwas (DESY, Hamburg)

Most models of Beyond-the-SM physics have a

## “FLAVOR PROBLEM”

In the Standard Model, large FCNCs are prevented by combo of CKM unitarity and small mixings between heavy and light quarks.

In most BTSM proposals, CKM mechanism fails by one loop. It fails because:

- Difficult to maintain CKM as only source of FCNCs if more states carry flavor (e.g. SUSY)
- If 3<sup>rd</sup> generation is “special”, it feeds back into all FCNCs (e.g. topcolor)
- If new gauge interactions differentiate flavors, then they directly mediate FCNCs.

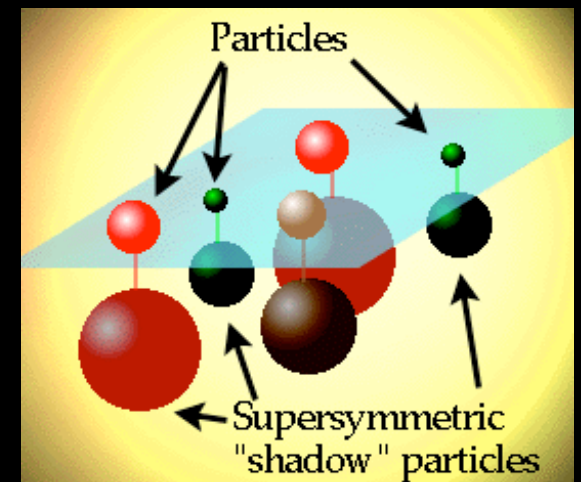
How does this work in Supersymmetry?

# How does this work in Supersymmetry?

First, **what is SUSY?**

SUSY is a predicted symmetry of nature:

- Only possible symmetry between fermions and bosons.
- Solves gauge hierarchy problem by canceling quadratic divergences, stabilizing weak scale.
- Consistent with GUT models, including coupling unification.
- Breaks EW symmetry dynamically thanks to large  $y_{\text{top}}$ .
- Predicts a superpartner for each SM particle, with spin different by  $1/2$ .
- Requires 2 Higgs doublets, with ratio of vevs =  $\tan\beta$  ( $\approx 1$  to  $60$ )



In SUSY:

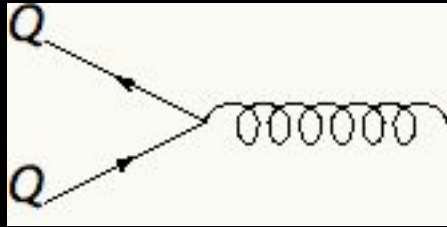
- Quark masses from Yukawa couplings/EWSB.
- Squark masses from Yukawas/EWSB and SUSY-breaking.
- Two sources uncorrelated, so squark mixing uncorrelated to quark mixing:

$$Q : \quad Q = V_{\text{CKM}} Q^0 \quad = V Q^0$$

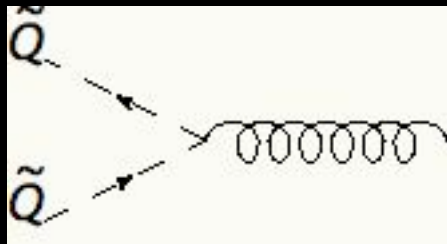
$$\tilde{Q} : \quad \tilde{Q} = V_{\text{CKM}} V_{\text{SUSY}} \tilde{Q}^0 \quad = \tilde{V} \tilde{Q}^0$$

For general SUSY models, squark mixing need not be anywhere near same as quark mixing.

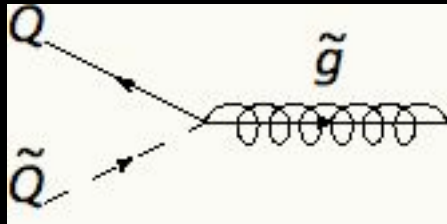
In SUSY:



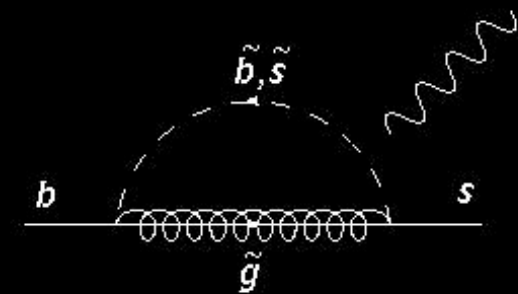
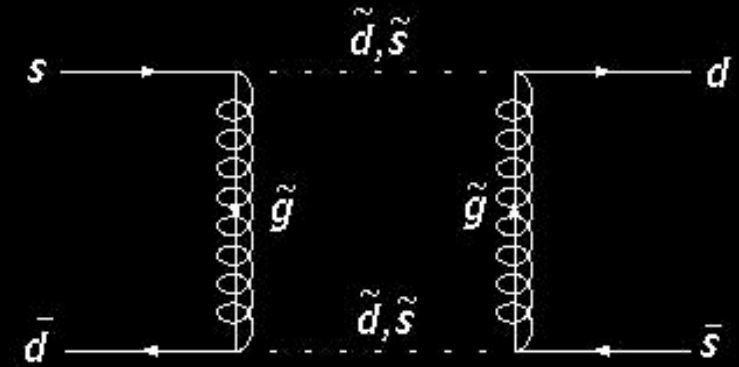
$$\sim V^\dagger V = 1$$



$$\sim \tilde{V}^\dagger \tilde{V} = 1$$



$$\sim V^\dagger \tilde{V} \neq 1$$

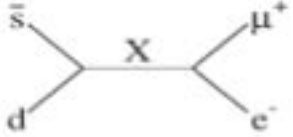
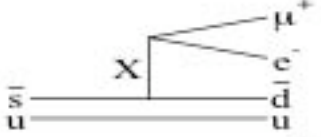
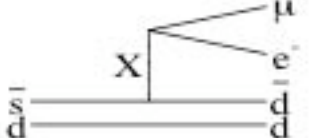
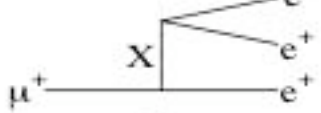
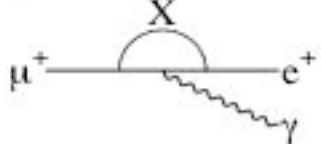
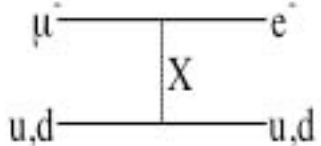


If two matrices not the same, large FCNCs result -- this is ruled out!

Note: flavor-changing always in loops.

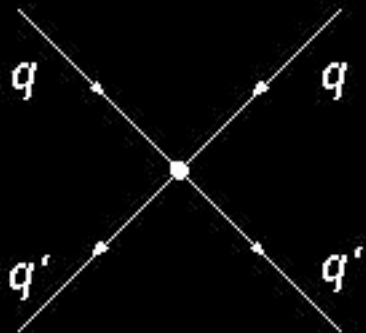
The “flavor problem” is even more general -- scale of physics required to solve hierarchy problem already ruled out by precision flavor studies:

$\Delta F=1$   
processes:

	$B(K_L \rightarrow \mu e) < 4.7 \times 10^{-12}$	$M_X > 150 \text{ TeV}/c^2$
	$B(K^+ \rightarrow \pi^+ \mu^+ e^-) < 4 \times 10^{-11}$	$M_X > 31 \text{ TeV}/c^2$
	$B(K_L \rightarrow \pi^0 \mu^+ e^-) < 3.2 \times 10^{-10}$	$M_X > 37 \text{ TeV}/c^2$
	$B(\mu^+ \rightarrow eee) < 1 \times 10^{-12}$	$M_X > 86 \text{ TeV}/c^2$
	$B(\mu^+ \rightarrow e^+ \gamma) < 1.2 \times 10^{-11}$	$M_X > 21 \text{ TeV}/c^2$
	Normalized Rate $< 6.1 \times 10^{-13}$	$M_X > 365 \text{ TeV}/c^2$

Even worse,  $\Delta F=2$  processes...

Given some new operator:



$$\frac{1}{\Lambda^2} (q\Gamma q')(q\Gamma q')$$

If  $\{q, q'\} =$

- $\{d, s\}$ ,  $\Delta M_K$  gives  $\Lambda < 1500 \text{ TeV}$ ;
- $\{d, b\}$ ,  $\Delta M_B$  gives  $\Lambda < 500 \text{ TeV}$ ;
- $\{s, b\}$ ,  $\Delta M_{Bs}$  gives  $\Lambda < 100 \text{ TeV}$ .

These are extremely strong constraints on the scale of new physics!! And CPV pushes them up another factor of 10!!!

Two questions raised by meson mixing data:

1. Is there any point to further high precision studies? Are meson-antimeson constraints so powerful that they rule out new physics in rare decays?

2. Is there any point in using flavor physics to probe (rather than just constrain) new physics?  
(Direct searches for 1500 TeV particles a long way off...)

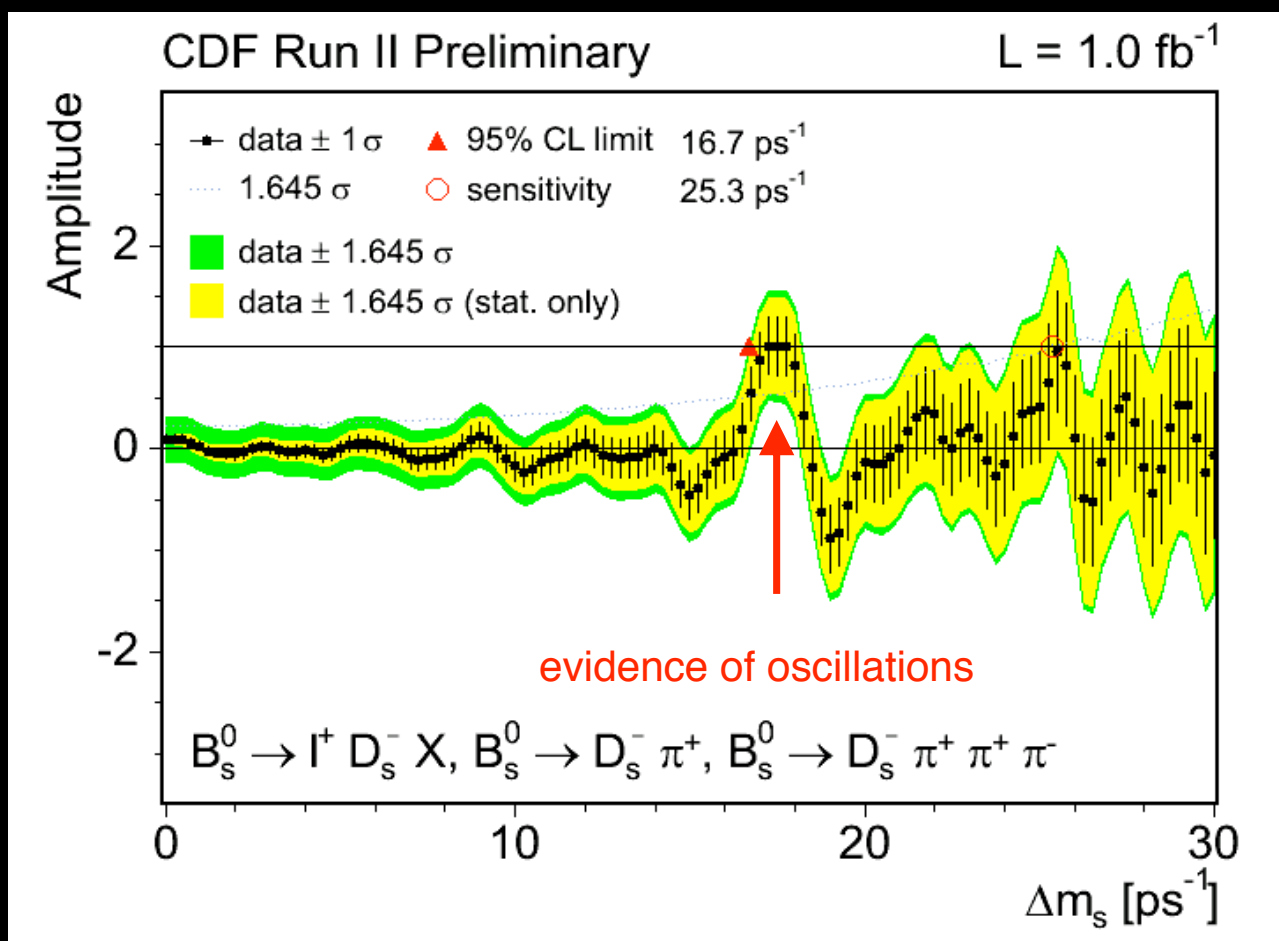
Two questions raised by meson mixing data:

1. Is there any point to further high precision studies? Are meson-antimeson constraints so powerful that they rule out new physics in rare decays?

Yes, there is a point. The case is harder to make for kaons, but easily made for B's. Scales probed by mixing and by rare decays very similar for B's: 10's to 100's of TeV.

Success of CKM picture at BaBar/Belle means new physics effects probably not huge in 1-3 sector, but what about 2-3 sector?

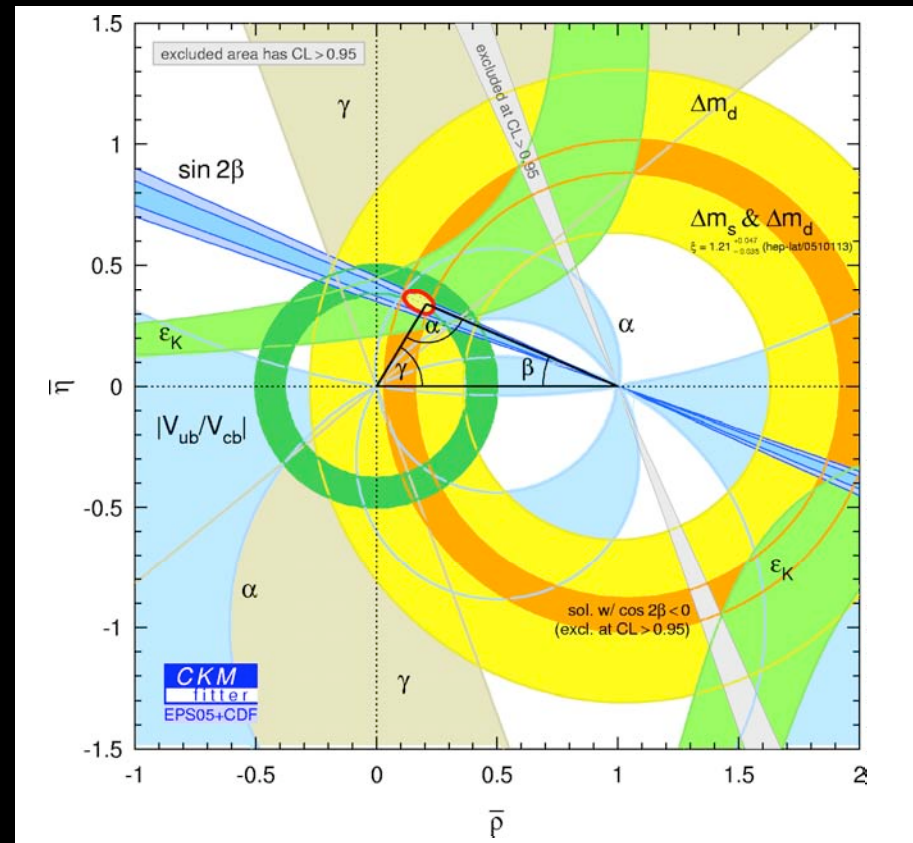
# In the news...



$$\Delta m_s = 17.33^{+0.42}_{-0.21} \pm 0.07 \text{ ps}^{-1} \text{ (CDF)}$$

$$\Delta m_s = \begin{cases} 21.7^{+5.9}_{-4.2} \text{ ps}^{-1} & \text{(CKMFitter)} \\ 21.5 \pm 2.6 \text{ ps}^{-1} & \text{(UTFit)} \end{cases}$$

This is great news for the Standard Model. The CKM picture is in impressive agreement with all the data!



The (somewhat) bad news for new physics:  
New physics could be hiding in **2-3** sector, but don't expect huge signals. There is no one "smoking gun" measurement to be done.

What lesson have the theorists learned from success of SM in flavor sector?

Either:

- New physics completely commutes with SM flavor structure (like universal  $Z'$ ).
- Or new physics is **Minimally Flavor Violating** (MFV).

What lesson have the theorists learned from success of SM in flavor sector?

Either:

- New physics completely commutes with SM flavor structure (like universal  $Z'$ ).
- Or new physics is **Minimally Flavor Violating** (MFV).

**Minimal Flavor Violation** is ansatz that only source of flavor violation in new physics is usual Yukawas, and that this is dominated by  $y_{\text{top}}$ .

All quark mixing encoded in CKM matrix!

N.B. This is not a model, it is a constraint placed on models. But results protected by approximate flavor symmetries of the SM, i.e.,  $U(3)^3$ .

MFV implies that:

- No FCNC operators not already in SM will appear.\*
- New contributions to FCNC operators suppressed by usual CKM factors.
- Existing operators will get  $O(1)$  corrections at best. Usually even smaller.\*
- No new source of CPV, so no new CPV asymmetries.\*
- Unitarity triangle expected to close approximately.

MFV implies that:

- No FCNC operators not already in SM will appear.\*
- New contributions to FCNC operators suppressed by usual CKM factors.
- Existing operators will get  $O(1)$  corrections at best. Usually even smaller.\*
- No new source of CPV, so no new CPV asymmetries.\*
- Unitarity triangle expected to close approximately.

In MFV models,  $\Delta M_K$  probes

$$\Lambda_{NP} = 1500 \text{ TeV} \times V_{td}V_{ts} = 500 \text{ GeV!}$$

Meanwhile,  $\Delta M_B$  probes

$$\Lambda_{NP} = 500 \text{ TeV} \times V_{td}V_{tb} = 4 \text{ TeV!}$$

(similarly for  $\Delta M_{Bs}$ )

If effects are 1-loop, then scales drop to 40 GeV and 350 GeV.

In MFV,  $\Delta F=1$  four-fermion operators are not so suppressed:

$$\begin{aligned}\Delta F = 2 : \quad \Lambda_{\text{MFV}} &= V_{ti} V_{tj} \times \Lambda_{\text{naive}} \\ \Delta F = 1 : \quad \Lambda_{\text{MFV}} &= \sqrt{V_{ti} V_{tj}} \times \Lambda_{\text{naive}}\end{aligned}$$

For the specific case of B-meson:

$$\begin{aligned}B^0 - \bar{B}^0 : \quad \Lambda_{\text{MFV}} &= V_{td} V_{tb} \times \Lambda_{\text{naive}} \simeq \frac{1}{100} \Lambda_{\text{naive}} \\ B^0 \rightarrow \ell^+ \ell^- : \quad \Lambda_{\text{MFV}} &= \sqrt{V_{td} V_{tb}} \times \Lambda_{\text{naive}} \simeq \frac{1}{10} \Lambda_{\text{naive}}\end{aligned}$$

So the constraints from  $\Delta F=2$  operators become weaker in relation to probative power of  $\Delta F=1$  rare decays!

**MFV works twice:** to reduce tight FCNC constraints on all forms of new physics and to increase phase space for new effects in rare decays.

There is one caveat (\*):

“Minimal” MFV assumes only  $y_{\text{top}}$  is large. Implicitly assumes single Higgs structure.

In models with 2+ Higgs,  $y_{\text{bot}}$  can be sizable  $\Rightarrow$  new operators can appear:

- New effects still  $\propto$  CKM elements.
- Largest effect is Higgs-mediated rare decays,  $B \rightarrow l^+ l^-$ ,  $B \rightarrow K l^+ l^-$ , etc.
- SM prediction too small for 50  $\text{ab}^{-1}$  SuperB, but NP can be **orders** greater.
- In SUSY, these are the well-known  $\tan^6 \beta$  effects that occur even in mSUGRA models.

MFV is commonly assumed in SUSY model building. In mSUGRA, gauge-mediation, anomaly-mediation, or any other model with squark degeneracy.\*

LHC will find the states, but will it really be MFV? Even if approximate degeneracy found, can there be non-CKM sources of flavor violation?

That's a vitally important question, and can't be answered at the LHC. One of the goals of SuperB must be to test MFV!

[\* SUSY models w/o degeneracy either require decoupling or alignment. The latter implies new, large contributions to  $D^0$ -mixing, another target for SuperB.]

One clean MFV prediction:

$$\frac{\text{Br}(B_d \rightarrow (X_s)\ell\ell)}{\text{Br}(B_s \rightarrow (X_s)\ell\ell)} = \left(\frac{V_{td}}{V_{ts}}\right)^2 \approx \frac{1}{25}$$

This is bad(?) news for a SuperB factory:

CDF finds  $\text{Br}(B_s \rightarrow \mu\mu) < 1 \times 10^{-7}$ , which means  $\text{Br}(B \rightarrow \mu\mu) < 4 \times 10^{-9}$  in MFV.

With  $50 \text{ ab}^{-1}$ , SuperB can get to  $\text{Br}(B \rightarrow \mu\mu) \approx 7 \times 10^{-9}$ , so not enough. (Why not  $10^{-10}$ ?)

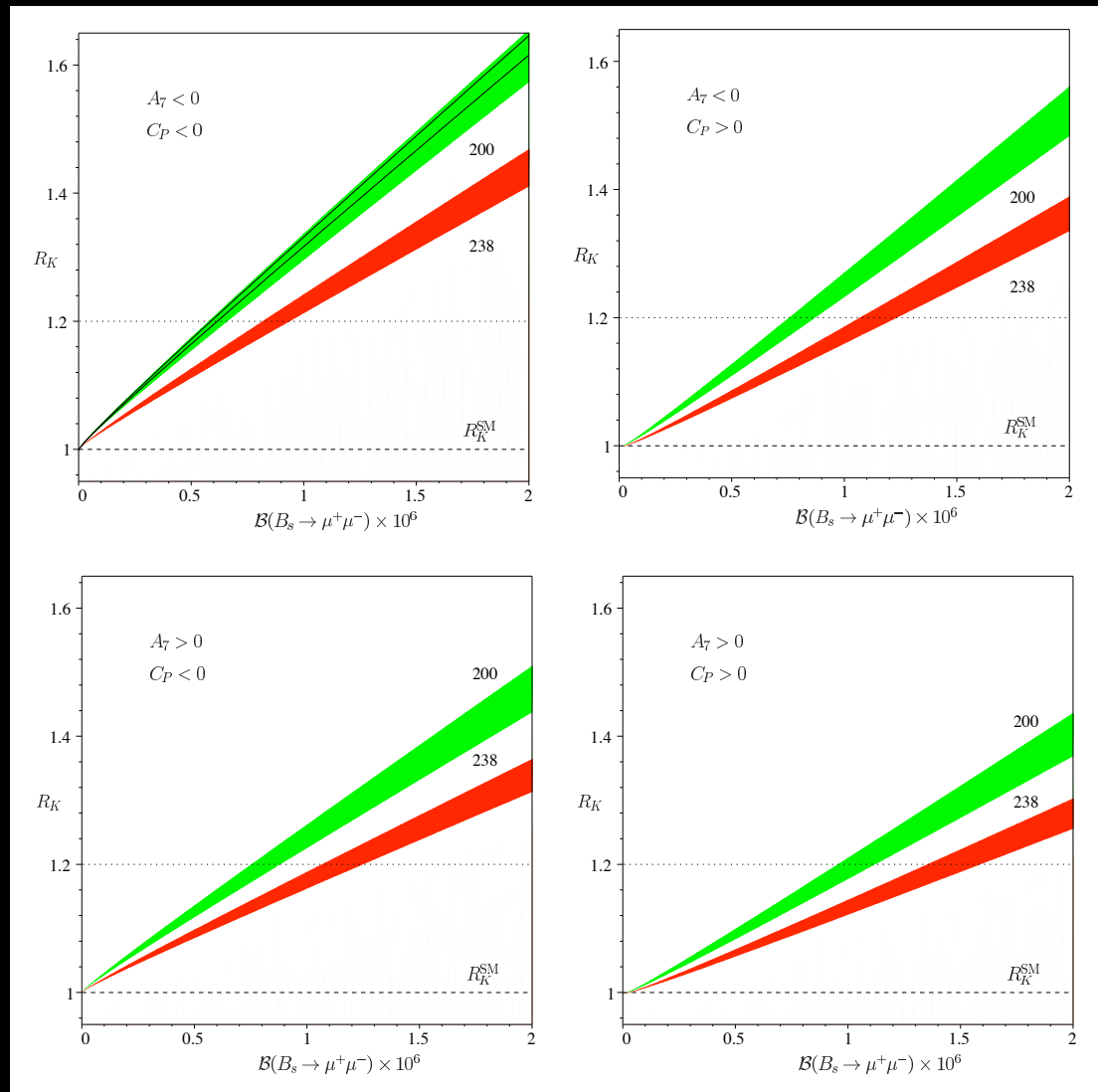
(Or need to get  $\text{Br}(B \rightarrow \tau\tau)$  down to  $10^{-6}$ . Unlikely!)

But SuperB can compete in 3-body final state:

$$\frac{\text{Br}(B \rightarrow X_s \mu\mu)}{\text{Br}(B \rightarrow X_s ee)} \sim 1 + \frac{1}{16\pi^2} \times (\text{factors}) \times \frac{m_\mu^2}{m_W^2}$$

Can be measured to 4% with  $10 \text{ ab}^{-1}$  or 2% with  $50 \text{ ab}^{-1}$ .

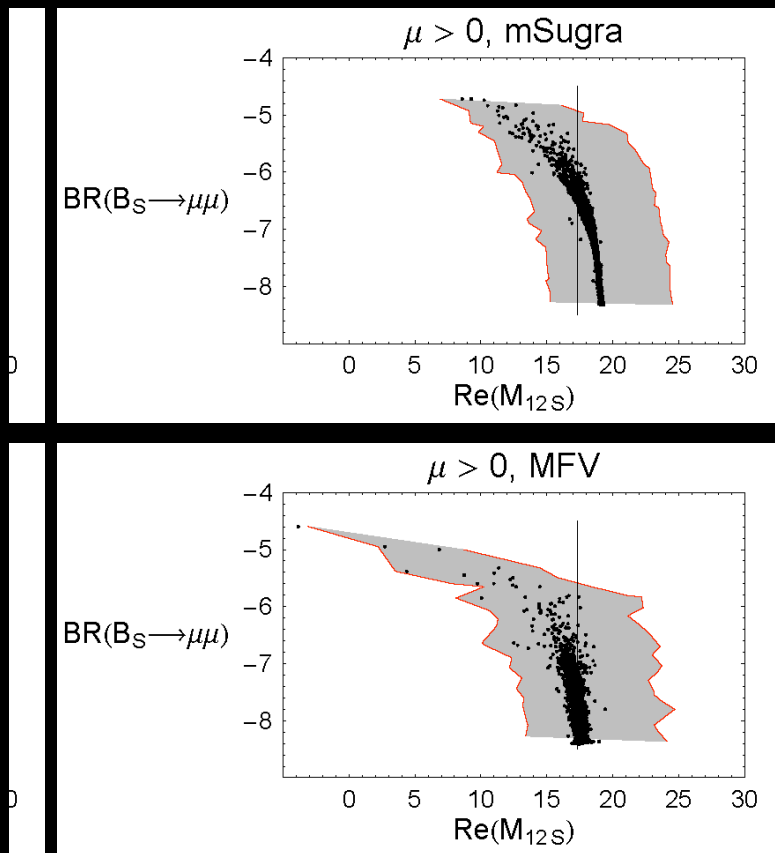
Current CDF bound on  $B_s \rightarrow \mu\mu$  of  $10^{-7}$  implies needed precision of at least 1-3%. So need that 50  $\text{ab}^{-1}$ !



Hiller, Krüger

In MFV, a correlation exists between  $B \rightarrow \mu\mu$  and  $B_s$ -mixing:  $\Delta M_s$  should be  $<$  SM value, more so the larger  $\text{Br}(B_s \rightarrow \mu\mu)$  is.

Recent CDF result could imply a new constraint:



But model dependence too big at present to constrain MFV scenario, except that  $\text{Br} < 10^{-6}$  predicted (OK!)

Lunghi, Porod, Vives

Another (unfortunate?) prediction of MFV:

Since there are no new sources of CPV, CP asymmetries will take their SM values, up to small corrections. Any zero/tiny asymmetry in SM will remain so.

So some new physics could appear in  $B \rightarrow \Phi K_S$ , but it won't be large. With  $50 \text{ ab}^{-1}$ , SuperB can get to 4% in  $S(B \rightarrow \Phi K_S)$  and compare to  $B \rightarrow \Psi K_S$ , a nice test of MFV, even if result is null.

Another way to test whether any observed sparticle degeneracy is really MFV or not:

## SQUARK MASS SUM RULES

Almost all SUSY models start from a degenerate spectrum:

$$\begin{aligned} \tilde{m}_{u,d_L}^2 &= \tilde{m}_{c,s_L}^2 = \tilde{m}_{t,b_L}^2 \\ \tilde{m}_{u_R}^2 &= \tilde{m}_{c_R}^2 = \tilde{m}_{t_R}^2 \\ \tilde{m}_{d_R}^2 &= \tilde{m}_{s_R}^2 = \tilde{m}_{b_R}^2 \end{aligned} \quad \text{at some scale}$$

Perfect degeneracy would mean no new FCNC's:

“super-GIM mechanism”

But loop corrections spoil degeneracy:

Can't both be  
diagonal or no  
CKM mixing

$$\begin{aligned} \frac{d}{d \log Q} (m_{\tilde{Q}}^2)_{ij} &= \frac{1}{16\pi^2} \left\{ -\frac{2}{15} g_1^2 M_1^2 - 6g_2^2 M_2^2 - \frac{32}{3} g_3^2 M_3^2 \right. \\ &\quad + 2 \left( Y_u^\dagger m_{\tilde{Q}}^2 Y_u + Y_d m_{\tilde{Q}}^2 Y_d^\dagger + Y_u^\dagger m_{\tilde{U}}^2 Y_u + Y_d m_{\tilde{D}}^2 Y_d^\dagger \right. \\ &\quad \left. \left. + Y_u^\dagger Y_u m_{H_u}^2 + Y_d Y_d^\dagger m_{H_d}^2 + A_u^\dagger A_u + A_d A_d^\dagger \right) \right\}_{ij} \end{aligned}$$

This generates squark mixing (i.e. partner of down quark is not down squark, but admixture).

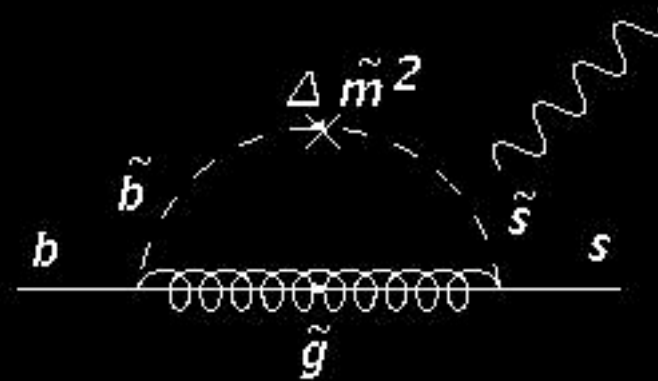
The resulting d-squark mass matrix is 6x6 (incl. LR mixing):

$$\begin{array}{cccccc}
 d_L & d_R & s_L & s_R & b_L & b_R \\
 \left( \begin{array}{cccccc}
 m_{\tilde{d}_L}^2 & m_d(A_d - \mu \tan \beta) & (\Delta_{12}^d)_{LL} & (\Delta_{12}^d)_{LR} & (\Delta_{13}^d)_{LL} & (\Delta_{13}^d)_{LR} \\
 & m_{\tilde{d}_R}^2 & (\Delta_{12}^d)_{RL} & (\Delta_{12}^d)_{RR} & (\Delta_{13}^d)_{RL} & (\Delta_{13}^d)_{RR} \\
 & & m_{\tilde{s}_L}^2 & m_s(A_s - \mu \tan \beta) & (\Delta_{23}^d)_{LL} & (\Delta_{23}^d)_{LR} \\
 & & & m_{\tilde{s}_R}^2 & (\Delta_{23}^d)_{RL} & (\Delta_{23}^d)_{RR} \\
 & [\tilde{m}_D^2]_{ij} & & & m_{\tilde{b}_L}^2 & m_b(A_b - \mu \tan \beta) \\
 & & & & & m_{\tilde{b}_R}^2
 \end{array} \right)
 \end{array}$$

LHC measures eigenvalues of this matrix. How do we learn the mixing angles?

By measuring rare FCNC's. And since this is MFV, look for signal in 3rd generation: SuperB!

Example:  $b \rightarrow s \gamma$



Rate is  $\propto [\Delta \tilde{m}_d^2]_{23}$ . Measure rate, measure the squark mixing parameters.

Then compare to MFV predictions. This can be done in a model-independent way using sum rules that connect squark mass matrix eigenvalues to their mixing angles.

Dudley & Kolda

At small  $\tan\beta$ , sum rules are simple and require few inputs from LHC:

$$(\Delta_{LL}^u)_{ij} = (\Delta_{RR}^u)_{ij} = (\Delta_{RR}^d)_{ij} = 0$$

$$\begin{aligned} (\Delta_{LL}^d)_{ij} &= V_{3i}^* V_{3j} [\tilde{m}_{b,1}^2 + \tilde{m}_{b,2}^2 - \tilde{m}_{d,L}^2 - \tilde{m}_{d,R}^2] \\ &= \frac{1}{3} V_{3i}^* V_{3j} [\tilde{m}_{t,1}^2 + \tilde{m}_{t,2}^2 - 2m_t^2 - \tilde{m}_{u,L}^2 + \tilde{m}_{u,R}^2] \end{aligned}$$

At moderate to high  $\tan\beta$ , sum rules are slightly more complicated and require more LHC input:

$$(\Delta_{RR}^u)_{ij} = (\Delta_{RR}^d)_{ij} = 0$$

$$\begin{aligned} (\Delta_{LL}^d)_{ij} &= \frac{V_{3i}^* V_{3j}}{8} [3 (\tilde{m}_{t,1}^2 + \tilde{m}_{t,2}^2 - \tilde{m}_{u_L}^2 - \tilde{m}_{u_R}^2 - 2m_t^2) \\ &\quad - \tilde{m}_{b,1}^2 - \tilde{m}_{b,2}^2 + \tilde{m}_{d_L}^2 + \tilde{m}_{d_R}^2] . \end{aligned}$$

The sum rules don't just test one particular SUSY scenario, but all SUSY scenarios with MFV.

Deviations from sum rules would be strong evidence for new sources of flavor violation not encoded in the CKM matrix.

I'm skipping much...

- $b \rightarrow s\gamma$  already puts tight constraints on SUSY parameter space, especially on  $H^\pm$  mass.
- $B \rightarrow TV$  places similar constraints on  $H^\pm$ .
- $S(B \rightarrow \Phi K_s)$  constrains many non-MFV SUSY models.

...not to mention important constraints from rare K decays,  $(g-2)_\mu$ , EDM measurements,  $\mu$ -e conversion, etc, not relevant to SuperB but still flavor physics.

But there's more: A SuperB factory is really a Super-Flavor factory. Will produce around  $10^{10}$   $\tau$ -pairs!

Lots of interesting New Physics in  $\tau$ -sector, thanks to neutrino mixing results.

But there's more: A SuperB factory is really a Super-Flavor factory. Will produce around  $10^{10}$   $\tau$ -pairs!

Lots of interesting New Physics in  $\tau$ -sector, thanks to neutrino mixing results.

Reminder:

- In neutrino sector, large mixings occur between  $\nu_\mu$  and  $\nu_\tau$ , and between  $\nu_\mu$  and  $\nu_e$ .
- This  $\nu$ LFV translates into  $c$ LFV in SM, but with amplitudes  $\propto (m_\nu/m_W)^2$ . Will never be seen!

- $\Delta m_{23}^2 = 2,4(1^{+0.21}_{-0.15}) \times 10^{-3} \text{ eV}^2$ ,  $\sin^2 \theta_{\mu 3} = 0.44(1^{+0.41}_{-0.22})$
  - $\Delta m_{12}^2 = 7.92(1 \pm 0.09) \times 10^{-5} \text{ eV}^2$ ,  $\sin^2 \theta_{e 2} = 0.314(1^{+0.18}_{-0.15})$
  - $\sin^2 \theta_{e 3} = 0.9^{+2.3}_{-0.9} \times 10^{-2}$
- (Fogli et al)

Large  $\nu$ -mixing



but no cLFV seen!

	2000PDG	current	future
$\tau \rightarrow \mu(e)\gamma$	$< 1.1(2.7) \times 10^{-6}$	$< 6.8(11) \times 10^{-8}$	$\sim 10^{-(8-9)}$
$\tau \rightarrow \mu(e)\eta$	$< 9.6(8.2) \times 10^{-6}$	$< 1.5(2.3) \times 10^{-7}$	$10^{-(9-10)}$
$\tau \rightarrow lll$	$< \sim 10^{-6}$	$< \sim 10^{-7}$	$< \sim 10^{-(9-10)}$
$\mu \rightarrow e\gamma$	$< 1.2 \times 10^{-11}$		$\sim 10^{-(13-14)}$ (MEG)
$\mu \rightarrow 3e$	$< 1.0 \times 10^{-12}$		$\sim 10^{-14}$ (?)
$\mu - e: \text{Ti}$	$< 4.3 \times 10^{-12}$		$\sim 10^{-18}$ (PRISM)

But there's more: A SuperB factory is really a Super-Flavor factory. Will produce around  $10^{10}$   $\tau$ -pairs!

Lots of interesting New Physics in  $\tau$ -sector, thanks to neutrino mixing results.

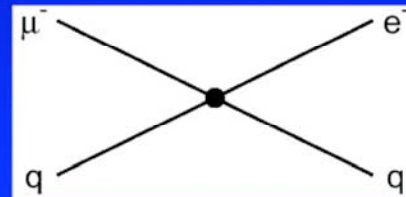
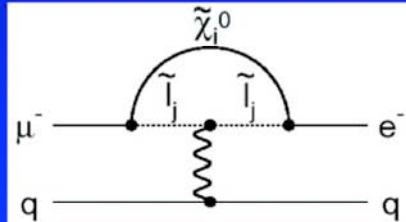
Reminder:

- In neutrino sector, large mixings occur between  $\nu_\mu$  and  $\nu_\tau$ , and between  $\nu_\mu$  and  $\nu_e$ .
- This  $\nu$ LFV translates into  $c$ LFV in SM, but with amplitudes  $\propto (m_\nu/m_W)^2$ . Will never be seen!
- New Physics model often have more direct ways to turn  $\nu$ LFV into  $c$ LFV.

# Example: $\mu$ -e conversion

## Supersymmetry

Predictions at  $10^{-15}$

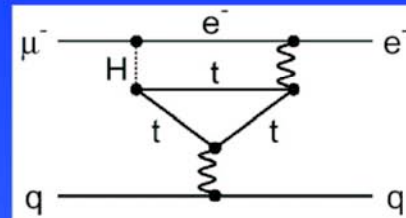
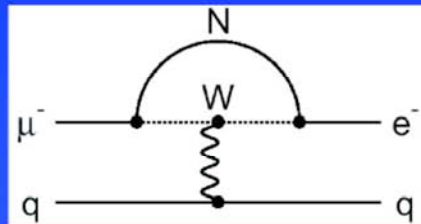


## Compositeness

$$\Lambda_c = 3000 \text{ TeV}$$

## Heavy Neutrinos

$$|U_{\mu N}^* U_{eN}|^2 = 8 \times 10^{-13}$$



## Second Higgs doublet

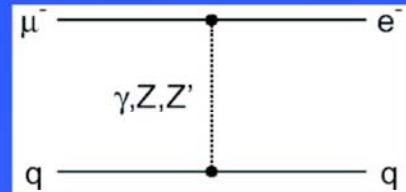
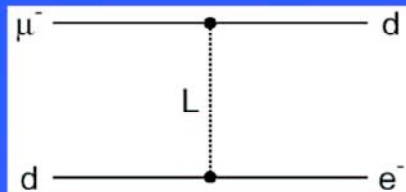
$$g_{H\mu e} = 10^{-4} \times g_{H\mu\mu}$$

## Leptoquarks

$$M_L =$$

$$3000 (\lambda_{\mu d} \lambda_{e d})^{1/2} \text{ TeV}/c^2$$

After W. Marciano



## Heavy $Z'$ , Anomalous $Z$ coupling

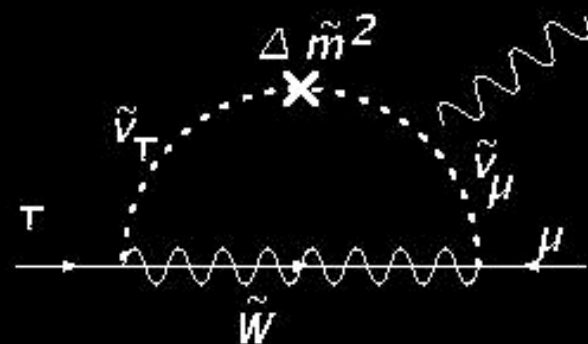
$$M_{Z'} = 3000 \text{ TeV}/c^2$$

$$B(Z \rightarrow \mu e) < 10^{-17}$$

SUSY has a simple way to turn vLFV into cLFV:  
SLEPTONS!

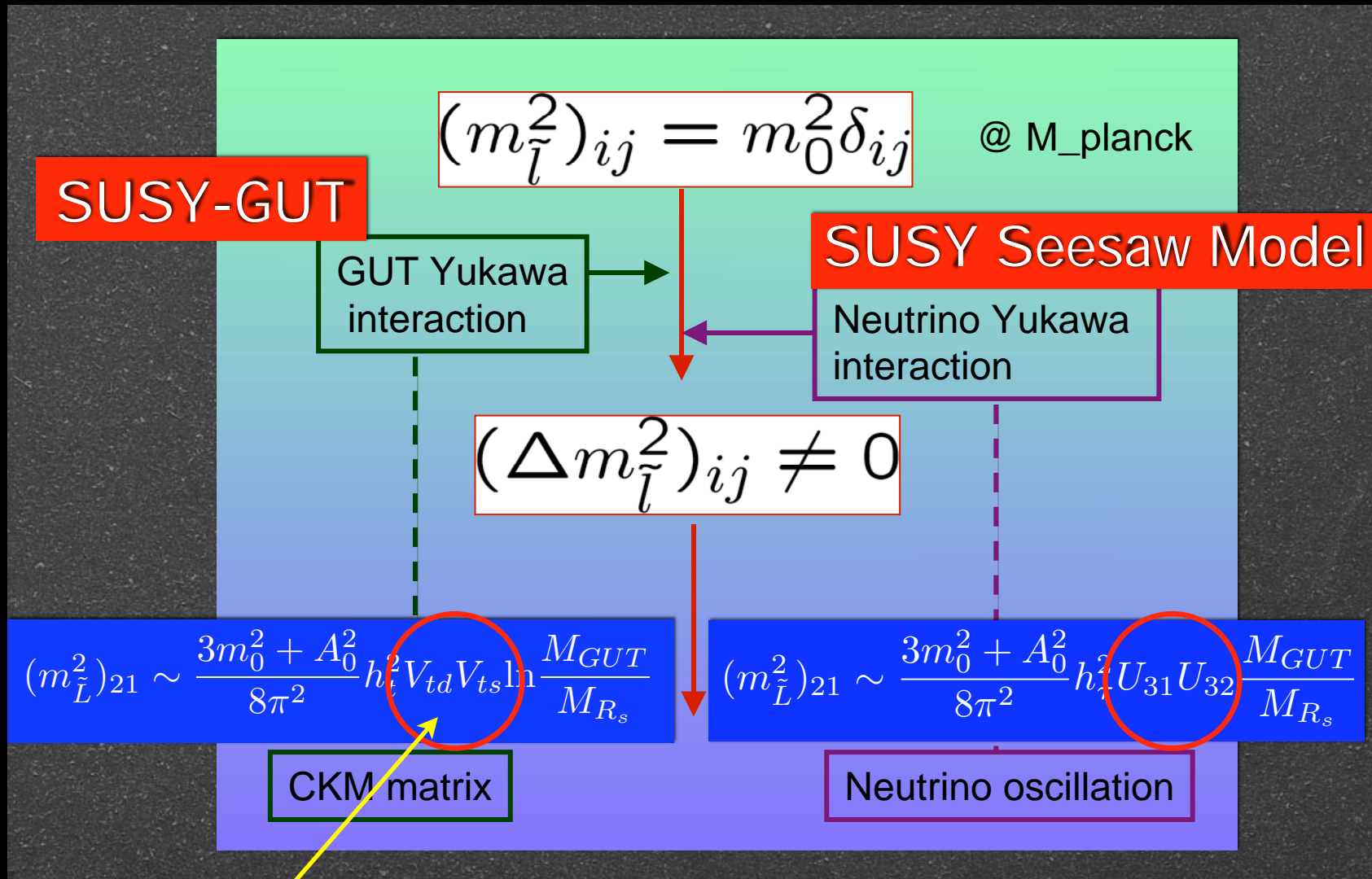
Sleptons encode the LFV in their mass matrices.

$$m_{\tilde{L}}^2 = \begin{pmatrix} \tilde{m}_e^2 & \Delta\tilde{m}_{e\mu}^2 & \Delta\tilde{m}_{e\tau}^2 \\ - & \tilde{m}_\mu^2 & \Delta\tilde{m}_{\mu\tau}^2 \\ - & - & \tilde{m}_\tau^2 \end{pmatrix}$$



The off-diagonal pieces generate cLFV at 1-loop.

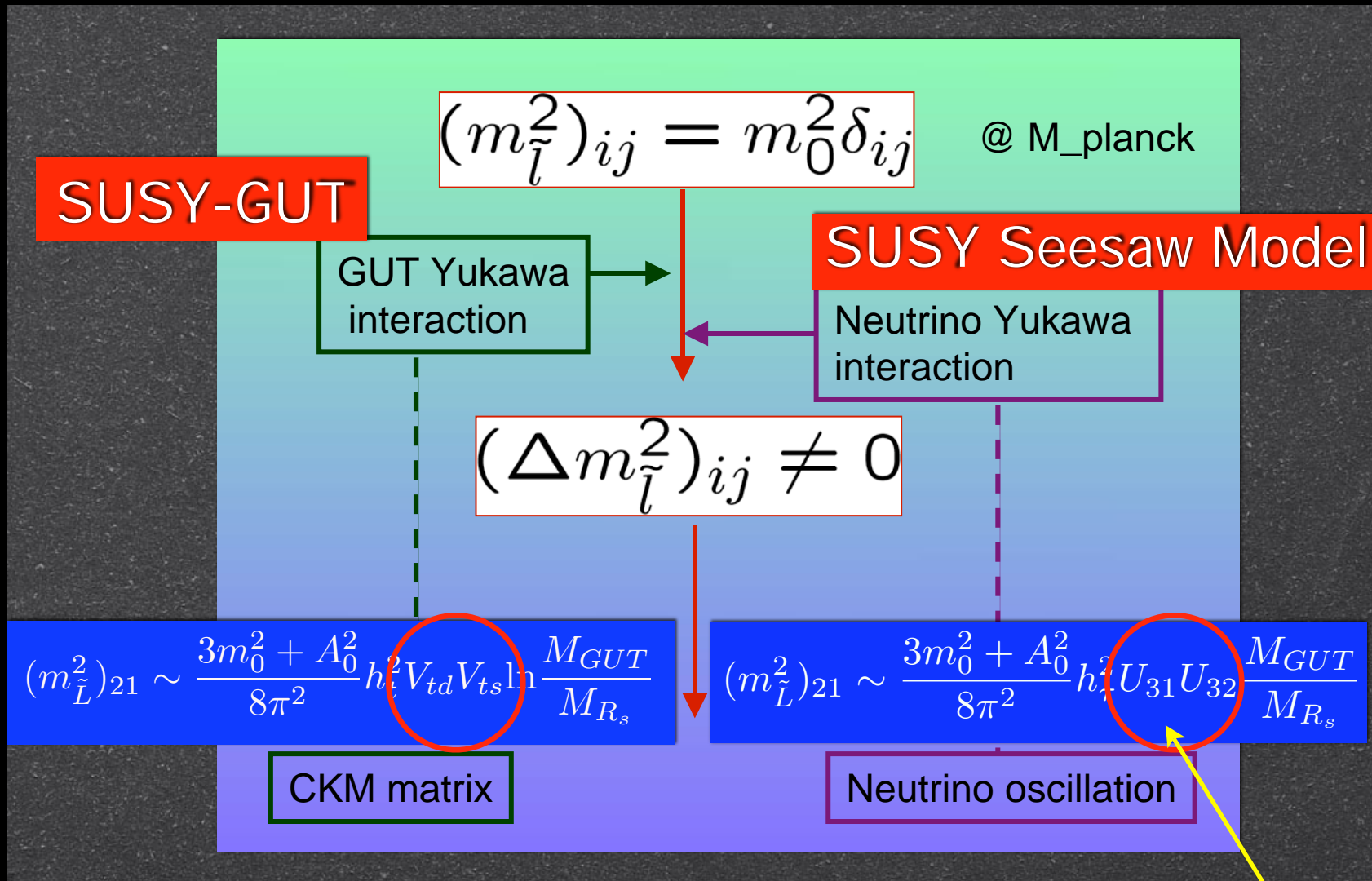
SUSY has 2 main mechanisms for generating off-diagonal slepton masses:



[Talk of Y. Kuno @ CIPANP '06]

Tied to quark mixing

SUSY has 2 main mechanisms for generating off-diagonal slepton masses:



Tied to neutrino mixing

How do  $\nu$ -masses affect sleptons?

In seesaw models, there are heavy RH neutrinos and neutrino Yukawa couplings at a scale  $M_R$ , where

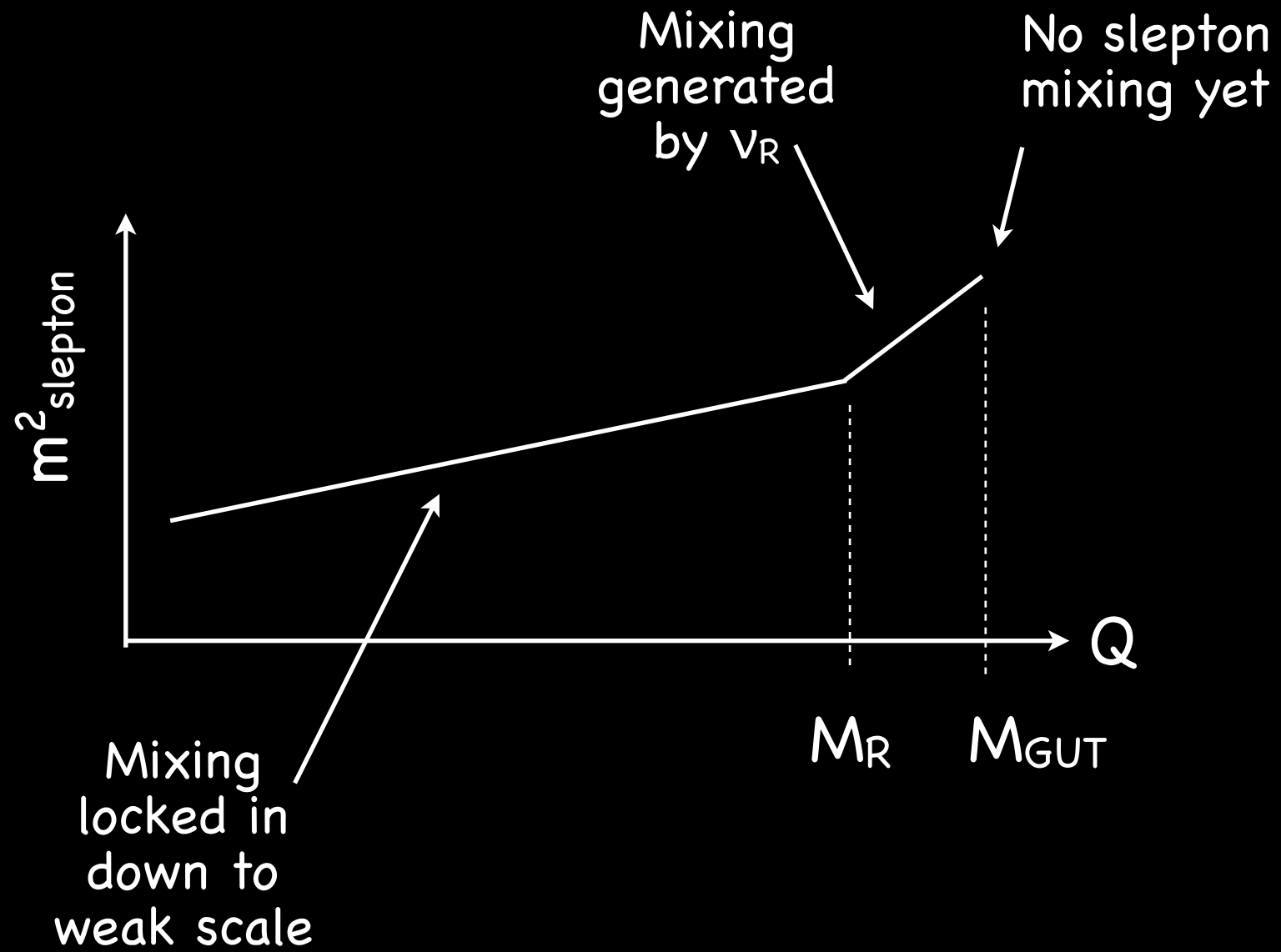
$$m_\nu = \frac{y_\nu^2 \langle H \rangle^2}{M_R}$$

At scales above  $M_R$ , the slepton mass RGE's include effects from  $y_\nu$ :

$$\frac{d\tilde{m}_\ell^2}{dt} = \dots + \frac{1}{8\pi^2} Y_\nu^\dagger Y_\nu (\tilde{m}_L^2 + \tilde{m}_E^2 + \tilde{m}_{H_u}^2 + A_\nu^2)$$

Large neutrino mixings imply  $Y_\nu$  possibly highly mixed.

$$(\Delta m_{\tilde{\ell}}^2)_{ij} \simeq -\frac{\log(M_X/M_R)}{16\pi^2} \left( 6m_0^2 (Y_\nu^\dagger Y_\nu)_{ij} + 2 (A_\nu^\dagger A_\nu)_{ij} \right)$$



Branching ratio  $\tau \rightarrow \mu \gamma$  depends on slepton mass mixing:

$$\text{Br}(\tau \rightarrow \mu \gamma) \simeq \frac{\alpha g^4}{64\pi^2} \frac{m_\tau^5 \tau_\tau}{M_{\text{SUSY}}^4} \frac{(\delta m^2)_{23}^2}{M_{\text{SUSY}}^4}$$

But there are 2 alternative  $\nu$ -mass matrices:

$$m_\nu \propto \begin{pmatrix} \epsilon & \epsilon & \epsilon \\ \epsilon & s_\theta^2 & c_\theta s_\theta \\ \epsilon & c_\theta s_\theta & c_\theta^2 \end{pmatrix}$$

Normal Hierarchy

$$m_\nu \propto \begin{pmatrix} \epsilon & c_\theta & s_\theta \\ c_\theta & \epsilon & \epsilon \\ s_\theta & \epsilon & \epsilon \end{pmatrix}$$

Inverted Hierarchy

Branching ratio  $\tau \rightarrow \mu \gamma$  depends on slepton mass mixing:

$$\text{Br}(\tau \rightarrow \mu \gamma) \simeq \frac{\alpha g^4}{64\pi^2} \frac{m_\tau^5 \tau_\tau}{M_{\text{SUSY}}^4} \frac{(\delta m^2)_{23}^2}{M_{\text{SUSY}}^4}$$

But there are 2 alternative  $\nu$ -mass matrices:

$$m_\nu \propto \begin{pmatrix} \epsilon & \epsilon & \epsilon \\ \epsilon & s_\theta^2 & c_\theta s_\theta \\ \epsilon & c_\theta s_\theta & c_\theta^2 \end{pmatrix}$$

Normal Hierarchy



$\tau \rightarrow \mu \gamma$

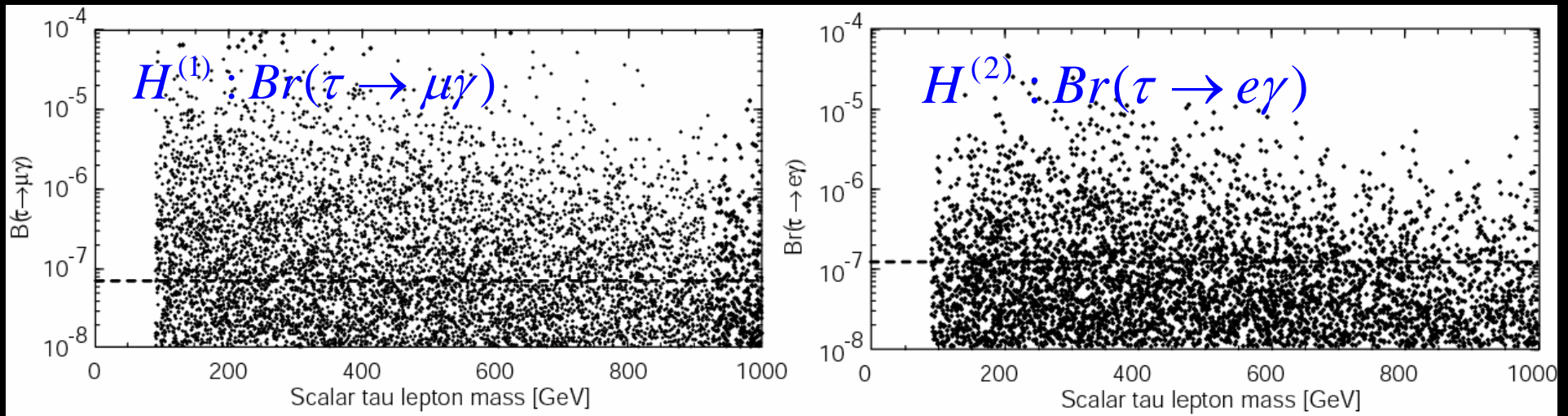
$$m_\nu \propto \begin{pmatrix} \epsilon & c_\theta & s_\theta \\ c_\theta & \epsilon & \epsilon \\ s_\theta & \epsilon & \epsilon \end{pmatrix}$$

Inverted Hierarchy



$\tau \rightarrow e \gamma$

# SuperB factory may solve an enduring problem in neutrino physics!



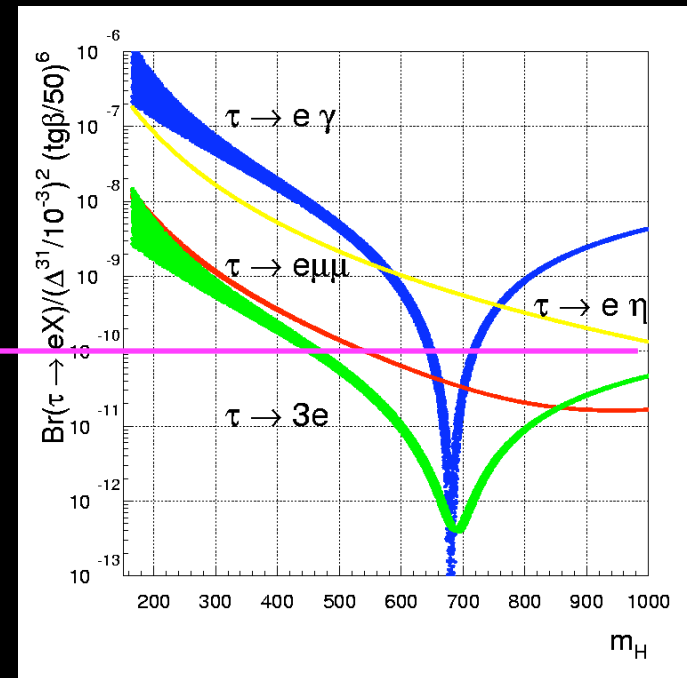
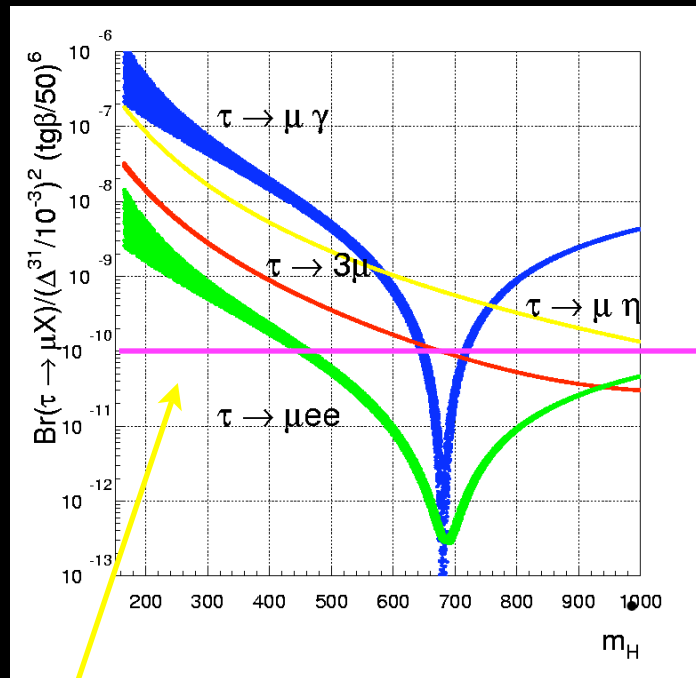
SuperB range ( $10^{-9}$  to  $10^{-10}$ )

Most (all?) reasonable mSUGRA parameter space probed by SuperB!

Higgs bosons can also mediate cLFV in models with 2+ Higgs bosons, if  $\tan\beta$  large.

Babu, CK

Modes are the same as standard cLFV, but ratios point to Higgs origin.



Paradisi

Approx SuperB  
limit

## Conclusions:

- The success of the Standard Model CKM scenario under severe tests by BaBar/Belle and CDF/D0 means that there is no one golden mode around which to sell a Super-B factory.
- But we KNOW the SM is incomplete from astrophysical data and theoretical consistency. New physics is expected at TeV scale.
- A SuperB factory is needed to constrain and test the kinds of new physics seen at LHC, particularly SUSY. Is nature minimally flavor violating or not?
- A SuperB factory is needed because our arguments might simply be wrong, and we'll never know if we don't check.