

$K^+ \rightarrow \pi^+ \nu \bar{\nu}$
at NNLO

Ulrich Haisch



in collaboration with A.J. Buras, M. Gorbahn & U. Nierste,
[hep-ph/0508165](#)

LoopFest IV, August 19, 2005, Snowmass, CO

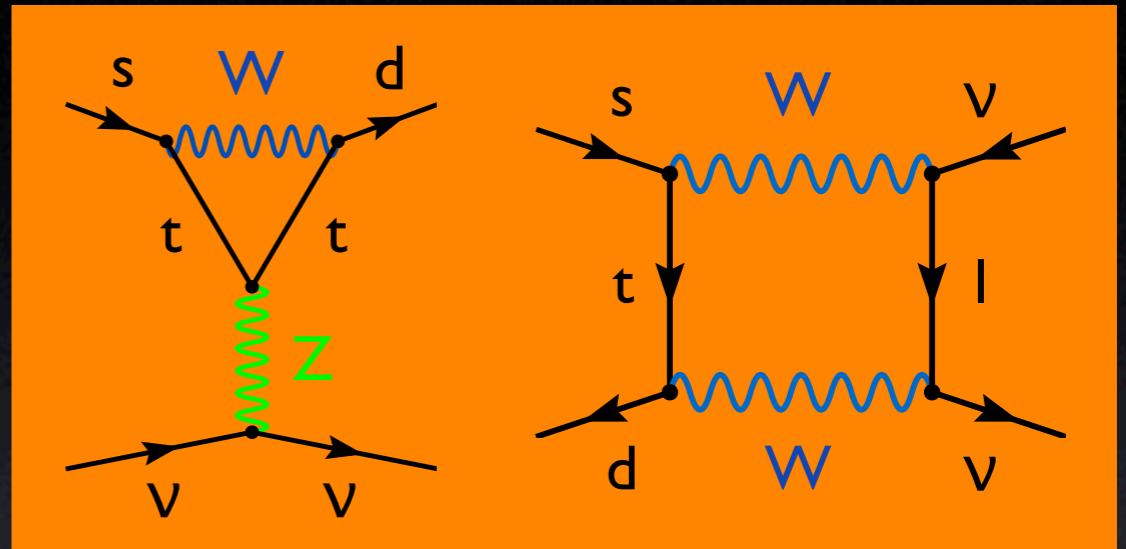
Next 29 minutes before



- Introduction
- Theoretical status of $K \rightarrow \pi \nu \bar{\nu}$
- NNLO calculation of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$
- Determination of unitarity triangle
- Conclusions

Introduction

- FCNC processes strongly suppressed in SM by loop and CKM factors
- SD effects are significant and calculable with high precision
- LD hadronic effects are small and under good theoretical control



precise
determination of flavor
structure of SM

Introduction

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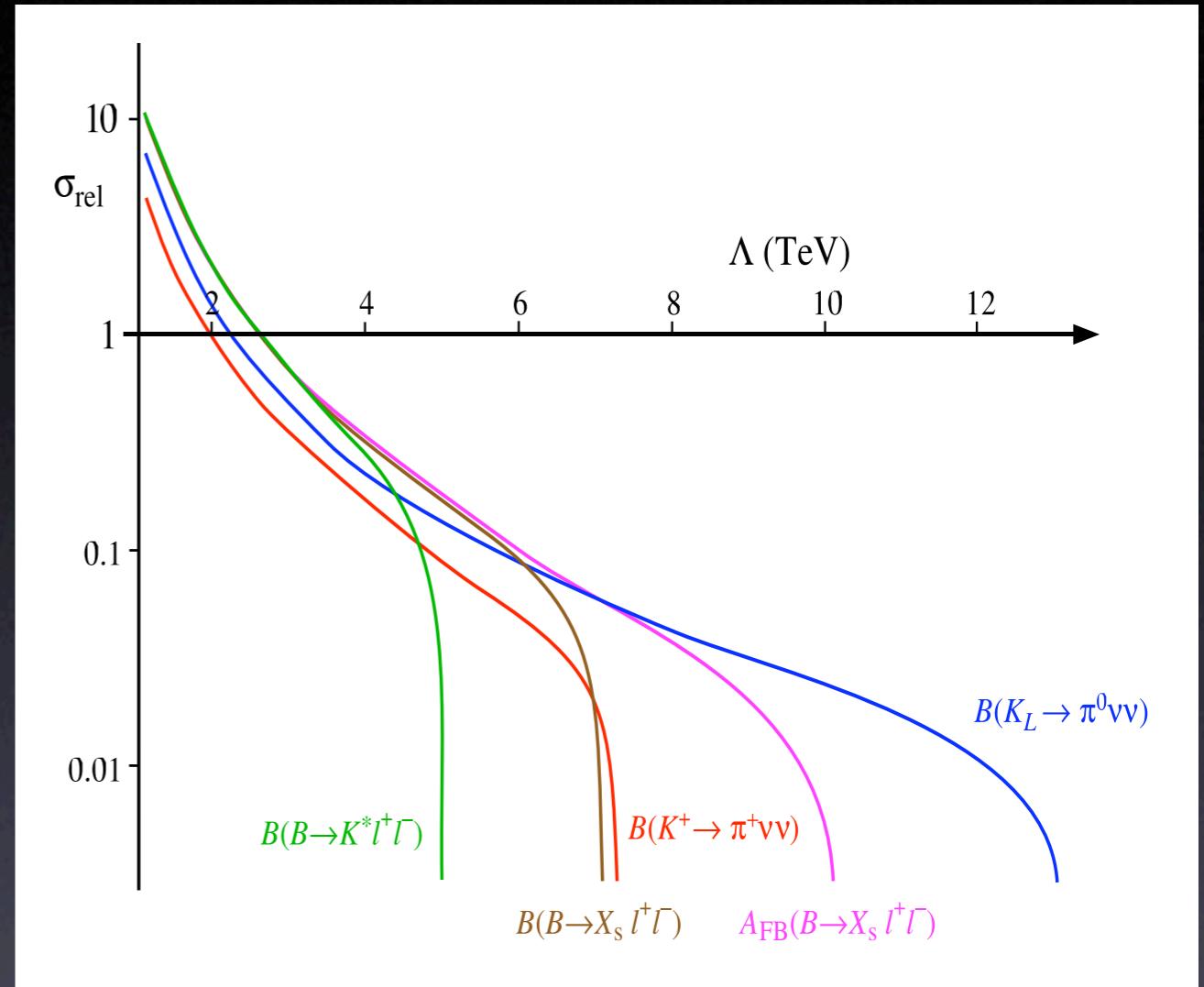
$$\begin{aligned}\sigma(\mathcal{B}(K_L \rightarrow \pi^0 \nu \bar{\nu})) &= \pm(1 - 2)\% \\ \sigma(\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})) &= \pm(3 - 4)\% \\ \sigma(\mathcal{A}_{\text{FB}}(B \rightarrow X_s l^+ l^-)) &= \pm(6 - 9)\% \\ \sigma(\mathcal{B}(B \rightarrow X_s \gamma)) &= \pm(8 - 14)\% \\ \sigma(\mathcal{B}(B \rightarrow X_s l^+ l^-)) &= \pm(10 - 17)\% \\ \sigma(\mathcal{B}(B_s \rightarrow \mu^+ \mu^-)) &= \pm 15\% \\ \sigma(\mathcal{A}_{\text{FB}}(B \rightarrow K^{(*)} l^+ l^-)) &= \pm 15\% \\ \sigma(\mathcal{B}(B \rightarrow (K^*, \rho, \omega) \gamma)) &= \pm(15 - 30)\% \\ \sigma(\mathcal{B}(B \rightarrow K^* l^+ l^-)) &= \pm(30 - 35)\% \\ \dots \end{aligned}$$

Introduction

- FCNC processes strongly suppressed in SM by loop and CKM factors
- SD effects are significant and calculable with high precision
- LD hadronic effects are small and under good theoretical control



enhanced
sensitivity to flavor
dynamics of NP



Bryman, Buras, Isidori & Littenberg '05

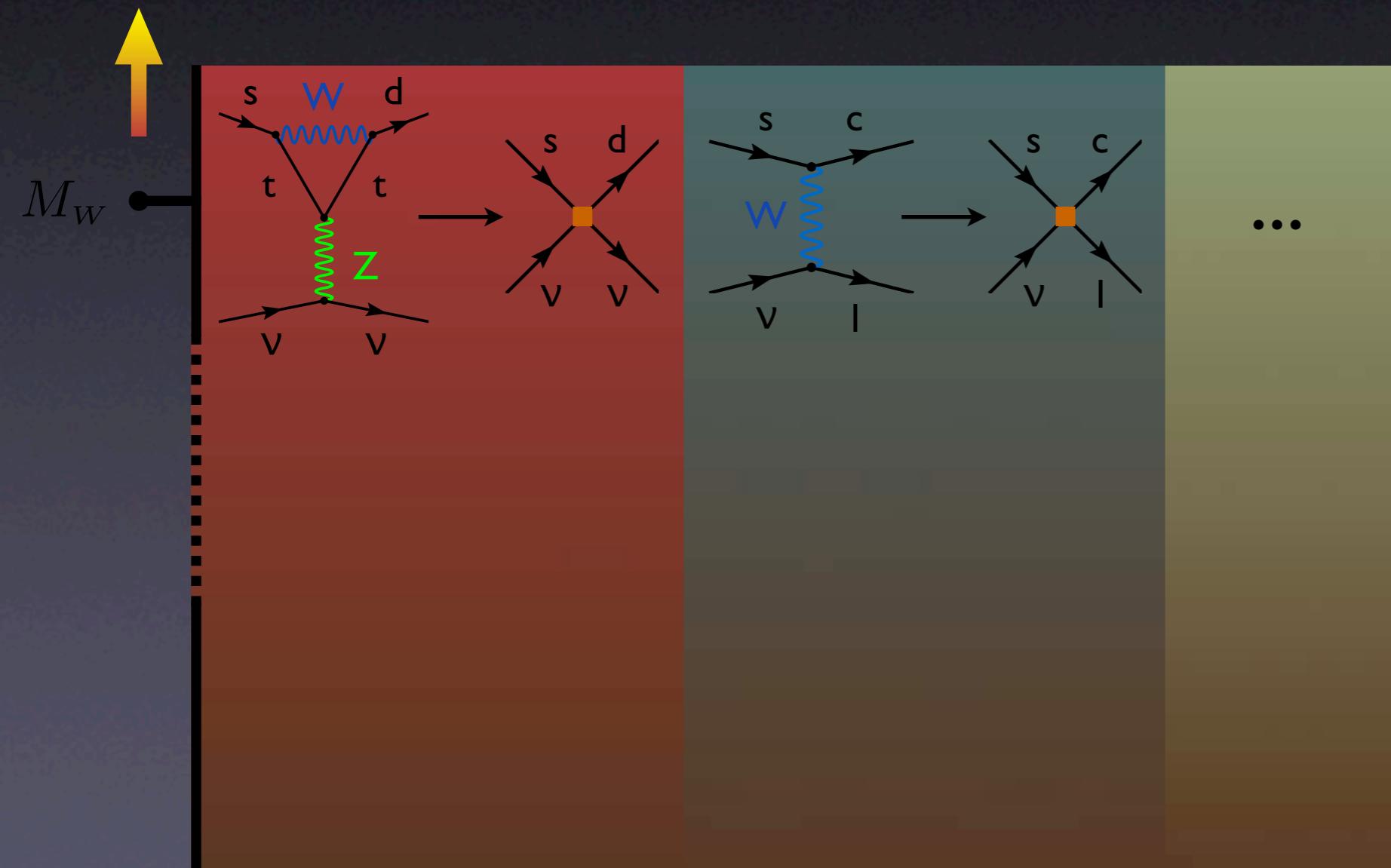
General properties of $K \rightarrow \pi \nu \bar{\nu}$

$$\mathcal{L}_{\text{eff}} = \frac{4G_F}{\sqrt{2}} \frac{\alpha}{2 \sin^2 \theta_W} \sum_{i=u,c,t} C^i(\mu) Q_\nu$$

$$Q_\nu = \sum_{l=e,\mu,\tau} (\bar{s}_L \gamma_\mu d_L)(\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$C^i(M_W) \propto m_i^2 V_{is}^* V_{id} \propto \begin{cases} \Lambda_{\text{QCD}}^2 \lambda \\ m_c^2 (\lambda + i\lambda^5) \\ m_t^2 (\lambda^5 + i\lambda^5) \end{cases}$$

u
c
t



- power-like GIM mechanism
- top quark contributions dominates
- QCD corrections are small
- large \mathcal{CP} phase

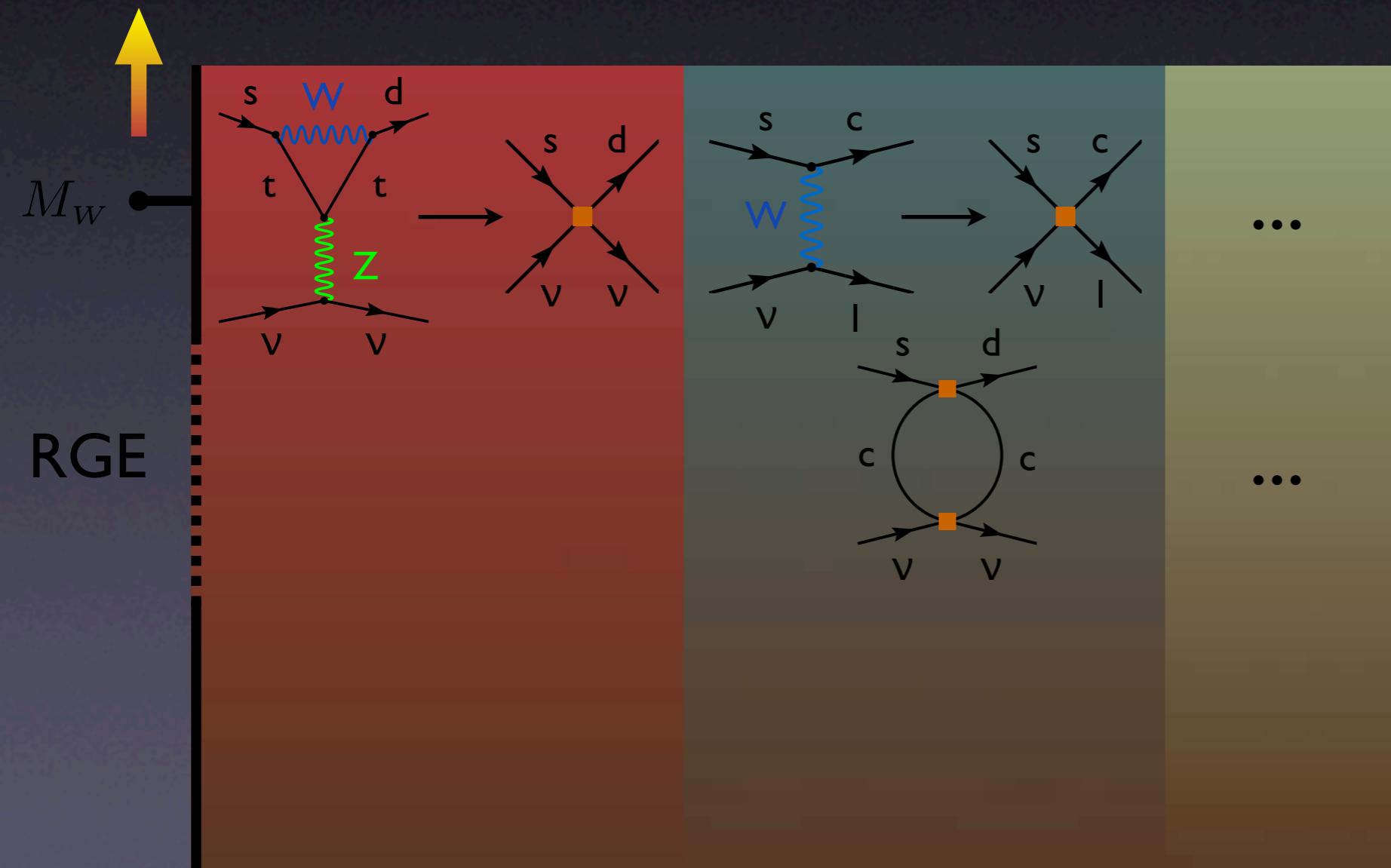
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$$Q_\nu = \sum_{l=e,\mu,\tau} (\bar{s}_L \gamma_\mu d_L)(\bar{\nu}_l L \gamma^\mu \nu_{lL})$$

$$\mu \frac{d}{d\mu} C_i(\mu) = \sum_j \underbrace{\gamma_{ji}(\mu)}_{\text{ADM}} C_j(\mu)$$

u
c



- top quark contribution does not evolve
- charm effects moderate for K^+ while negligible for K_L

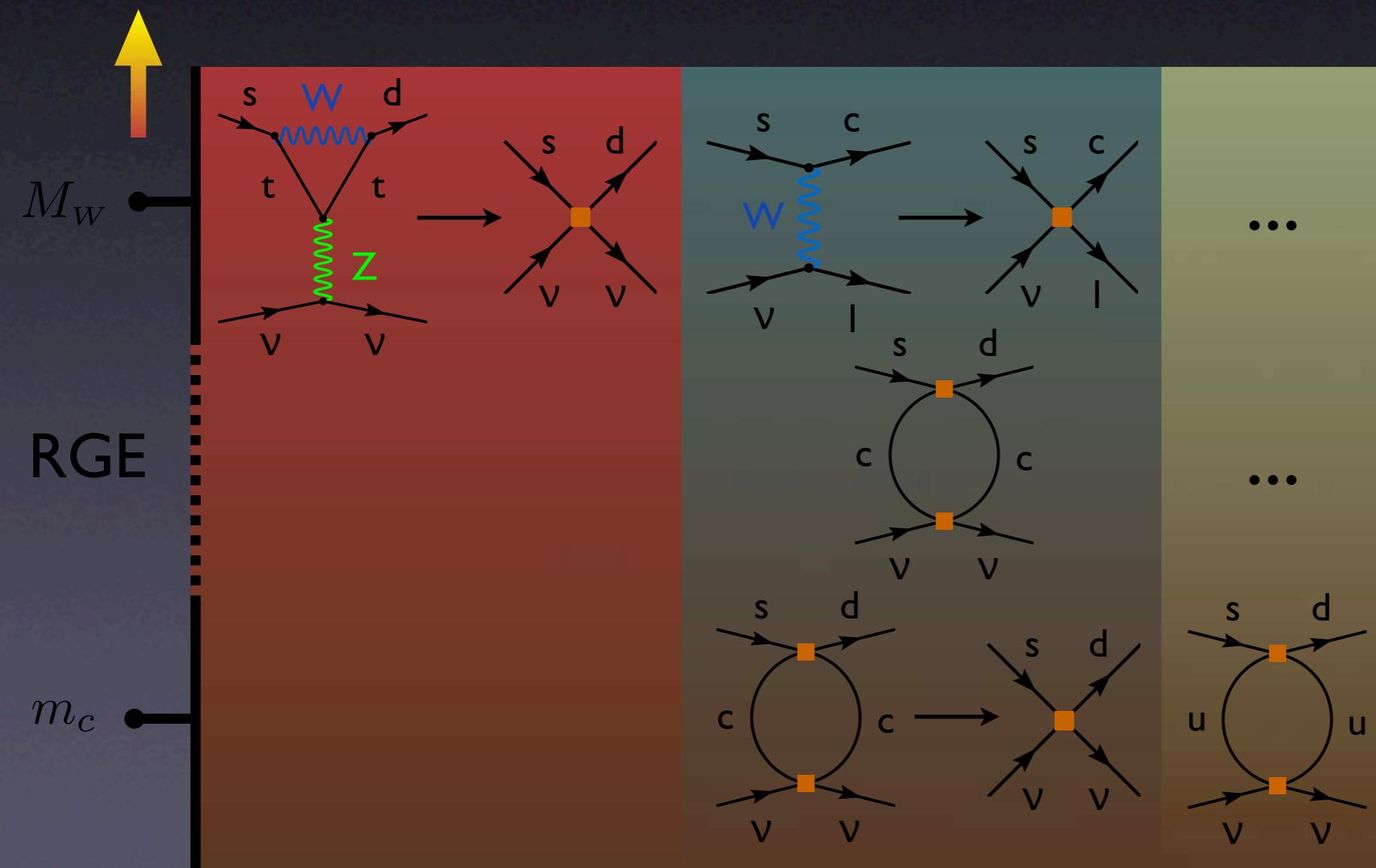
General properties of $K \rightarrow \pi \nu \bar{\nu}$

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$$Q_\nu = \sum_{l=e,\mu,\tau} (\bar{s}_L \gamma_\mu d_L)(\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$\langle \pi | \bar{s}_L \gamma_\mu d_L | K \rangle \propto \langle \pi^0 | \bar{s}_L \gamma_\mu u_L | K^+ \rangle$$

u



- hadronic matrix elements precisely known from K_{e3}^+
- neutrino pair in CP eigenstate
- K_L decay purely CP

NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im} \lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re} \lambda_t}{\lambda^5} X + \frac{\text{Re} \lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right]$$

$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_W}$$

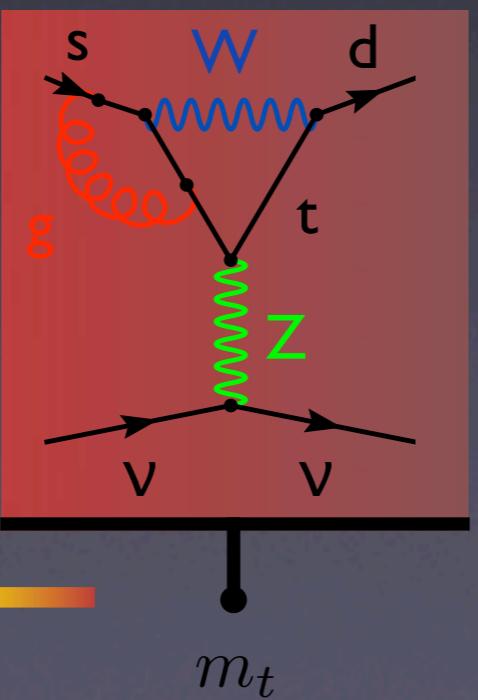
$$X = 1.46 \pm 0.04 \text{ (NLO)}$$

t

Buchalla & Buras '93, '99;
Misiak & Urban '99

$$\lambda_i = V_{is}^* V_{id}$$

$$\lambda = |V_{us}| \approx 0.22$$



NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

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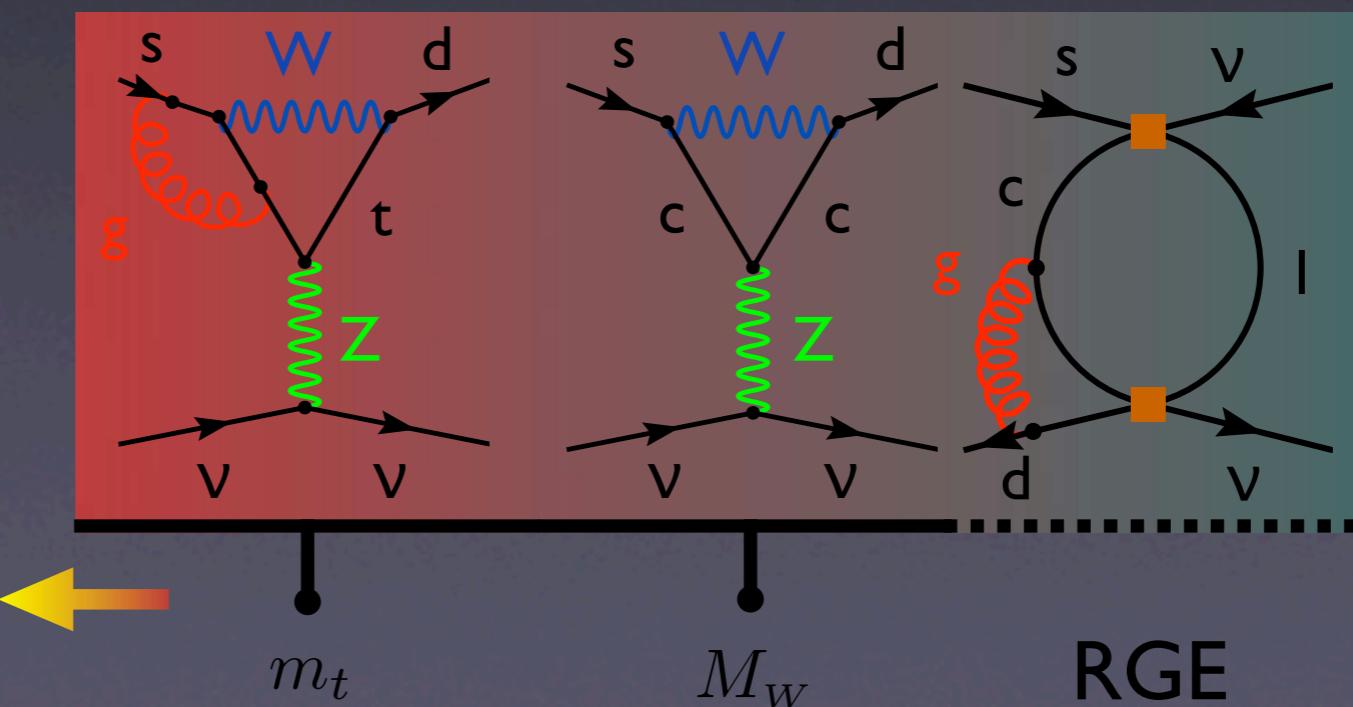
$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_W}$$

$$P_c = 0.37 \pm 0.06 \text{ (NLO)}$$

$$\lambda_i = V_{is}^* V_{id}$$

Buchalla & Buras '94, '99;
Buras, Gorbahn, Nierste & UH '05

$$\lambda = |V_{us}| \approx 0.22$$



NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

u
c
t

$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im} \lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re} \lambda_t}{\lambda^5} X + \frac{\text{Re} \lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right]$$

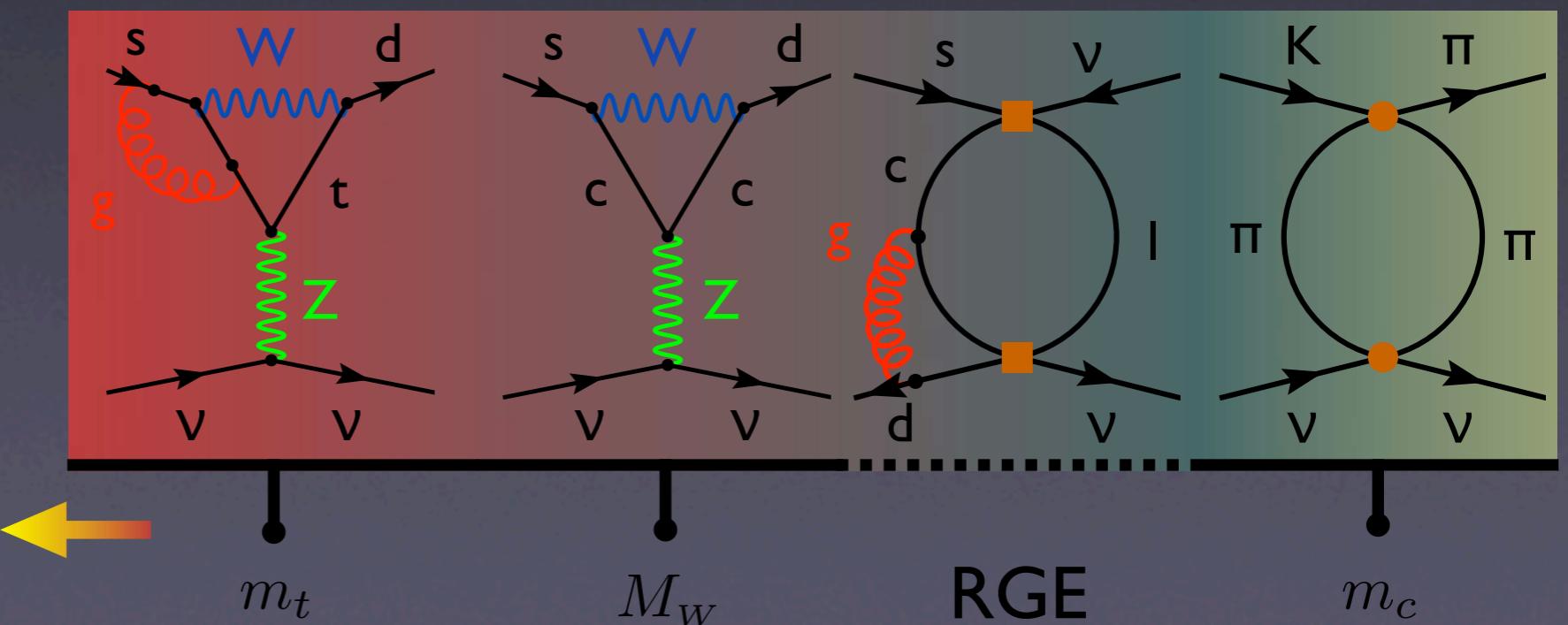
$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_W}$$

$$\delta P_c = 0.04 \pm 0.02 \text{ (CHPT)}$$

$$\lambda_i = V_{is}^* V_{id}$$

Isidori, Mescia & Smith '05

$$\lambda = |V_{us}| \approx 0.22$$



NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

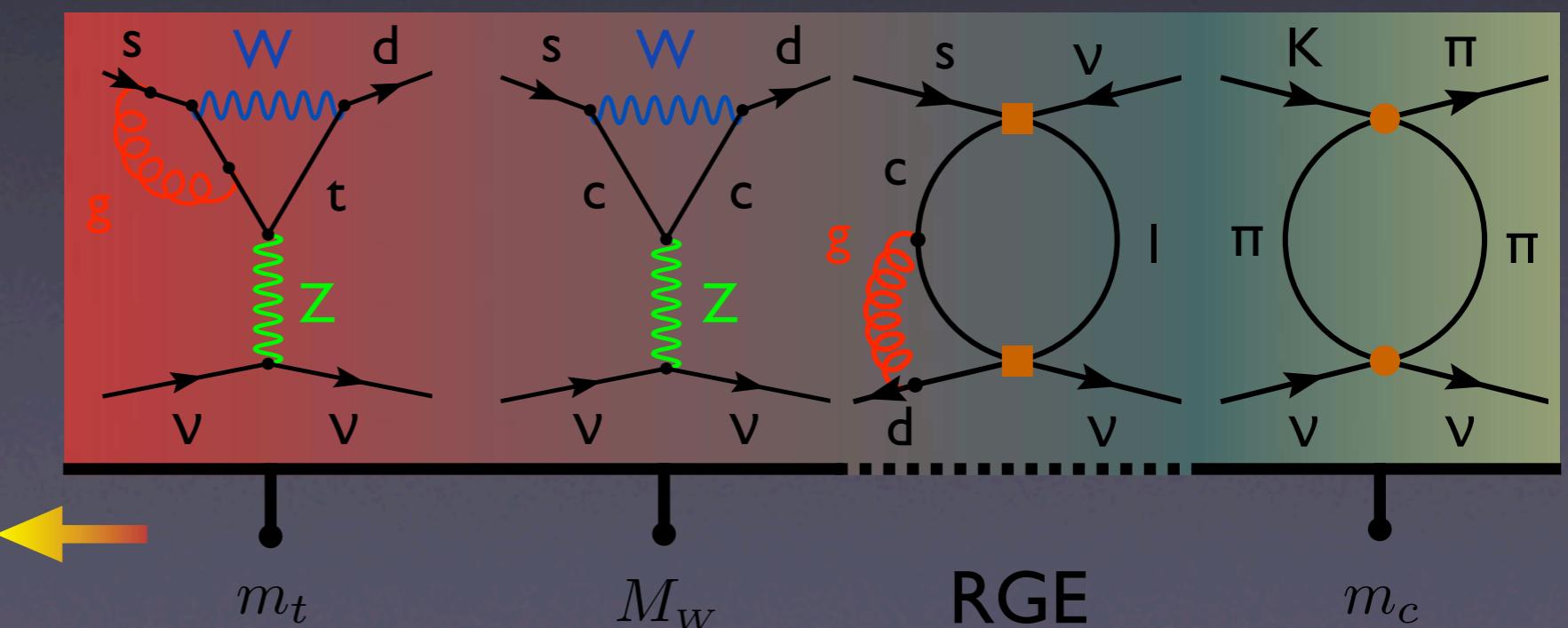
u
c
t

$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im} \lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re} \lambda_t}{\lambda^5} X + \frac{\text{Re} \lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right]$$

$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_W}$$

$$r_{K^+} = 0.90 \pm 0.03 \text{ (SU(2))}$$

Marciano & Parsa '96



NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

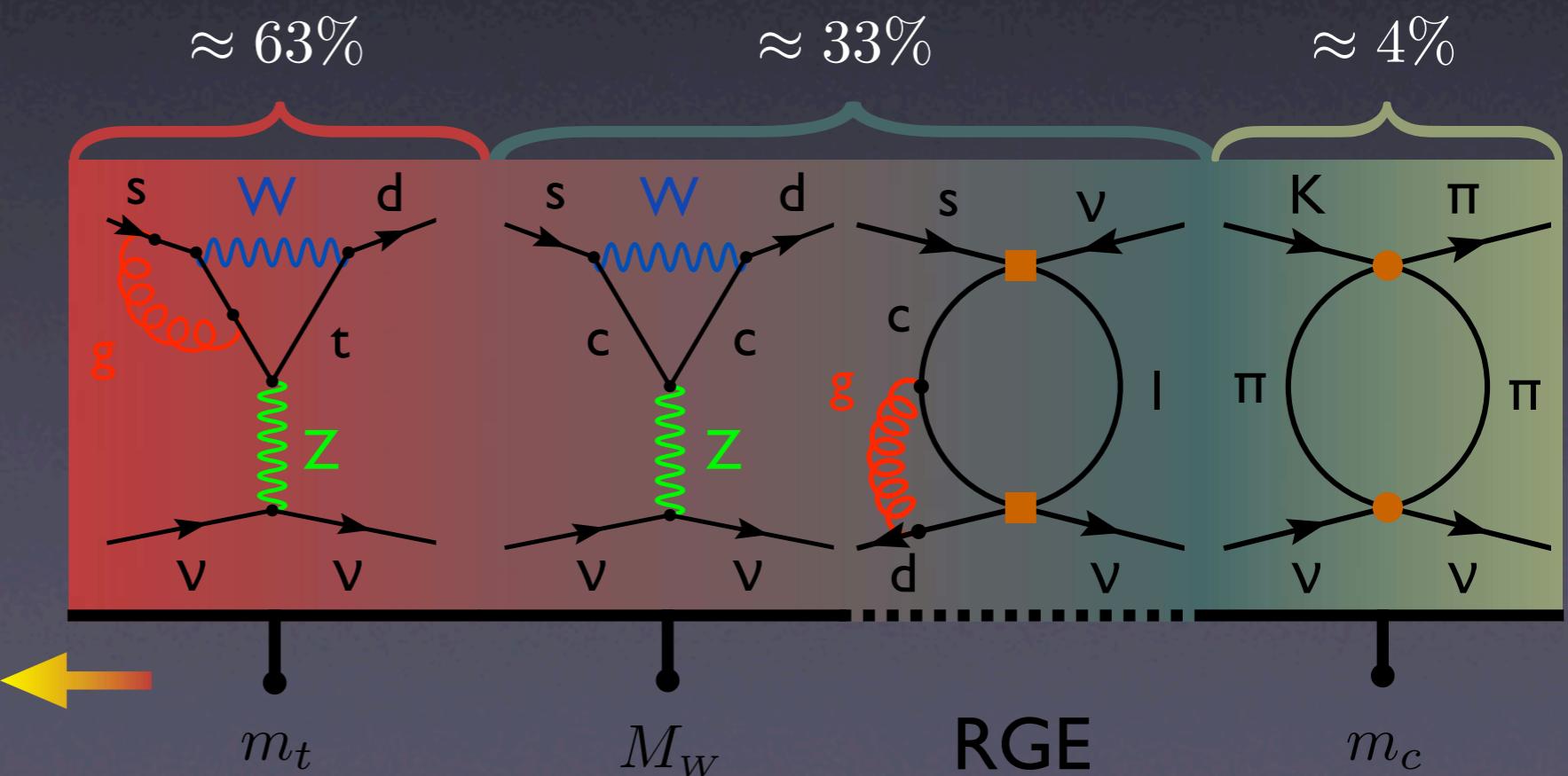
$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im} \lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re} \lambda_t}{\lambda^5} X + \frac{\text{Re} \lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right] = (7.9 \pm 1.3) \times 10^{-11}$$

u
c
t

$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_W}$$

Buras, Gorbahn, Nierste & UH '05

- dominant theoretical error of $\approx 6\%$ due to perturbative charm contribution P_c
- $\approx 50\%$ of total error of $\approx 16\%$ due to m_c and CKM elements

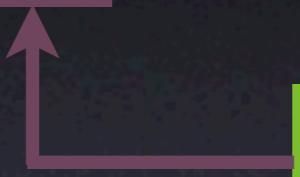


NLO SM prediction of $K_L \rightarrow \pi^0 \nu \bar{\nu}$

$$\mathcal{B}(K_L) = \kappa_L \left(\frac{\text{Im} \lambda_t}{\lambda^5} X \right)^2$$

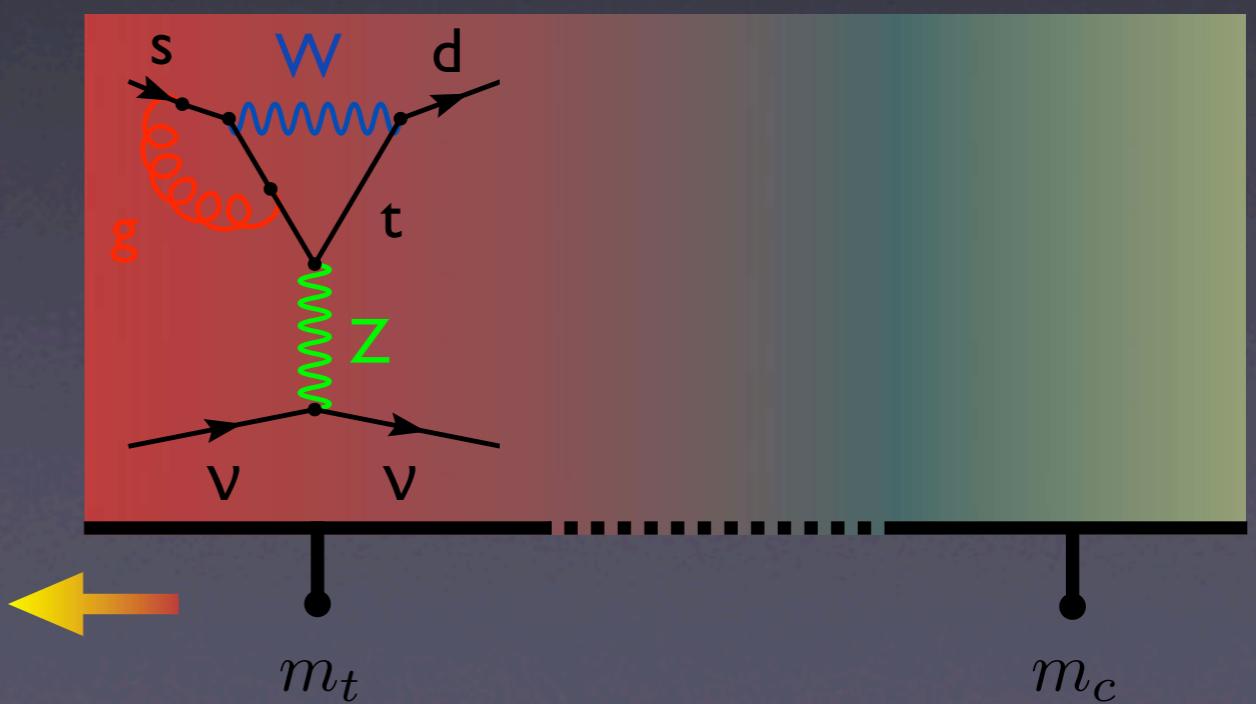
t

$$\kappa_L = \kappa_+ \frac{r_{K_L}}{r_{K^+}} \frac{\tau(K_L)}{\tau(K^+)}$$



$$r_{K_L} = 0.94 \pm 0.03 \text{ (SU(2))}$$

Marciano & Parsa '96



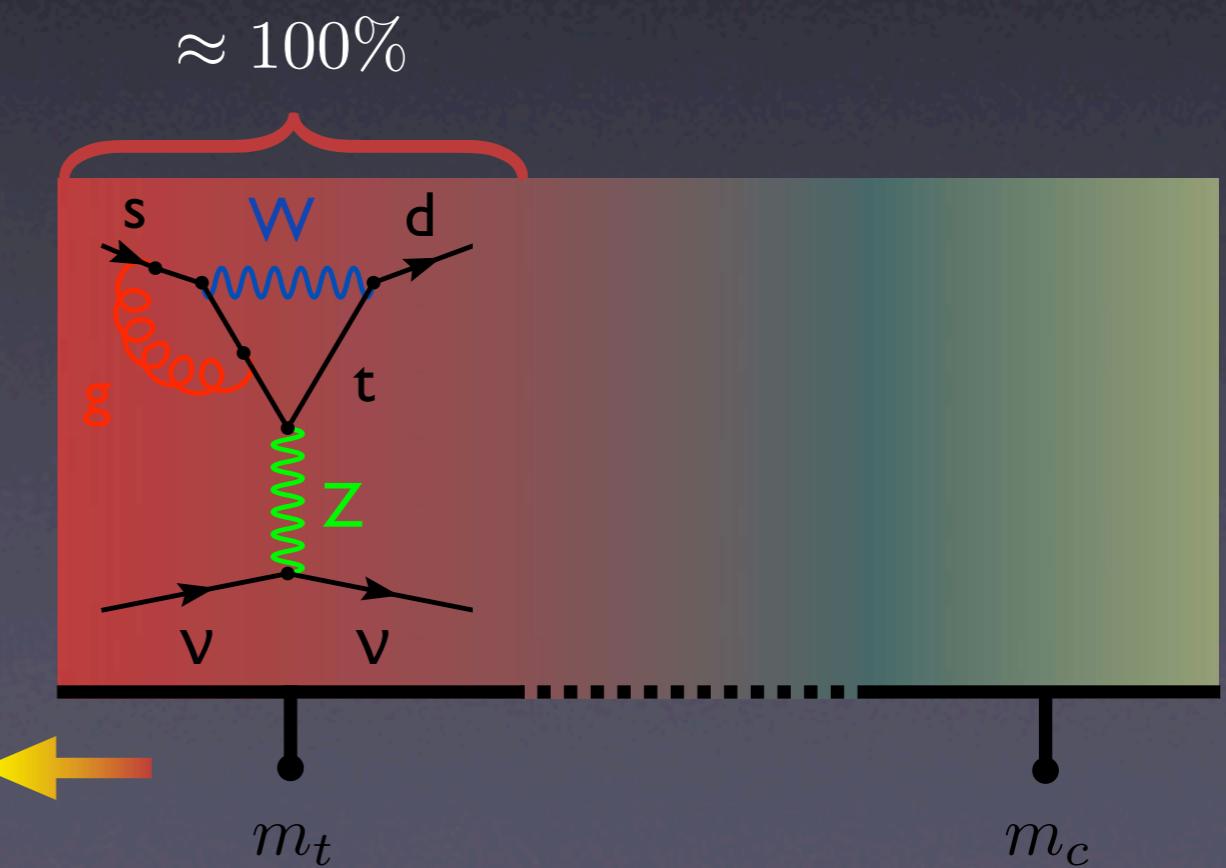
NLO SM prediction of $K_L \rightarrow \pi^0 \nu \bar{\nu}$

$$\mathcal{B}(K_L) = \kappa_L \left(\frac{\text{Im} \lambda_t}{\lambda^5} X \right)^2 = (2.9 \pm 0.4) \times 10^{-11}$$

$$\kappa_L = \kappa_+ \frac{r_{K_L}}{r_{K^+}} \frac{\tau(K_L)}{\tau(K^+)}$$

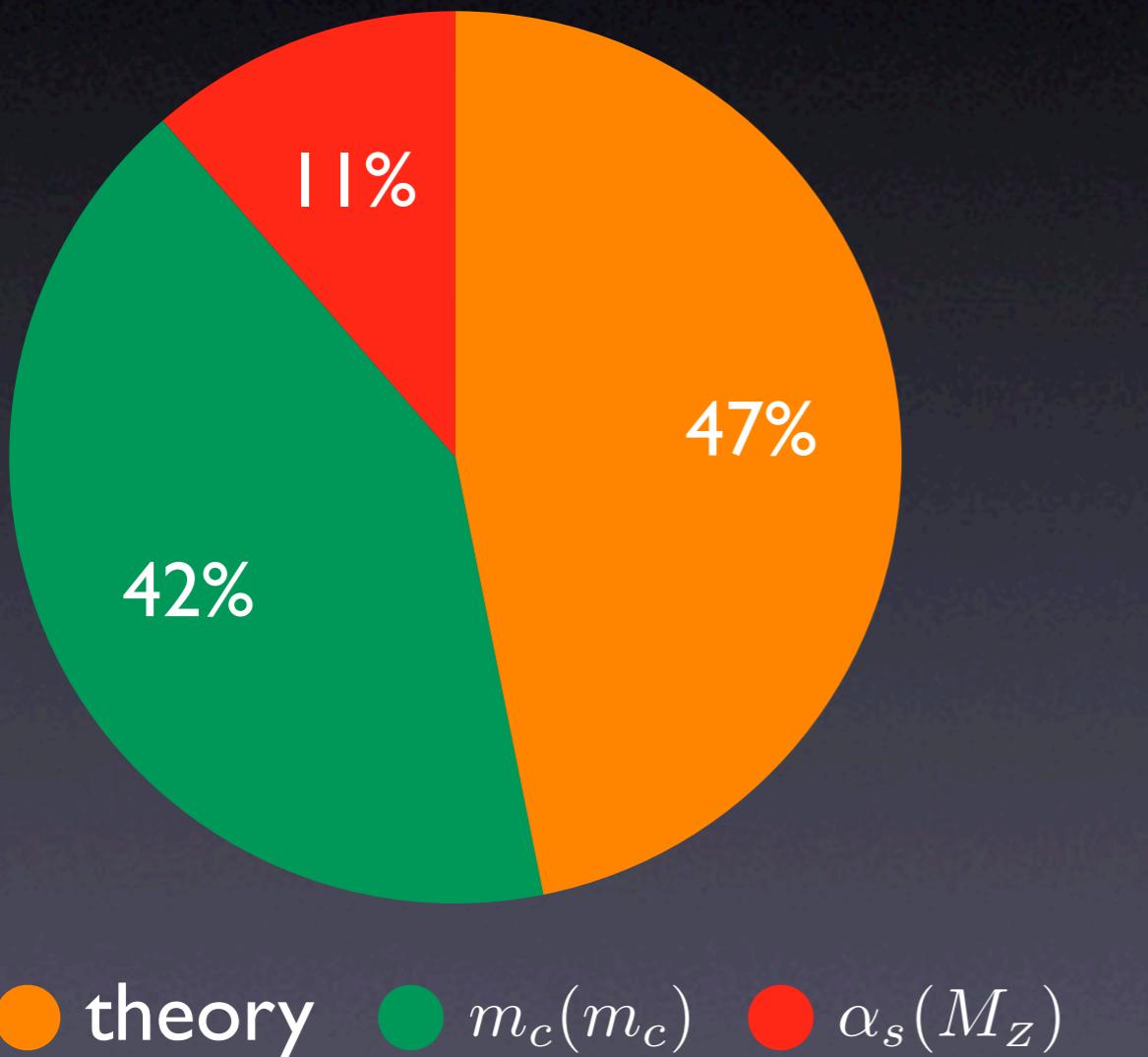
Littenberg '89;
Buchalla & Buras '96;
Buchalla & Isidori '98;
Buras, Gorbahn, Nierste & UH '05

- very small theoretical error of only $\approx 1\%$
- $\approx 90\%$ of total error of $\approx 15\%$ due to CKM parameters
- within SM amount of \mathcal{CP} can in principle be determined with unmatched precision

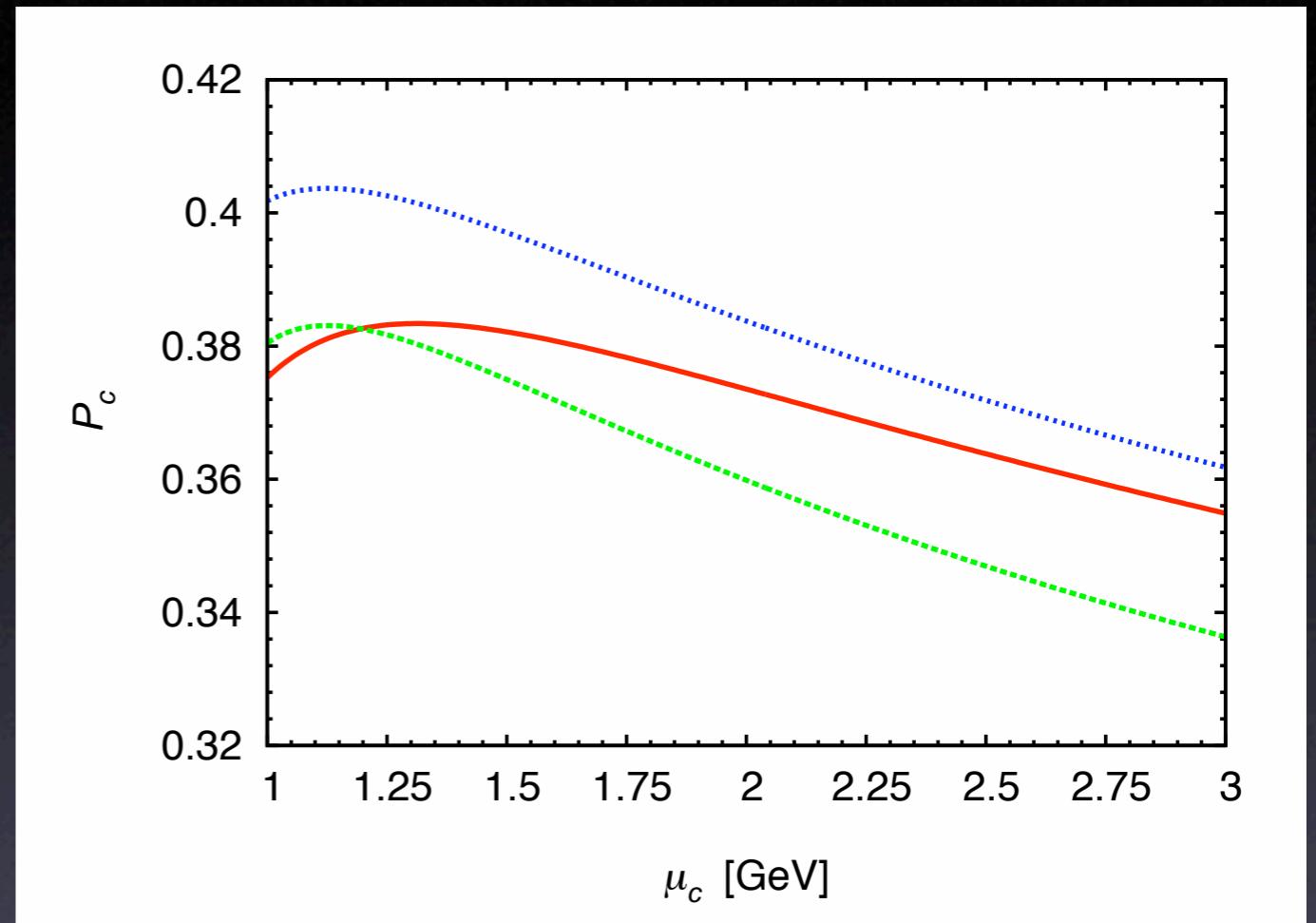
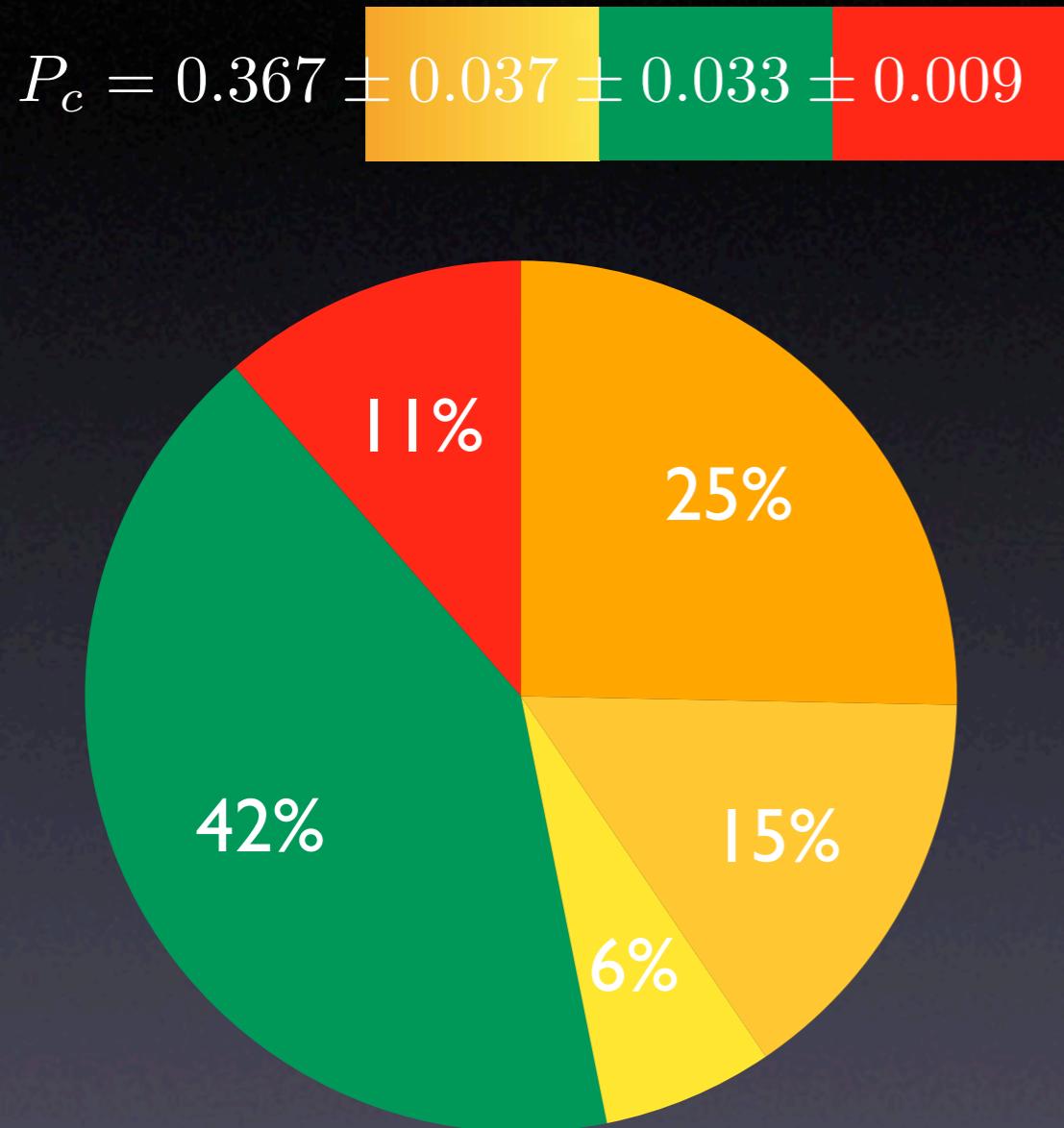


Error budget of P_c at NLO

$$P_c = 0.367 \pm 0.037 \pm 0.033 \pm 0.009$$



Error budget of P_c at NLO

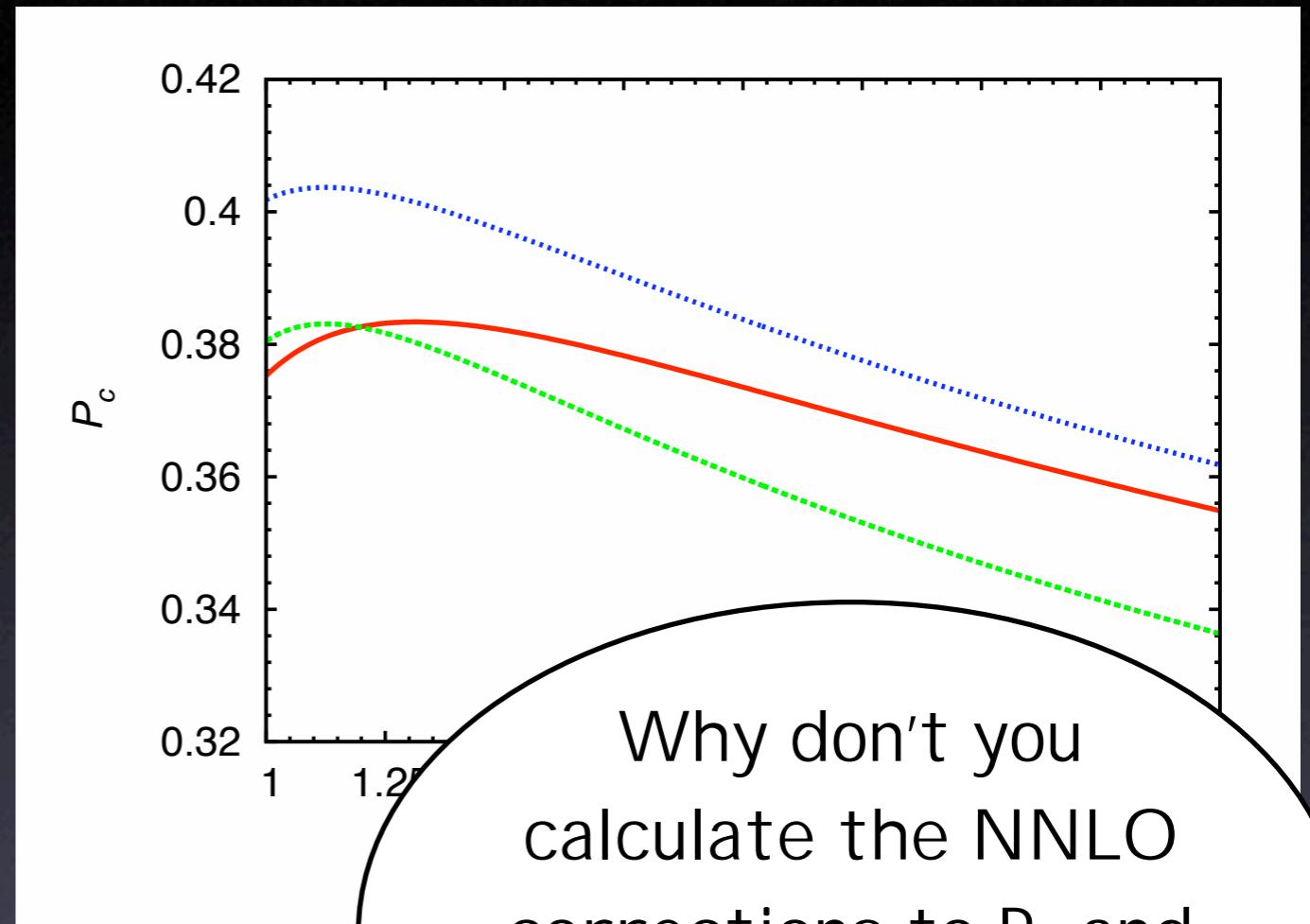
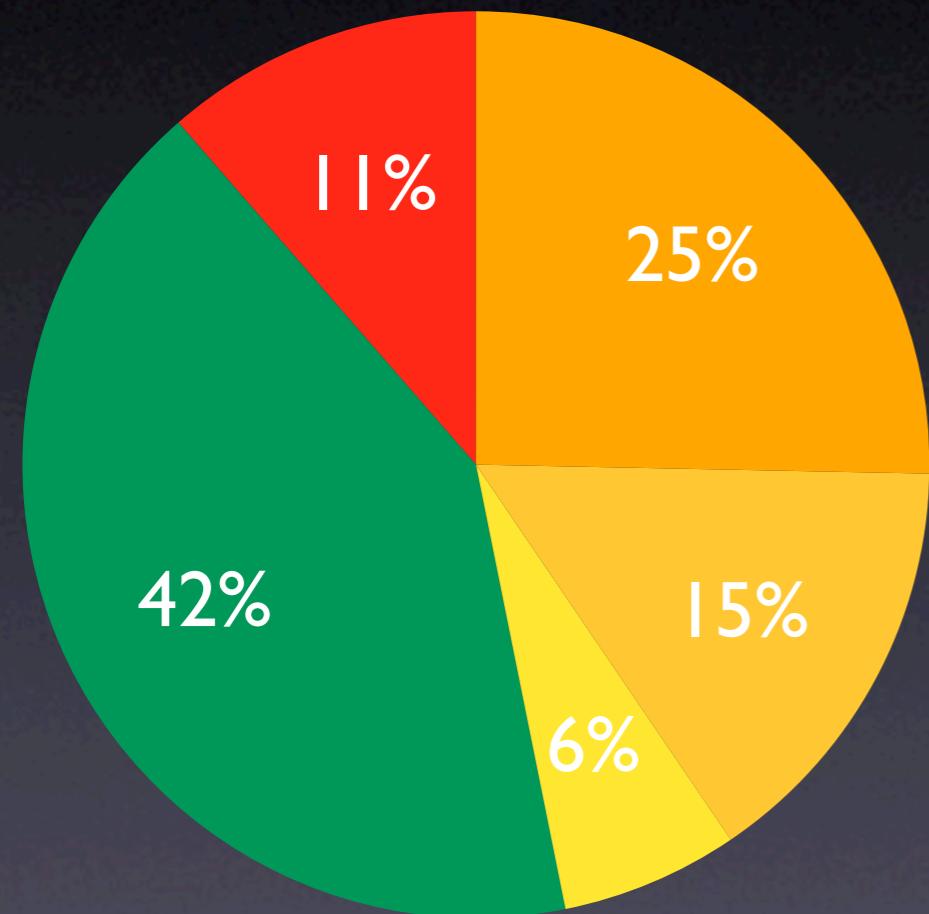


Buras, Gorbahn, Nierste & UH '05

- μ_c
- $\alpha_s(\mu_c)$
- $\mu_b + \mu_W$
- $m_c(m_c)$
- $\alpha_s(M_Z)$

Error budget of P_c at NLO

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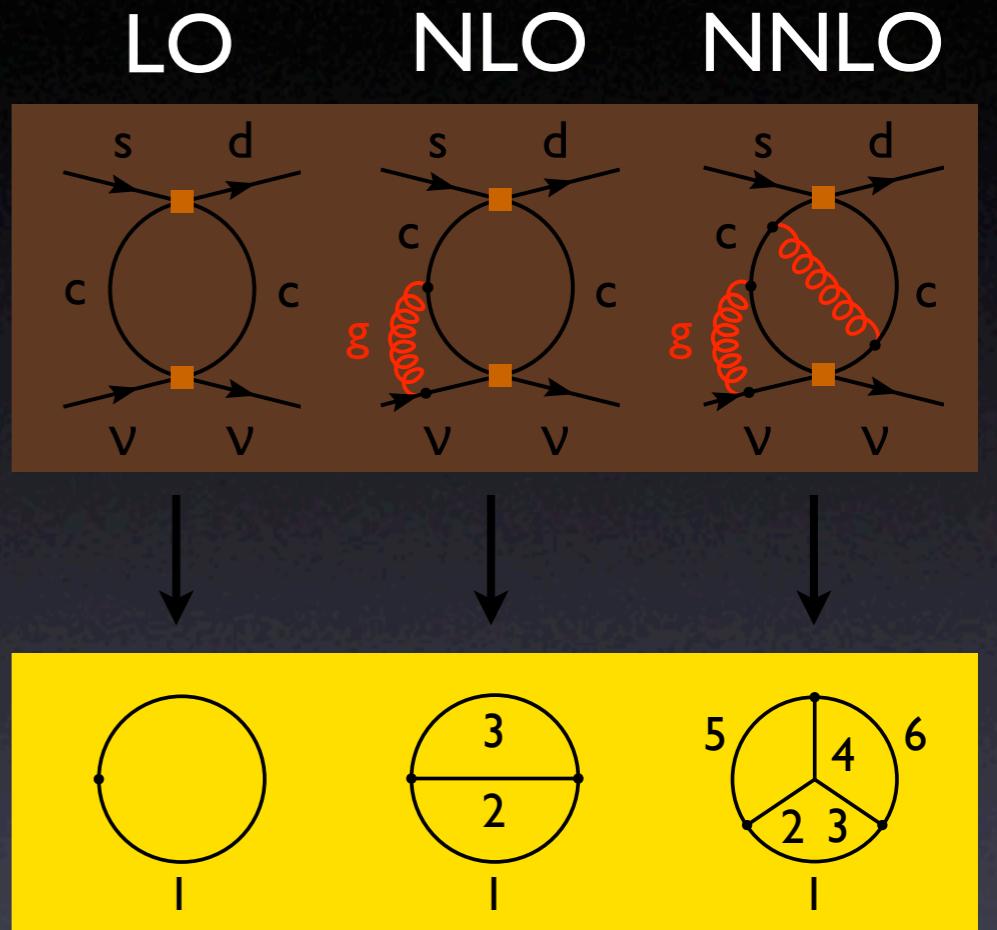
Why don't you calculate the NNLO corrections to P_c and reduce the theoretical error?



- μ_c
- $\alpha_s(\mu_c)$
- $\mu_b + \mu_W$
- $m_c(m_c)$
- $\alpha_s(M_Z)$

NNLO ADM calculation of P_c

- ADM is determined from I/ϵ_{UV} of I-, 2- and 3-loop diagrams
- integrals have I/ϵ_{IR} and I/ϵ_{UV} that are indistinguishable in DR
- in $\overline{\text{MS}}$ scheme I/ϵ_{UV} are polynomial in masses and momenta after subtraction of subdivergences
- calculation of counterterms reduces to computation of massive tadpoles



$$\frac{1}{(k+p)^2 - m^2} = \frac{1}{k^2 - M^2} - \frac{p^2 + 2k \cdot p - m^2 + M^2}{k^2 - M^2} \frac{1}{(k+p)^2 - m^2}$$

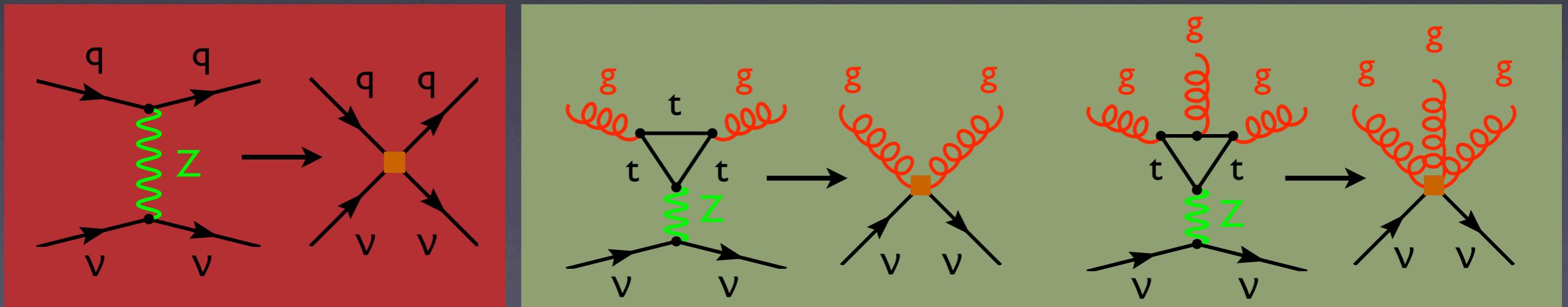
New features in NNLO calculation of P_c

$$Q_Z = \sum_q ((I_q^3 - 2e_q \sin^2 \theta_W) Q_V^q - I_q^3 (Q_A^q + Q_{CS})) \quad a_{CS} = \begin{cases} 2 & \text{HV} \\ \frac{2}{3} & \text{DRED} \end{cases}$$

$$Q_V^q = \sum_{l=e,\mu,\tau} (\bar{q} \gamma_\mu q) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$Q_A^q = \sum_{l=e,\mu,\tau} (\bar{q} \gamma_\mu \gamma_5 q) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$Q_{CS} = \frac{g^2}{16\pi^2} a_{CS} \epsilon^{\mu\nu\lambda\kappa} \left(G_\mu^a \partial_\nu G_\lambda^a + \frac{1}{3} g f^{abc} G_\mu^a G_\nu^b G_\lambda^c \right) \sum_{l=e,\mu,\tau} (\bar{\nu}_{lL} \gamma_\kappa \nu_{lL})$$



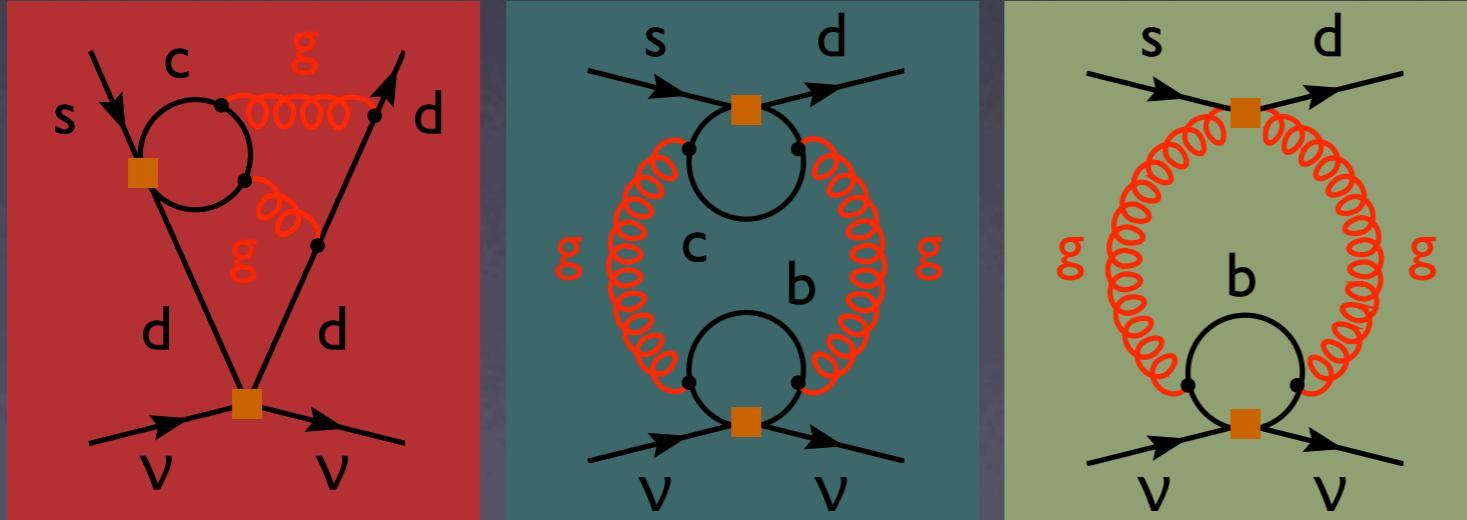
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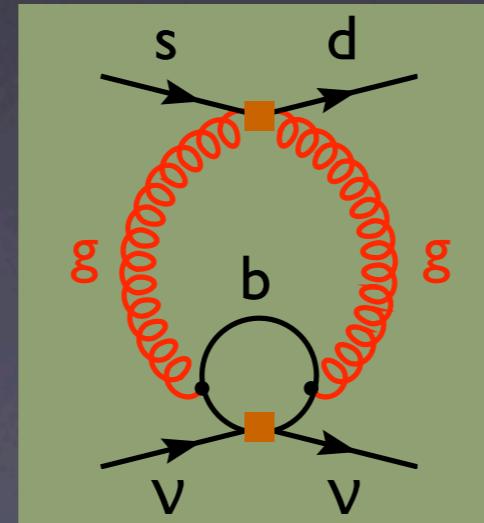
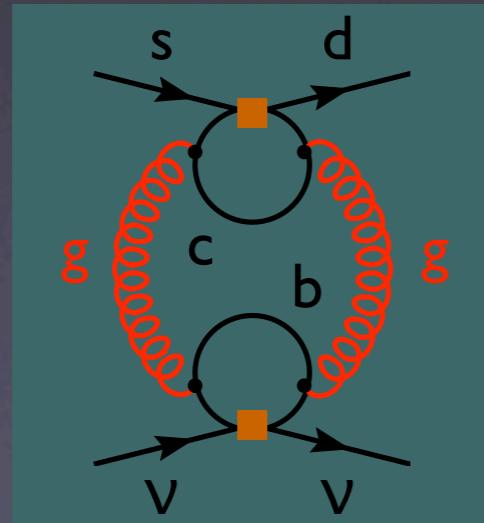
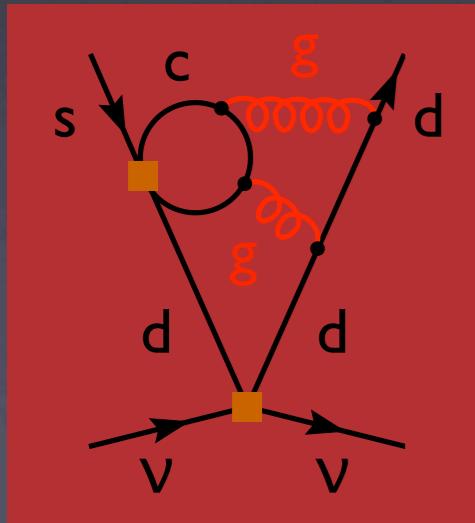
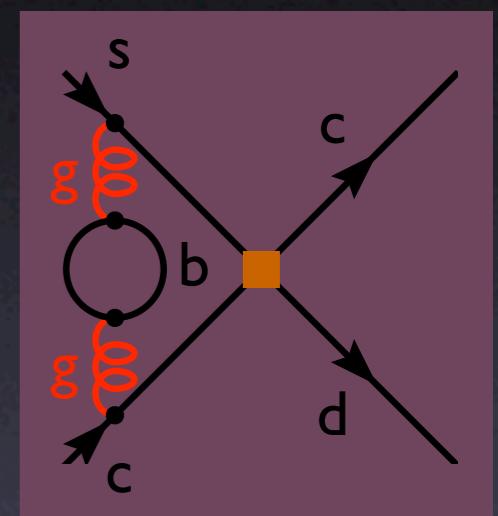
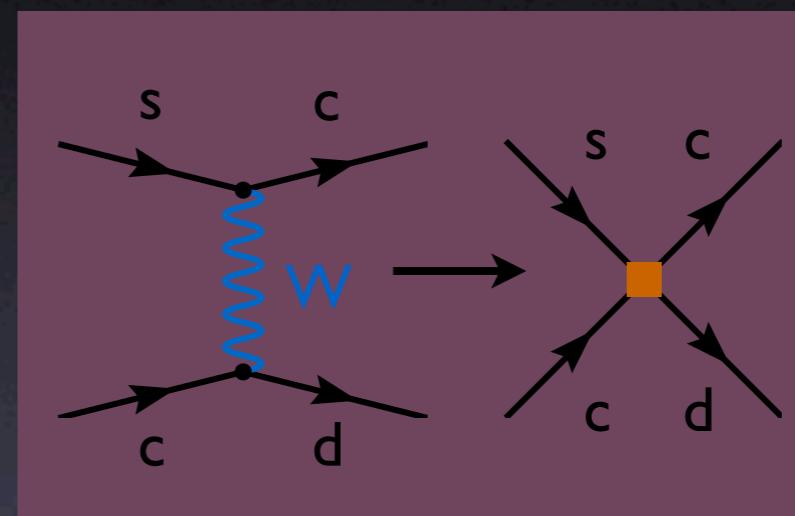
- vector part of neutral-current coupling starts to contribute
- contributions from axial-vector and Chern-Simons operator cancel exactly

New features in NNLO calculation of P_c

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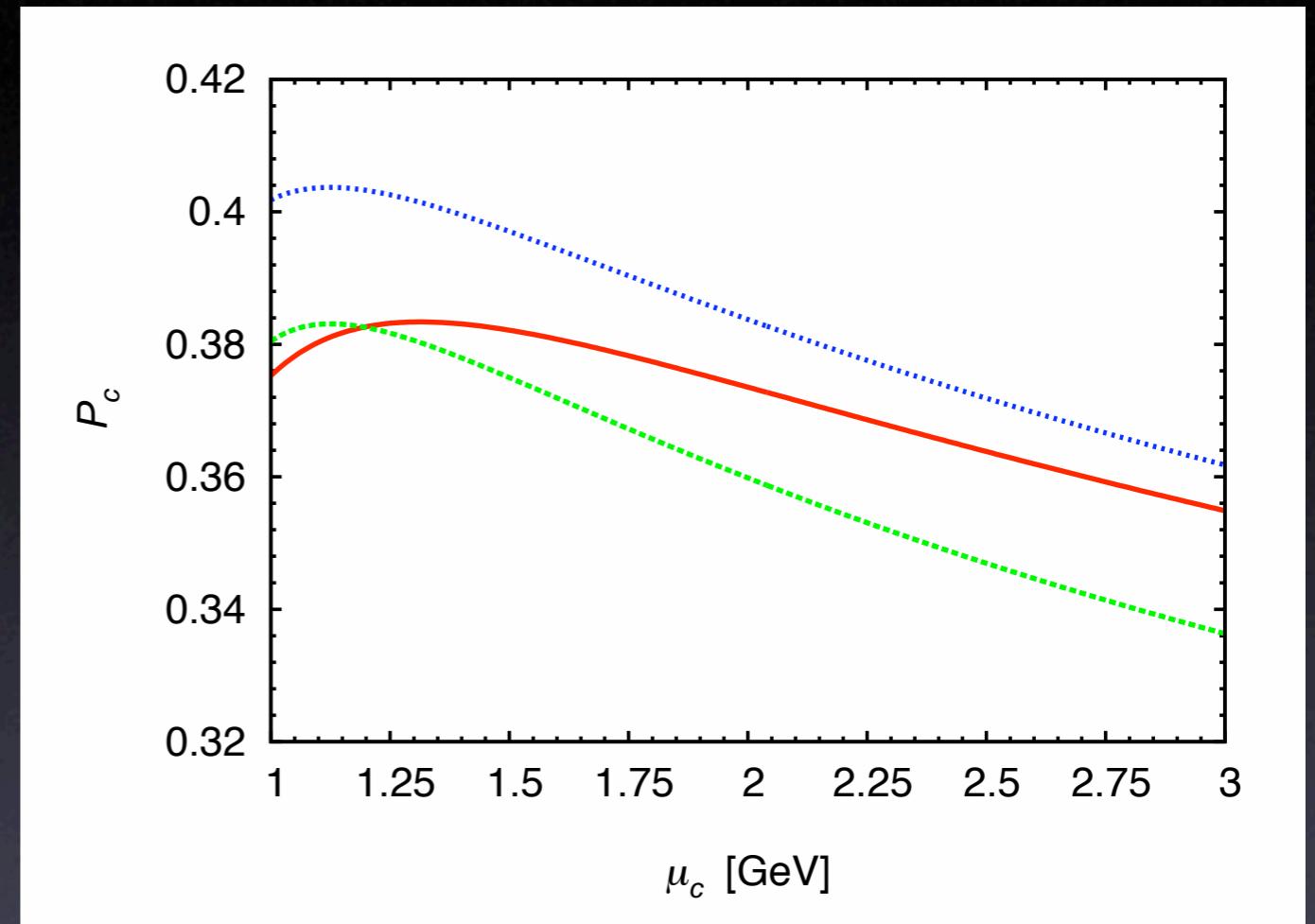
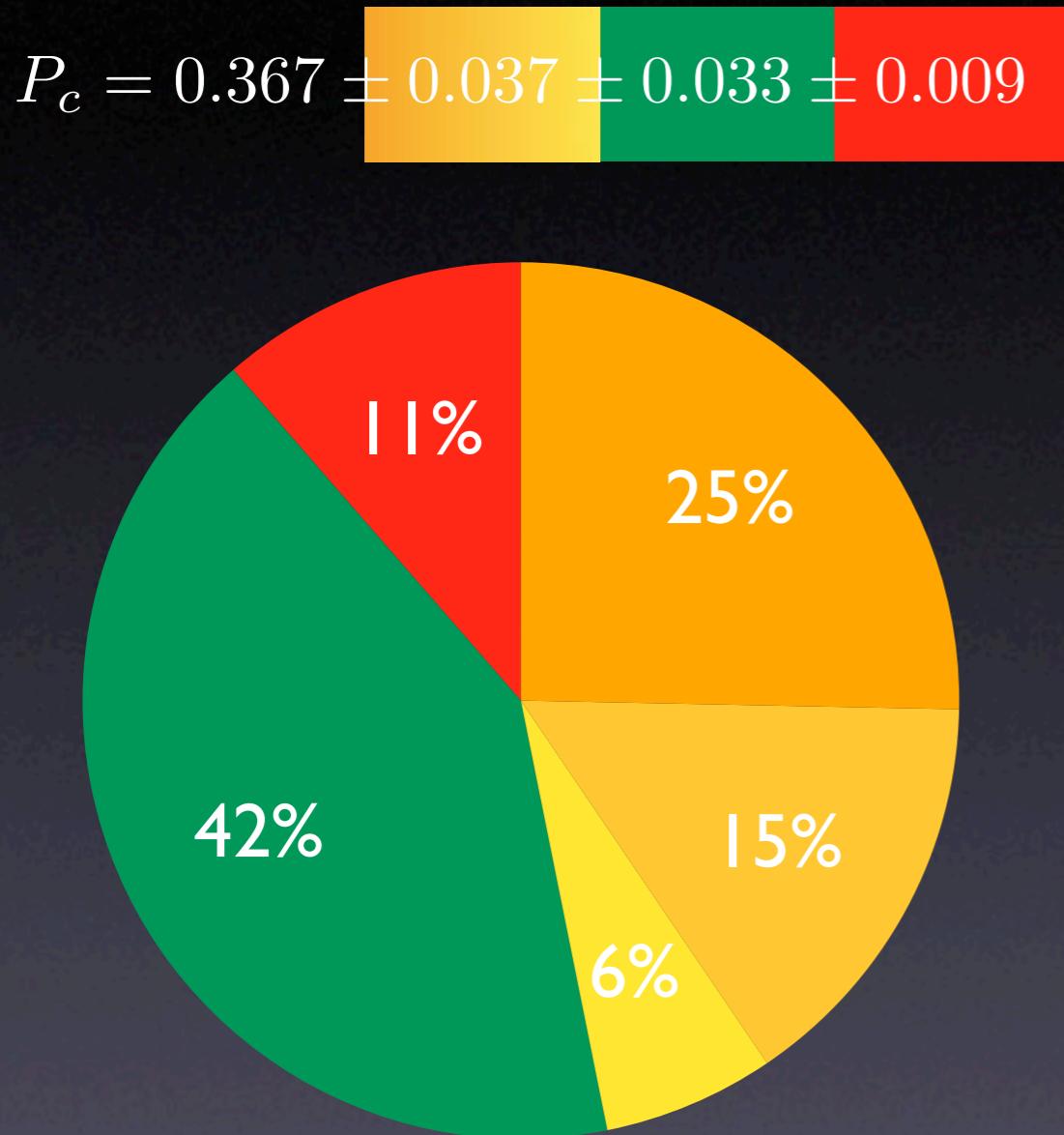
$$Q_1 = (\bar{s}_L \gamma_\mu c_L)(\bar{c}_L \gamma^\mu d_L)$$

$$Q_2 = (\bar{s}_L \gamma_\mu T^a c_L)(\bar{c}_L \gamma^\mu T^a d_L)$$



- non-trivial matching corrections at bottom quark threshold arise for current-current operators

Error budget of P_c and $B(K^+)$ at NLO

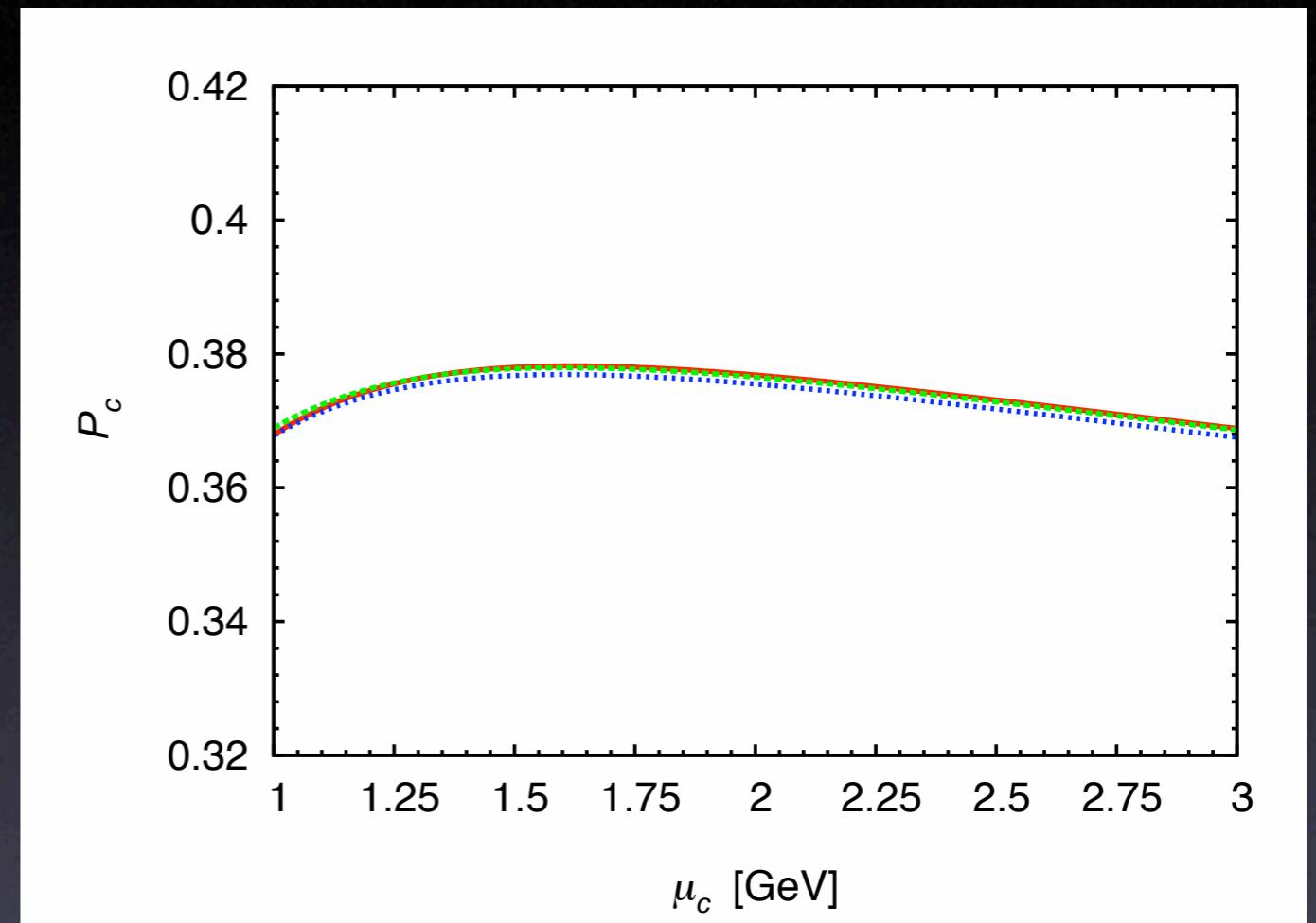
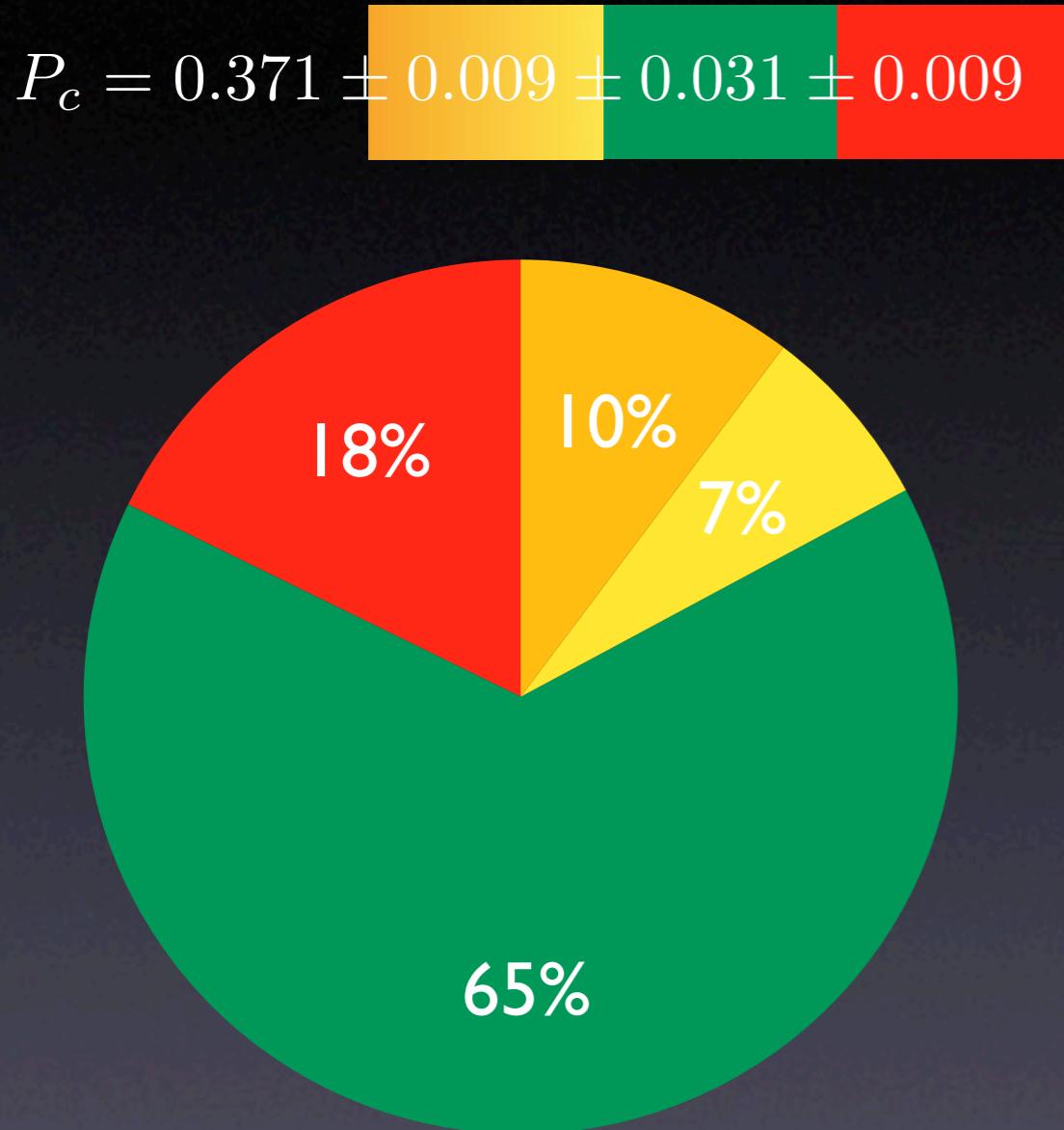


$$\mathcal{B}(K^+) = (7.93 \pm 0.77_{P_c} \pm 0.84_{\text{other}}) \times 10^{-11}$$

- μ_c
- $\alpha_s(\mu_c)$
- $\mu_b + \mu_W$
- $m_c(m_c)$
- $\alpha_s(M_Z)$

Buras, Gorbahn, Nierste & UH '05

Error budget of P_c and $B(K^+)$ at NNLO



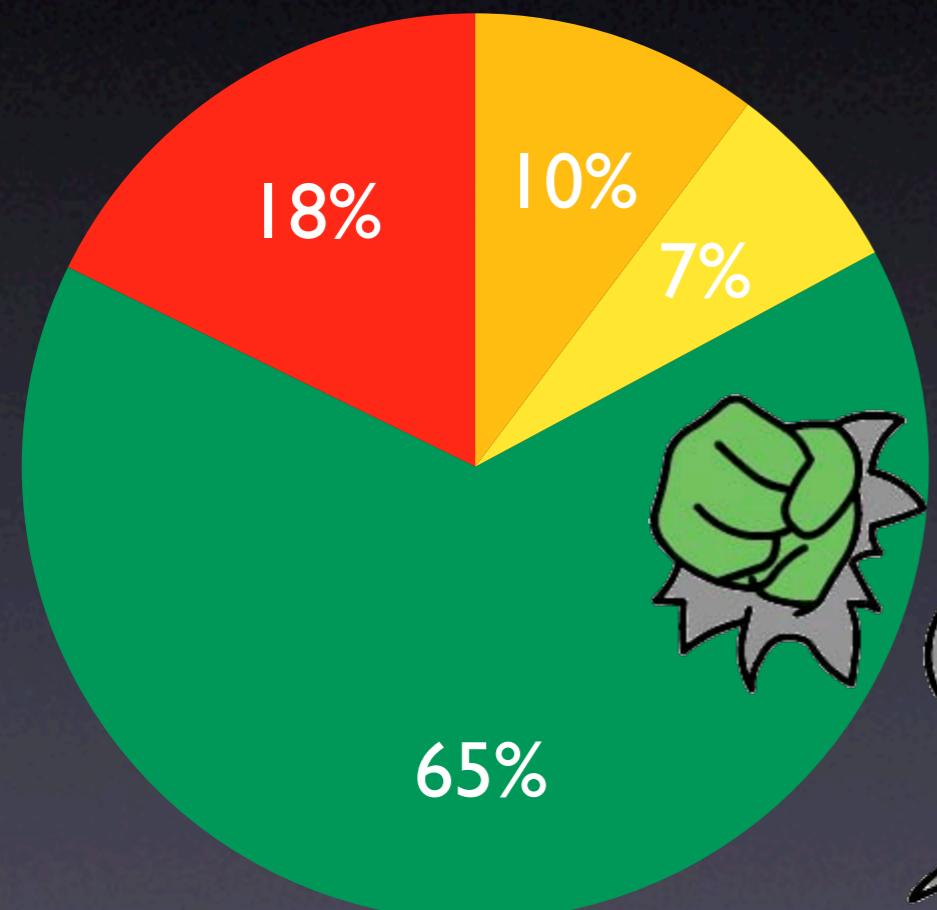
$$\mathcal{B}(K^+) = (7.96 \pm 0.49_{P_c} \pm 0.84_{\text{other}}) \times 10^{-11}$$

- $\mu_c + \alpha_s(\mu_c)$
- $m_c(m_c)$
- $\mu_b + \mu_W$
- $\alpha_s(M_Z)$

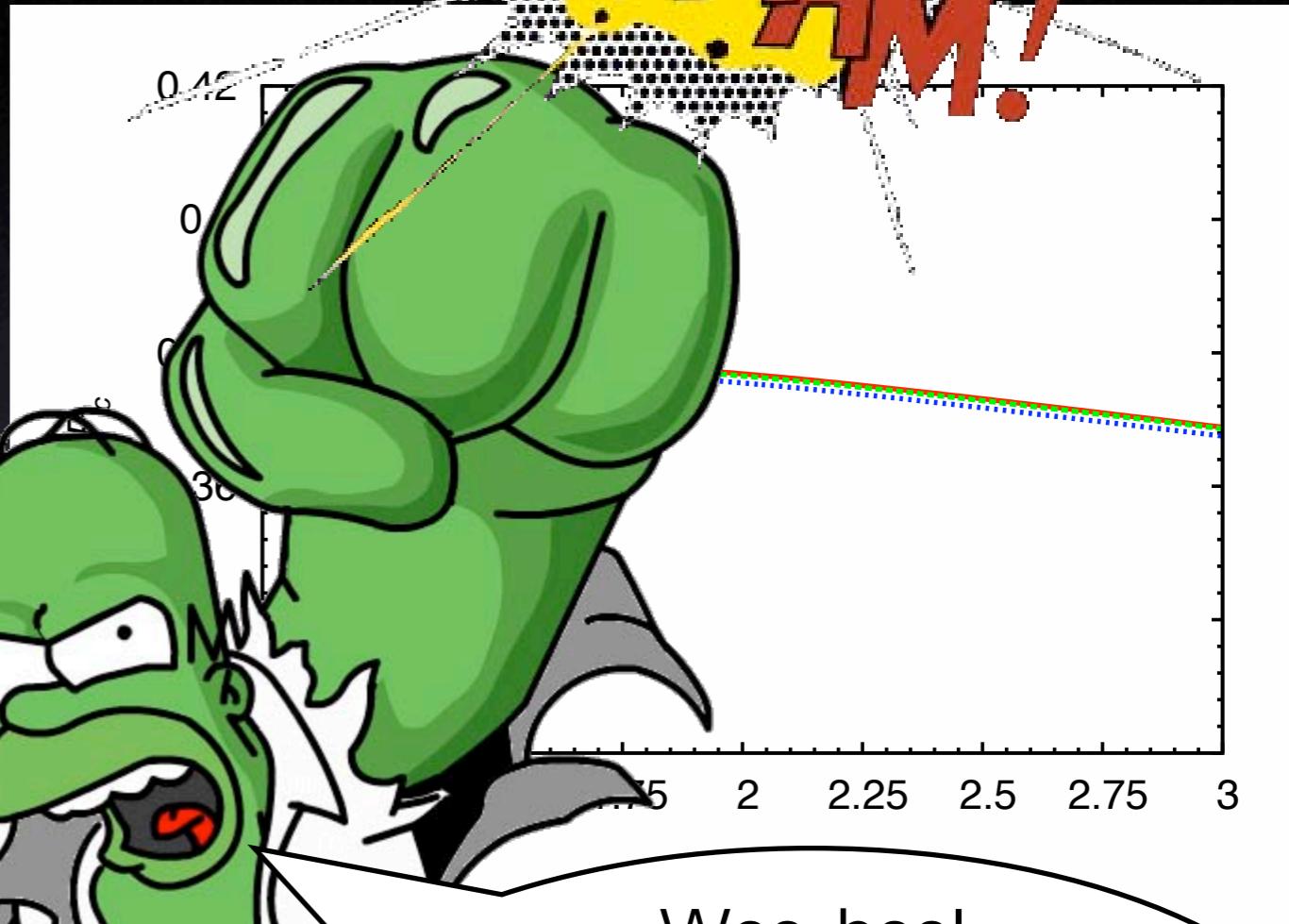
Buras, Gorbahn, Nierste & UH '05

Error budget of P_c and $B(K^+)$

$$P_c = 0.371 \pm 0.009 \pm 0.031 \pm 0.009$$



- $\mu_c + \alpha_s(\mu_c)$
- $m_c(m_c)$
- $\mu_b + \mu_W$
- $\alpha_s(M_Z)$



Woo-hoo!
NNLO calculation reduces
theoretical error of P_c and
 $B(K^+)$ by factor 4!

Unitarity triangle from $K \rightarrow \pi \nu \bar{\nu}$

d	s	b		
	$1 - \frac{\lambda^2}{2}$	λ	V_{ub}	u
	$-\lambda$	$1 - \frac{\lambda^2}{2}$	V_{cb}	c
	V_{td}	V_{ts}	1	t

$$A \approx 0.82$$

$$\bar{\rho} = \rho \left(1 - \frac{\lambda^2}{2} \right)$$

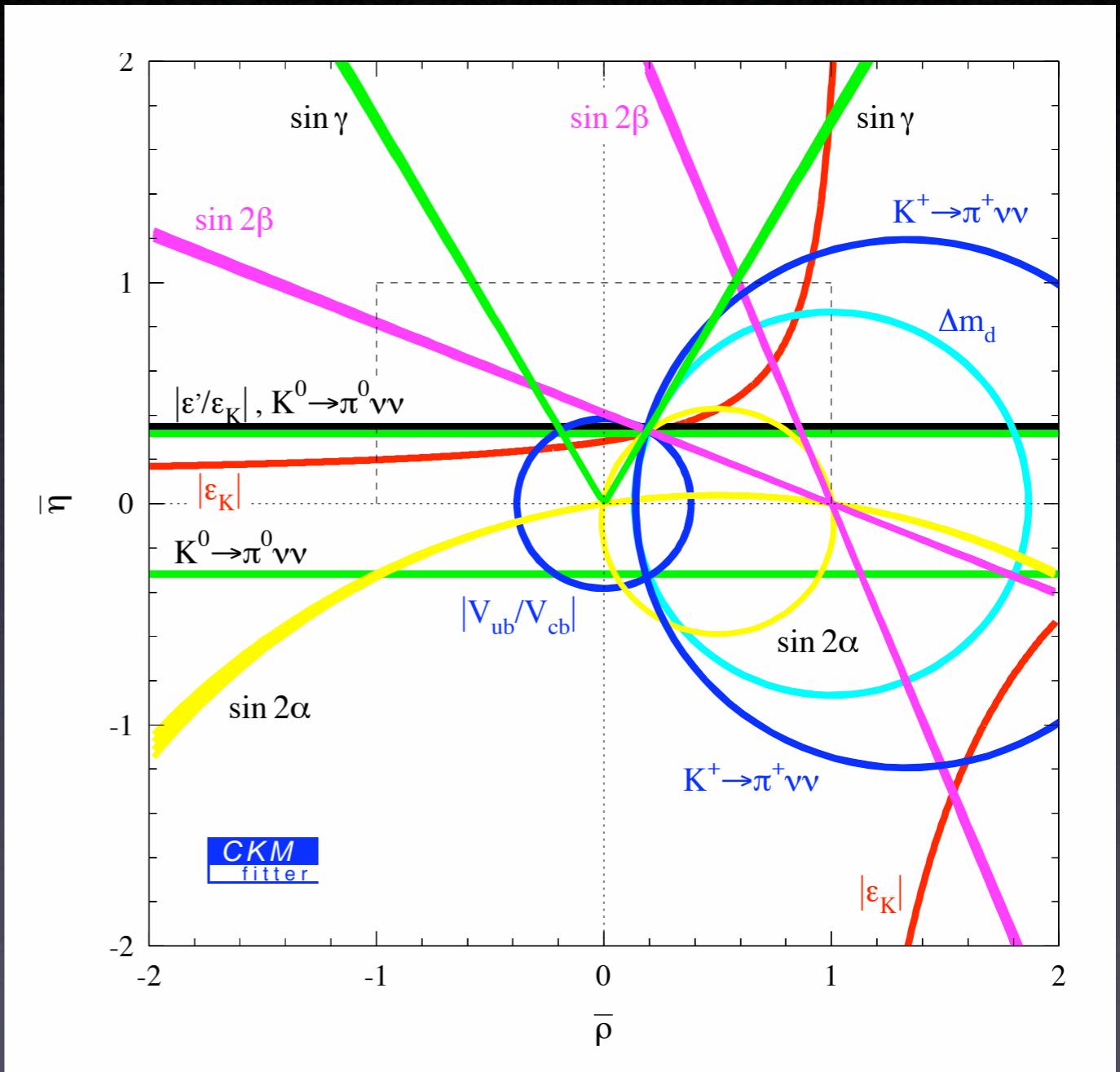
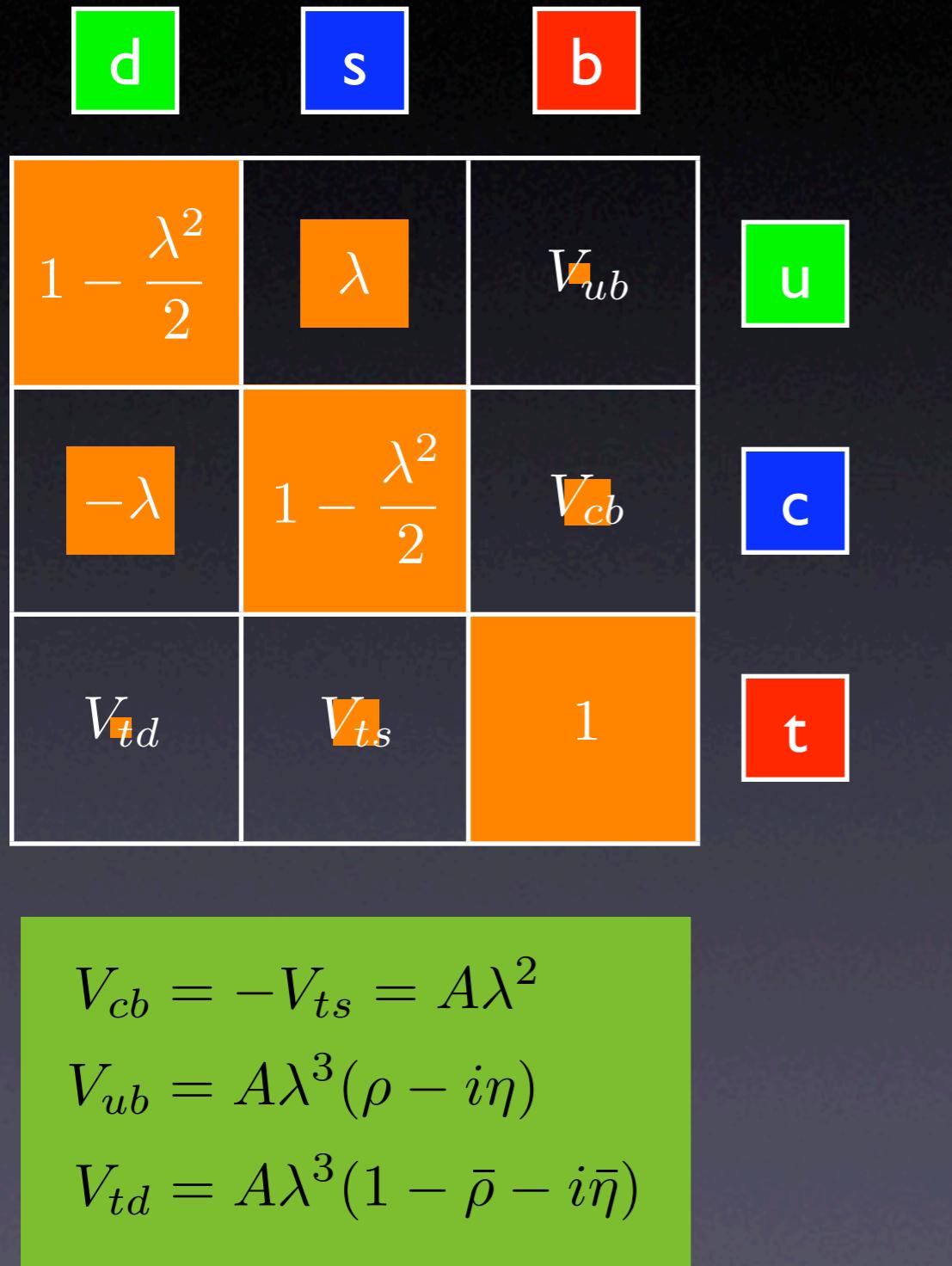
$$\bar{\eta} = \eta \left(1 - \frac{\lambda^2}{2} \right)$$

$$V_{cb} = -V_{ts} = A\lambda^2$$

$$V_{ub} = A\lambda^3(\rho - i\eta)$$

$$V_{td} = A\lambda^3(1 - \bar{\rho} - i\bar{\eta})$$

Unitarity triangle from $K \rightarrow \pi \nu \bar{\nu}$



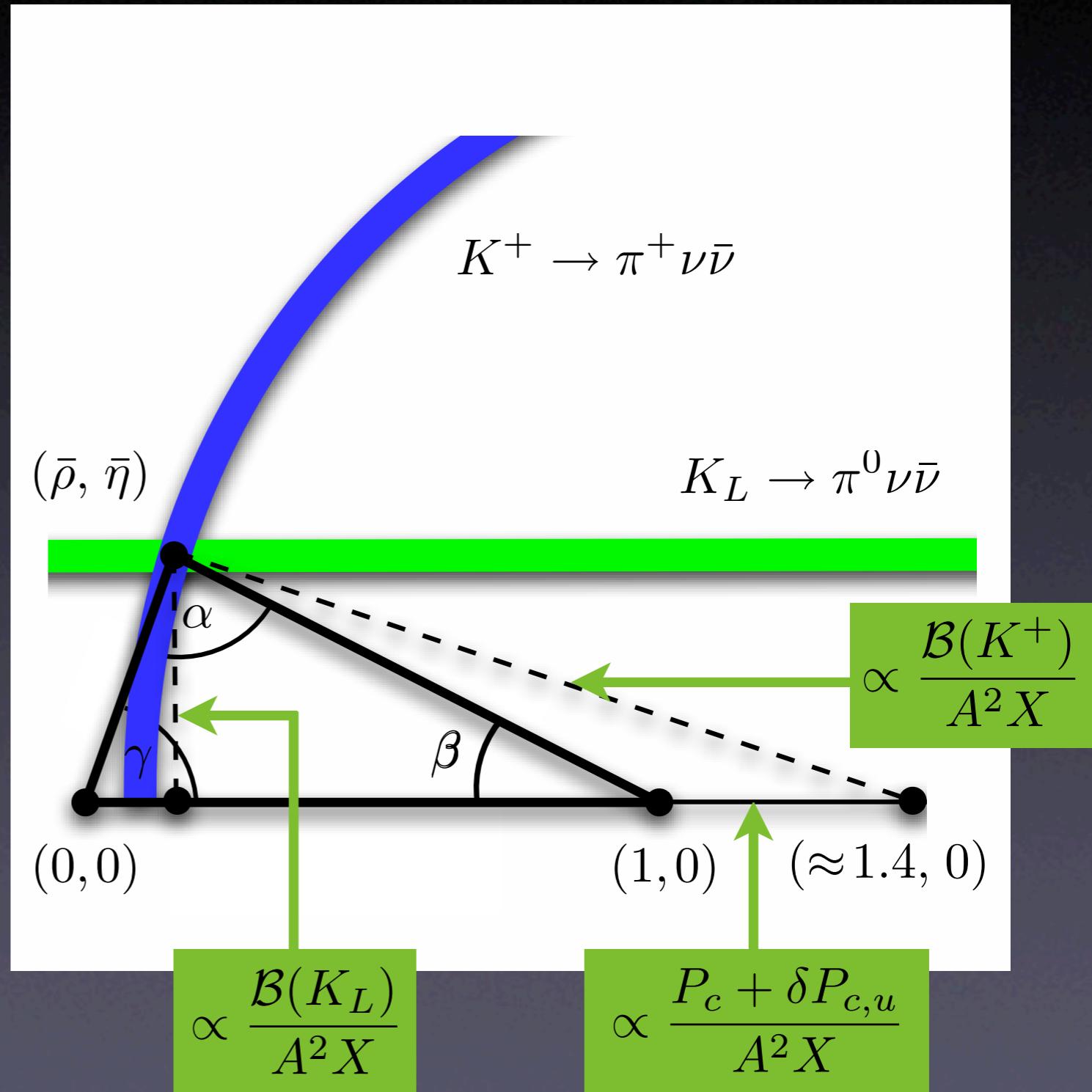
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d	s	b	
$1 - \frac{\lambda^2}{2}$	λ	V_{ub}	u
$-\lambda$	$1 - \frac{\lambda^2}{2}$	V_{cb}	c
V_{td}	V_{ts}	1	t

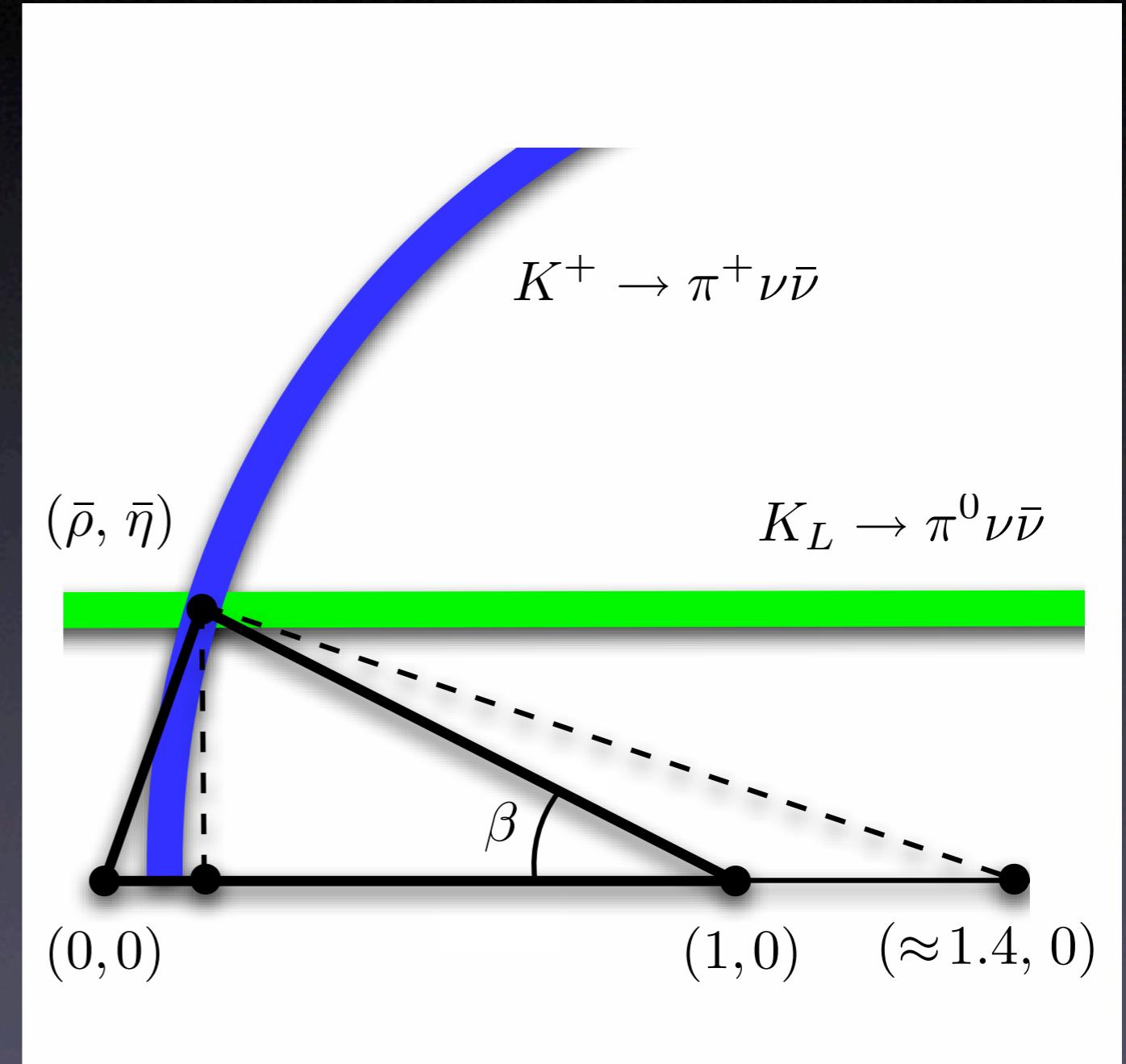
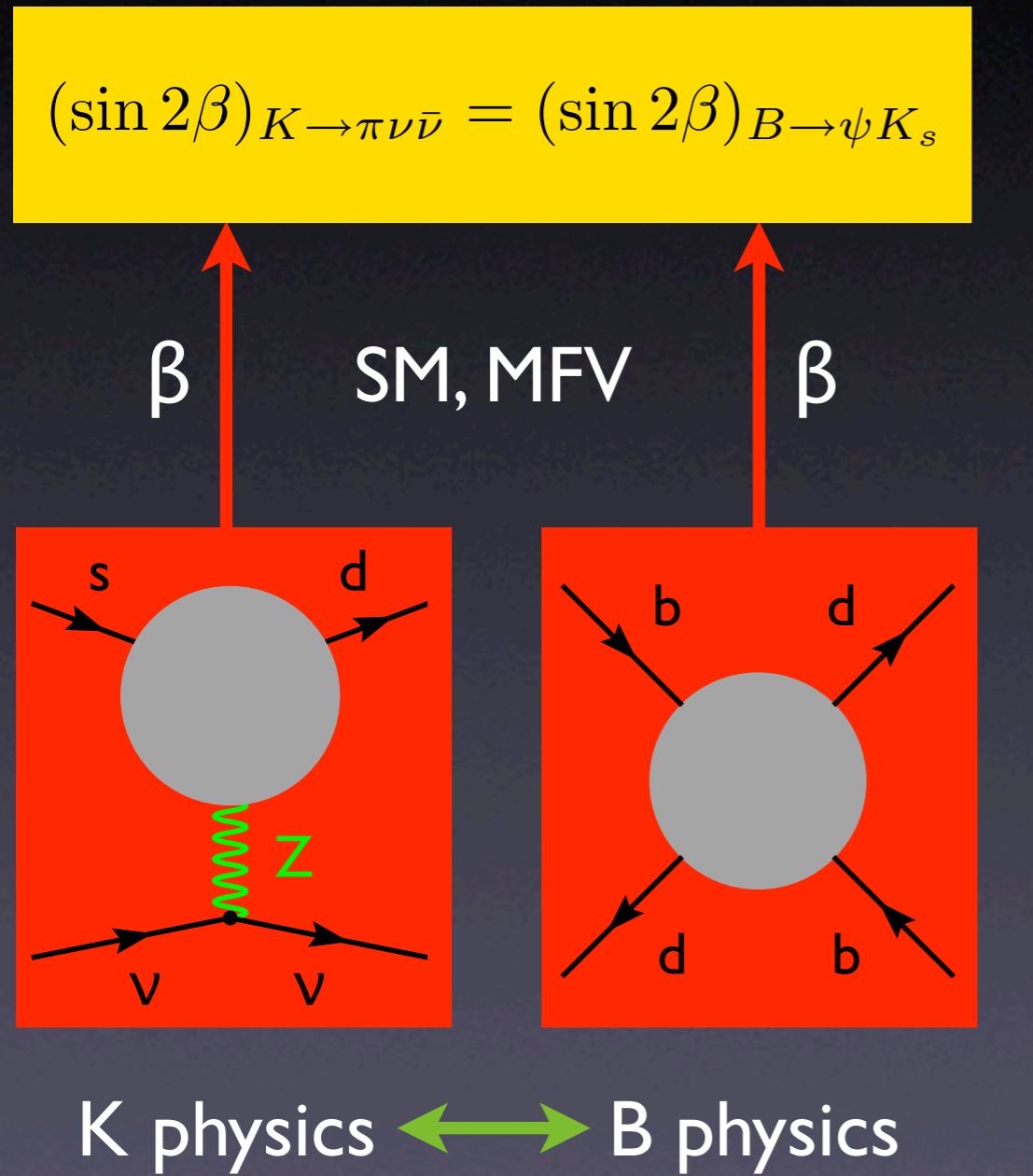
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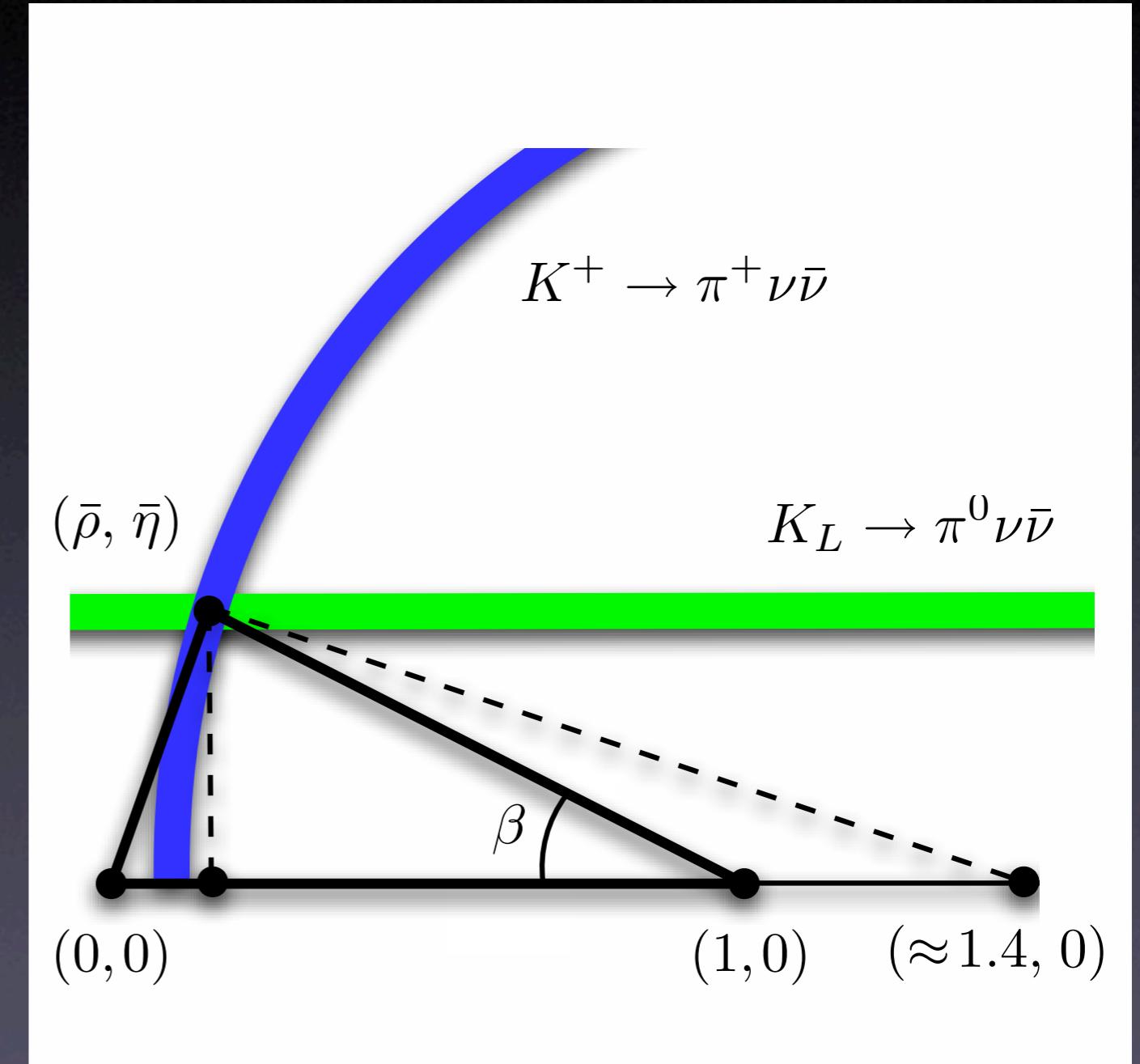
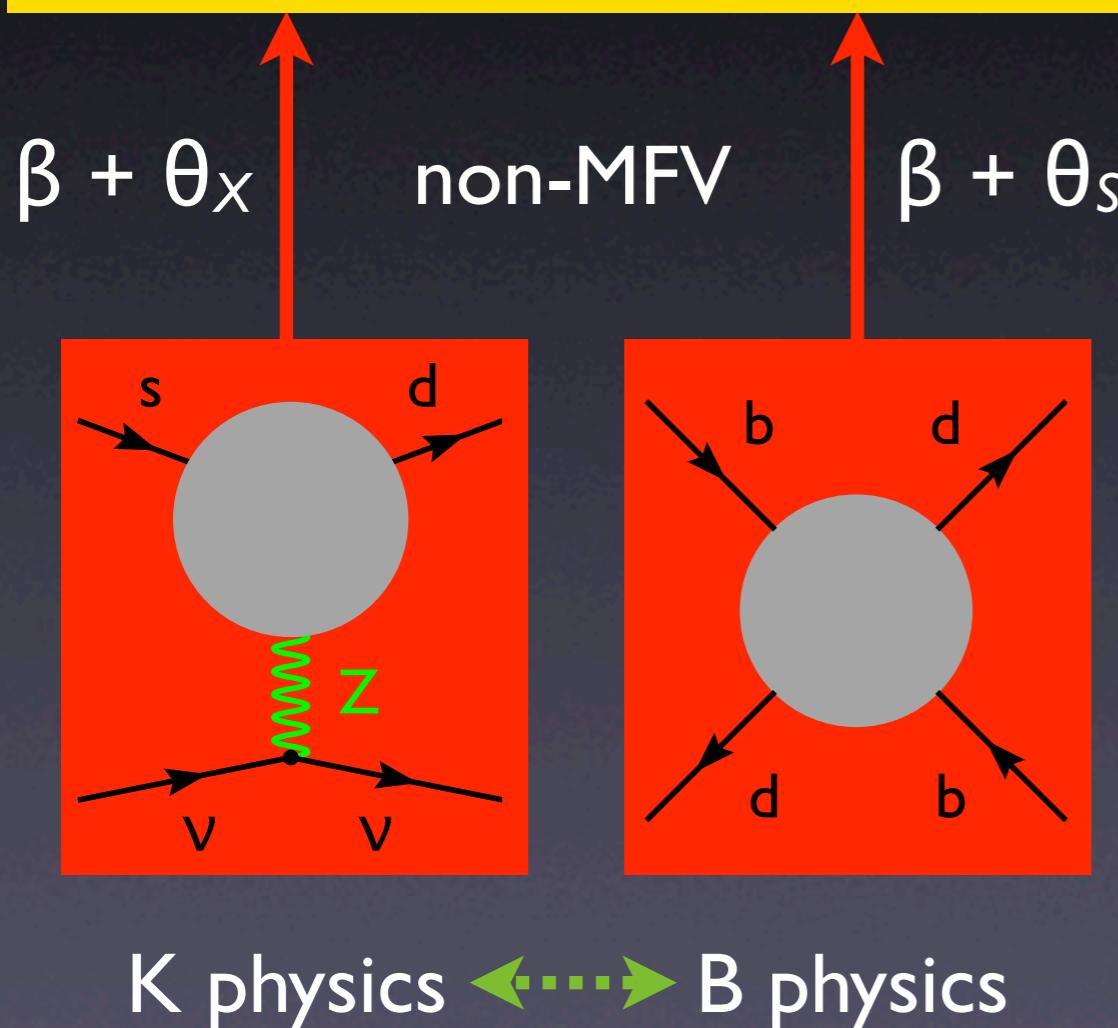


Unitarity triangle from $K \rightarrow \pi \nu \bar{\nu}$



Unitarity triangle from $K \rightarrow \pi \nu \bar{\nu}$

$$(\sin 2\beta_X)_{K \rightarrow \pi \nu \bar{\nu}} \neq (\sin 2\beta_S)_{B \rightarrow \psi K_S}$$



Unitarity triangle fit from $K \rightarrow \pi\nu\bar{\nu}$

$$\mathcal{B}(K^+) = (14.7^{+13.0}_{-8.9}) \times 10^{-11}$$

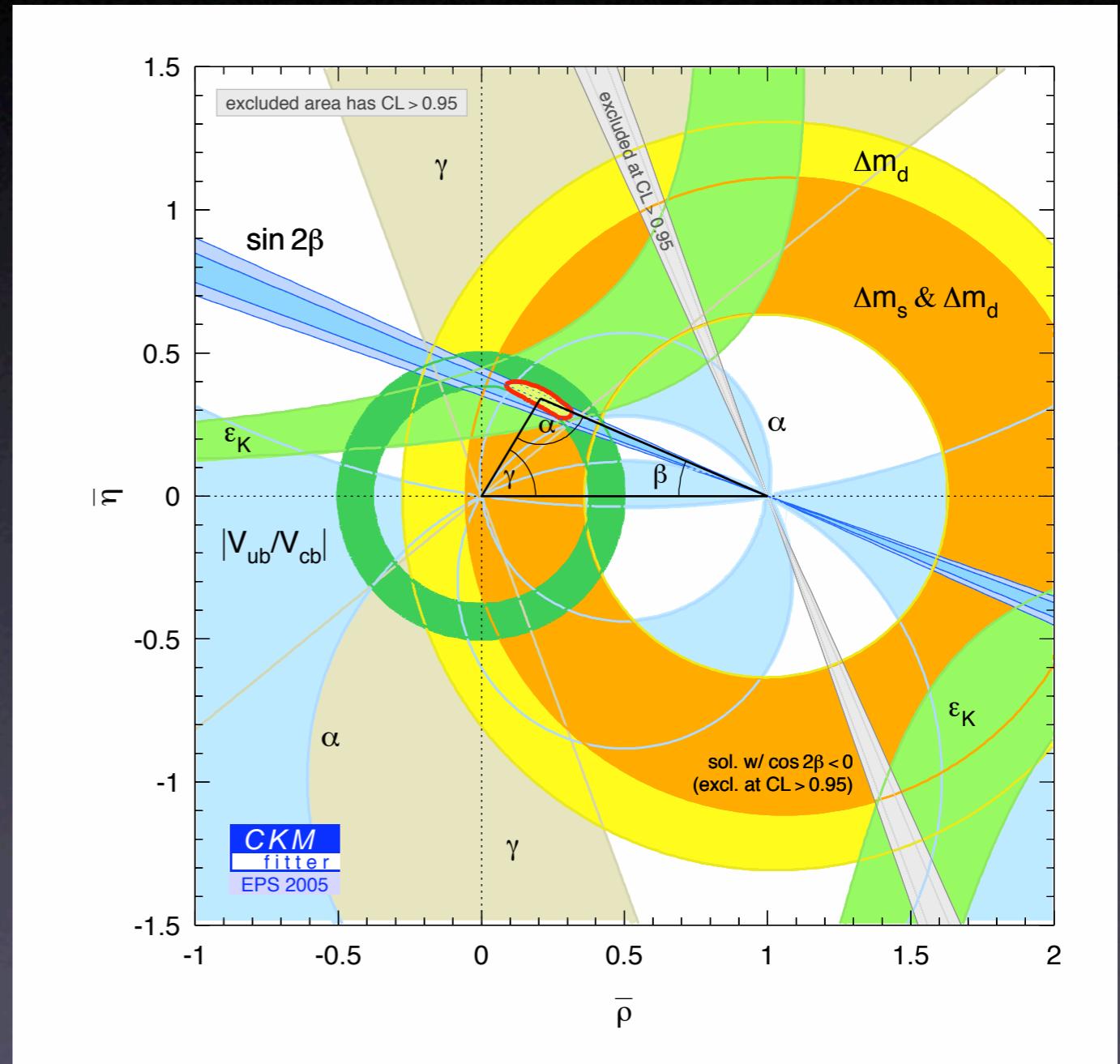
BNL AGS E787 & E949 '04

$$\mathcal{B}(K_L) < 5.9 \times 10^{-7} \text{ (90\% CL)}$$

FNAL KTeV E799-II '00

$$\mathcal{B}(K_L) < 2.86 \times 10^{-7} \text{ (90\% CL)}$$

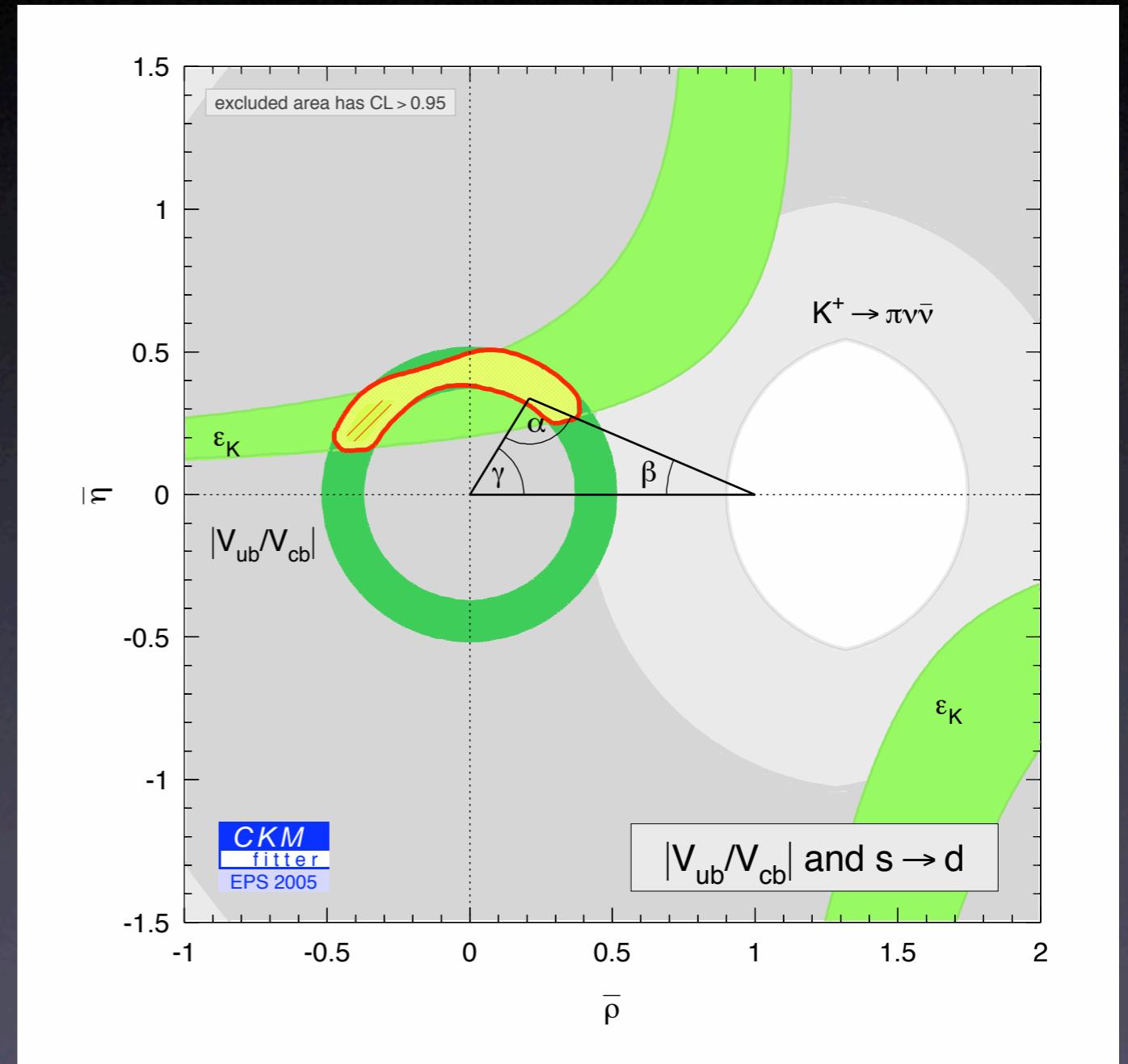
KEK PS E39 | a '05



Unitarity triangle fit from $K \rightarrow \pi\nu\bar{\nu}$

$$\mathcal{B}(K^+) = (14.7^{+13.0}_{-8.9}) \times 10^{-11}$$

BNL AGS E787 & E949 '04



Unitarity triangle fit from $K \rightarrow \pi\nu\bar{\nu}$

$$\mathcal{B}(K^+) = (7.50 \pm 0.75) \times 10^{-11}$$

$$\mathcal{B}(K_L) = (2.67 \pm 0.27) \times 10^{-11}$$

Future (?)

$$\frac{\sigma(|V_{td}|)}{|V_{td}|} = \pm 4.1\%$$

$$\sigma(\sin 2\beta) = \pm 0.025$$

$$\sigma(\gamma) = \pm 4.9^\circ$$

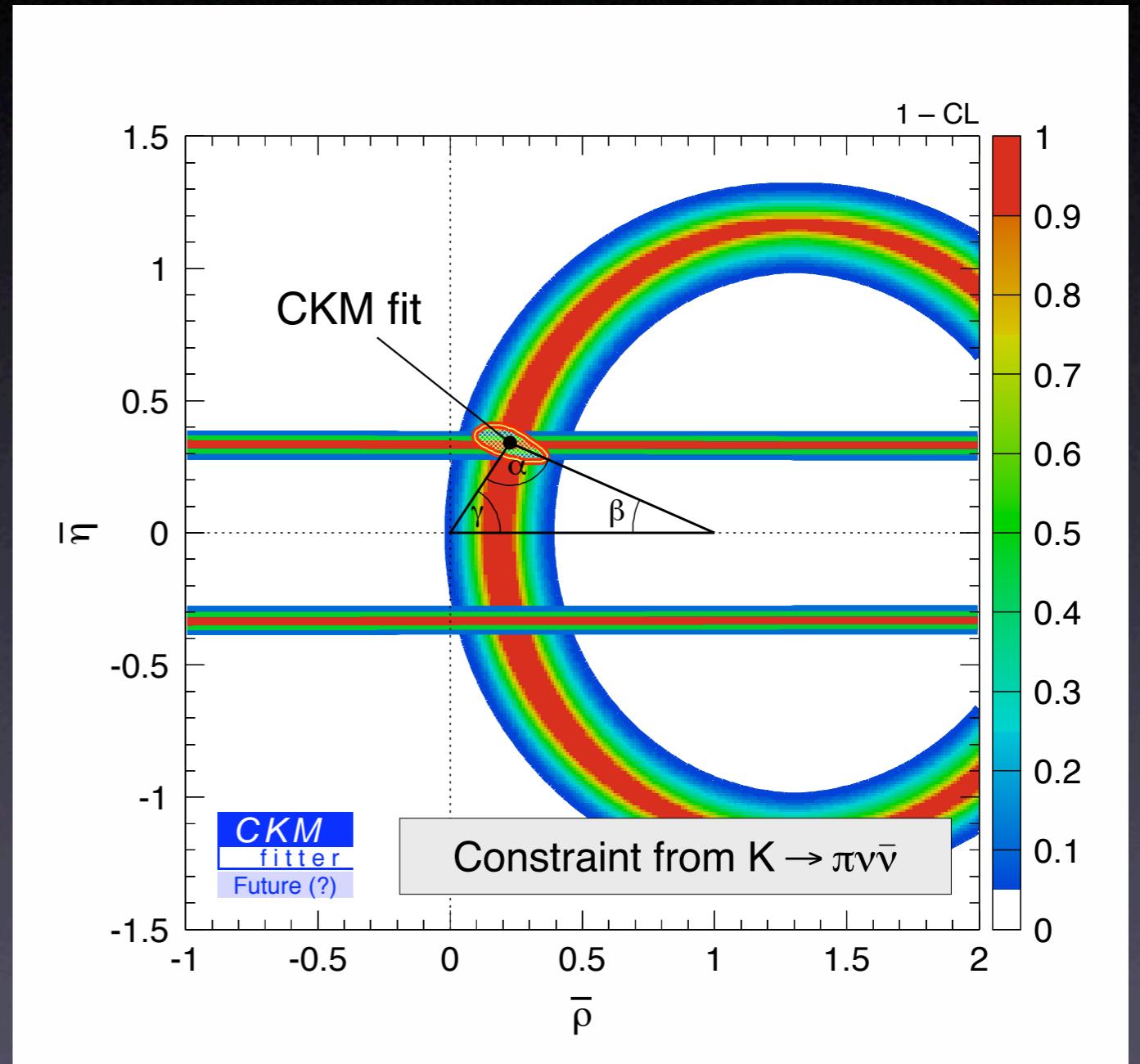
$$\frac{\sigma(|V_{td}|)}{|V_{td}|} = \pm 1.0\%$$

$$\sigma(\sin 2\beta) = \pm 0.006$$

$$\sigma(\gamma) = \pm 1.2^\circ$$

NLO
(theory error only)

NNLO
(theory error only)



Unitarity triangle fit from $K \rightarrow \pi\nu\bar{\nu}$

$$\mathcal{B}(K^+) = (7.50 \pm 0.75) \times 10^{-11}$$

$$\mathcal{B}(K_L) = (2.67 \pm 0.27) \times 10^{-11}$$

Future (?)

$$\frac{\sigma(|V_{td}|)}{|V_{td}|} = \pm 1.0\%$$

$$\sigma(\sin 2\beta) = \pm 0.006$$

$$\sigma(\gamma) = \pm 1.2^\circ$$

$$\frac{\sigma(|V_{td}|)}{|V_{td}|} = \pm 8.4\%$$

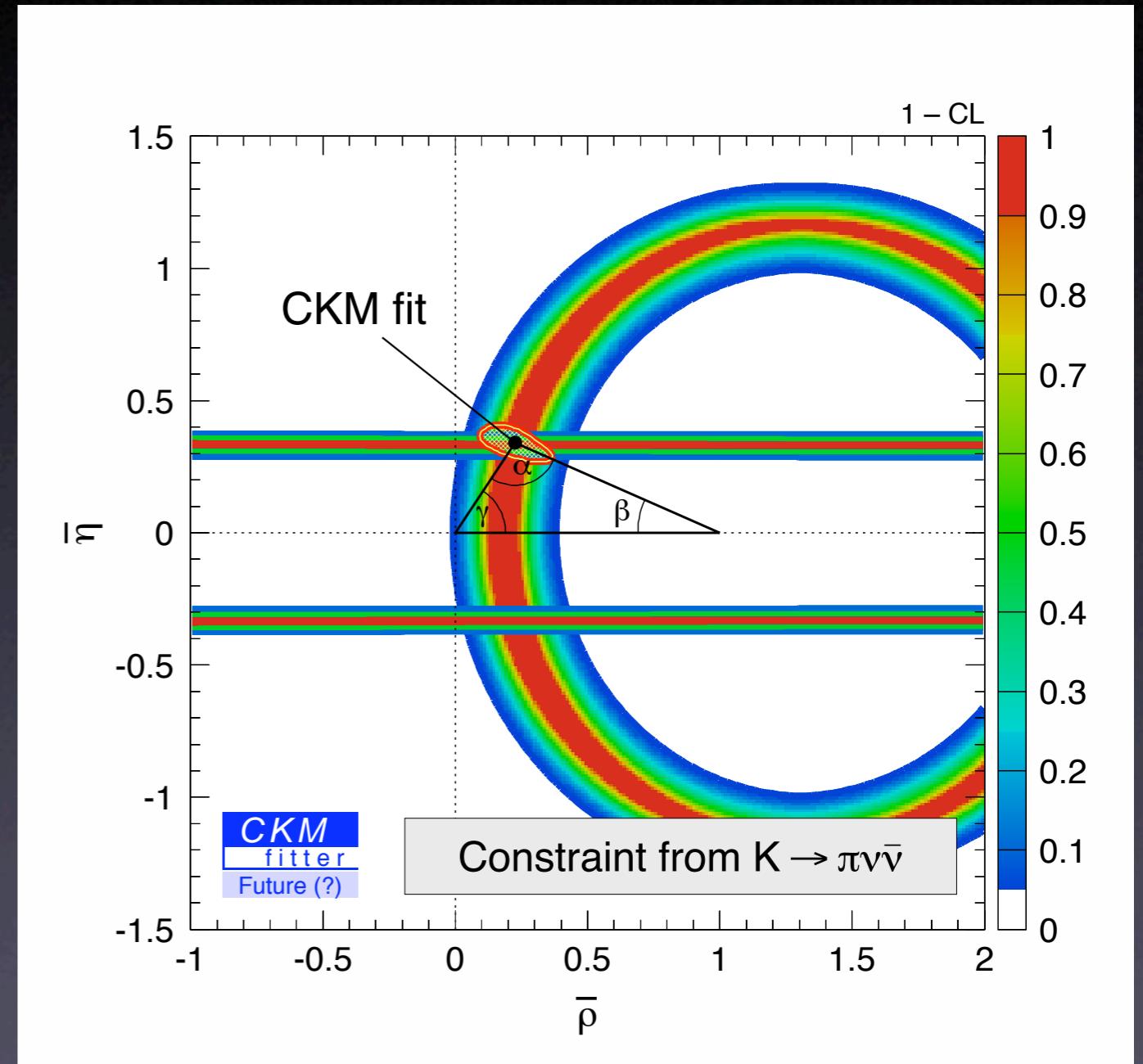
$$\sigma(\sin 2\beta) = \pm 0.056$$

$$\sigma(\gamma) = \pm 11^\circ$$

Future (?)

NNLO
(theory error only)

NNLO
(all uncertainties)



Conclusions

- SD dominated rare K^+ and K_L decays offer powerful and complementary test of flavor sector of SM
- NNLO calculation of charm contribution to K^+ is now available
- there is sizeable room for NP in this golden modes and their clean theoretical character remains valid in essential all extensions of SM
- measurements of branching ratios at $\approx 10\%$ would substantially improve our understanding of flavor dynamics at TeV scale



Credits

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