

# Measurement of the Muon Anomalous Magnetic Moment to 0.7 ppm at BNL

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on behalf of the g-2 Collaboration

Results from the Data of 2000

# Muon g-2 Collaboration

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● *Spokesperson*

● *Project Manager*

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# Outline

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- Motivation of BNL  $g-2$  Experiment
- Method
- Experiment
- 2000 Data Analysis
- Results and Summary

# Motivation of BNL $g-2$ Experiment

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- Precision Test of the Standard Model

ElectroWeak Contributions  
Muon Substructure,  
W Compositeness and Anomalous  
Couplings, etc...

- Probe Physics Beyond the Standard Model

Supersymmetry, etc..

# g-factor

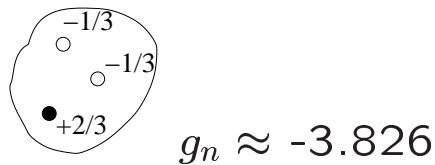
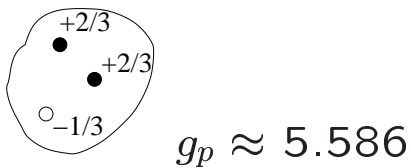
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A particle with spin has a magnetic moment :

$$\vec{\mu} = g \frac{e\hbar}{2mc} \vec{S}$$

1930's : For Point particles with spin  $\frac{1}{2}$ ,  $g=2$  (Dirac)

**Hadrons:**



**Leptons:**  $g=2$  ?

1940's :  $g_e \approx 2.002$  (Lamb shift), later on  $g_\mu \approx 2.002$

Anomalous magnetic moment :  $a = \frac{g-2}{2}$

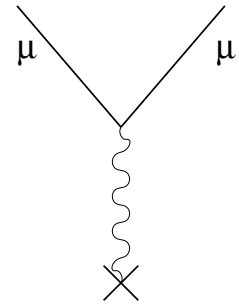
$g \neq 2$  (QED)  $a = \frac{\alpha}{2\pi} + \dots$

# Forces

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Dirac Particles

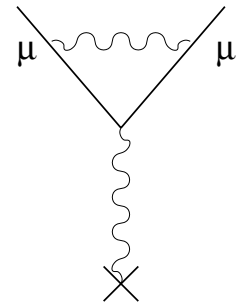
$$g = 2$$



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Electromagnetic (QED)

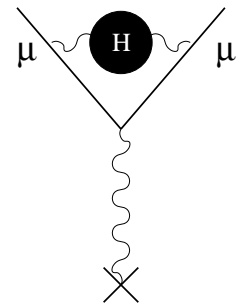
$$\Delta g = 0.00233169411(6)$$



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Strong Force (Hadronic)

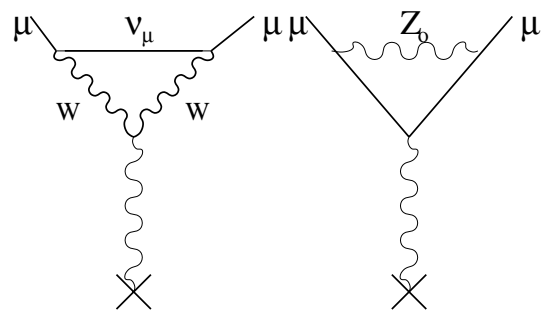
$$\Delta g = 0.00000013818(70)$$



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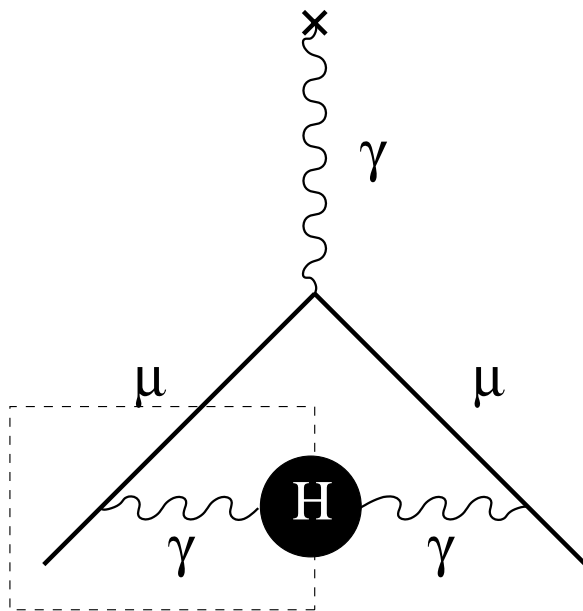
Weak Force

$$\Delta g = 0.00000000302(8)$$

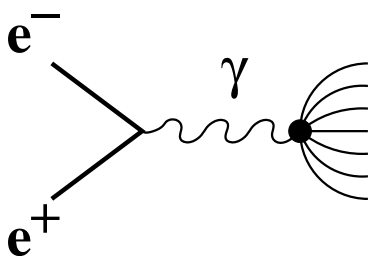


# Hadronic Vacuum Polarization

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HADRONIC CONTENT of the PHOTON



# Hadronic Contribution

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Dispersion Relation :

$$a_{\mu}^{had} = \frac{\alpha^2(0)}{3\pi^2} \int_{4m_{\pi}^2}^{\infty} ds \frac{K(s)}{s^2} R(s)$$

s : square of center of mass energy  
K(s) : smooth weight function (0.63-1)  
strong weight to low energy  
 $\sqrt{s} < 1.8$  GeV, 92% of the  $a_{\mu}$ )

$$R(s) = \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)}$$



# Method

In the presence of the vertical field,  $\mathbf{B}$   
for  $g=2$

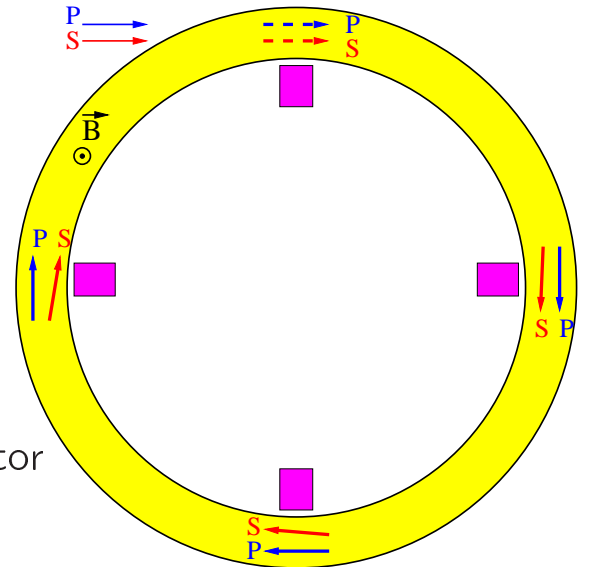
$$\omega_c = \frac{eB}{m_\mu c \gamma} = \frac{2\mu_B B}{\hbar \gamma} \quad \omega_c = \omega_s$$

Due to anomaly on the magnetic moment,  
Spin precess faster than the momentum

$$\omega_s = \frac{eB}{m_\mu c \gamma} (1 + a_\mu \gamma)$$

Spin precession relative to momentum vector

$$\omega_a = \omega_s - \omega_c = \frac{e}{m_\mu c} a_\mu B$$



Plus weak focusing electric field  $\vec{E}$  :

$$\vec{\omega}_a = \frac{d\vec{\vartheta}_a}{dt} = -\frac{e}{m_\mu c} \left[ a_\mu \vec{B} - \left( a_\mu - \frac{1}{\gamma^2 - 1} \right) \vec{\beta} \times \vec{E} \right]$$

Choose  $\gamma = 29.3$ , so the contribution of  $\vec{\beta} \times \vec{E}$  is removed.  
 $\Rightarrow P_\mu = 3.1\text{GeV}$

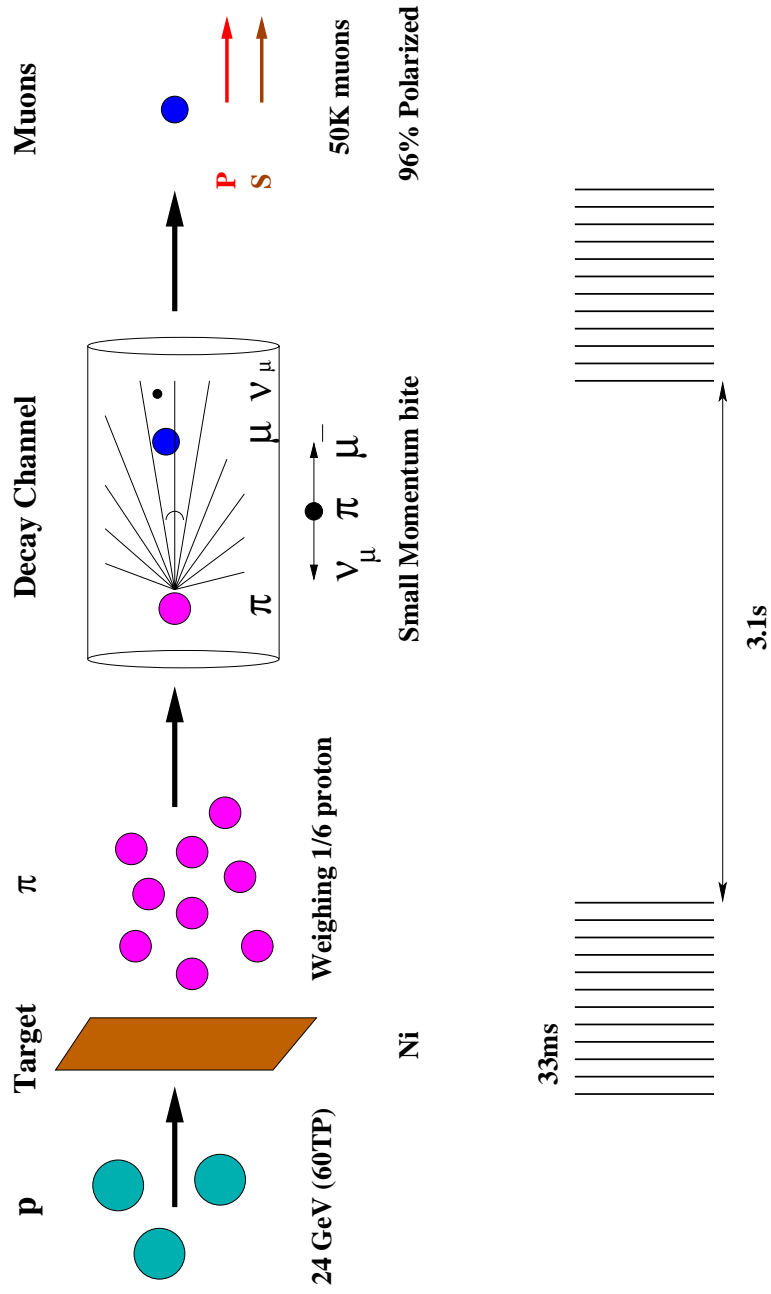
$\Rightarrow$  Measure  $|\vec{B}|$  and  $\omega_a$  and calculate  $a_\mu$  !

# Why This Experiment was Possible

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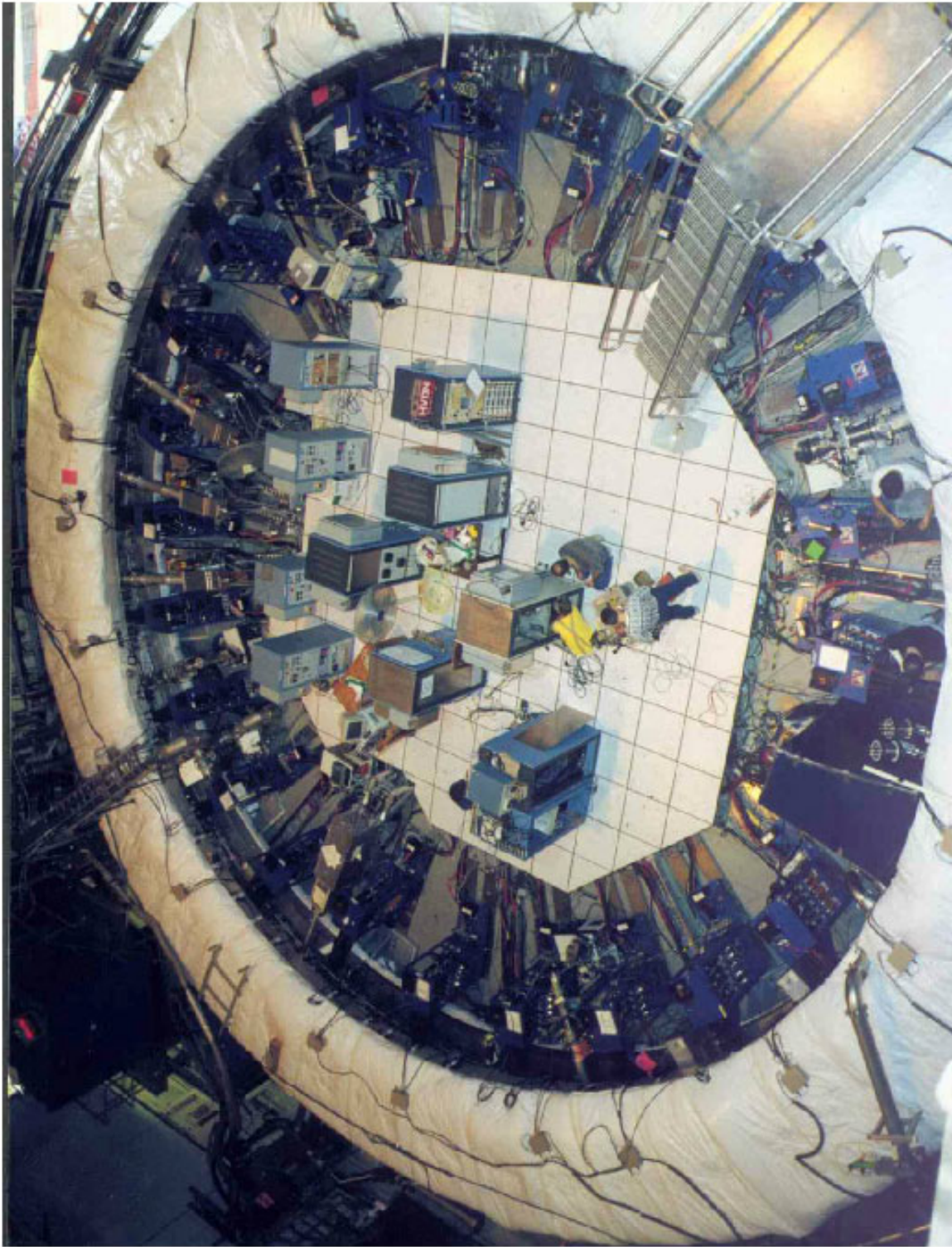
- AGS, the Highest Intensity Proton Source,
- Store Directly Polarized Muons,
- The Largest Single-Coiled Superconducting Magnet ( $r=7.11\text{m}$ ), Provides 1.45T Uniform Magnetic Field.

# The Beam



# The g-2 Ring

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# Magnet

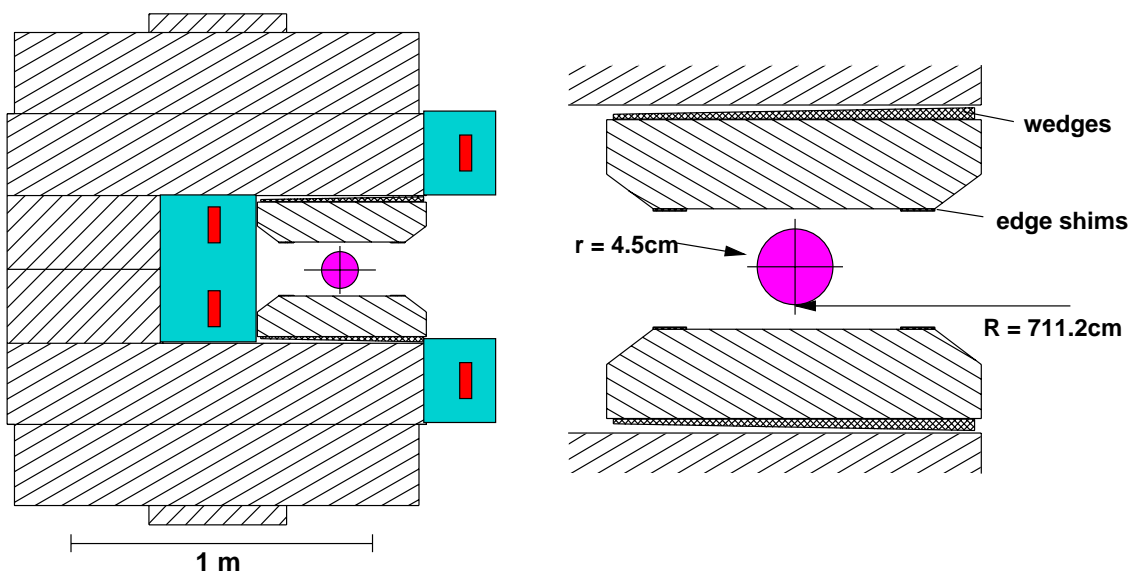
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1.45 Tesla uniform field

7.112m radius

700 Ton

-  Cryostat
-  Superconducting coils
-  Storage Region



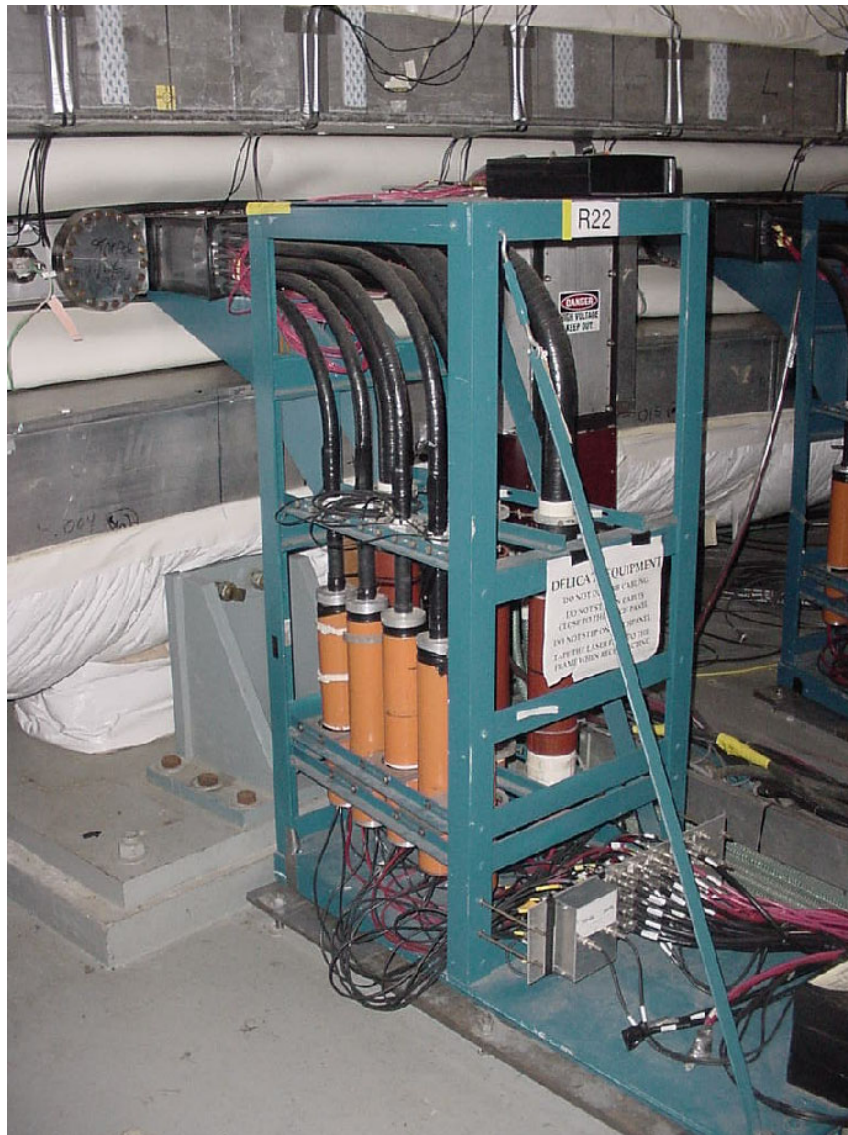


# Detectors

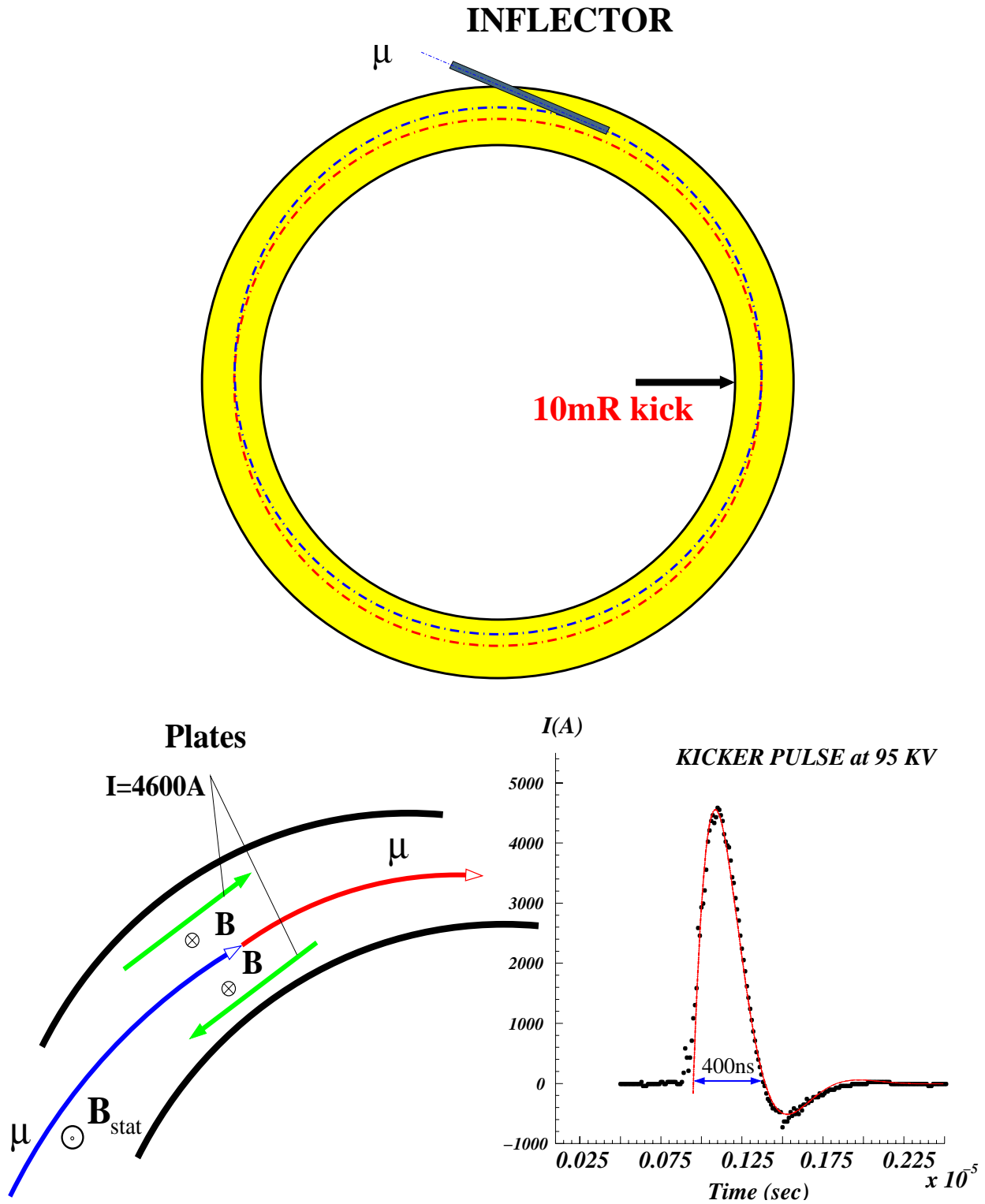
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24 Pb/Scintillating fiber electromagnetic calorimeters

Energy resolution of  $\sigma \approx 10\% \sqrt{E(\text{GeV})}$

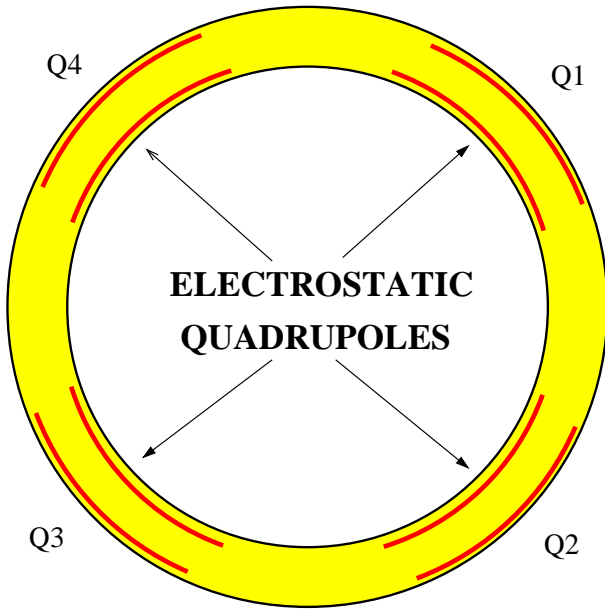


# Inflector and Muon Kicker

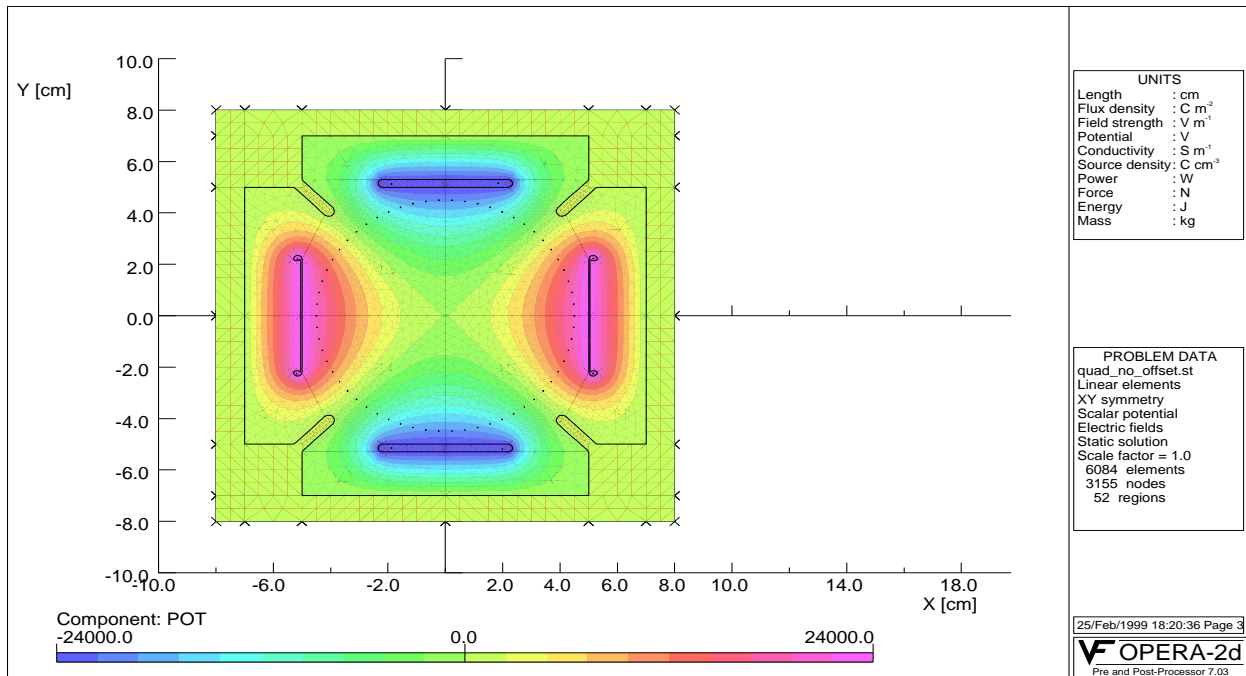
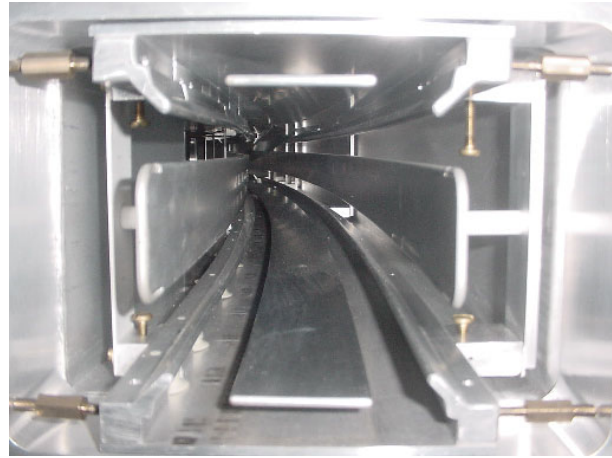


# Electrostatic Quadrupoles

The Weak Focusing Provided by



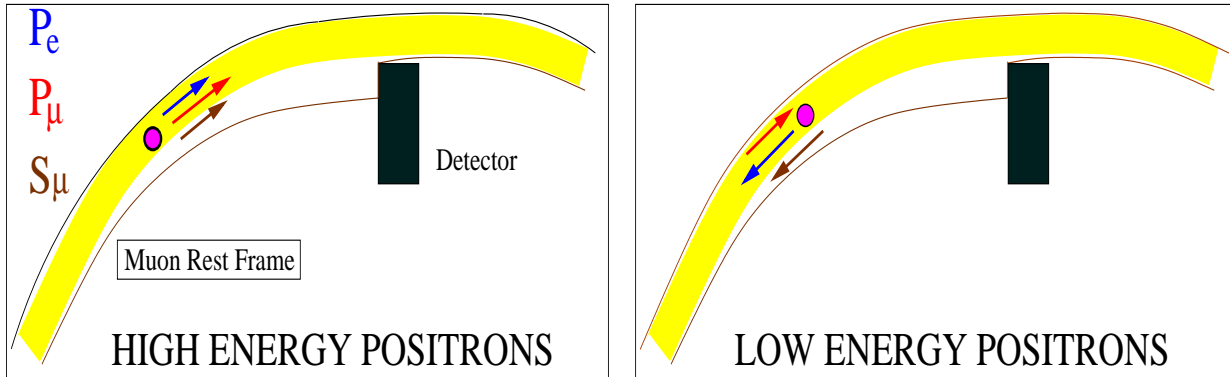
43% of circumference of the Ring





# Measurement of $\omega_a$

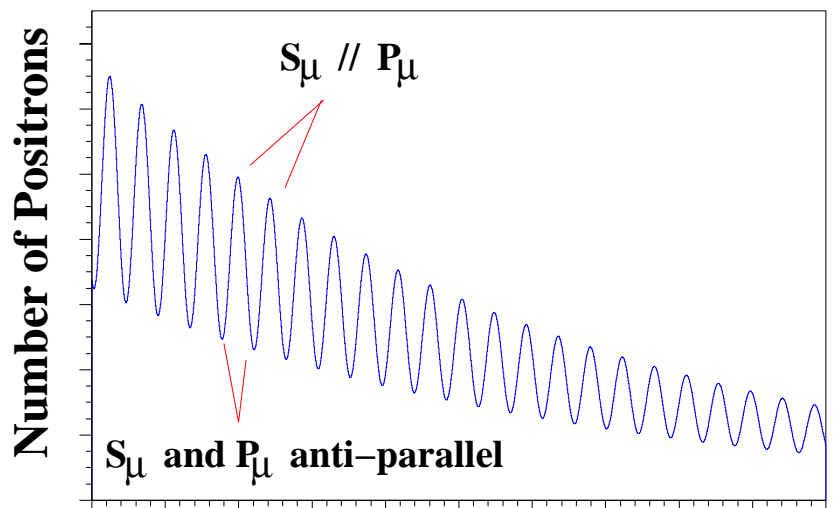
Parity violating 3 body decay :  $\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$



$S_\mu$  is precessing in the field  $\rightarrow e^+$  counting rate is modulated with the  $\mu$  spin precession

Decay Positron Spectrum :

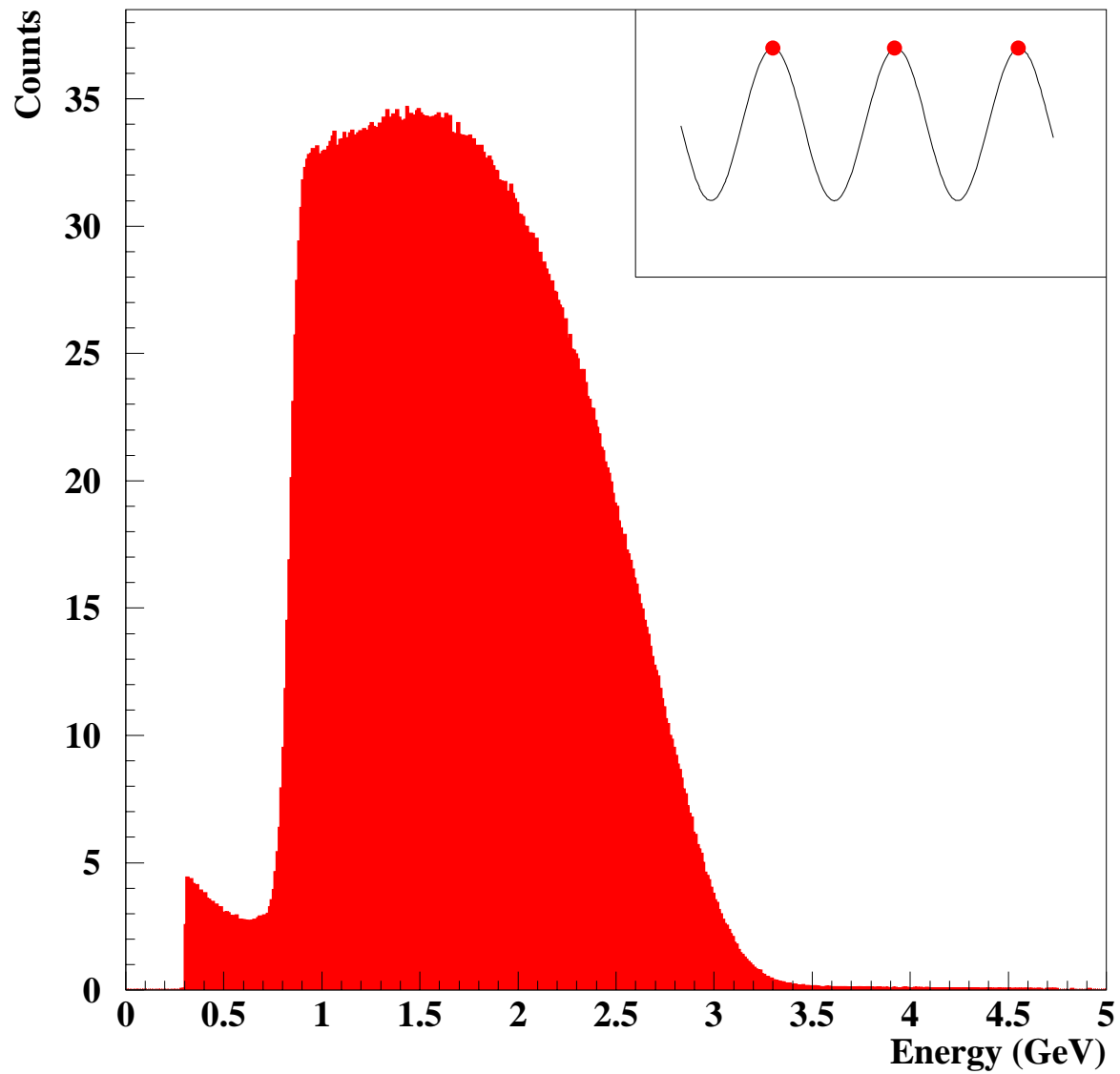
$$N(t) = N_0 e^{-t/\tau} (1 + A \cos[\omega_a t + \phi])$$



$$\frac{\Delta\omega_a}{\omega_a} = \frac{\sqrt{2}}{A\omega_a\gamma\tau\sqrt{N(E)}}$$

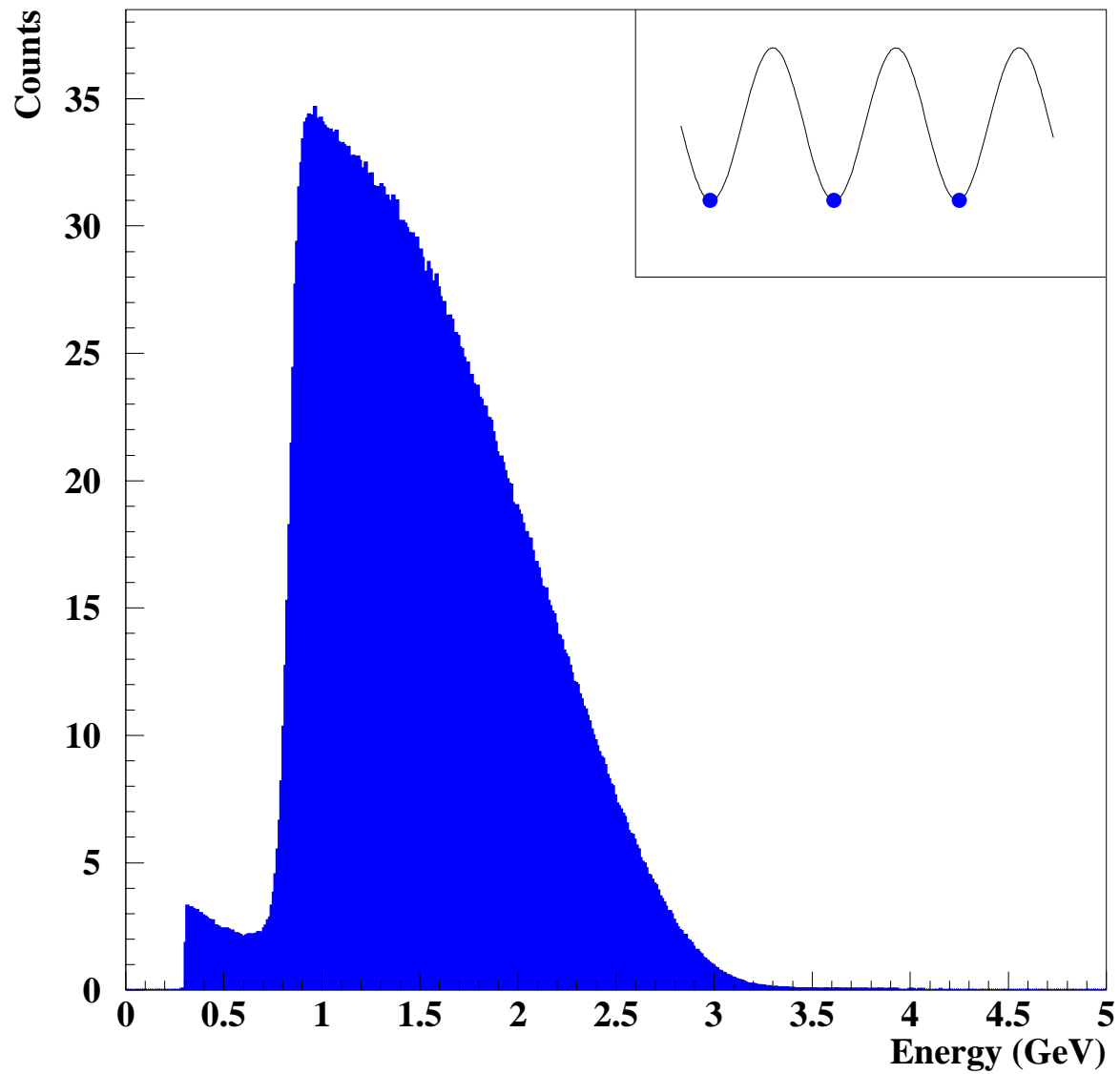
# Energy Spectrum

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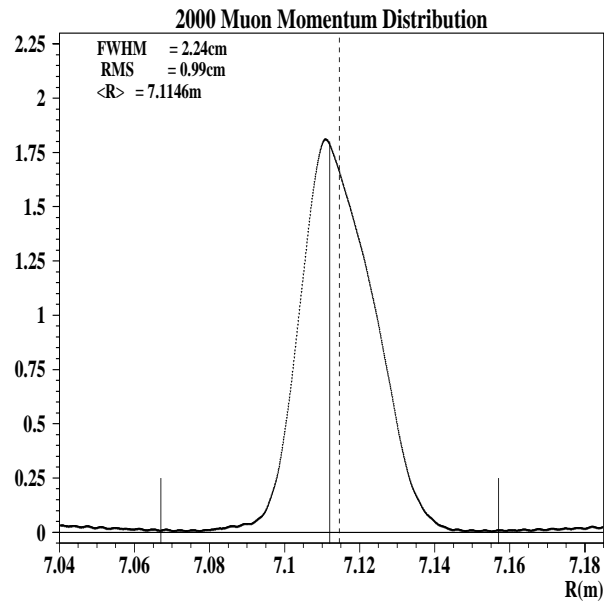
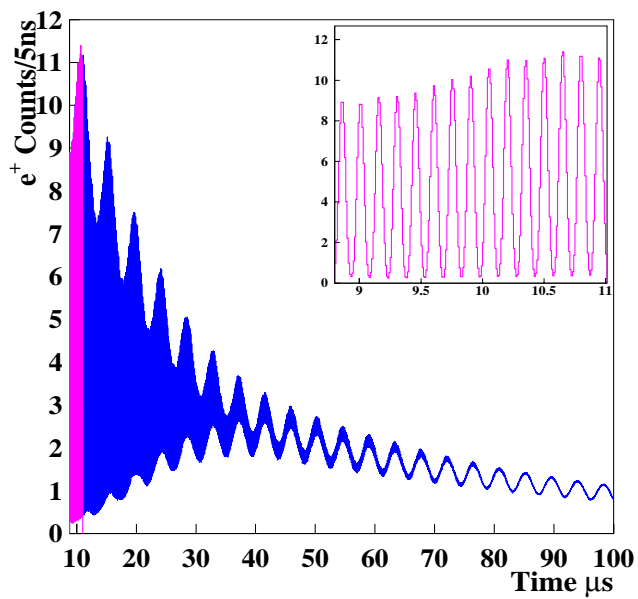
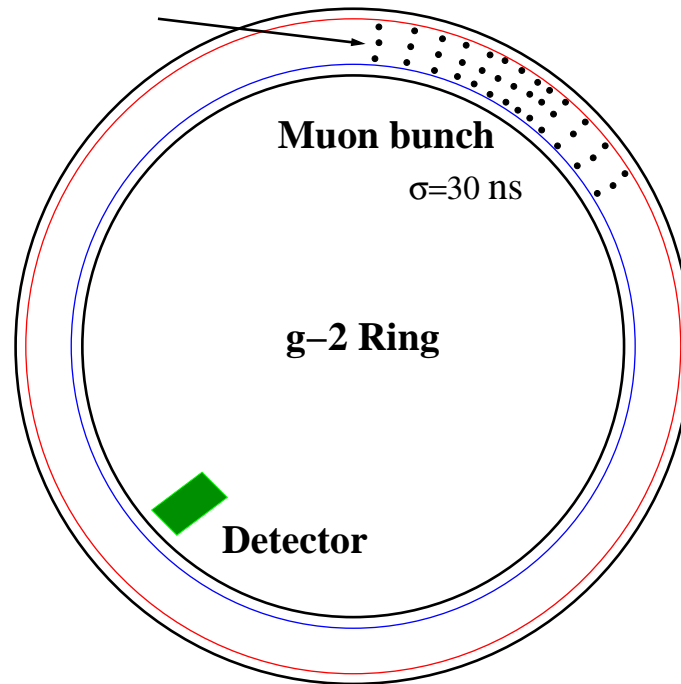


# Energy Spectrum

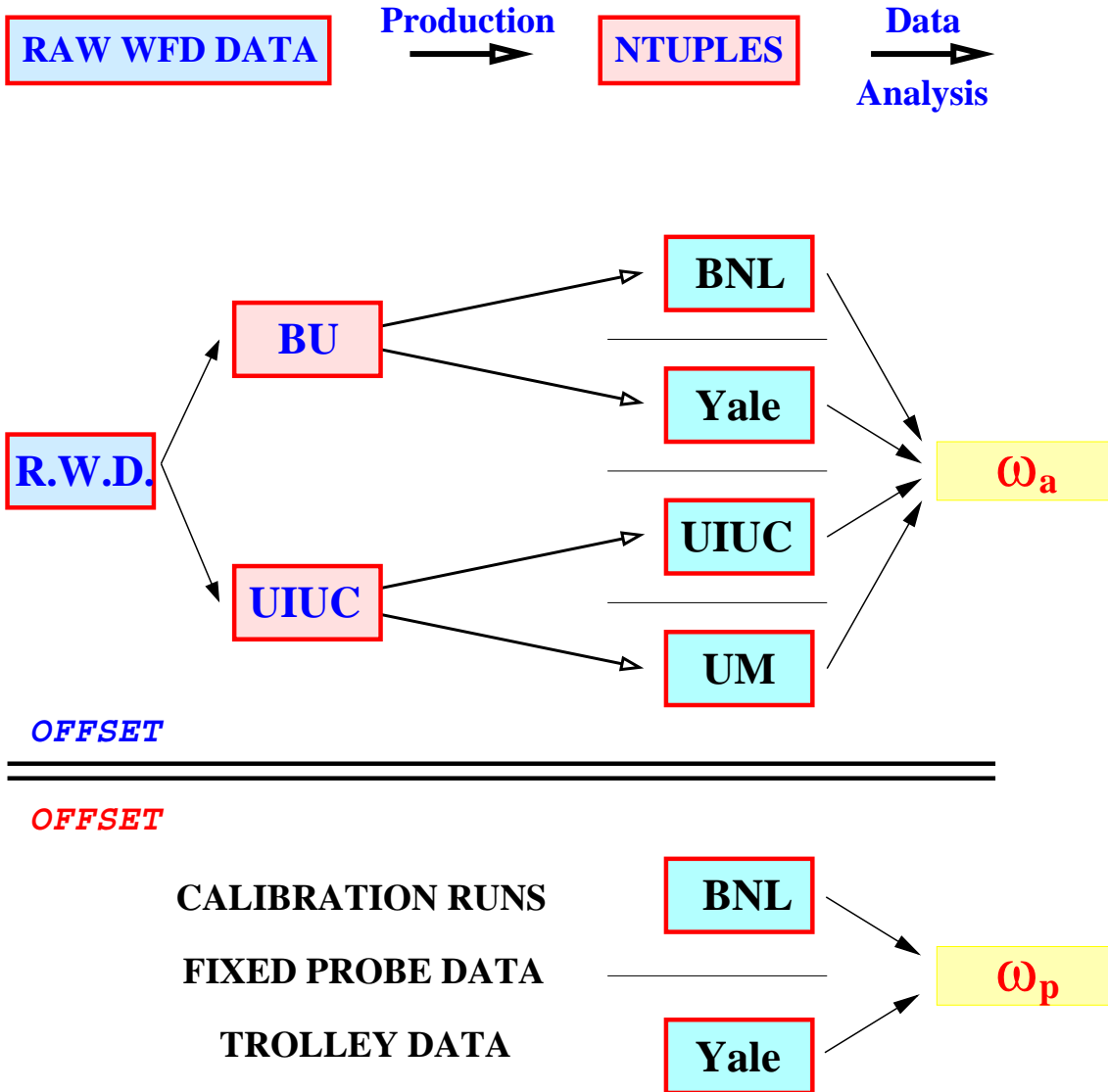
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# Fast Rotation



# Analysis of the 2000 Data

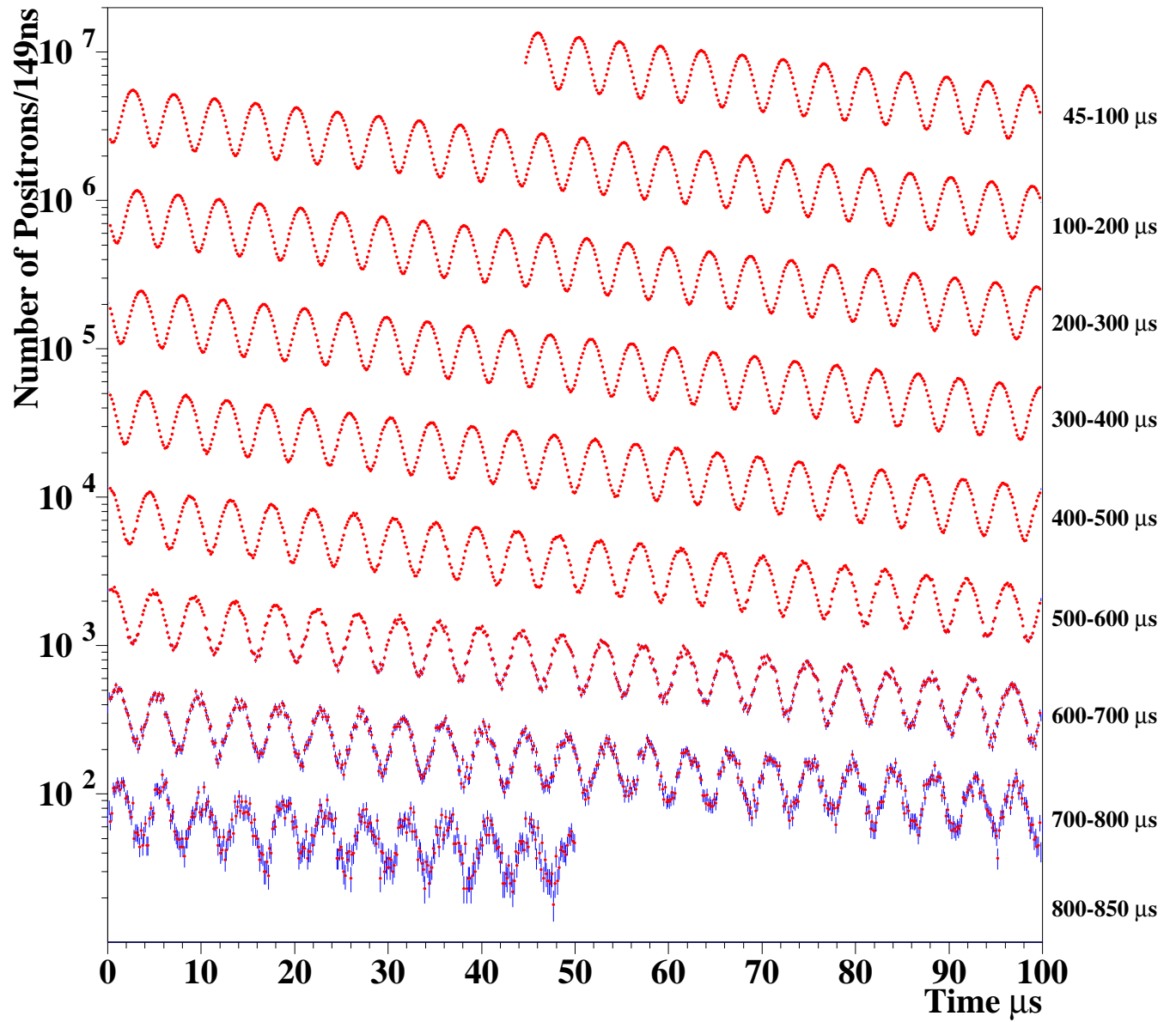


$$a_\mu = \frac{\frac{\omega_a}{\omega_p}}{\lambda - \frac{\omega_a}{\omega_p}} \quad \lambda = \frac{\mu_\mu}{\mu_p} = 3.183\ 345\ 39(10)^*$$

\* W. Liu et al. Phys. Rev. Lett. **82**, 711 (1999);  
D. E. Groom et al. Eur. Phys. J. **C15**, 1 (2000).

# 2000 Data

## 4 Billion Positrons with $E > 2$ GeV



# Fitting Function

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Ideal Fitting Function :

$$dN/dt = N_0 e^{-t/\tau} (1 + A \cos[\omega_a t + \phi_a])$$

10M e<sup>+</sup>/bin requires to include small effects!!

CBO Related Effects :

- $N_0 \rightarrow N_0(t)$
- $A \rightarrow A(t)$
- $\phi_a \rightarrow \phi(t)$

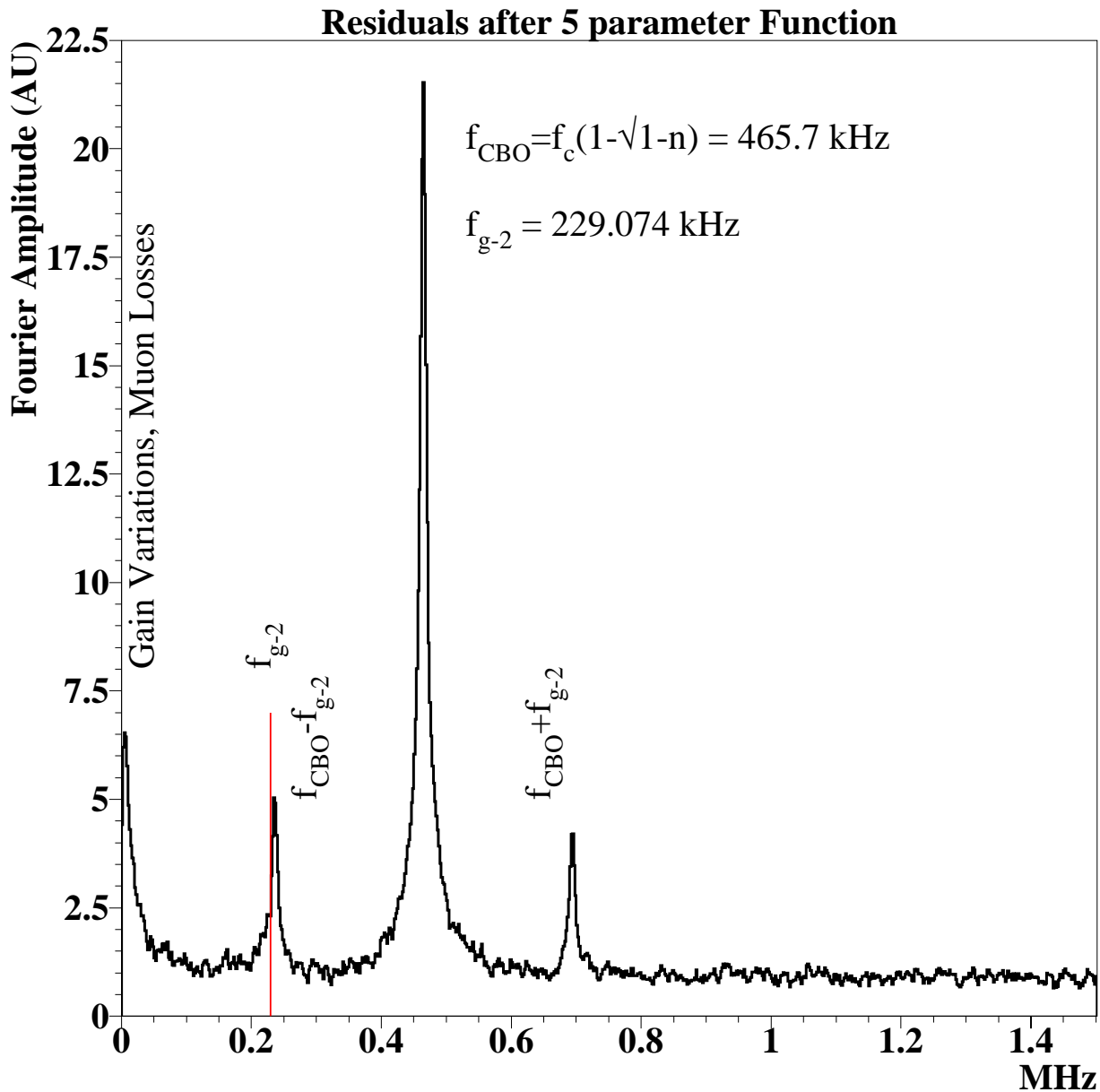
$$N_0(t) = N_0 \{1 + A_N e^{-t/\tau_{cbo}} \cos(2\pi f_{cbo} t + \phi_N)\}$$

$$A(t) = A \{1 + A_A e^{-t/\tau_{cbo}} \cos(2\pi f_{cbo} t + \phi_A)\}$$

$$\phi_a(t) = \phi + A_\phi e^{-t/\tau_{cbo}} \cos(2\pi f_{cbo} t + \phi_\phi)$$

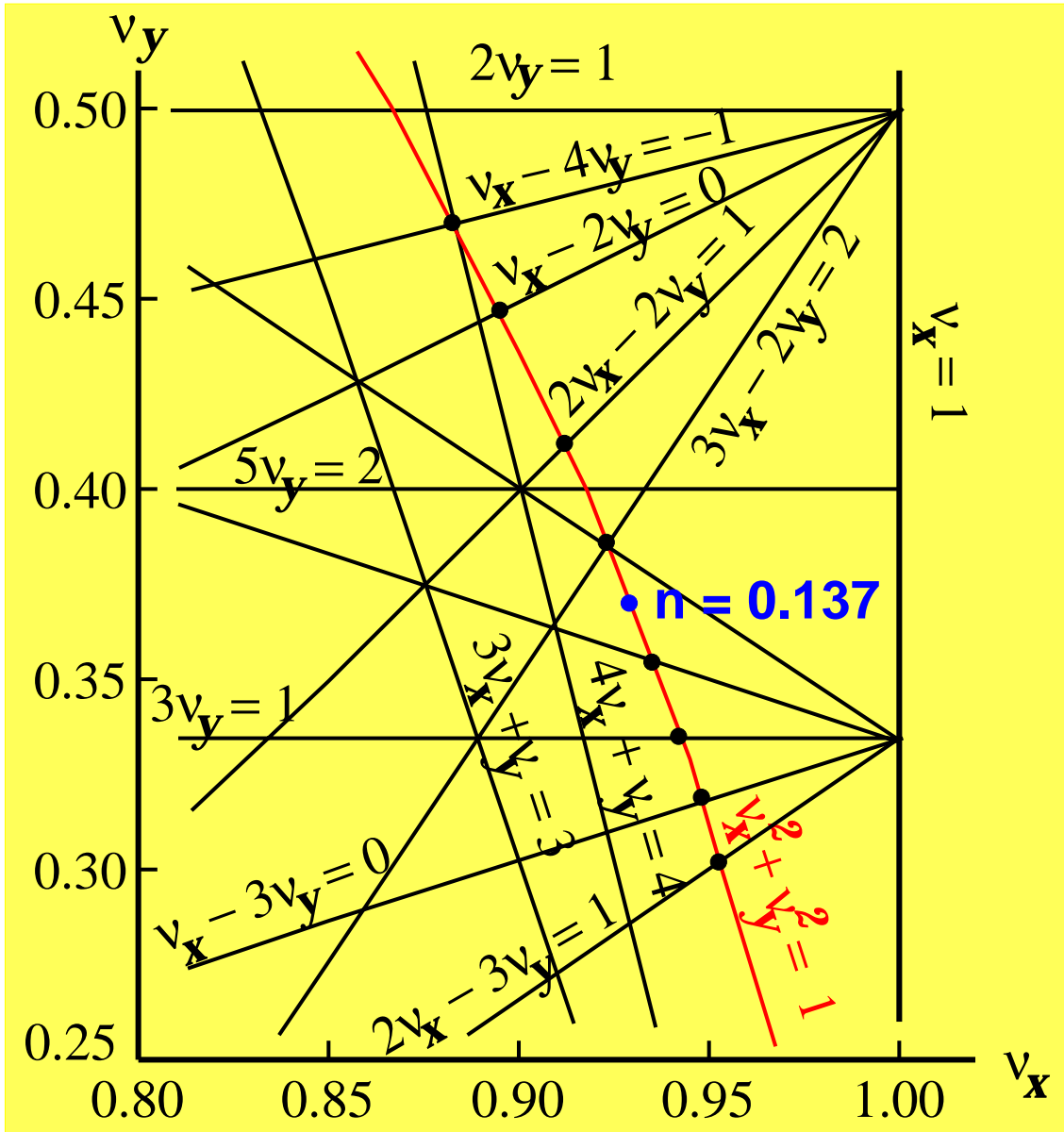
$A_N, A_A$  and  $A_\phi$  amplitudes are consistent with MC results. ( $A_N = 1\%$ ,  $A_A \approx 0.1\%$ ,  $A_\phi \approx 1\text{mR}$ )

# Fourier Analysis

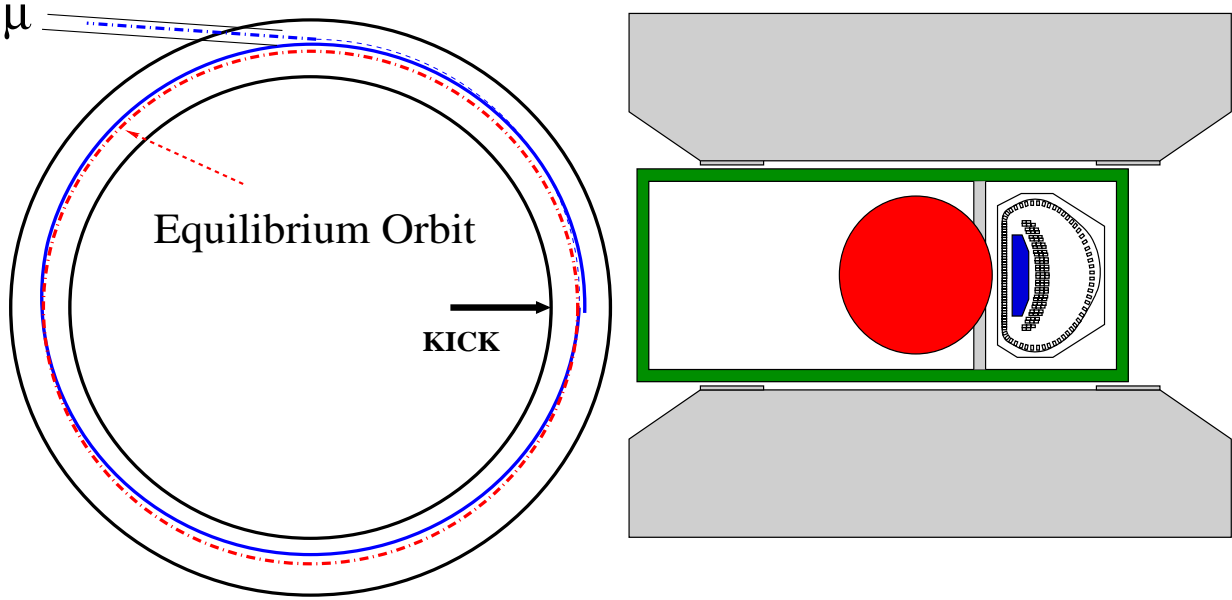




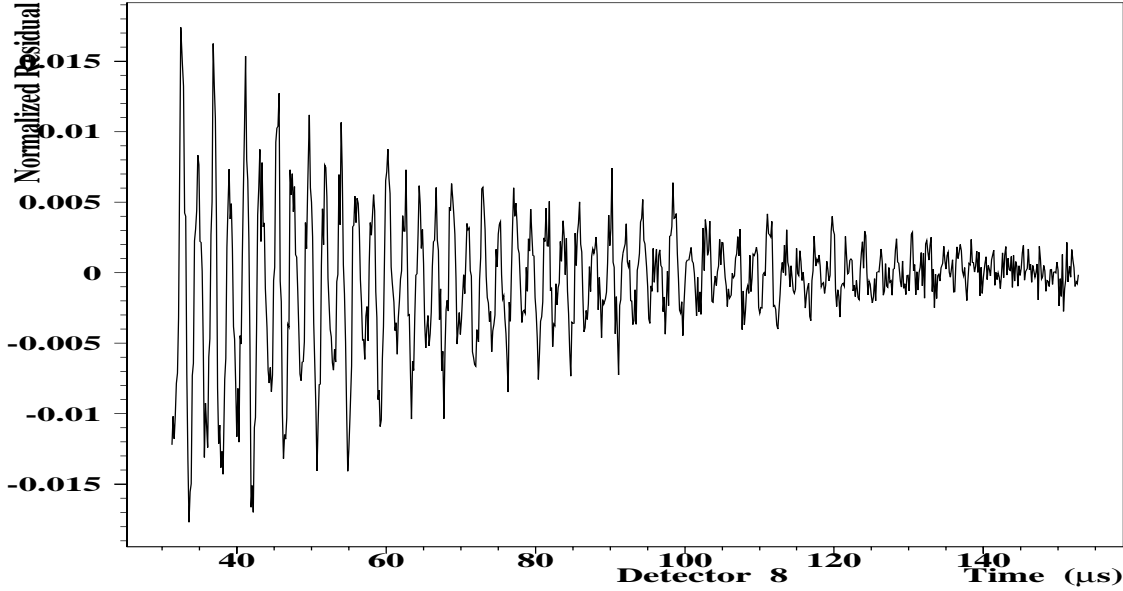
# Beam Tune



# CBO(Coherent Betatron Oscillations)

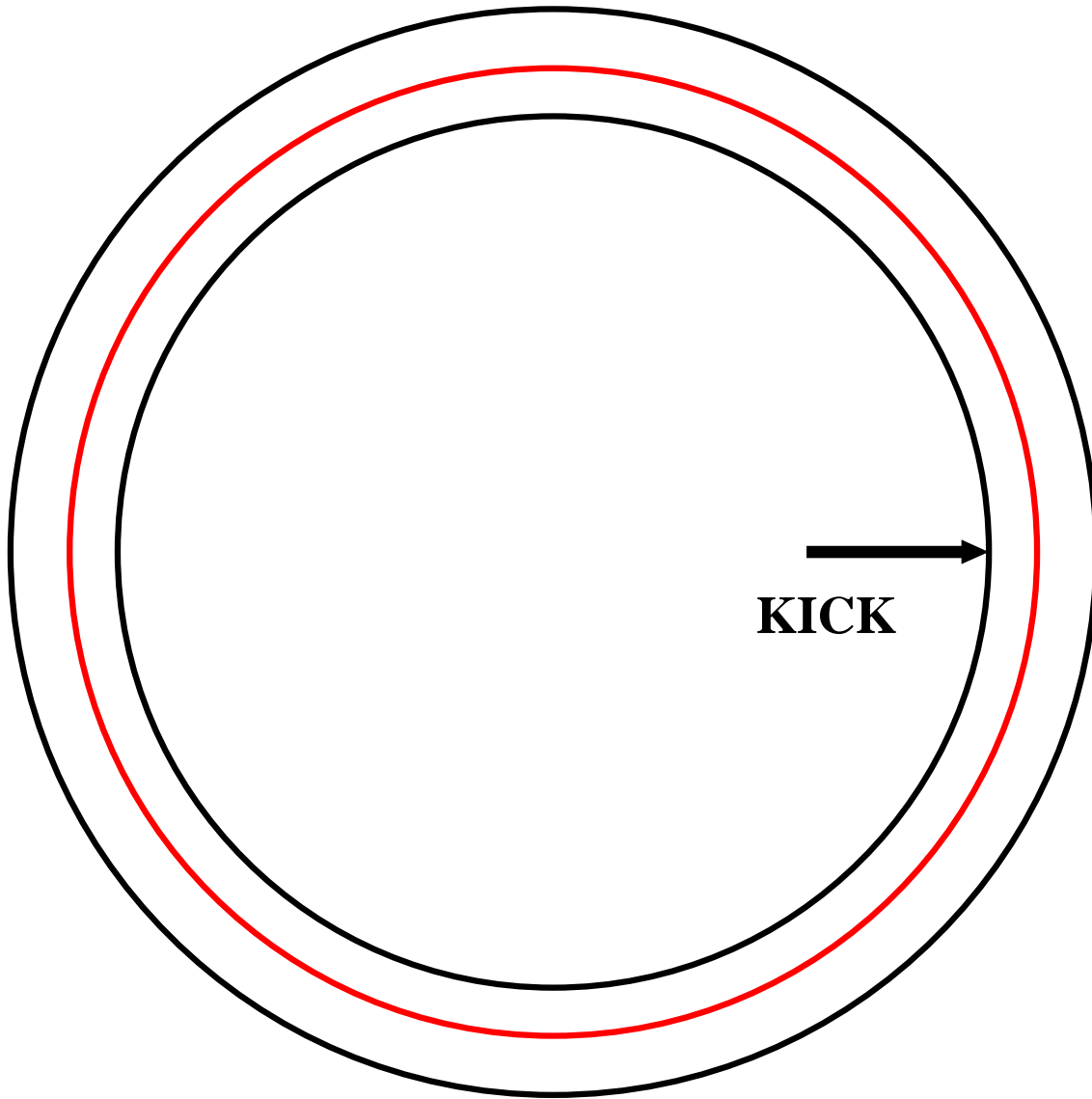


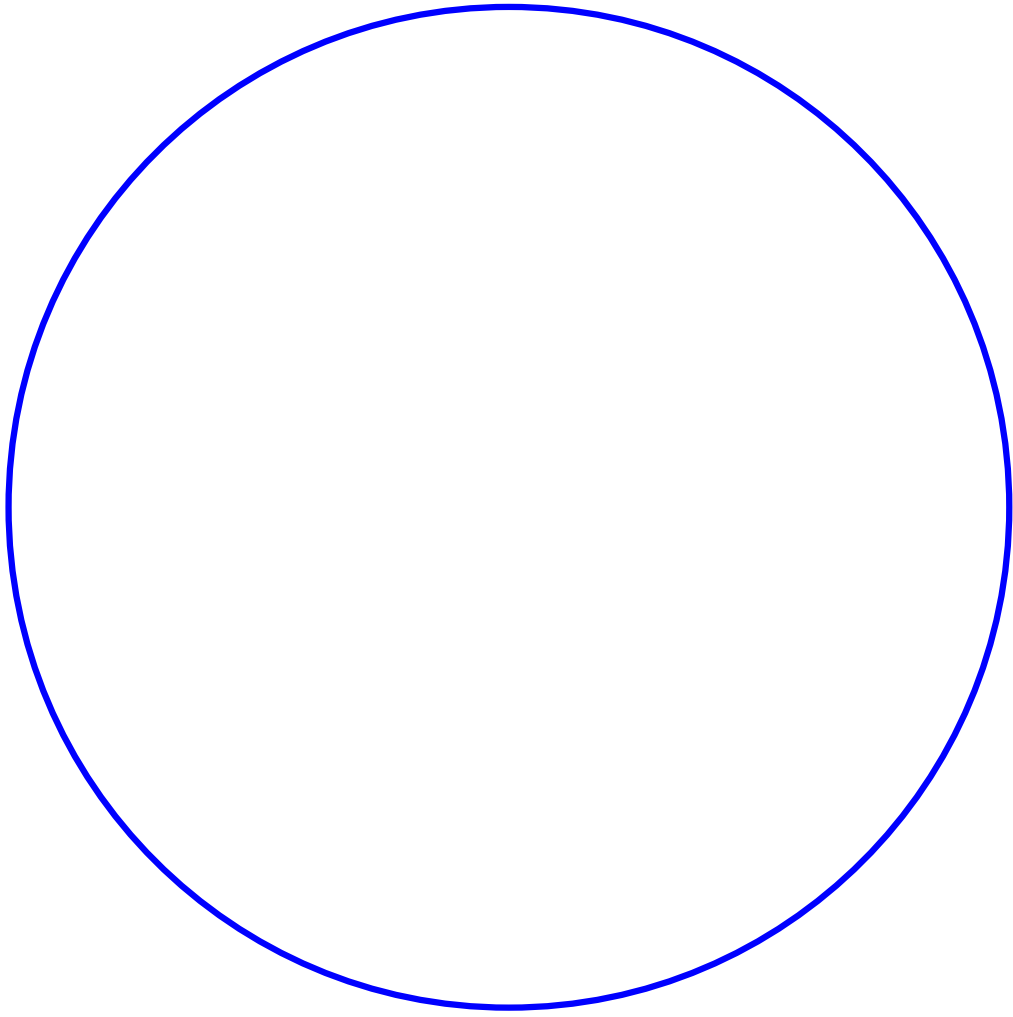
Look at the residuals (DATA - 5 Parameter Fit)  
Fourier Analysis of this residual  $\rightarrow$  CBO



# Coherent Betatron Oscillations (CBO)

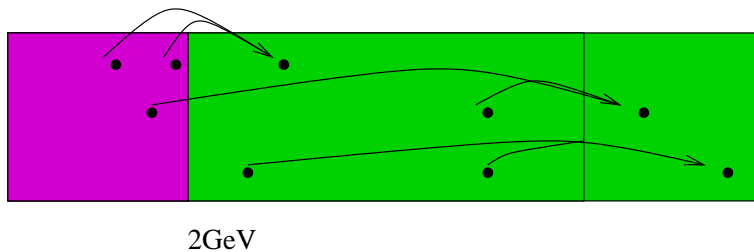
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# Pileup

Pileup  $e^+$  carry wrong energy and phase information for  $g-2$



An early to late effect caused by the change on muon intensity (due to decay) and enhanced by the Fast Rotation Factor

It can be accounted for by adding the related terms to the fitting function

$$N_0 e^{-t/\tau} (1 + A \cos(\omega_a + \phi)) \times [n_p e^{-t/\tau} (1 + A_p \cos(\omega_a + \phi + \phi_{pu}))]$$

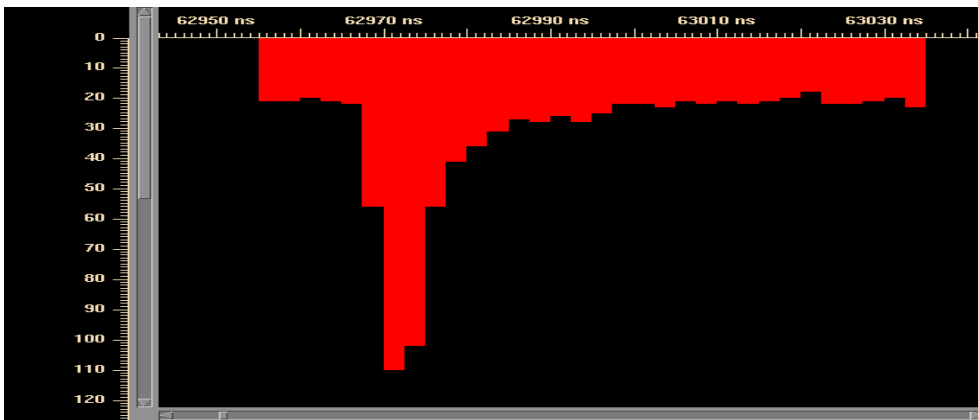
Factor two loss in the statistical error of  $\omega_a$  !

Therefore a different solution to the problem is necessary

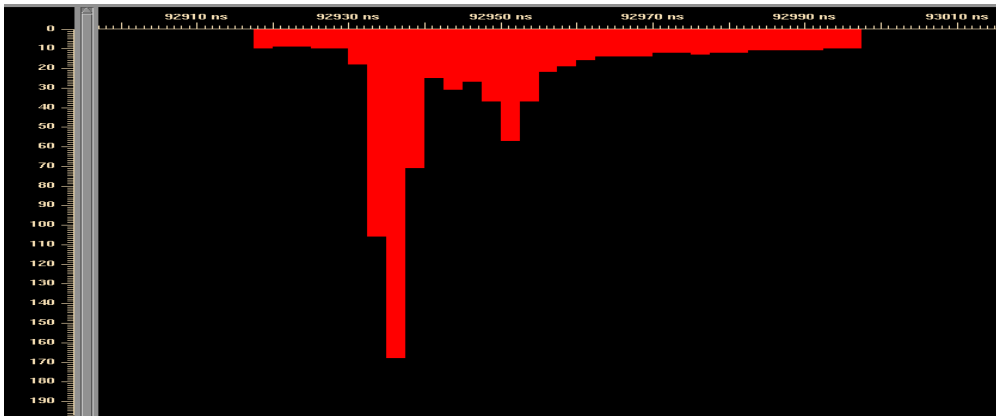
# Pileup

Each positron pulse above 1GeV is digitized with a 400MHz WFD

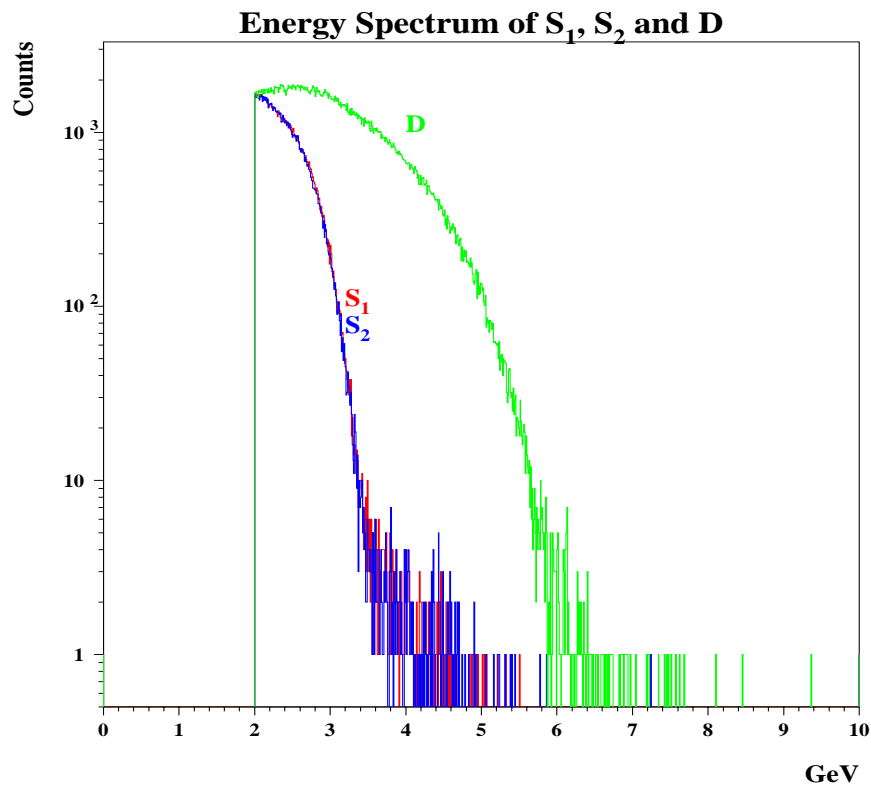
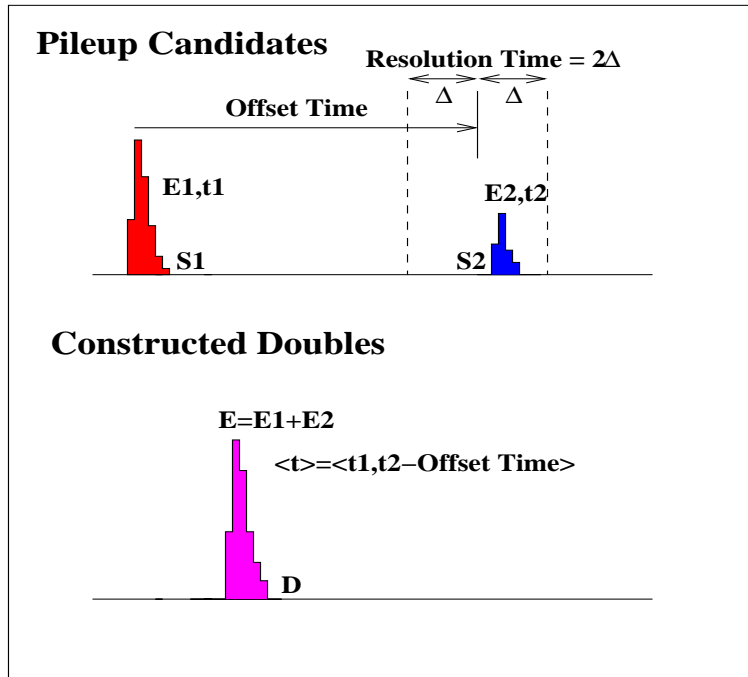
If the pulse height is above the threshold, start digitizing ("WFD Island"  $\approx 80$ ns)



If there are pulses after the main triggering pulse on the same island (shadow pulses).

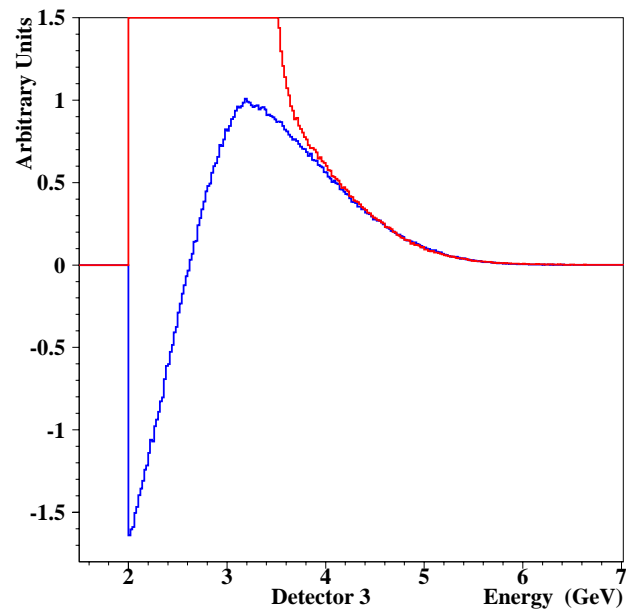
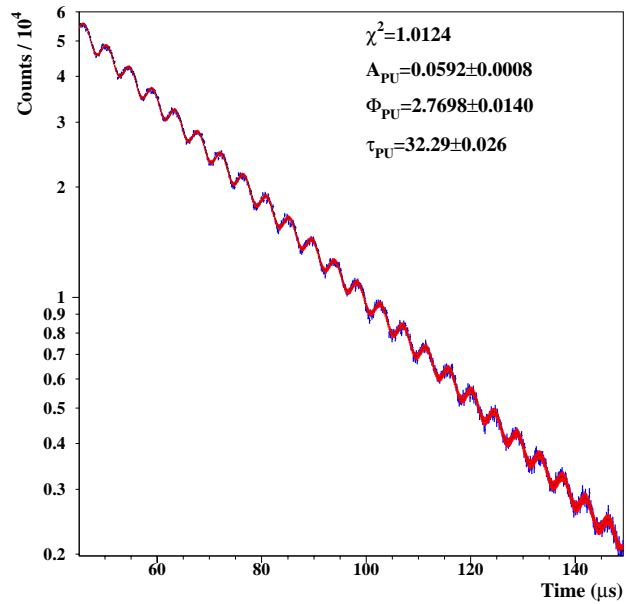


# Pileup



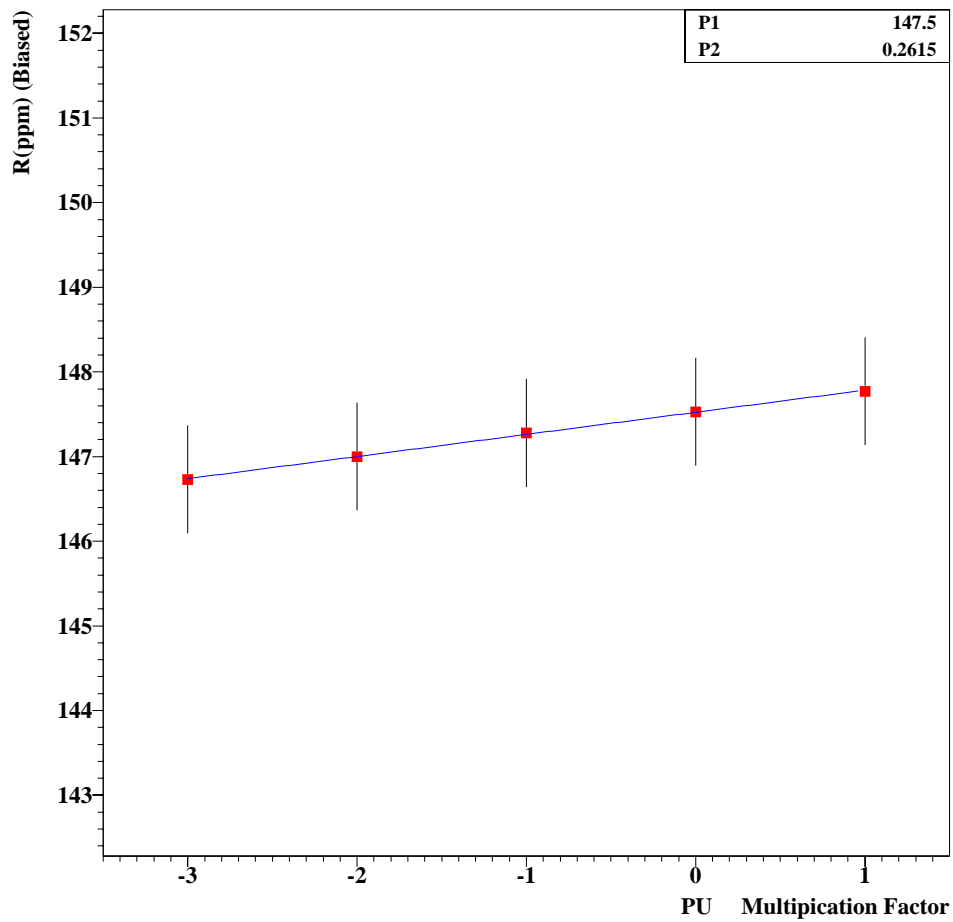
# Pileup

Fit to Constructed Pileup ( $D - S_1 - S_2$ )





# Pileup Systematic



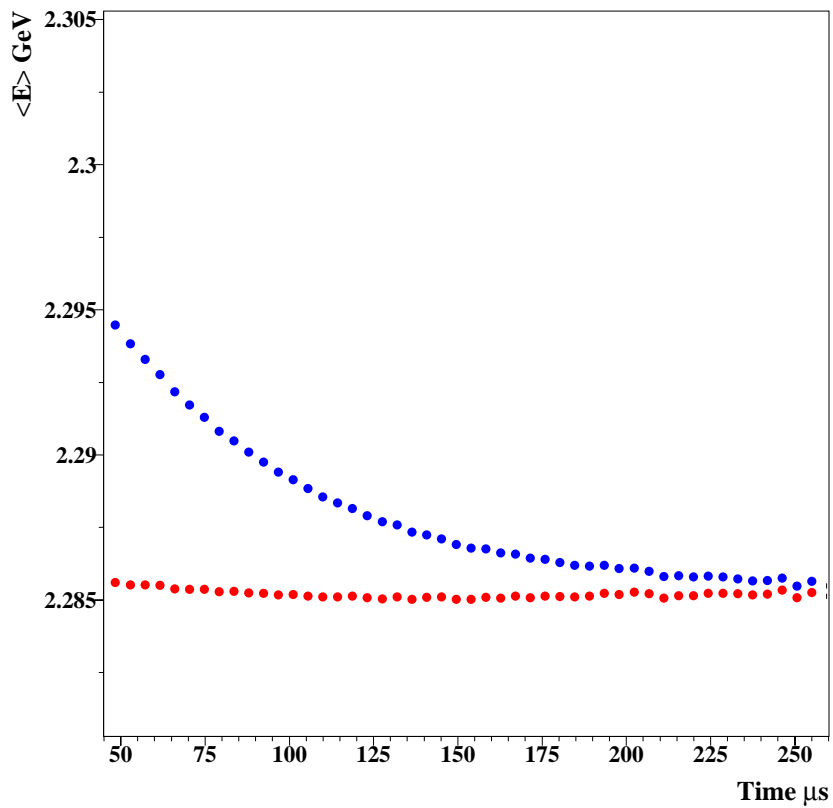
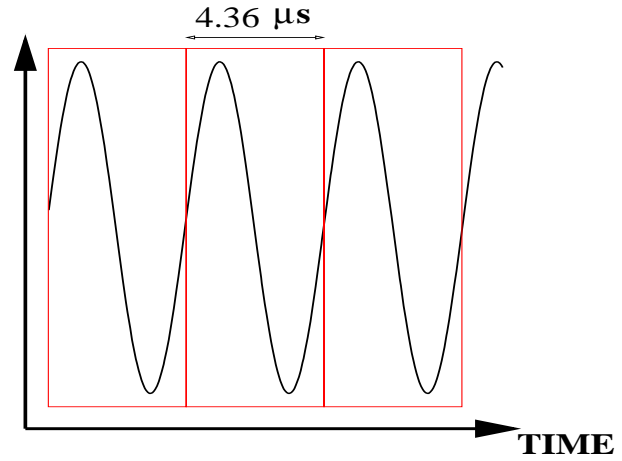
$$Data \times (1 + \lambda \times PU)$$

No pileup subtraction  $\rightarrow 0.26\text{ppm}$

# Gain Variations

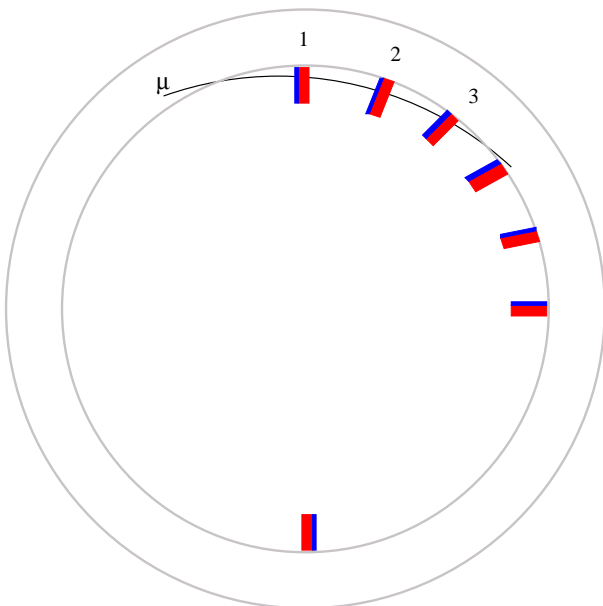
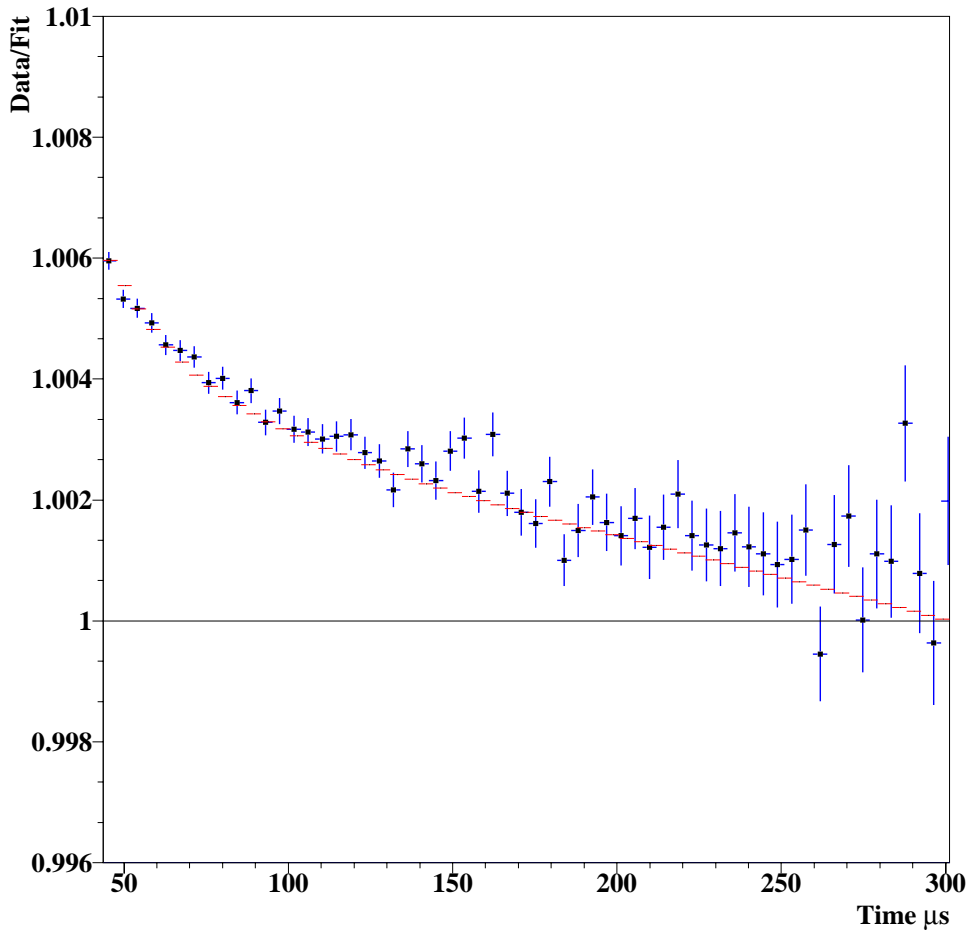
Average Energy

$$\bar{E} = \frac{\sum_i E_i n_i}{\sum_i n_i}$$

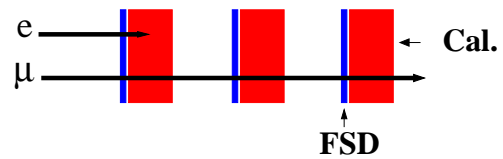


# Muon Losses

After Pileup Subtraction and including the CBO

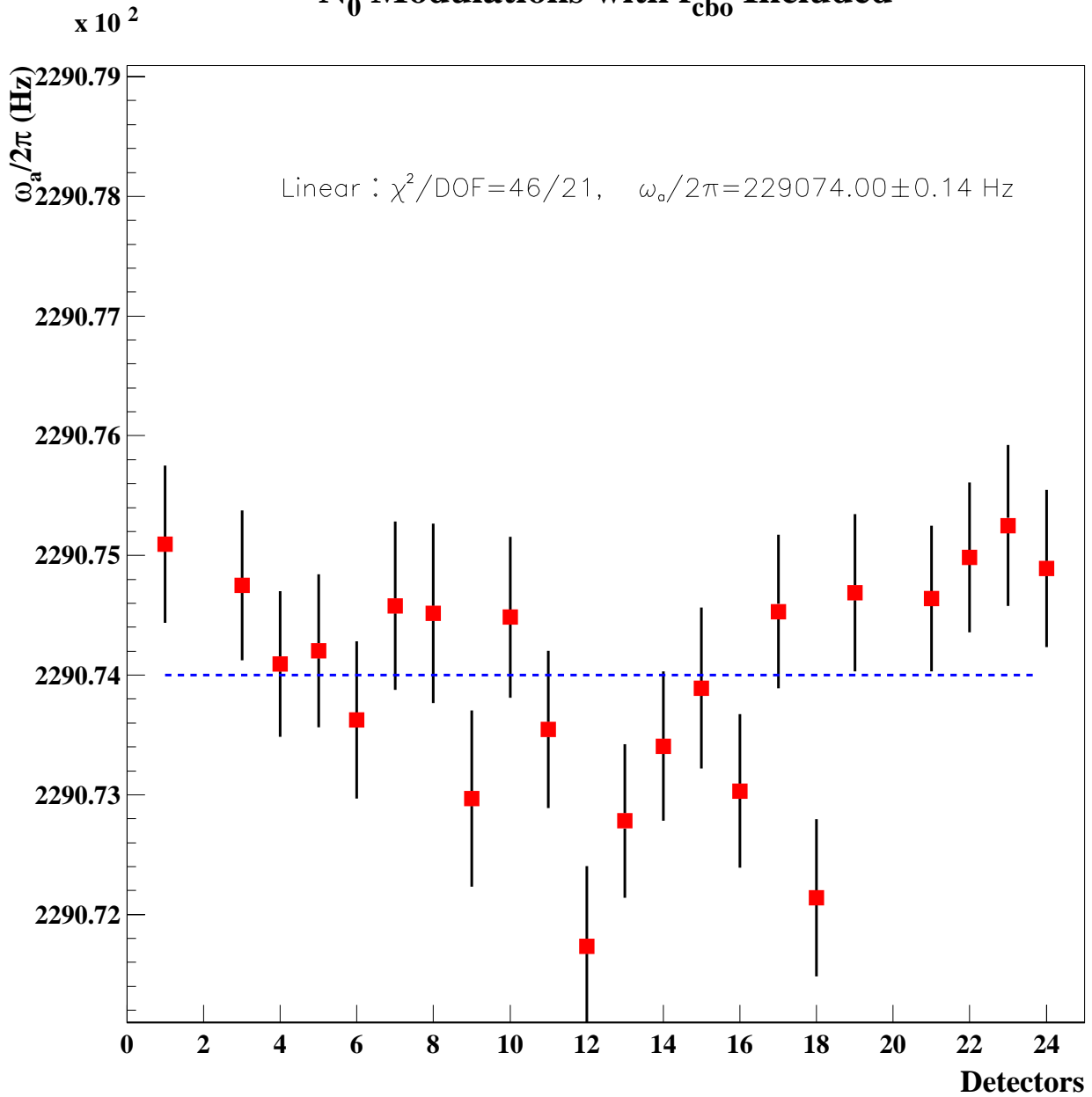


**Calorimeters : almost no energy deposition**  
**FSD's : Hits to any finger**



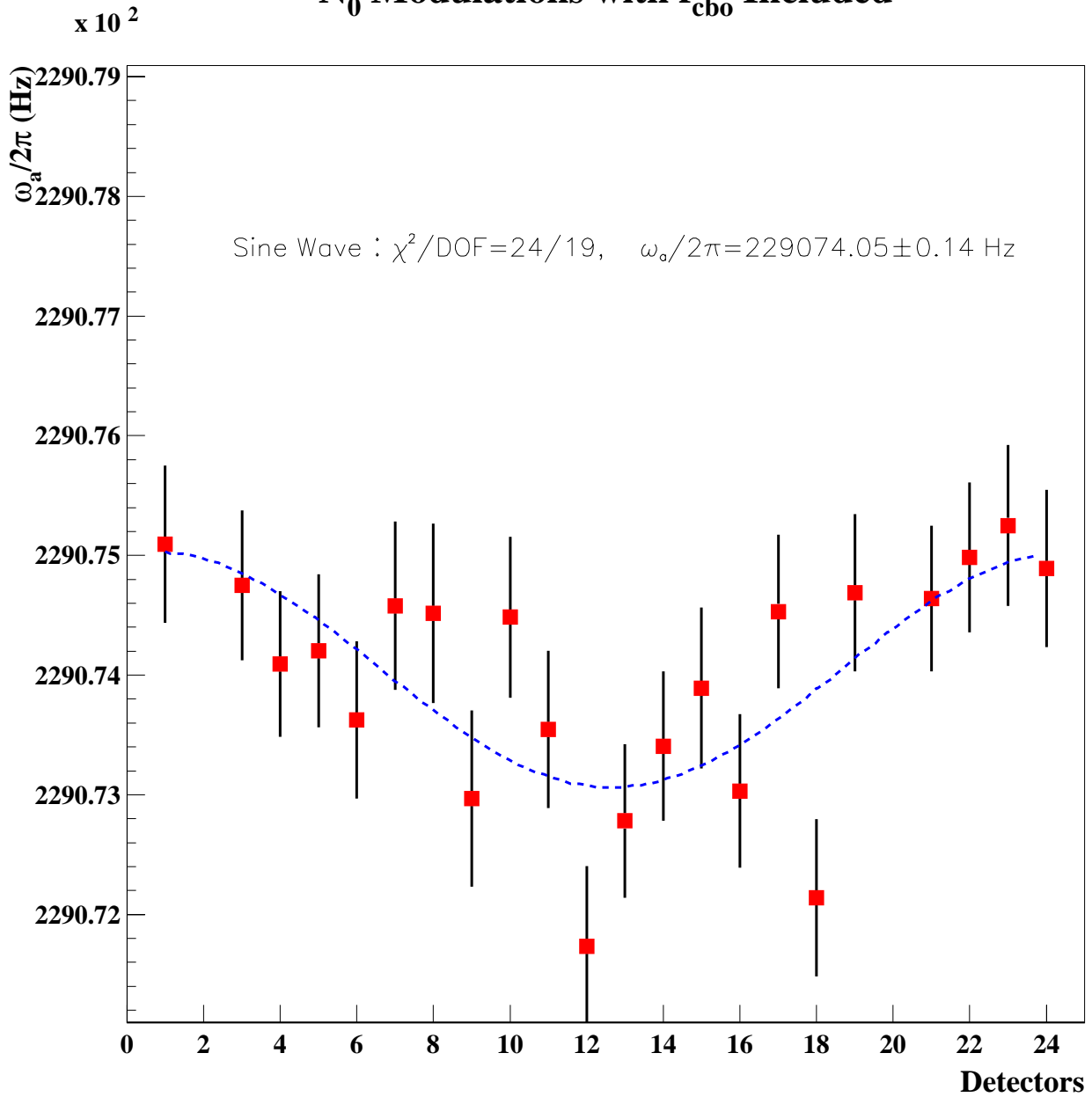
# 2000 Data

## $N_0$ Modulations with $f_{cbo}$ Included



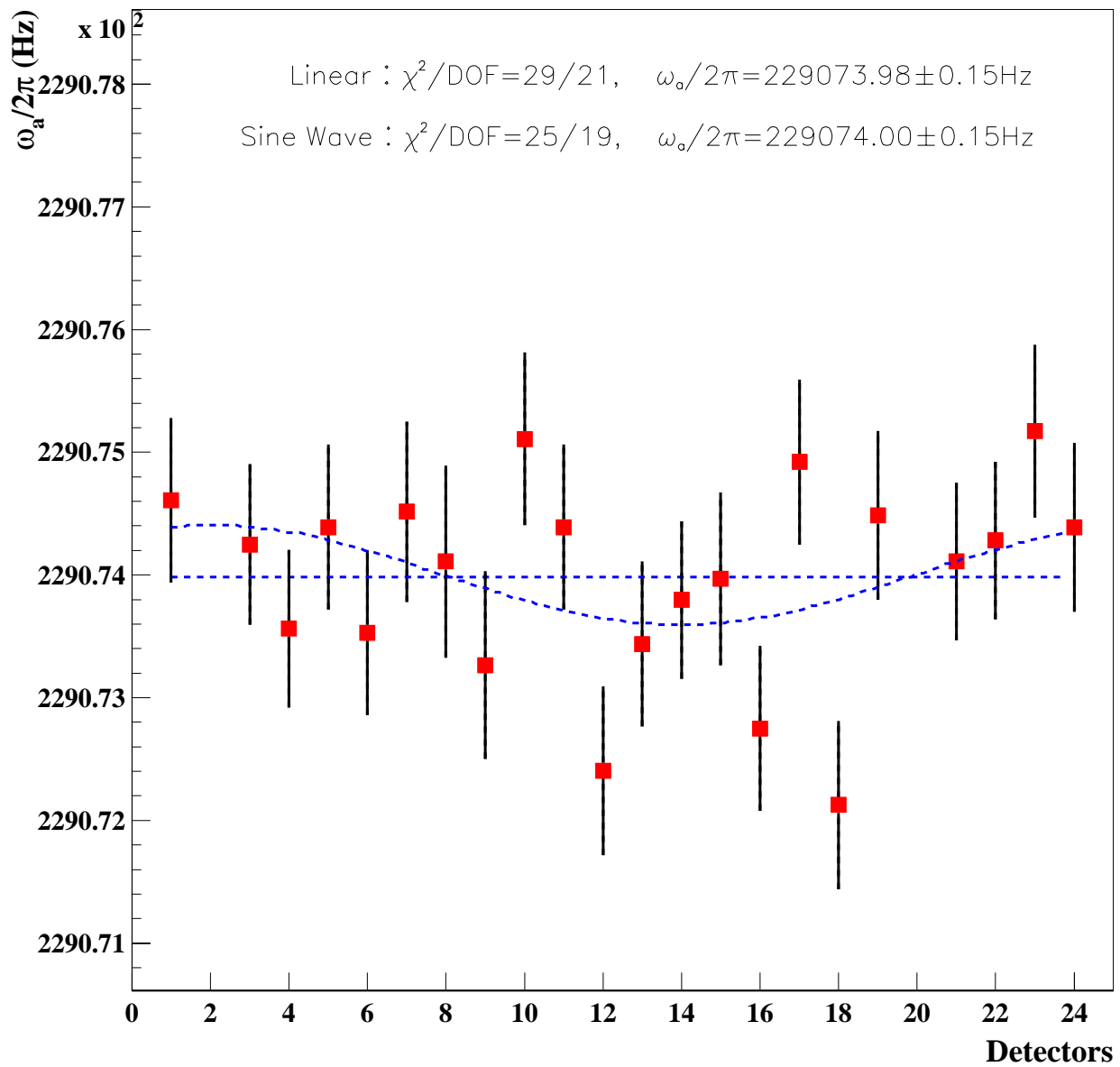
# 2000 Data

## $N_0$ Modulations with $f_{cbo}$ Included



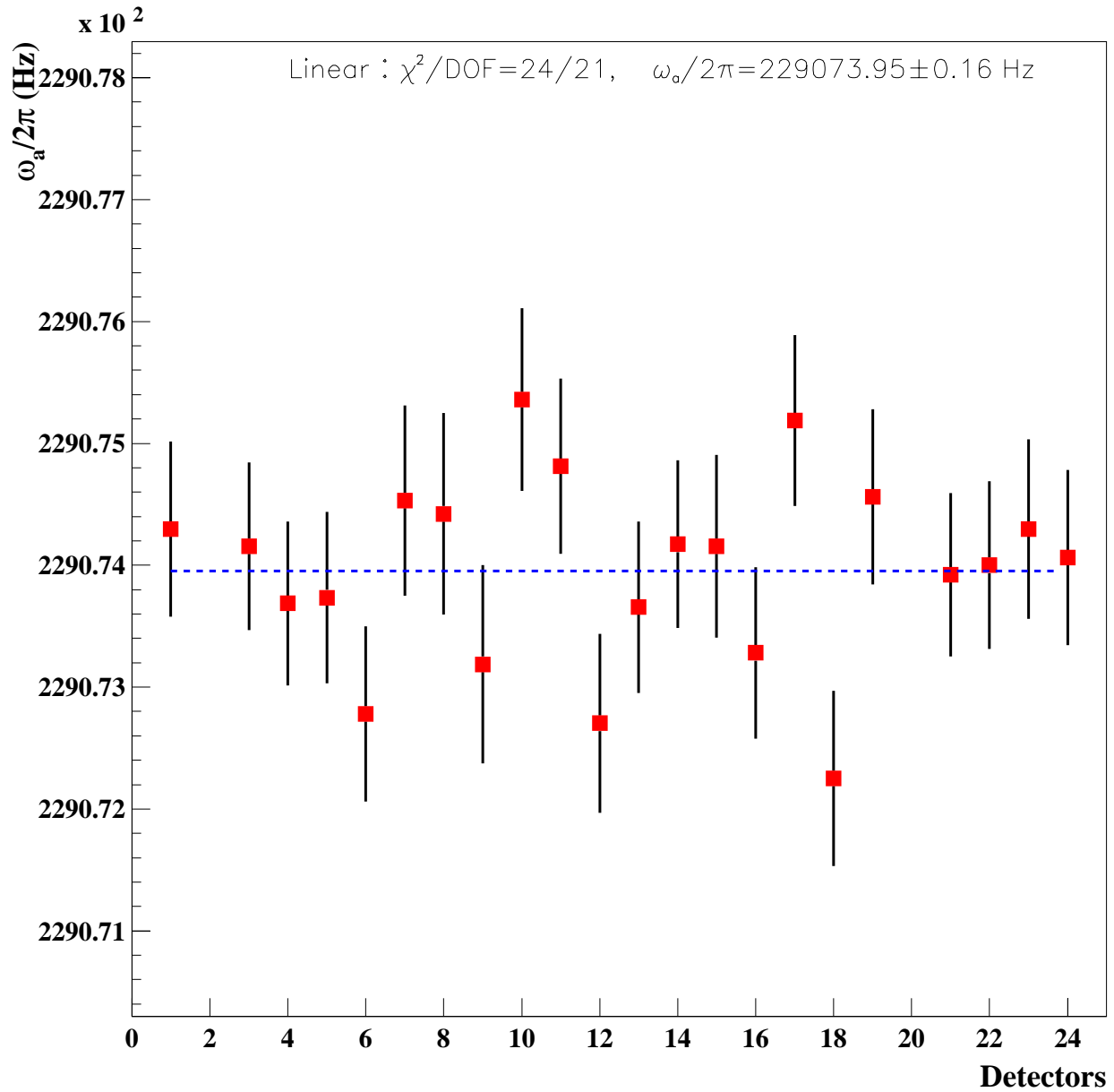
# 2000 Data

## $N_0$ and A Modulations with $f_{\text{cbo}}$ Included



# 2000 Data

## $N_0$ , $A$ and $\phi$ Modulations with $f_{cbo}$ Included



## 4 Independent Analysis of $\omega_a$

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- Function Modulating  $N_0$ ,  $A$ ,  $\phi_a$  with  $f_{cbo}$ .
- Function Modulating  $N_0$ ,  $A$  with  $f_{cbo}$ .
- Strobing the data @  $f_{cbo}$  so  $\omega_a$  becomes independent of  $f_{cbo}$ .
- Ratio Method;  $\omega_a$  becomes independent of slow effects, e.g. muon losses.

### Systematic Uncertainties on $\omega_a$

Source of Errors	Size (ppm)
CBO	0.21
Pileup	0.13
Gain Changes	0.13
Lost Muons	0.10
Binning and Fitting Procedure	0.06
Others	0.06
<b>Total</b>	<b>0.31</b>

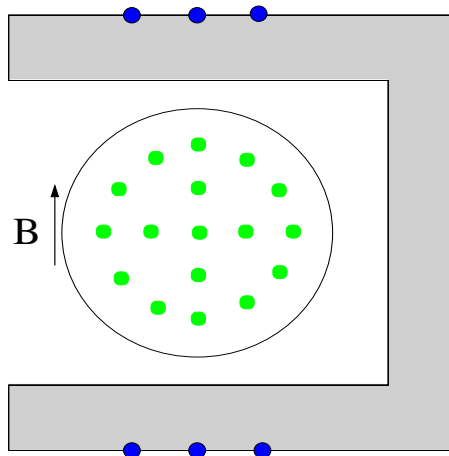
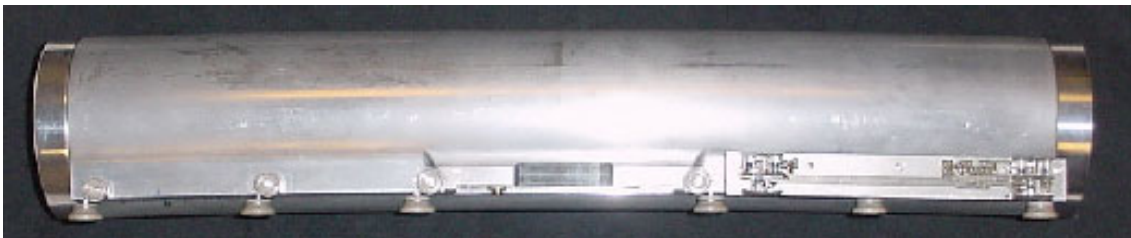


# B Field Measurements

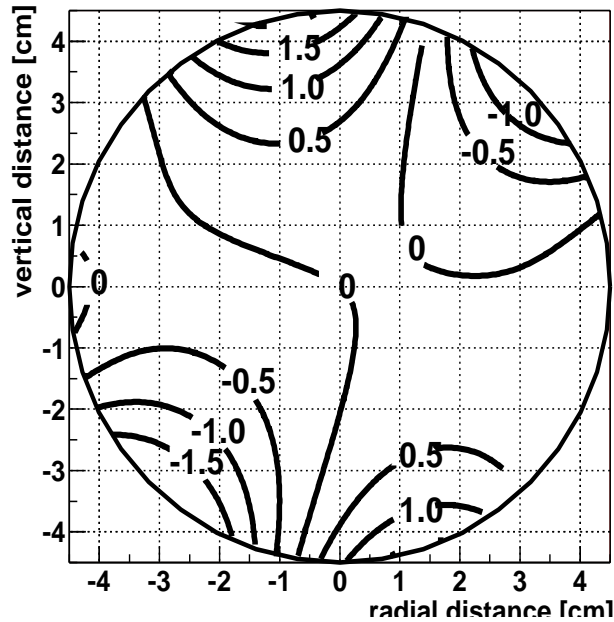
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Two complementary methods for the field measurement

- Fixed NMR probes  
366 Fixed NMR probes  
Distributed around the ring  
Continuous measurement
- Beam Tube Trolley  
17 NMR probes  
Field measurements in the vacuum  
2-3 times a week

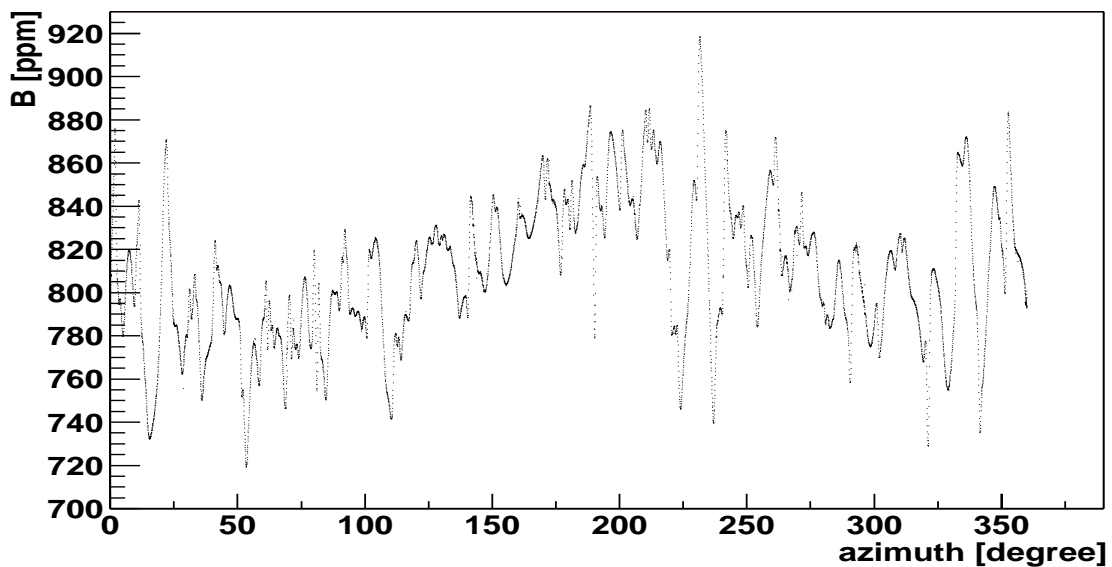


# B Field Measurements



## Multipoles (ppm)

	normal	skew
Quad	0.24	0.29
Sext	-0.53	-1.06
Octu	-0.10	-0.15
Decu	0.82	0.54



Total 0.24 ppm systematic uncertainty for the  $\omega_p$ .

## Computation of $a_\mu$ :

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Independent/blind analysis of  $\omega_a$  and  $\omega_p$  provides  $a_\mu$  :

$$a_\mu = \frac{\omega_a}{\frac{e}{m_\mu} \langle B \rangle} = \frac{\omega_a / \omega_p}{\mu_\mu / \mu_p - \omega_a / \omega_p}$$

# Results

$$a_\mu(SM) = 11659178(7) \times 10^{-10} (0.6 \text{ ppm})^*$$

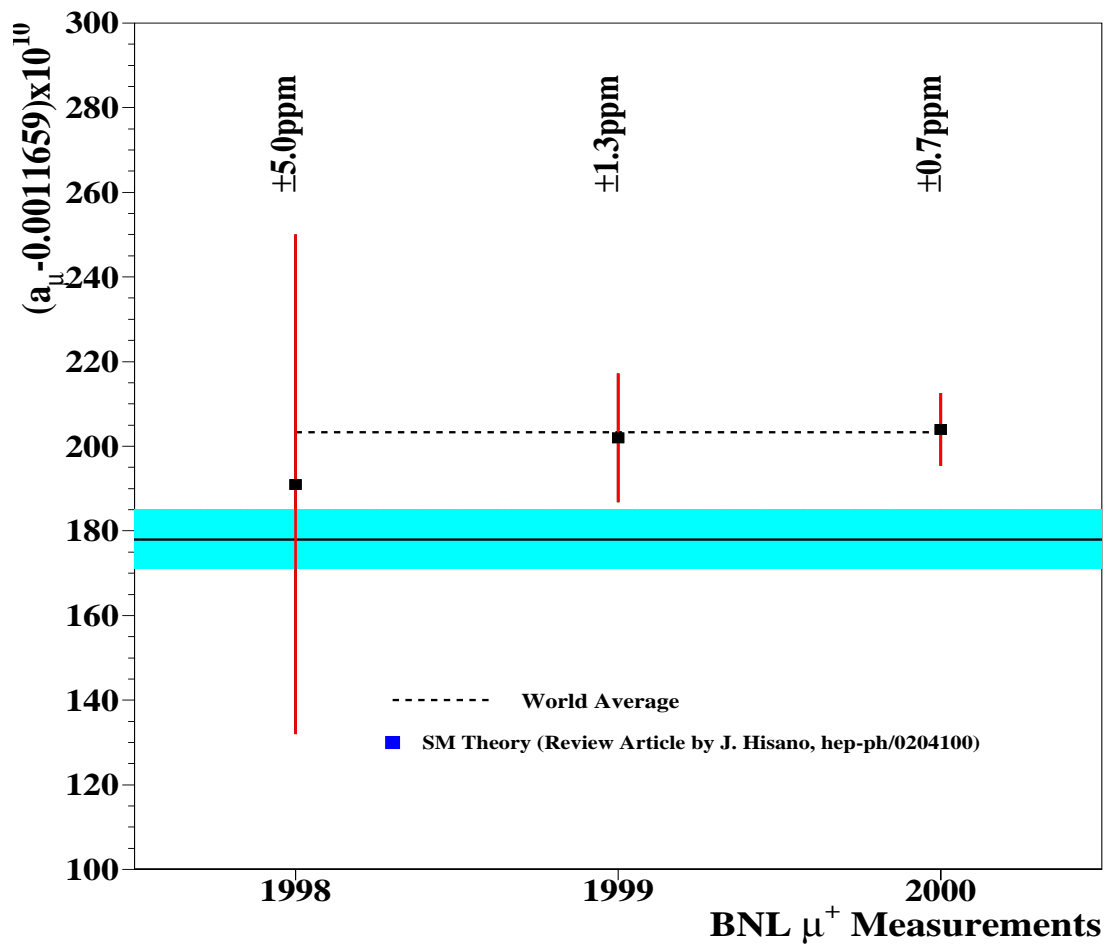
More  $a_\mu(SM)$  available !

BNL Data of 2000 :

$$a_\mu(exp) = 11659204(7)(5) \times 10^{-10} (0.7 \text{ ppm})$$

Exp. World Average :

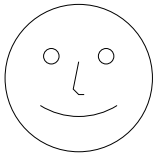
$$a_\mu(exp) = 11659203(8) \times 10^{-10} (0.7 \text{ ppm})$$



# 2001 at a Glance

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## Excellent Run



$\approx 3.5\text{B } e^-$  collected ( $E > 2 \text{ GeV}$ ,  $t > 30\mu\text{s}$ )

**Production**, parallel with the data taking

**Purpose** : To monitor as many parameters as possible during the data taking

Beam quality,  
Muon losses,  
Detector thresholds,  
Gate-on times continuously monitored

**Energy thresholds** ; All detectors  $1\text{GeV} \leq E_{th}$

**Gate on Times** ; All detectors  $t \leq 20\mu\text{s}$

**Two different focusing index** ;  $n=0.122$   $n=0.142$  will provide a better understanding of the CBO