

Atmospheric Neutrinos and Long Baseline Neutrino Oscillation Experiments

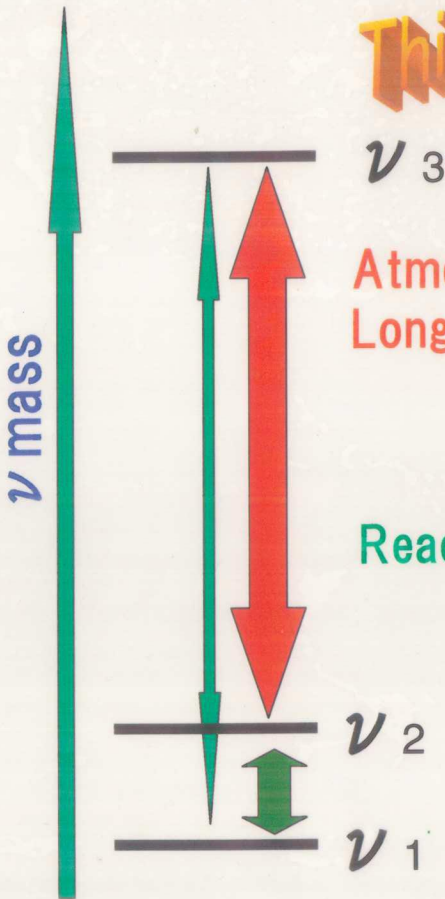
Takaaki Kajita

Institute for Cosmic Ray Research, Univ. of Tokyo

outline

- Introduction
- Atmospheric neutrinos
 - ⇒ data and $\nu_{\mu} \rightarrow \nu_{\tau}$ analysis
 - ⇒ 3 flavor oscillation and reactor exp's
 - ⇒ Search for CC ν_{τ} events
- Long-baseline accelerator experiments
 - ⇒ Status of K2K
 - ⇒ Future experiments
- Summary

Introduction



This talk

Kamiokande
Soudan2
Super-Kamiokande
MACRO

Atmospheric neutrinos
Long baseline exp.

K2K

....

This talk

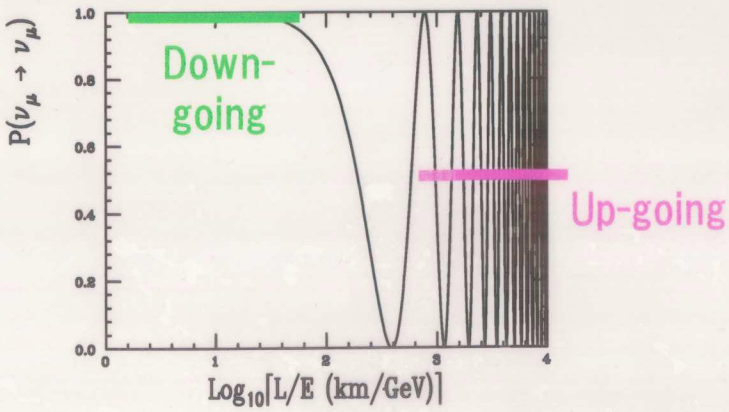
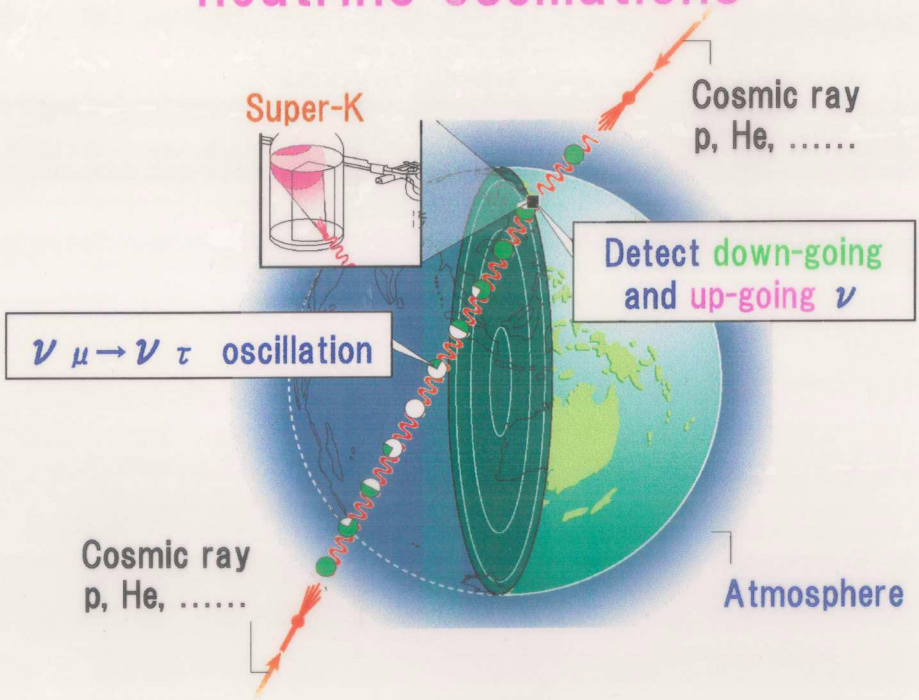
Reactor exp.

CHOOZ
Palo Verde

This talk

Solar neutrinos
(Reactor exp.)

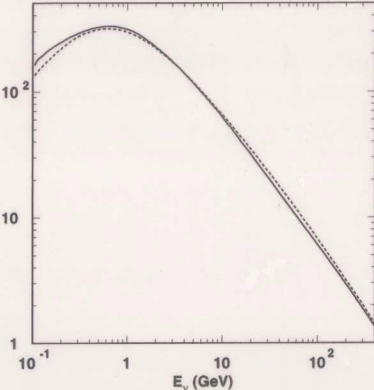
Atmospheric neutrinos and neutrino oscillations



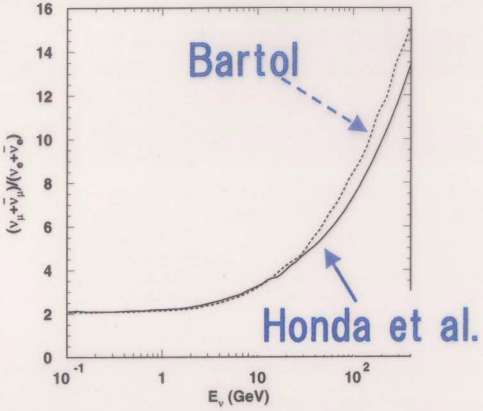
Atmospheric neutrino flux

Overall flux

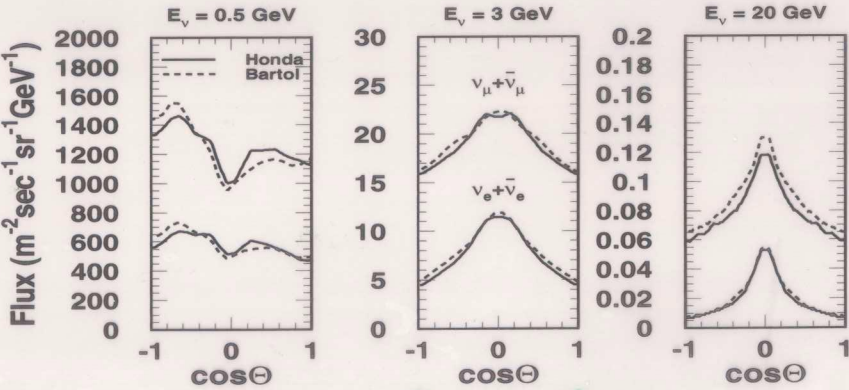
$(\nu_\mu + \bar{\nu}_\mu)$ flux $\times E \nu^2$



$$\frac{(\nu_\mu + \bar{\nu}_\mu)}{(\nu_e + \bar{\nu}_e)}$$



Zenith angle (flight length)



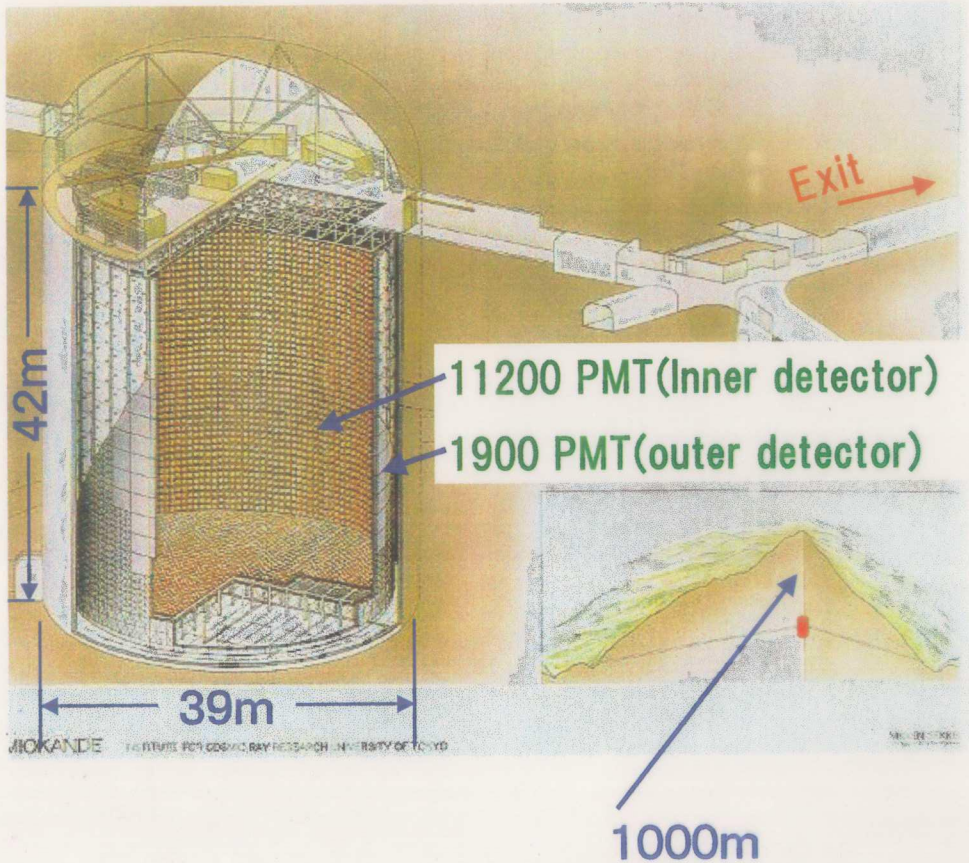
$L \nu = 12800$

$\sim 20\text{km}$



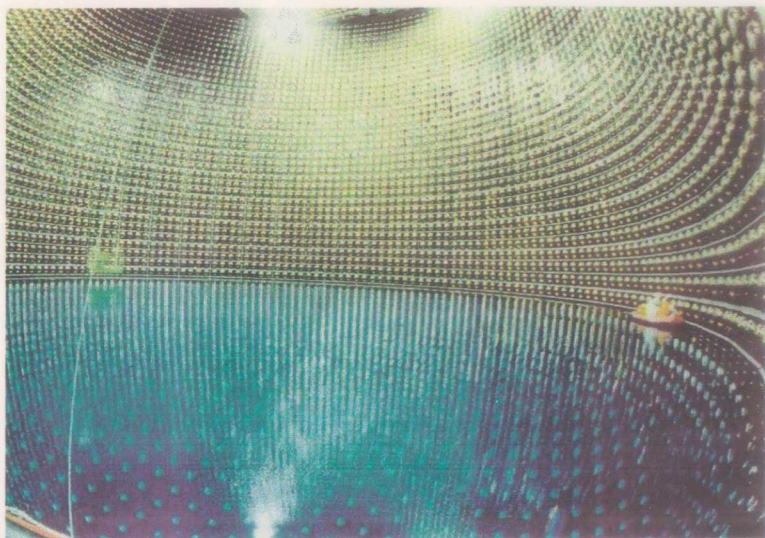
Super-Kamiokade detector

50,000 ton water Cherenkov detector
(22.5 kton fiducial volume)

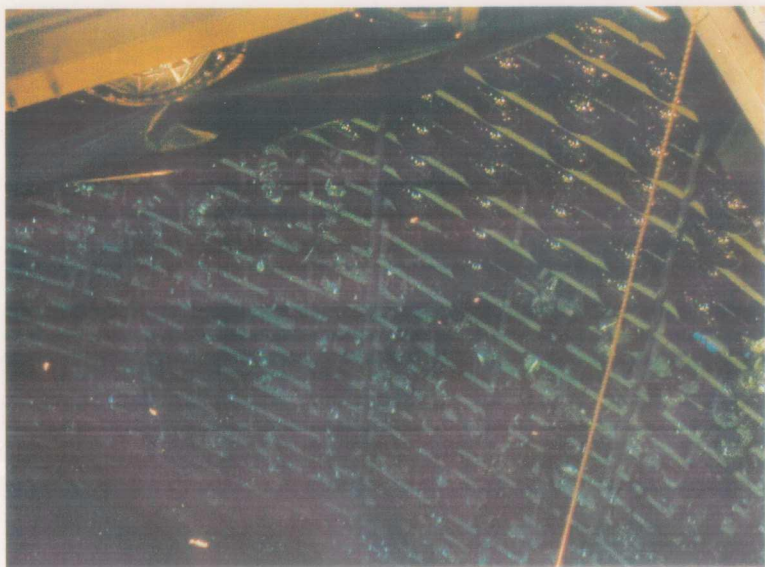


Super-Kamiokande accident...

Before..



After..



Damage, cause and future...

- **Damage:**

6777/11146 broken ID PMTs

1100/1885 broken OD PMTs

Electronics, HV ... are almost OK

- **Cause:**

A PMT at the bottom imploded.

(reason: not sure...)



Shock wave was generated.



Chain reaction.

- **Future: (We hope...)**

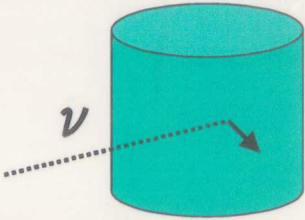
1) Restart experiment late 2002 with about $\frac{1}{2}$ of PMTs.

2) Full recovery in 2006 before JHF-SK neutrino oscillation experiment.

We need strong and continuous support and encouragement from all of you.

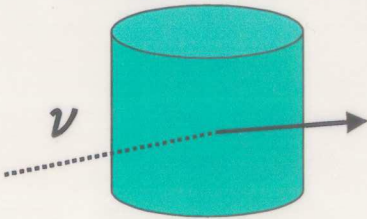
Various types of atmospheric ν events

FC (fully contained)



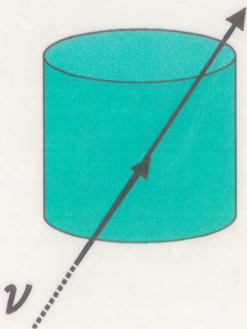
- Both CC ν_e and ν_μ (+NC)
- Need particle identification to separate ν_e and ν_μ
- ~ 10300 events (Super-K)

PC (partially contained)



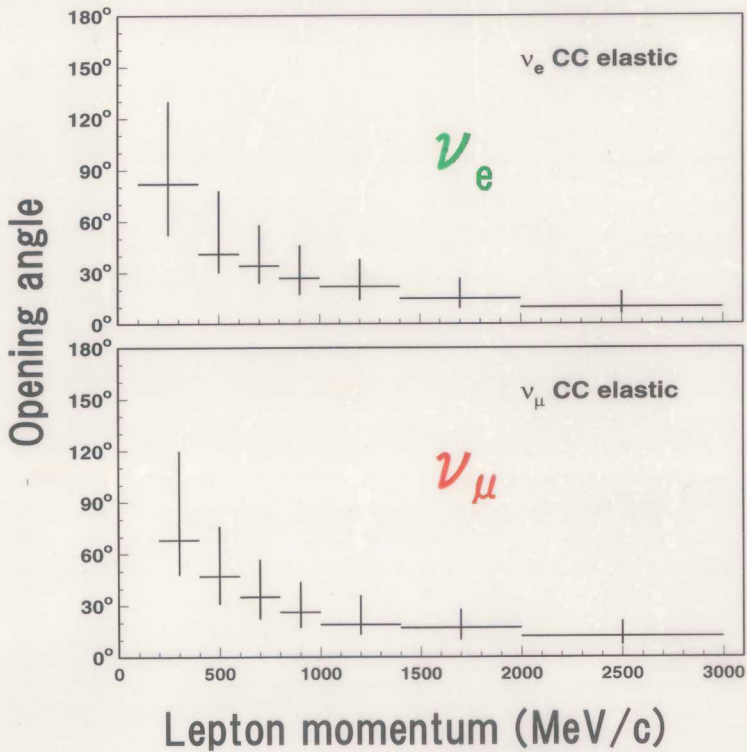
- \simeq CC ν_μ
- ~ 750 events

Upward going muon

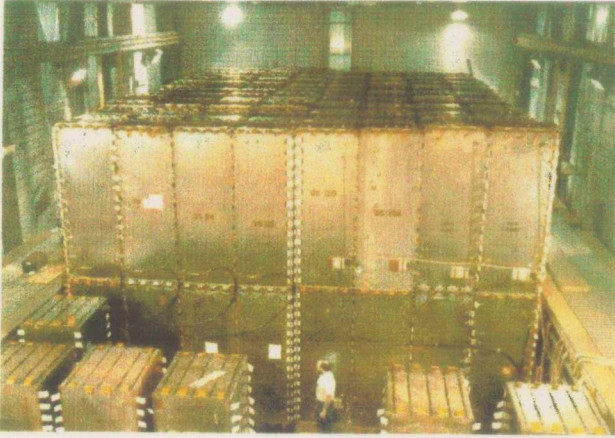


- \simeq CC ν_μ
- ~ 1400 events (through)
- ~ 320 events (stopping)

Angular correlation between neutrino and lepton

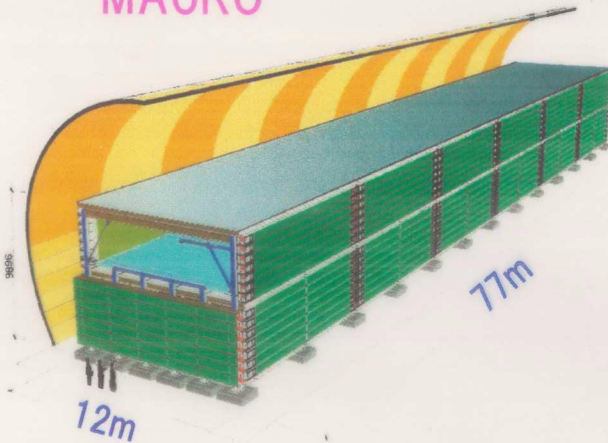


Soudan-2 (1kton tracking calorimeter)



Fully contained events
(~ 320 ev.)

MACRO

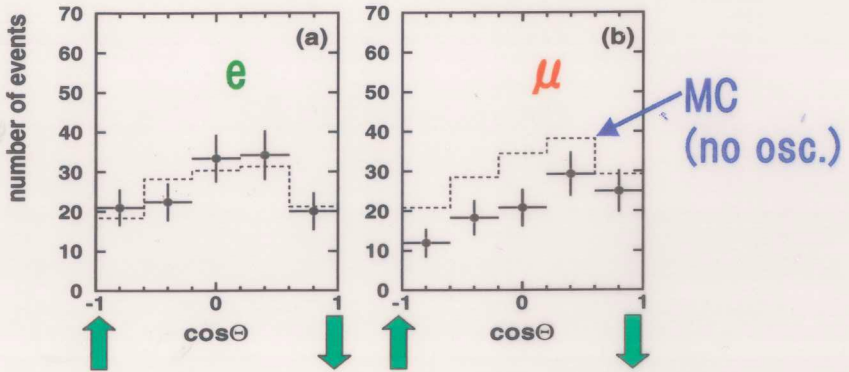


Upward going muons,
partially contained events
(≈ 1000 ev.)

Soudan-2 data

High { energy
angular } resolution sample

(5.1 kton·yr)

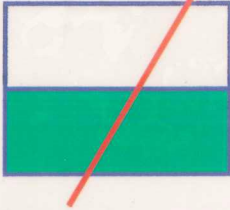


Deficit of upward going muon events.

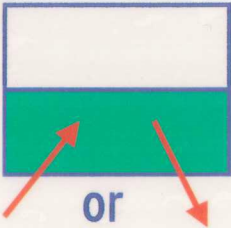
⇒ Consistent with Super-K.

MACRO data

Upward through-going muons

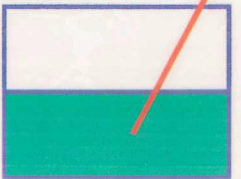


Upward stopping muons + Downward PC events

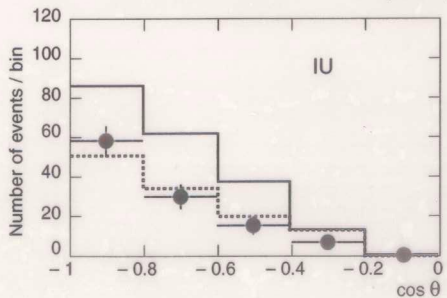
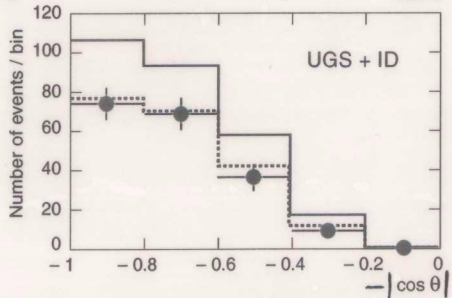
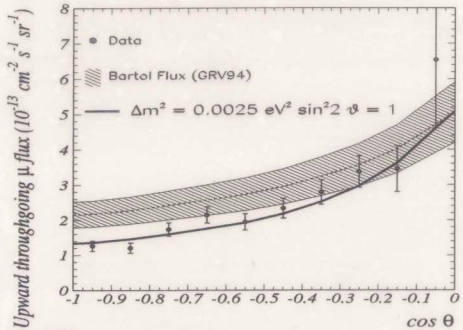


or

Upward PC events

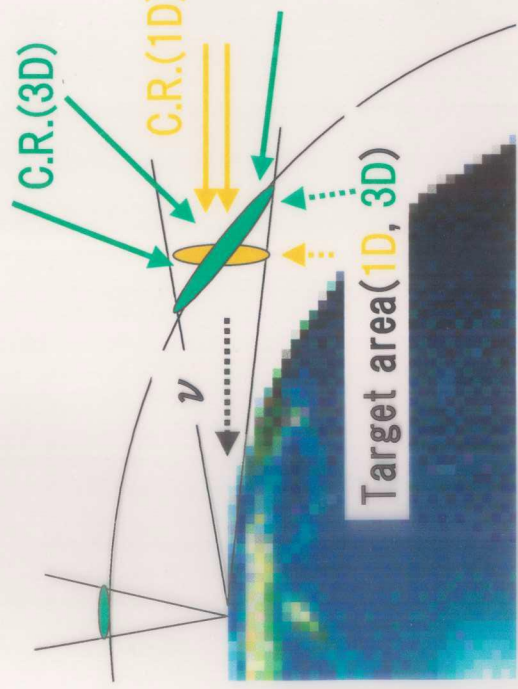


..... $\Delta m^2 = 2.5 \cdot 10^{-3}$

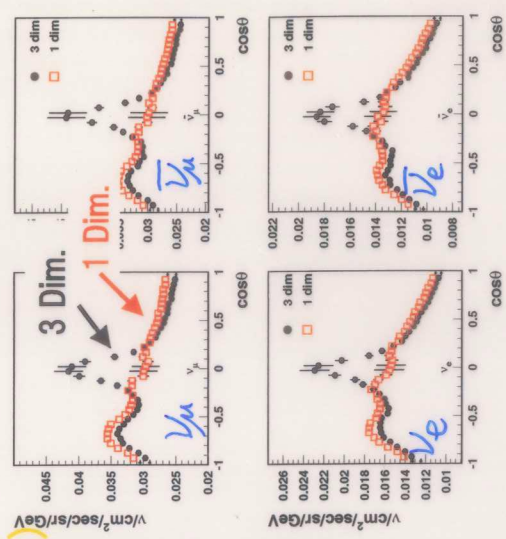


Consistent with $\nu_\mu \rightarrow \nu_\tau$ osc.

New 3 dimensional calculation of the flux



Sub-GeV flux at Kamioka (Fluka group)

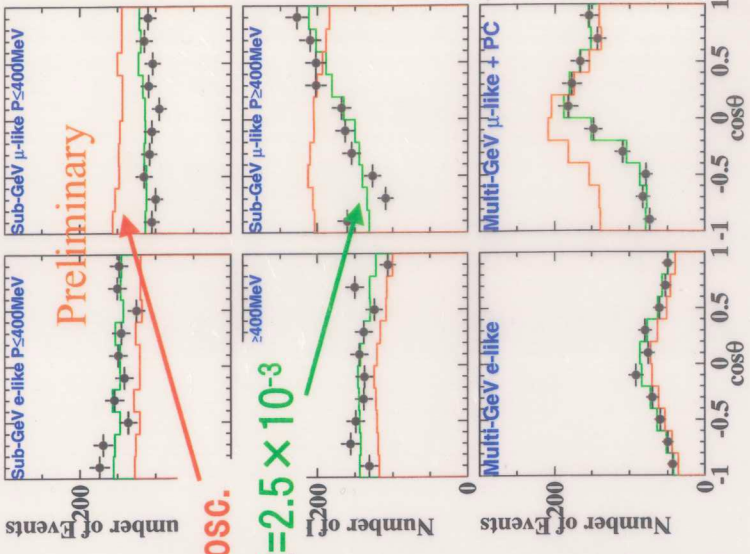
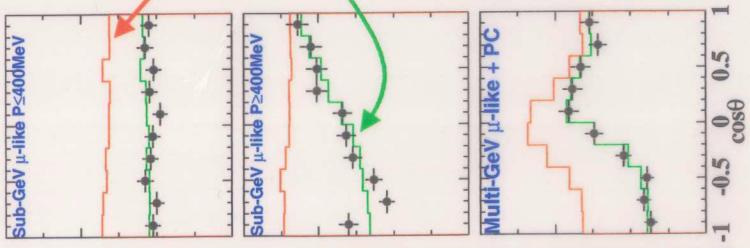
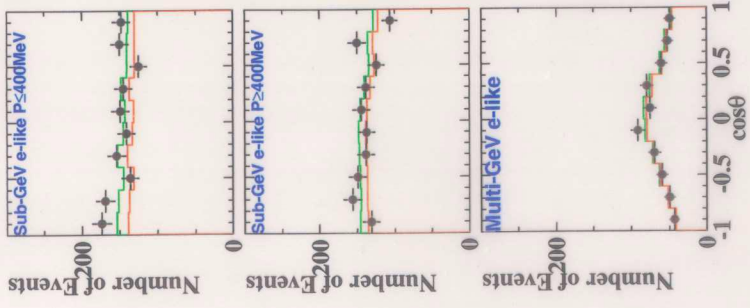


(No big difference for Multi-GeV flux)

Zenith angle distributions (1D vs. 3D)

1D (Honda et al.,)

3D (Fluka group)



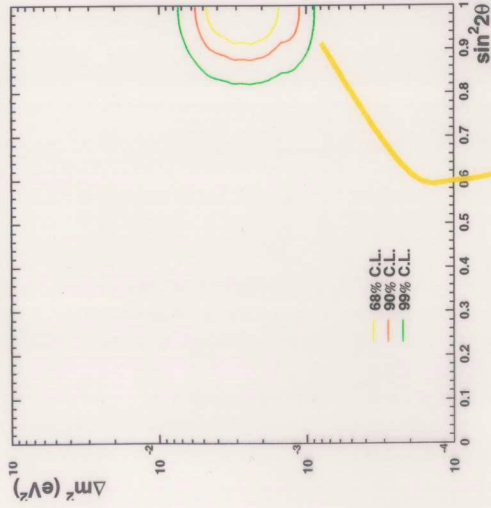
Preliminary

No OSC.

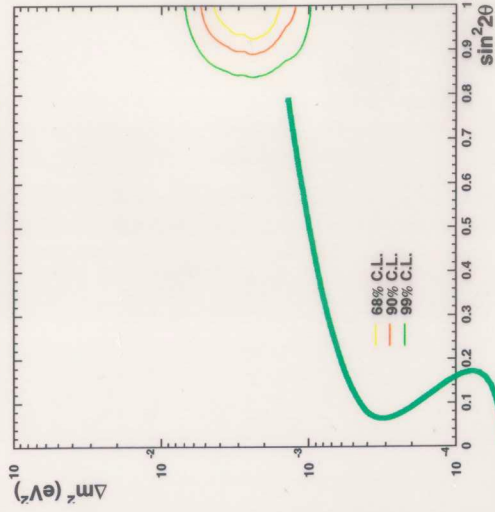
$$\Delta m^2 = 2.5 \times 10^{-3}$$

Allowed parameter region (1D vs. 3D) based on FC+PC events

1D (Honda et al.,)



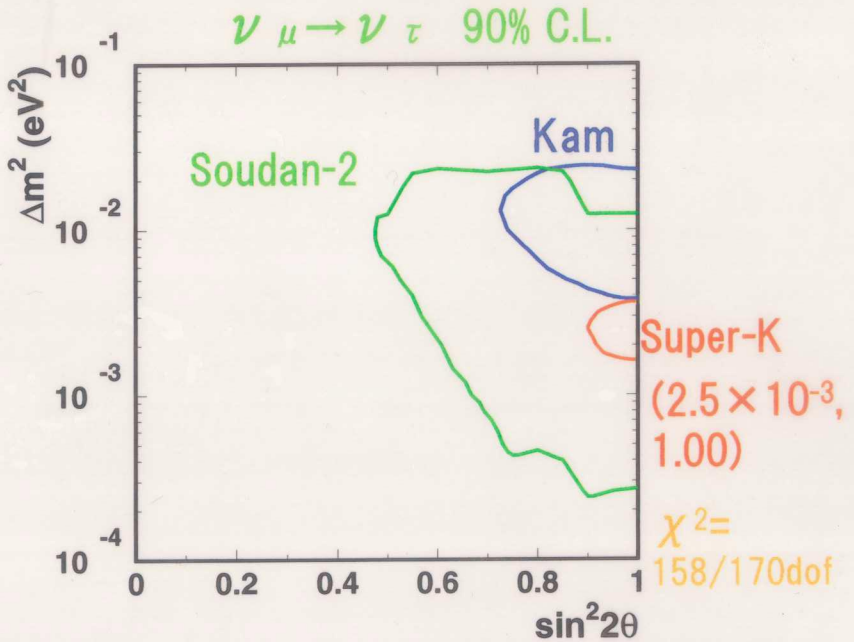
3D (Fluka group)



No big difference! (1D flux in this talk)

(Future: 3D flux will be used.)

Allowed parameter region (contained events + up-going μ 's)



$$\begin{cases} \Delta m^2 = (1.6 \sim 3.6) \times 10^{-3} \text{eV}^2 \\ \sin^2 2\theta > 0.90 \quad (90\% \text{ C.L., Super-K}) \end{cases}$$

3 flavor oscillations

★ Approximation

————— $m_{\nu 3}$

$$\Delta m_{23}^2 = \Delta m_{13}^2 \equiv \Delta m^2$$

===== $m_{\nu 2}$
===== $m_{\nu 1}$

$$\Delta m_{12}^2 = 0$$



$$\theta_{13}, \theta_{23}, \Delta m^2$$

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = 4 |U_{\alpha 3}|^2 |U_{\beta 3}|^2 \sin^2(1.27 \Delta m^2 L/E)$$

$$P(\nu_{\alpha} \rightarrow \nu_{\alpha}) = \frac{1 - 4 |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) \sin^2(1.27 \Delta m^2 L/E)}{1 - 4 |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) \sin^2(1.27 \Delta m^2 L/E)}$$

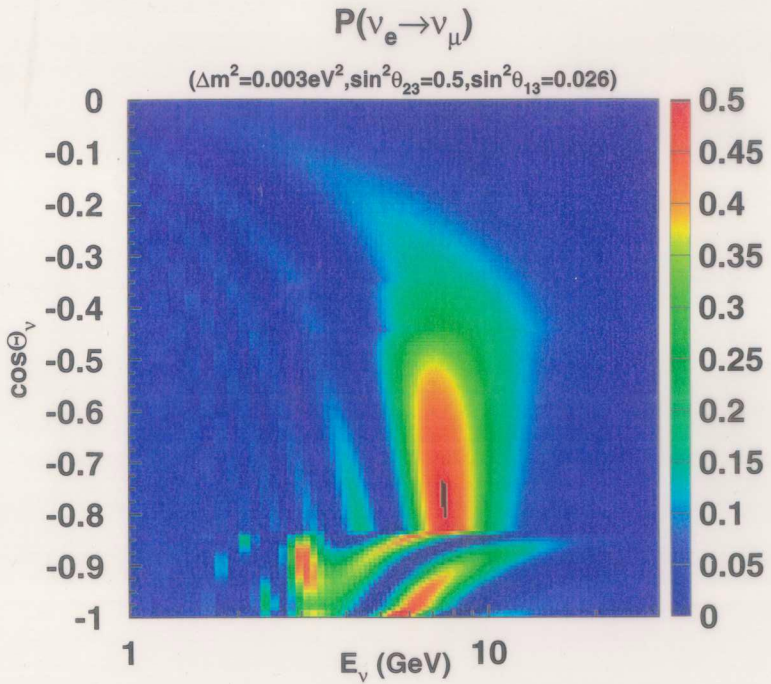
$$|U_{e3}|^2 = \sin^2 \theta_{13}$$

$$|U_{\mu 3}|^2 = \cos^2 \theta_{13} \sin^2 \theta_{23}$$

$$|U_{\tau 3}|^2 = \cos^2 \theta_{13} \cos^2 \theta_{23}$$

3 flavor analysis

➔ Matter effect

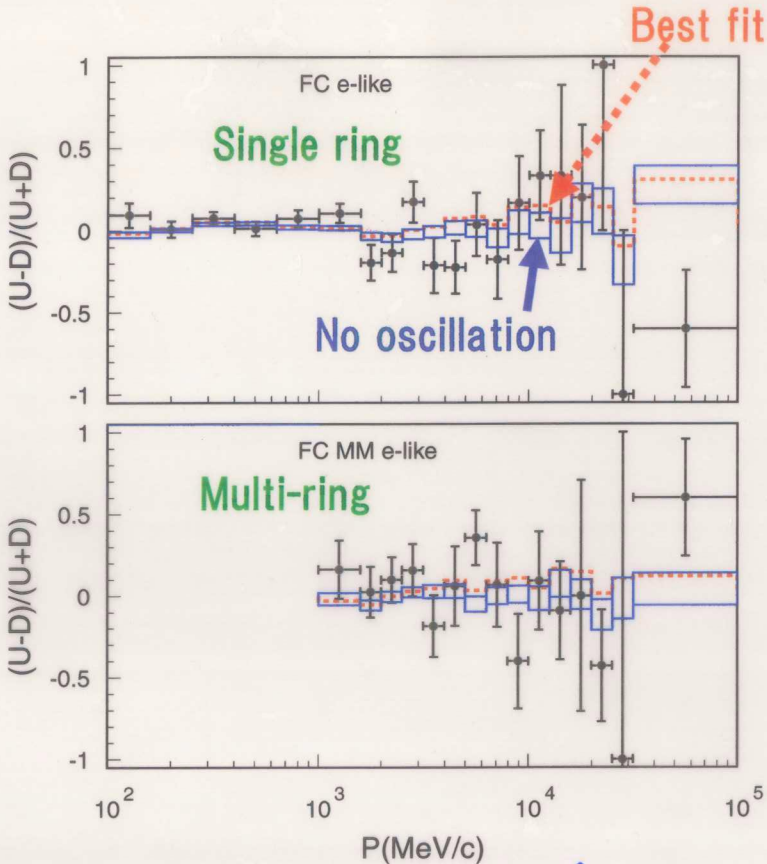


For $\theta_{23} = 45^\circ$;

$$P(\nu_\mu \rightarrow \nu_e) = 1/2 \times P(\nu_e \rightarrow \nu_\mu)$$

➔ No visible oscillation effect for

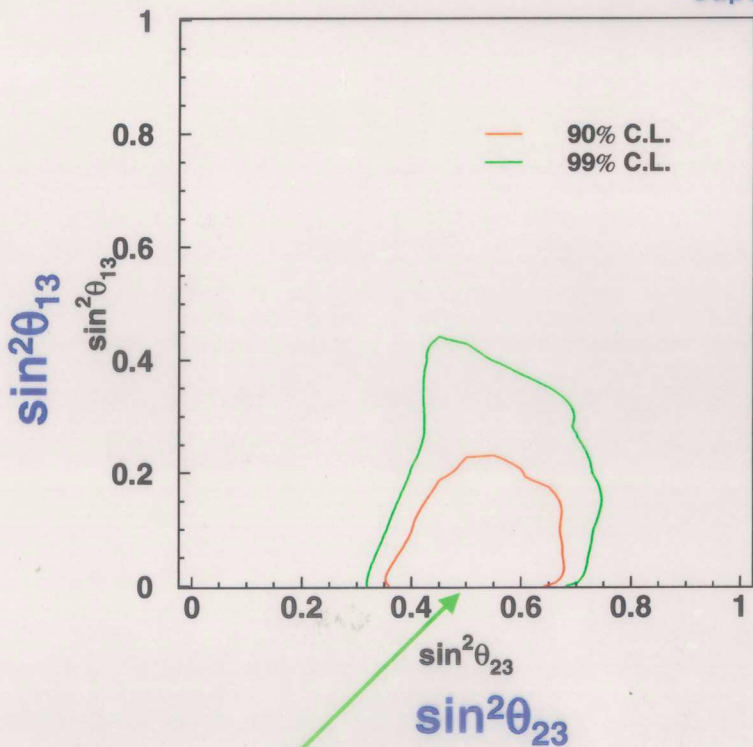
e-like data (up-down asymmetry)



➡ No evidence for ^{excess of} upward-going e-like events in the 10 GeV energy region.

Allowed region for $\sin^2 \theta_{13}$ & $\sin^2 \theta_{23}$

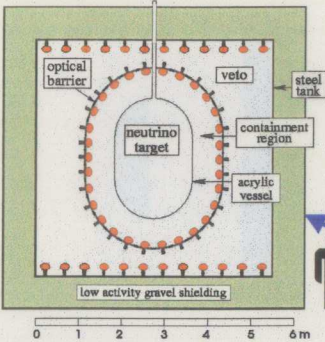
Super-K



Pure maximal $\nu_{\mu} \rightarrow \nu_{\tau}$

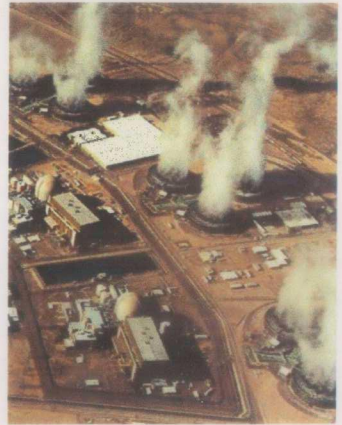
Reactor experiments

CHOOZ (5ton)

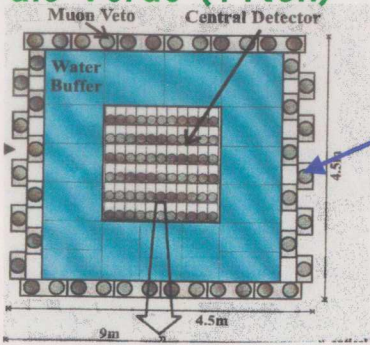


1km

$$\bar{\nu}_e$$



Palo Verde (11ton)



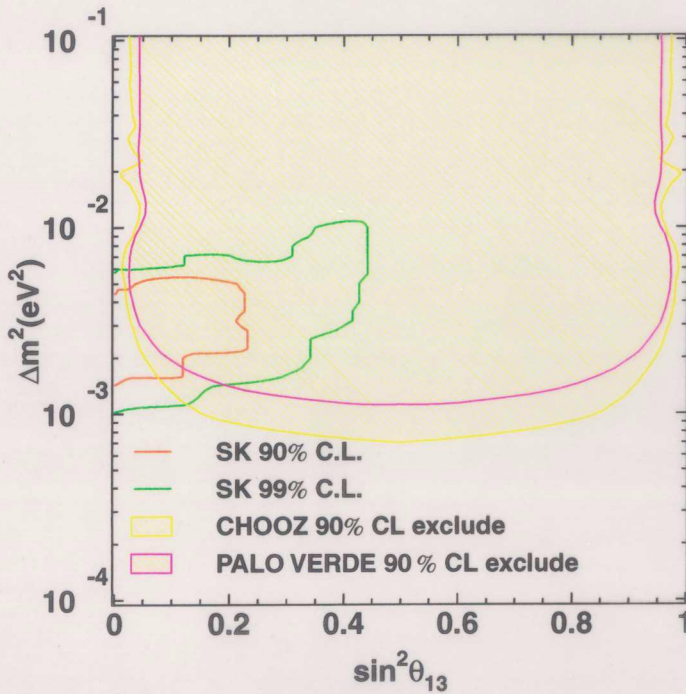
$$\bar{\nu}_e$$

890m

$L/E \sim 300 \rightarrow$ sensitivity: $\Delta m^2 \gtrsim 10^{-3}$

Information on θ_{13}

3 flavor analysis and reactor experiments




Other proposed scenarios

- Oscillations to sterile neutrinos
- Oscillations generated by new physics such as violation of Lorentz invariance
- Flavor Changing Neutral Current
- Neutrino decay



Disfavored at > 99% C.L.



Only $\nu_{\mu} \rightarrow \nu_{\tau}$ oscillations remain

If so,

Neutrino decay ?

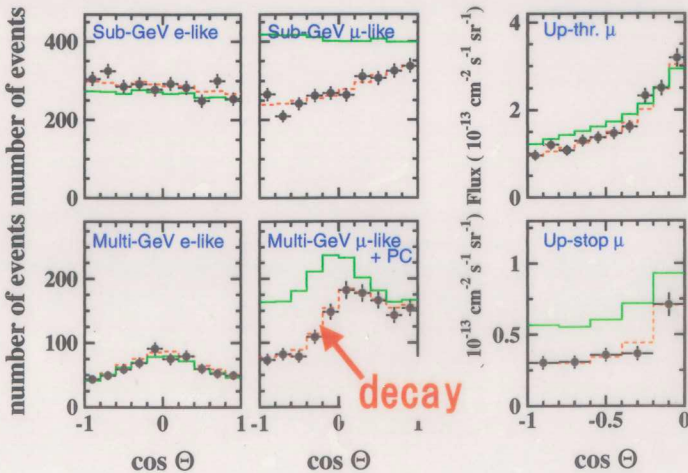
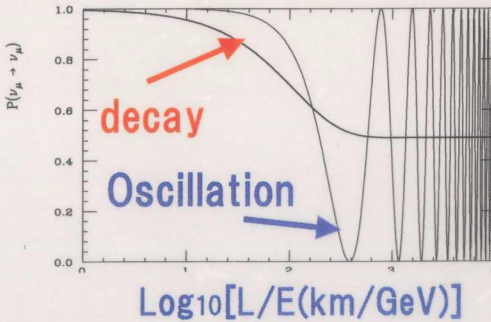
★ Scenario (V.Barger et al., PLB 462 (1999) 109):

$$\nu_\mu = \cos\theta \nu_2 + \sin\theta \nu_3 \quad (\nu_\tau = -\sin\theta \nu_2 + \cos\theta \nu_3)$$

decay \rightarrow X

For $\Delta m^2 \rightarrow 0$; $P(\nu_\mu \rightarrow \nu_\mu) = (\sin^2\theta + \cos^2\theta e^{-\alpha L/2E})^2$

$$\alpha = m/\tau$$



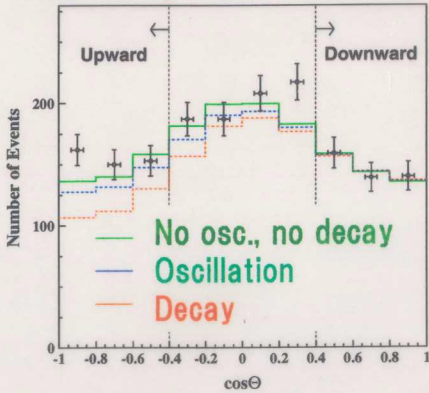
Decay scenario explains the CC data well.

Neutrino decay ?

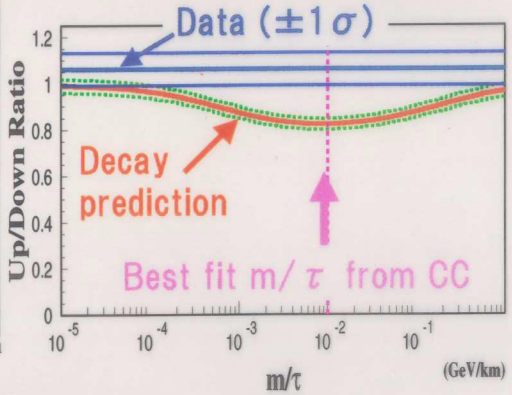
However, **NC** events should also decrease due to decay.

New!

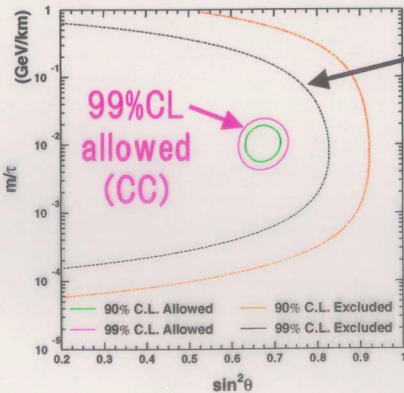
Multi-ring e-like (29% NC)



Up/down ratio vs. m/τ



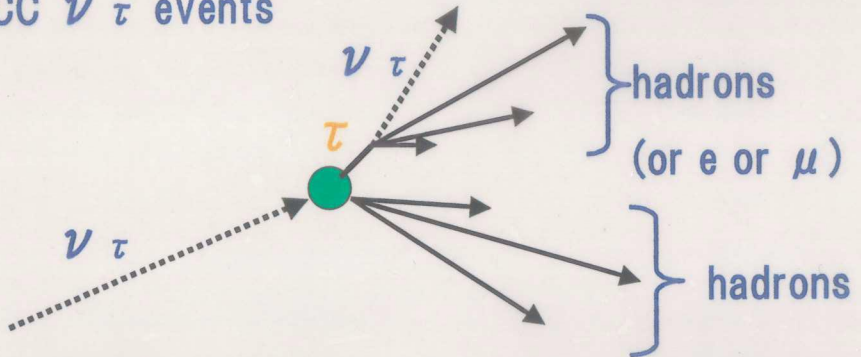
Excluded & allowed regions



Neutrino decay cannot explain CC and NC events simultaneously.

Search for CC ν_τ

CC ν_τ events



● Many hadrons
(But no big difference with other events .)

↳ BAD ⇨ τ - likelihood analysis

● Upward going only

↳ GOOD ⇨ Zenith angle

Only ~ 1.0 CC ν_τ FC events/kton \cdot yr

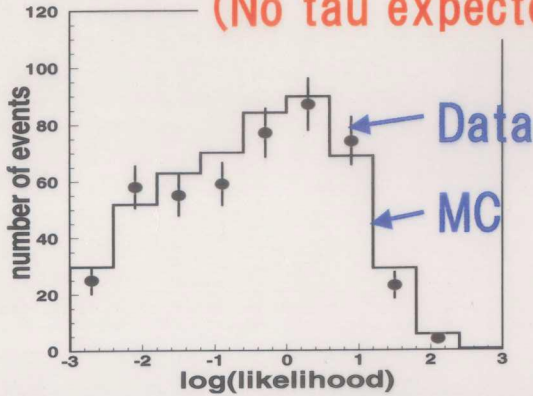
(BG (other ν events) ~ 130 ev./kton \cdot yr)

Tau likelihood distribution

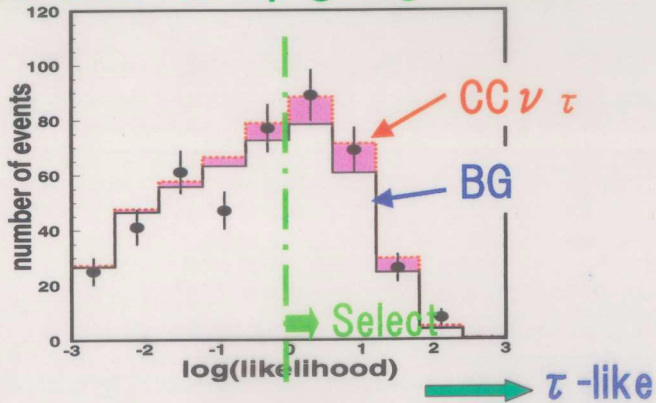
Likelihood dist. for down going events

(No tau expected)

OK!



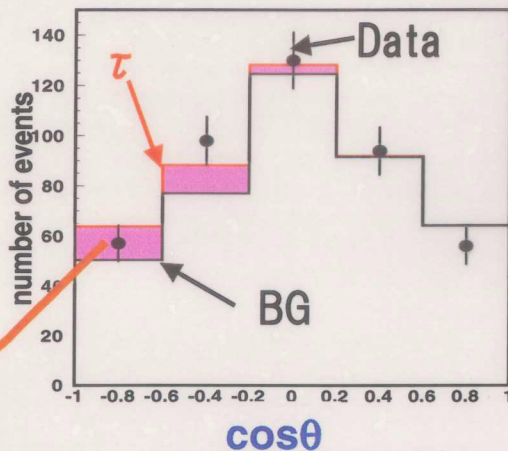
Likelihood dist. for up going events



Zenith angle distribution

Zenith angle for τ enriched sample

Super-K



$\epsilon = 44\%$

τ production: 66 ± 41 $^{+25}_{-18}$ Likelihood method (1)

Other analyses

79 $^{+44}_{-40}$ Likelihood method(2)
(Energy flow etc)

92 ± 35 $^{+17}_{-23}$ Neural network method

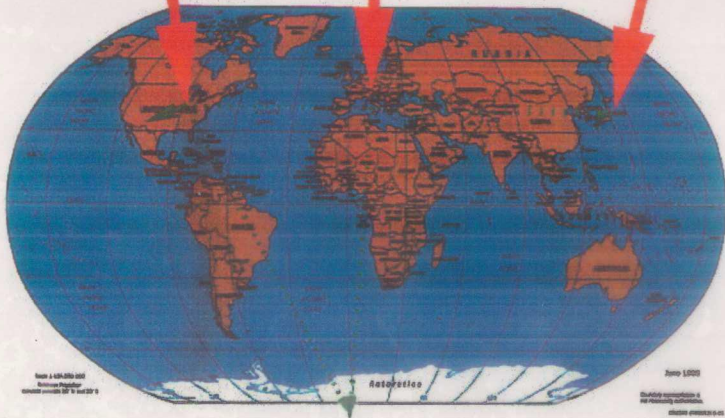
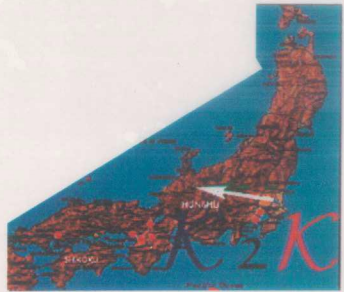
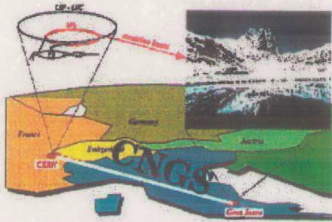


(Expected tau ev. = 74)

Consistent with τ production at the $1.5 \sim 2\sigma$ level (VERY preliminary).

Long baseline experiments

CERN to Gran Sasso Neutrino Beam



Map © 2000-2001
National Geographic
www.nationalgeographic.com

June 2000
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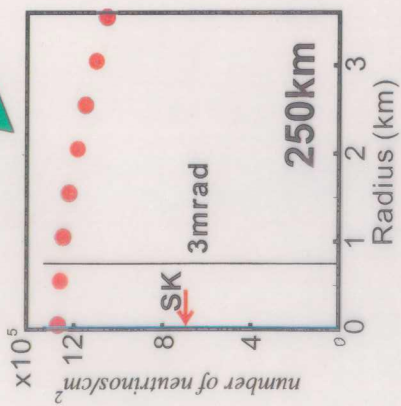
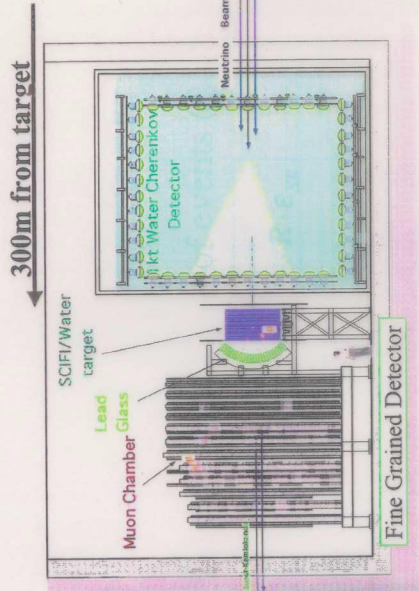
Color

K2K long baseline experiment



$L=250\text{km}, \langle E_\nu \rangle = 1.5\text{GeV}$

Front Neutrino Detector (FD)



Status of the K2K long baseline experiment

4.8×10^{19} p.o.t. (50% of design, 1999 – Jul. 2001)

New

$\Delta m^2 (\times 10^{-3} eV^2)$

Obs. No Ocsi. 3 5 7

FC 22.5kt 56 80.6 $+7.3$ 52.4 34.6 29.2
 -8.0

1-ring 32 48.4 \pm 6.7 28.1 17.8 16.6

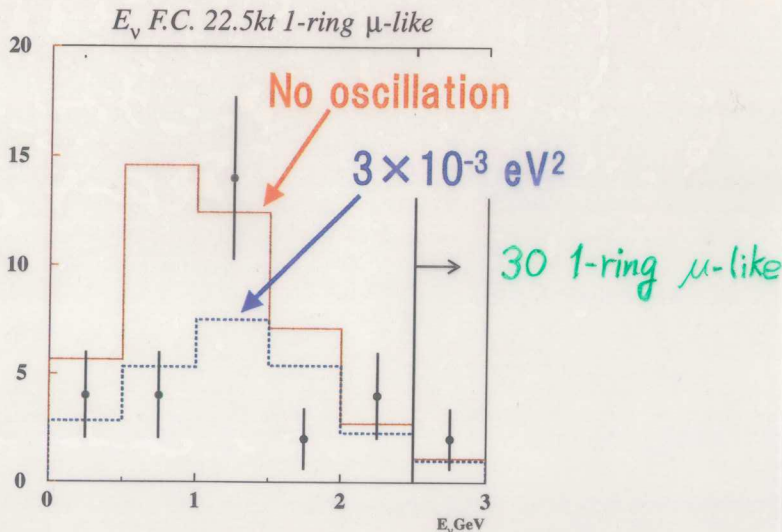
μ -like 30 44.0 \pm 6.8 24.4 14.6 13.5

e-like 2 4.4 \pm 1.7 3.7 3.2 3.0

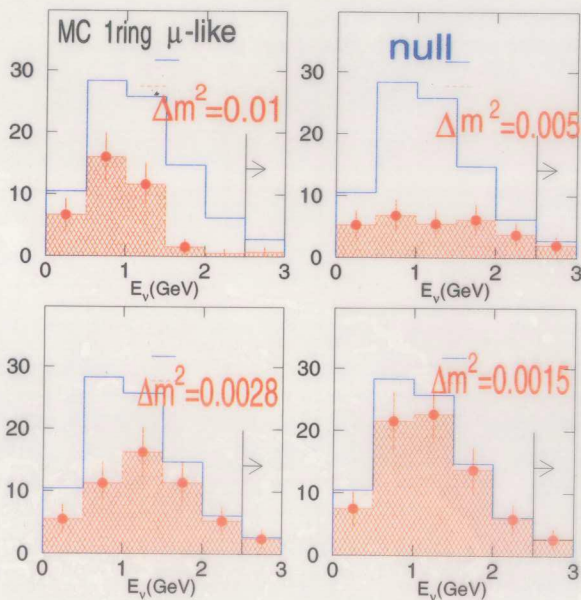
multi ring 24 32.2 \pm 5.3 24.3 16.8 12.6

Consistent with $\Delta m^2 = 3 \times 10^{-3} eV^2$

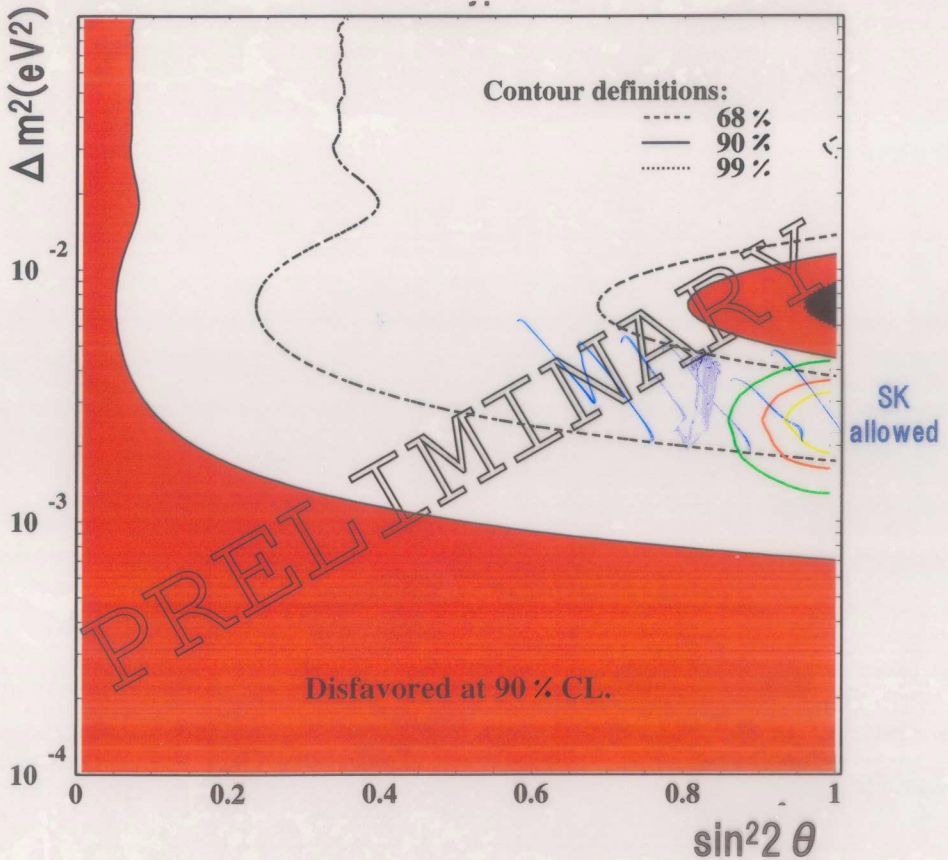
Reconstructed neutrino energy distribution



How about the expected distributions....
(1×10^{20} pot)



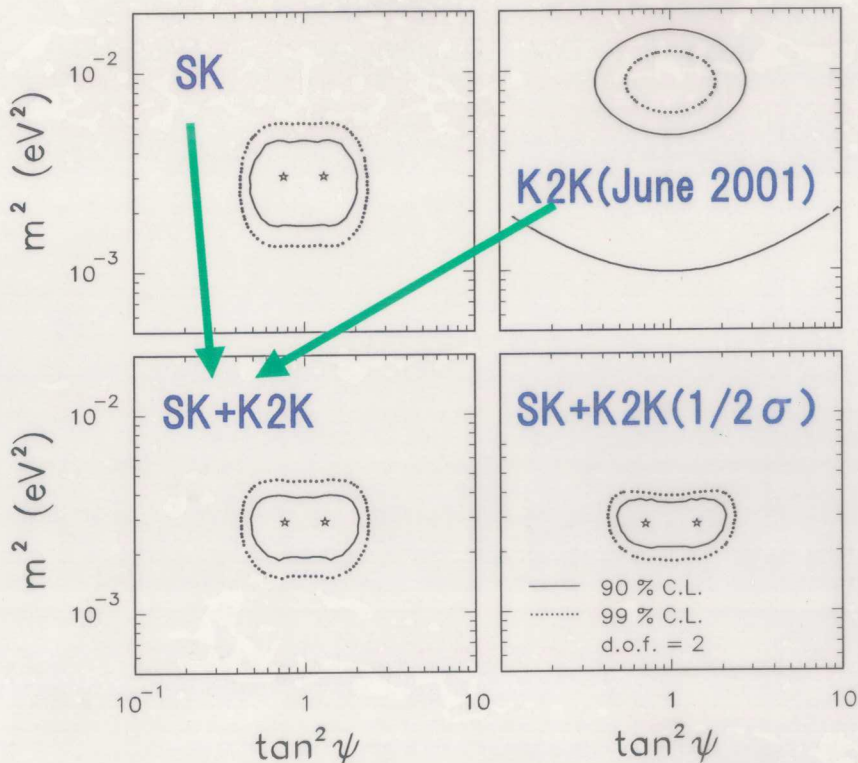
K2K allowed parameter region (Based on the number of events)



SK + K2K combined analysis

G.L.Fogli et al., hep-ph/0110039

SK + K2K, 2ν oscillations



➡ SK+K2K reduces the Δm^2 uncertainty by 20-30%.

Important to reduce the K2K total rate uncertainty.

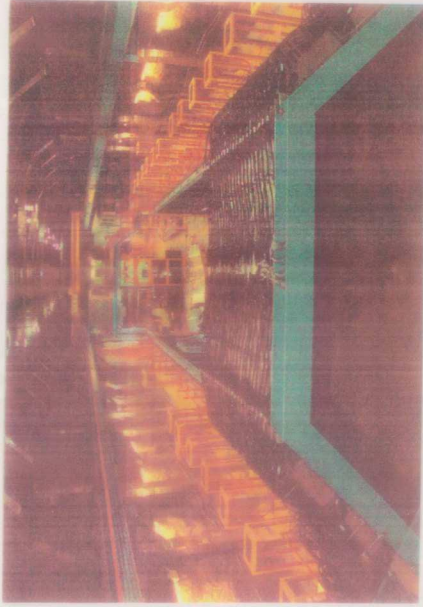
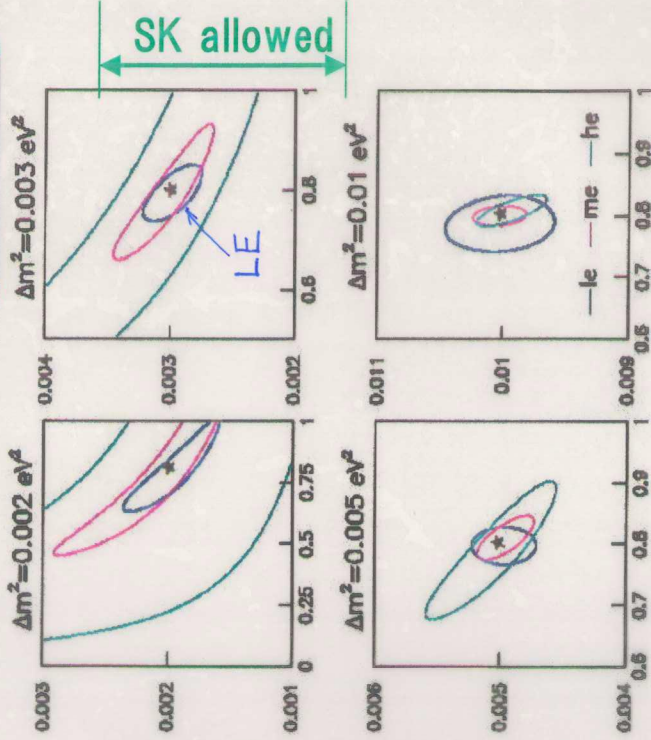
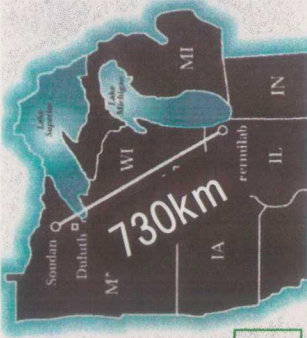
Status of future LBL ν osc. exp's (1)

MINOS

Fermilab \rightarrow Soudan (start: 2004 ?)

5.4 kton Fe/Plastic scinti.
(w/ WLS) detector

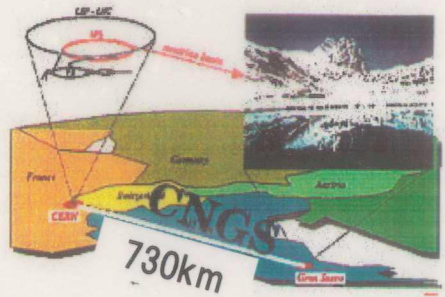
Sensitivity (2yrs)



Status of future LBL ν osc. exp's (2)

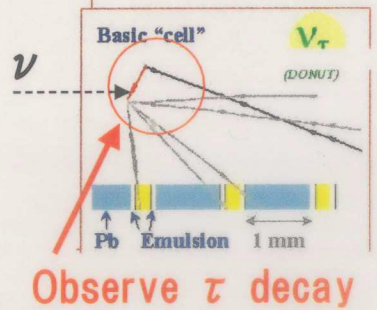
CNGS main physics goal: Direct detection of CC ν τ events

CERN to Gran Sasso Neutrino Beam



OPERA (Start: 2005)

μ spectrometer



5yr Sensitivity (target mass : 1.8kton)

$\Delta m^2 (\times 10^{-3})$	1.5	3.0	5.0	Background
Number of ev.	4.1	18.3	44.1	0.57

μ bundle by Icarus T600

ICARUS

600ton detector was successfully tested on surface.



Multi-kton detector was recently proposed. Waiting for approval...

Status of future LBL ν osc. exp's (3)

JHF - Kamioka

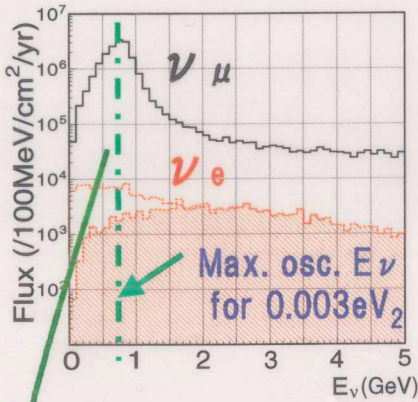
(start 2007 ?)



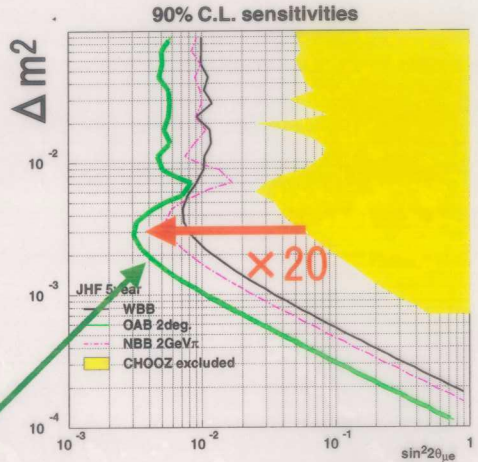
× 100 more intensity than K2K,
 $E_\nu < 1\text{GeV}$



Precise measurements of Δm^2 , θ_{23} , θ_{13}



Sensitivity to θ_{13}



$0.5 \cdot \sin^2 2 \theta_{13}$



(2nd phase: CP violation (w/ Hyper-Kamiokande))

Summary

- Neutrino oscillation signal is consistent with **pure $\nu_{\mu} \rightarrow \nu_{\tau}$** . (atmospheric and reactor experiments)
- Allowed parameter region:
$$\left\{ \begin{array}{l} \sin^2 2\theta_{23} > 0.90 \\ 1.6 < \Delta m_{23}^2 < 3.6 \times 10^{-3} \text{ eV}^2 \end{array} \right.$$
- Many other possibilities are disfavored.
- Search for CC ν_{τ} events: in progress.
- K2K data are consistent with SK atmospheric neutrino data.
- In the near future, long baseline experiments will give us much more precise information on neutrino mass and mixing.