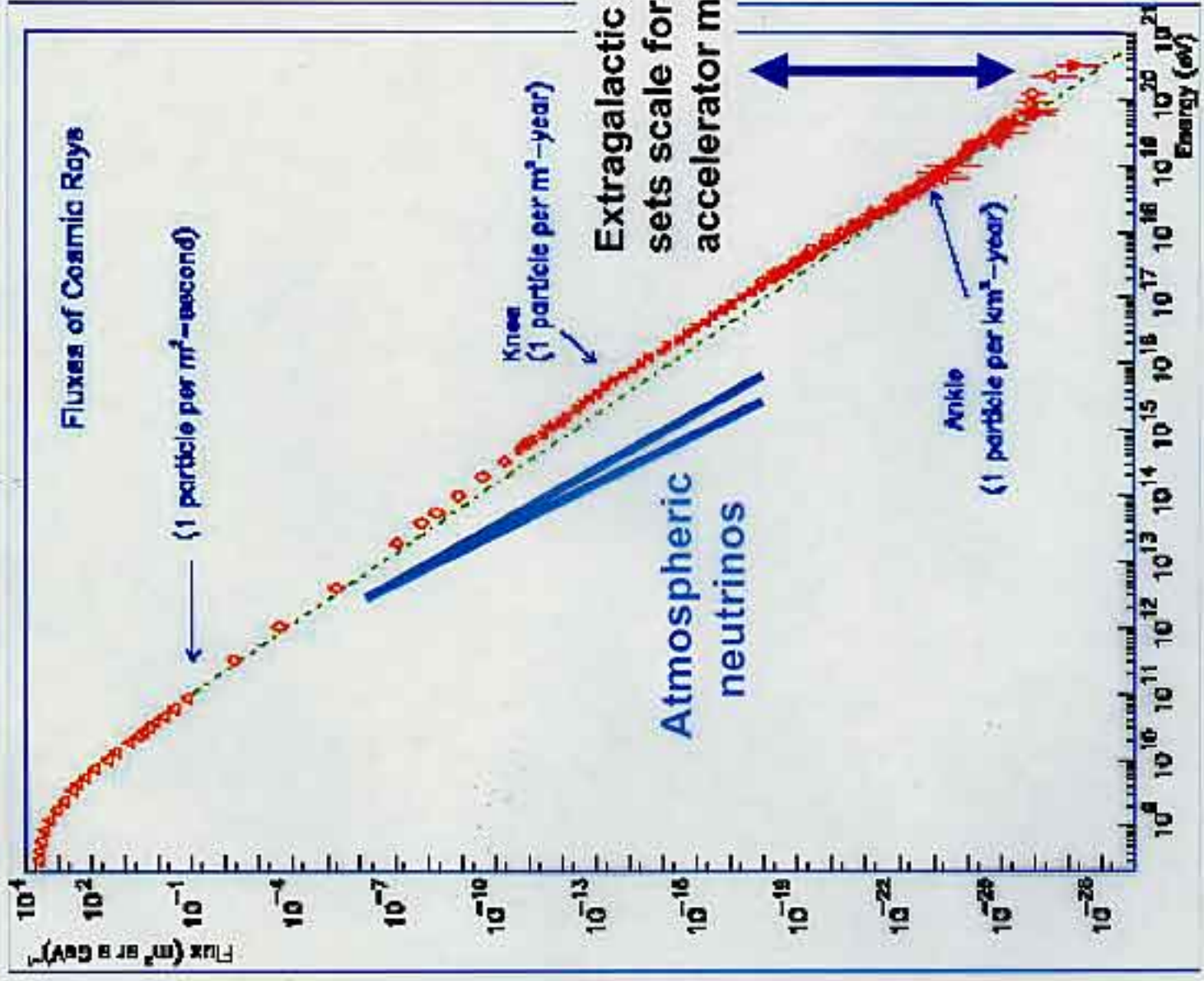


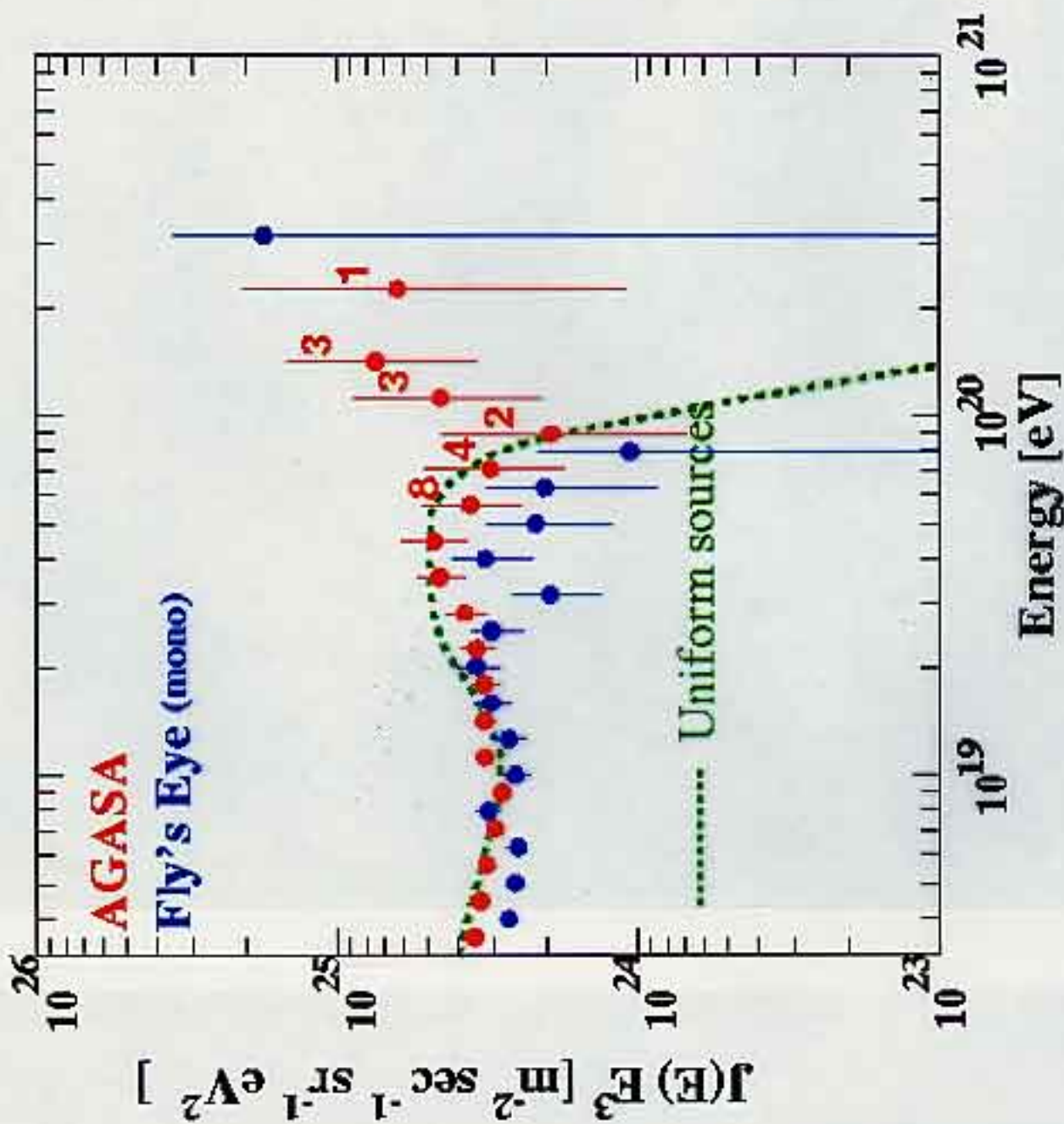
ULTRA HIGH ENERGY NEUTRINOS

- Motivate ν astronomy
- UHE ν N cross sections
- Transparency of earth $\rightarrow \nu_{\tau}$
- Detection of UHE ν_{τ} 's

Alan Martin
WIN02

Cosmic Ray spectrum





Origin of UHE cosmic rays? a mystery

Bottom Up

V. powerful astrophysical objects may accelerate particles to $\gtrsim 10^{20}$ eV, but

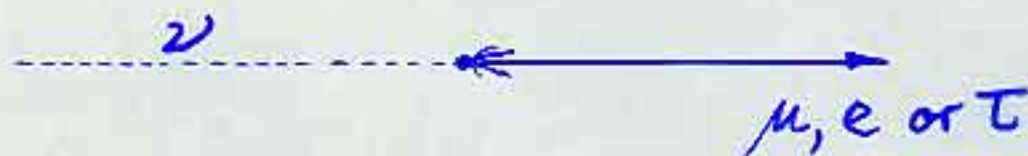
- source environment degrades energy
- GZK cut-off: γ_{CMB} prevents p, nuclei, γ 's reaching us if $E. \gtrsim 10^{20}$ eV (unless source 'nearby')

Top Down

Collapse of topological defects } attractive
Decay of super massive relic particles } possibilities, but

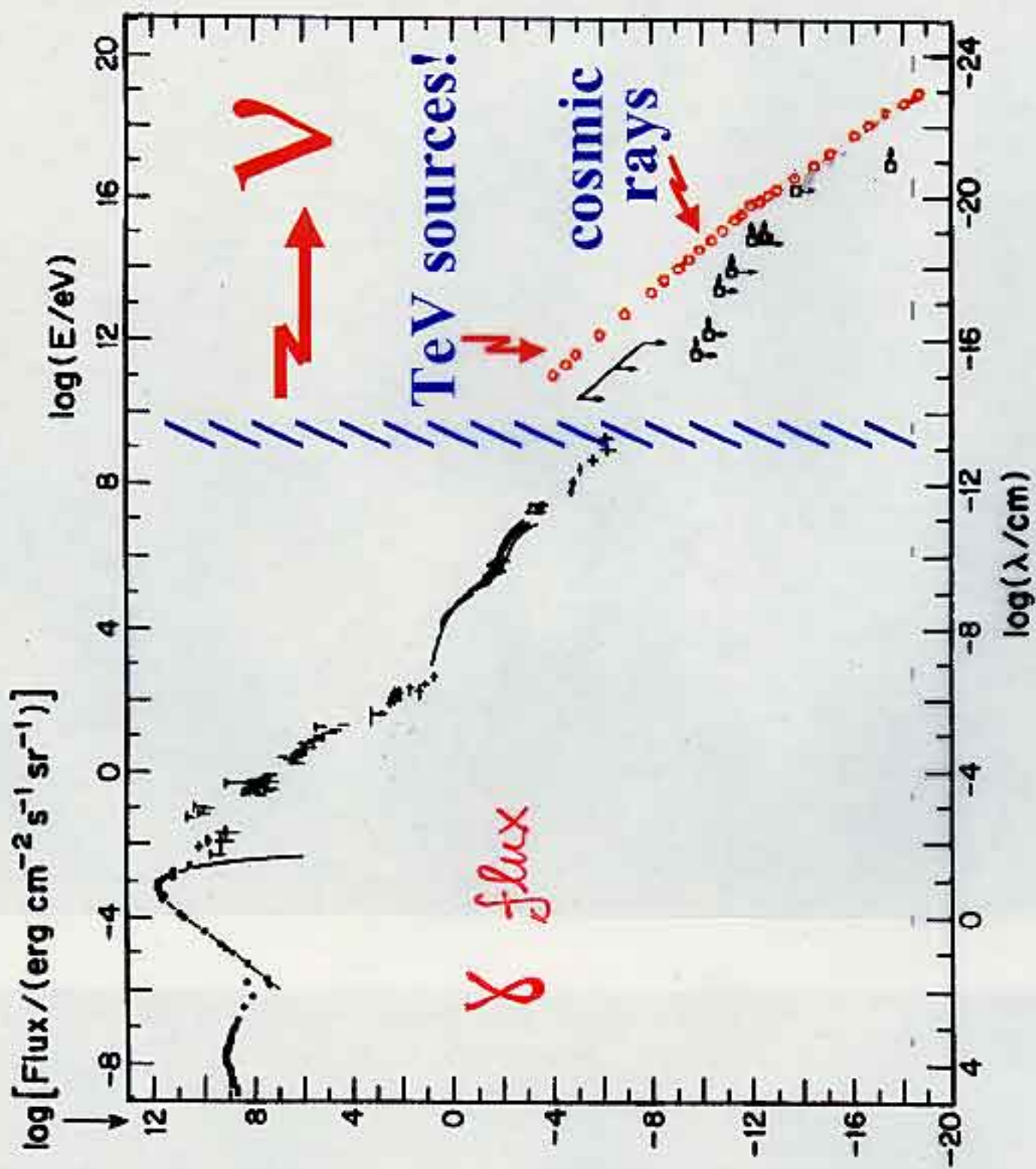
severe constraints from observed γ flux.

→ UHE \gtrsim astronomy

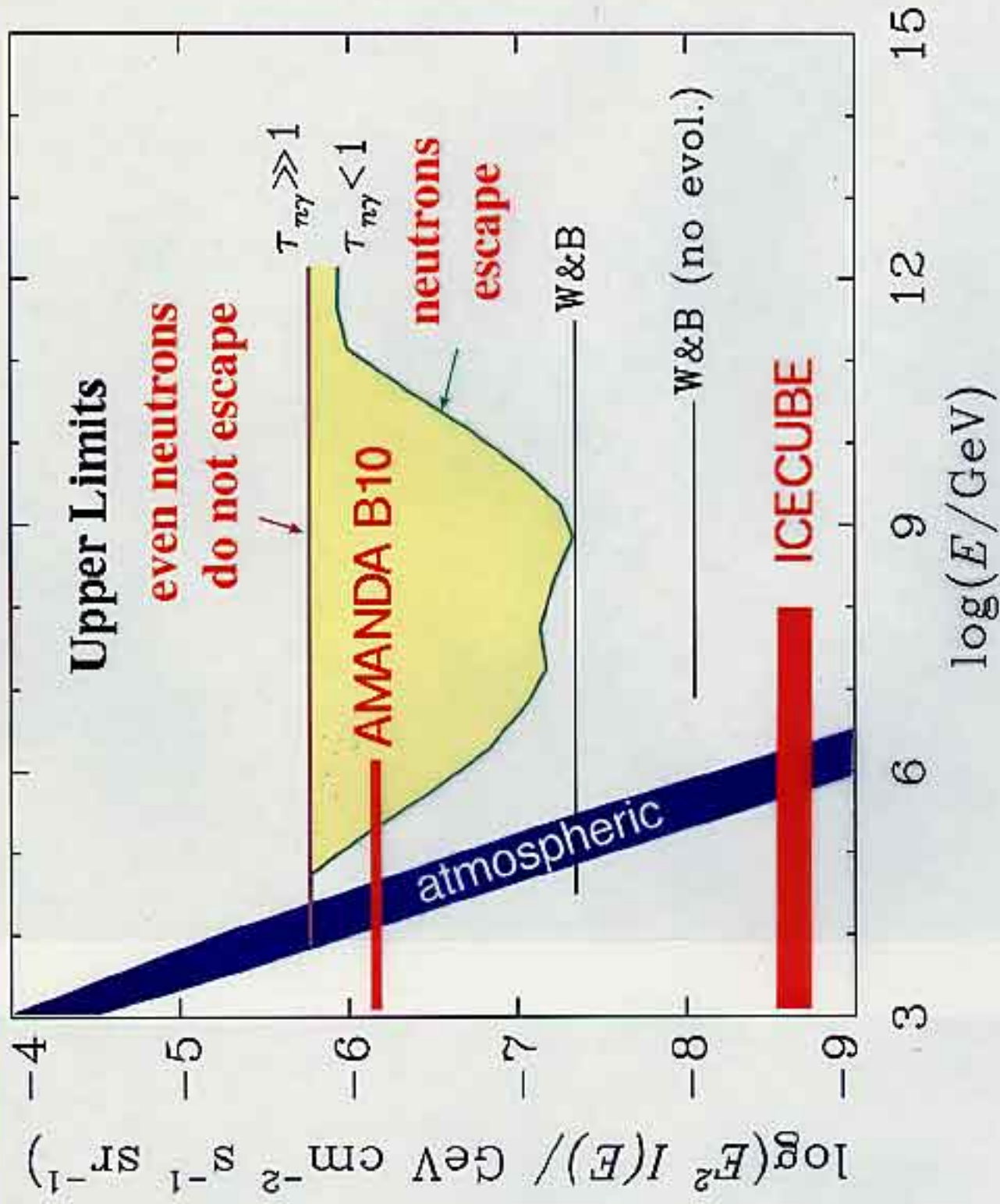


$\gtrsim 1 \text{ TeV}$

direction \sim conserved

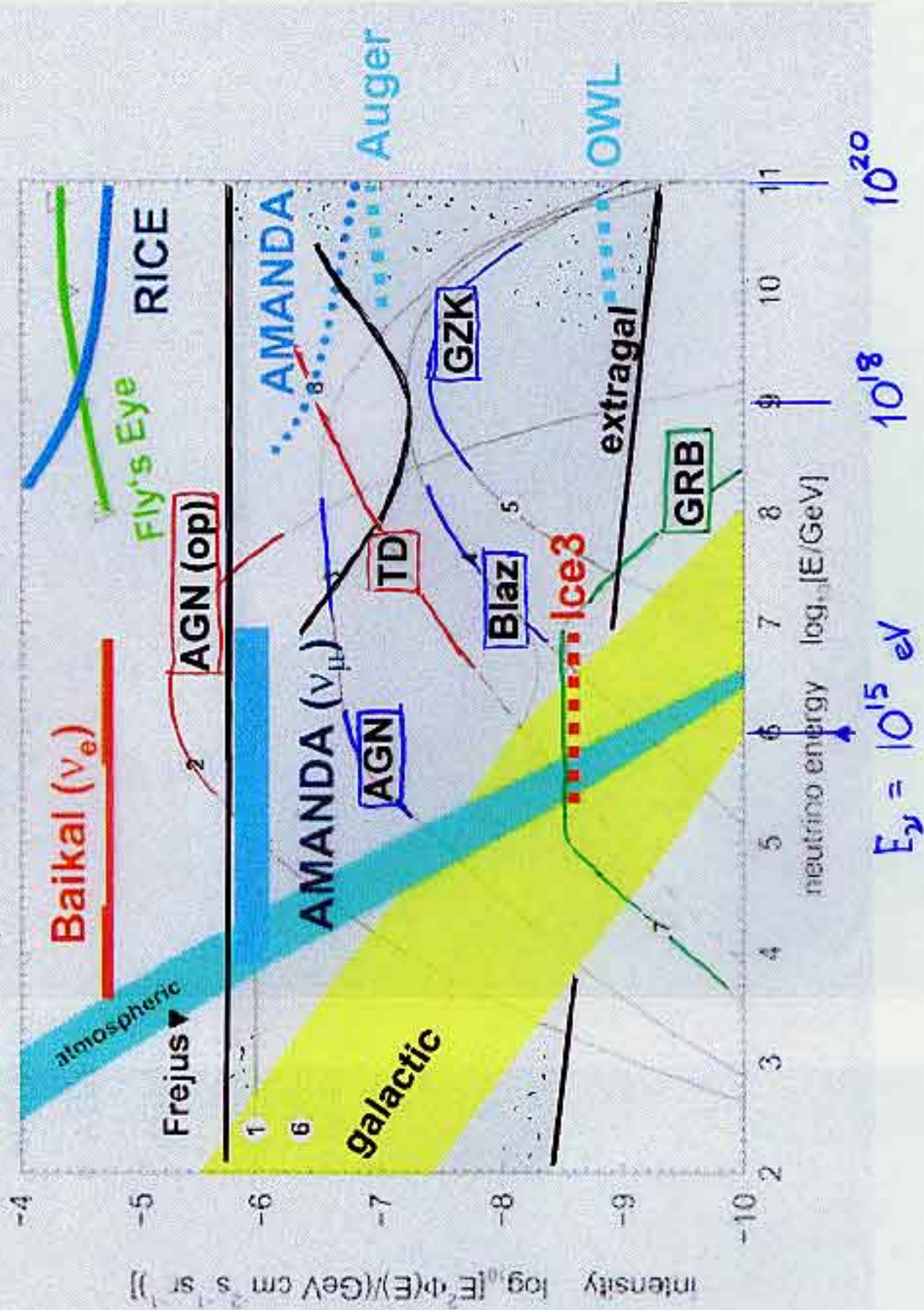


neutrinos associated with the source of the cosmic rays?

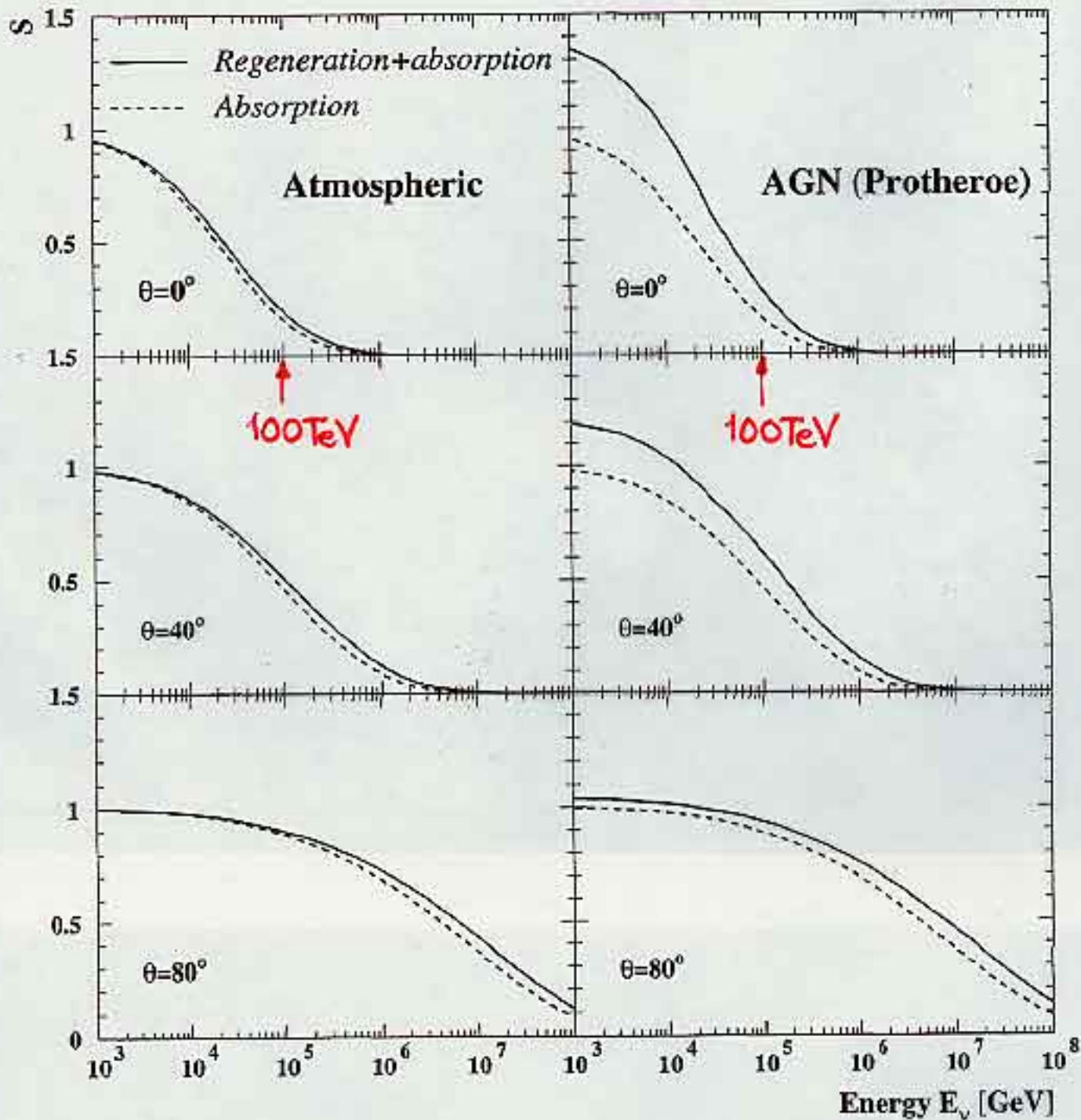


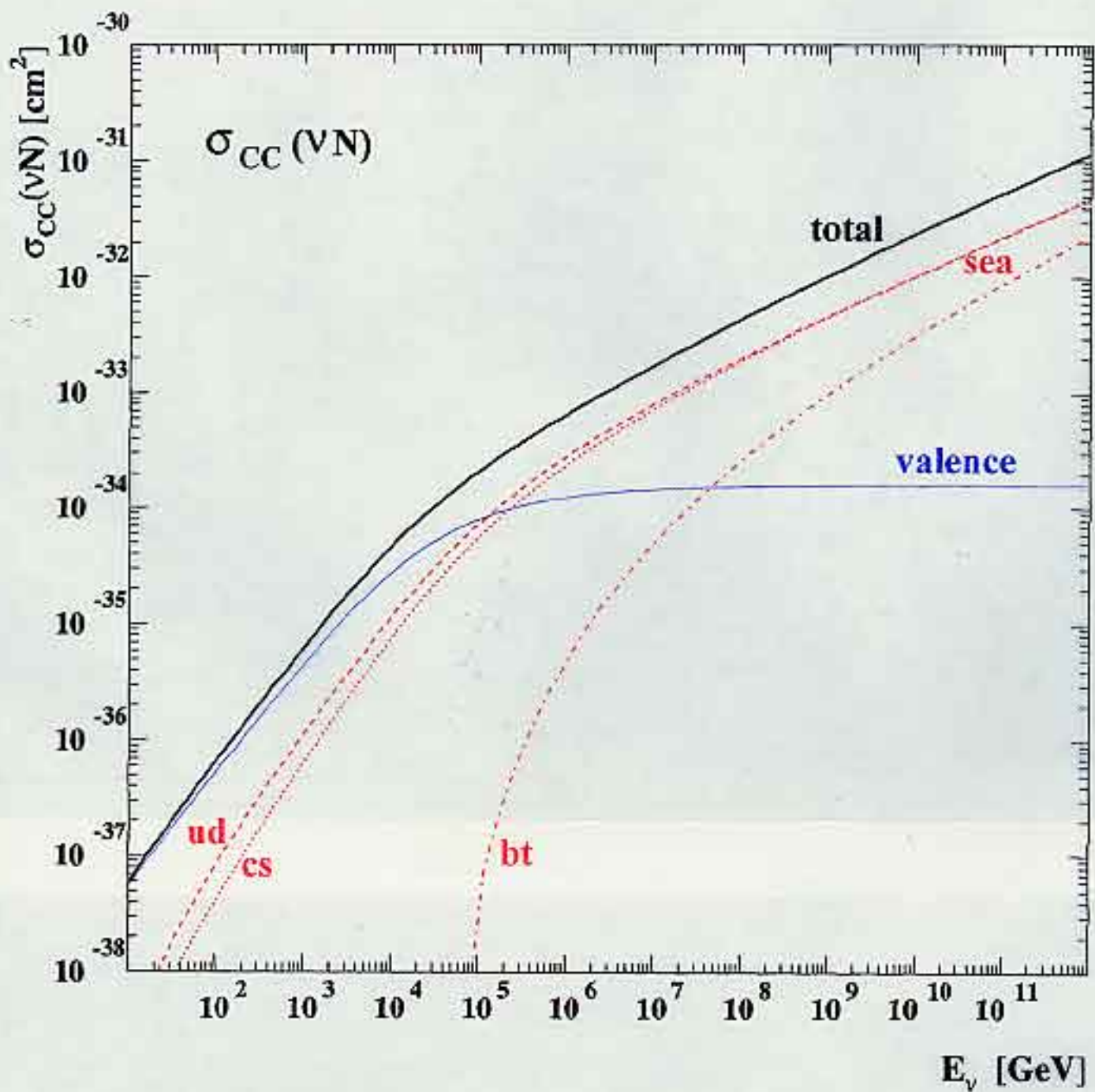
$\nu_\mu + \bar{\nu}_\mu$ from diffuse sources

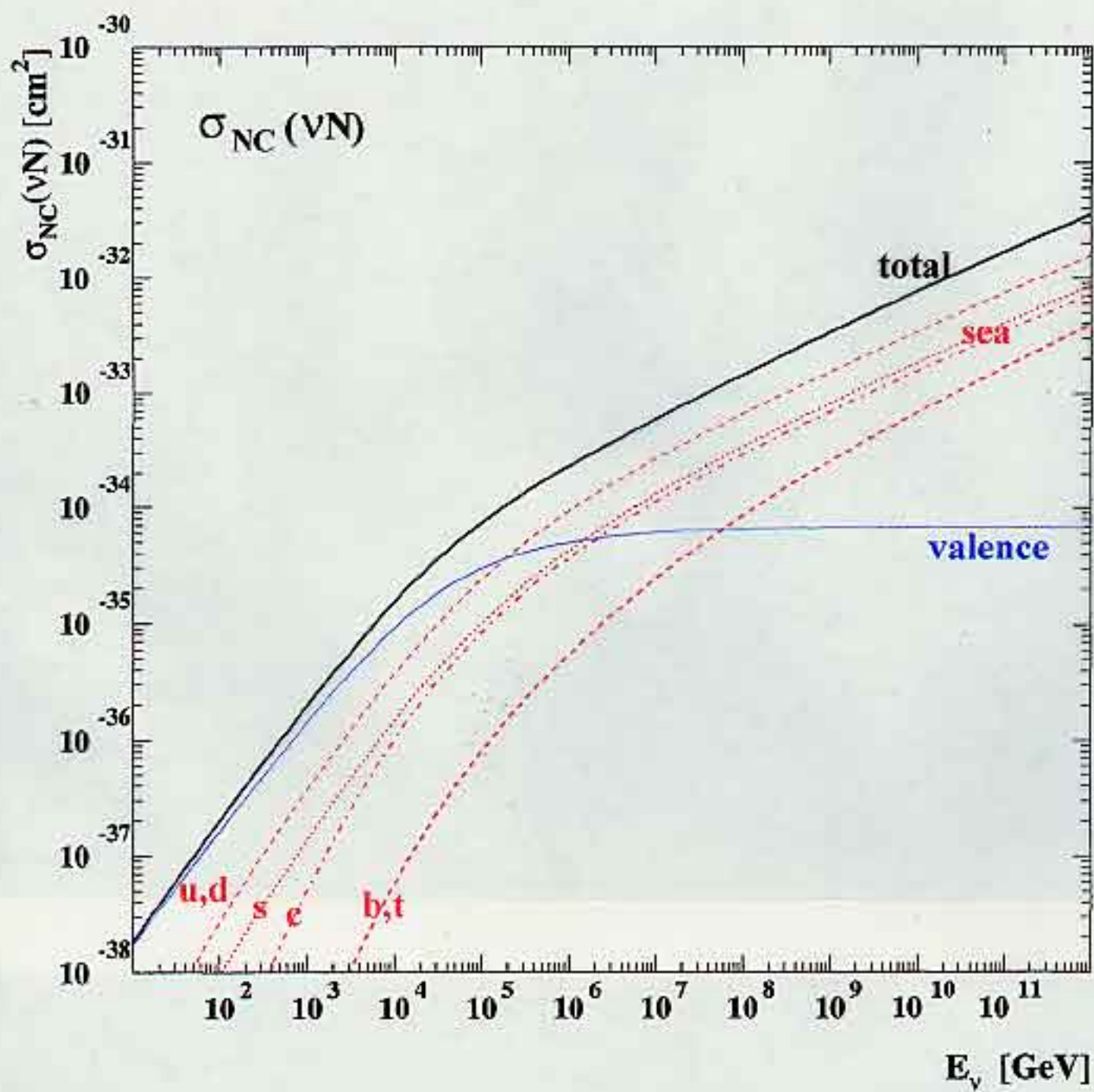
from Learned, Mannheim/Barwick

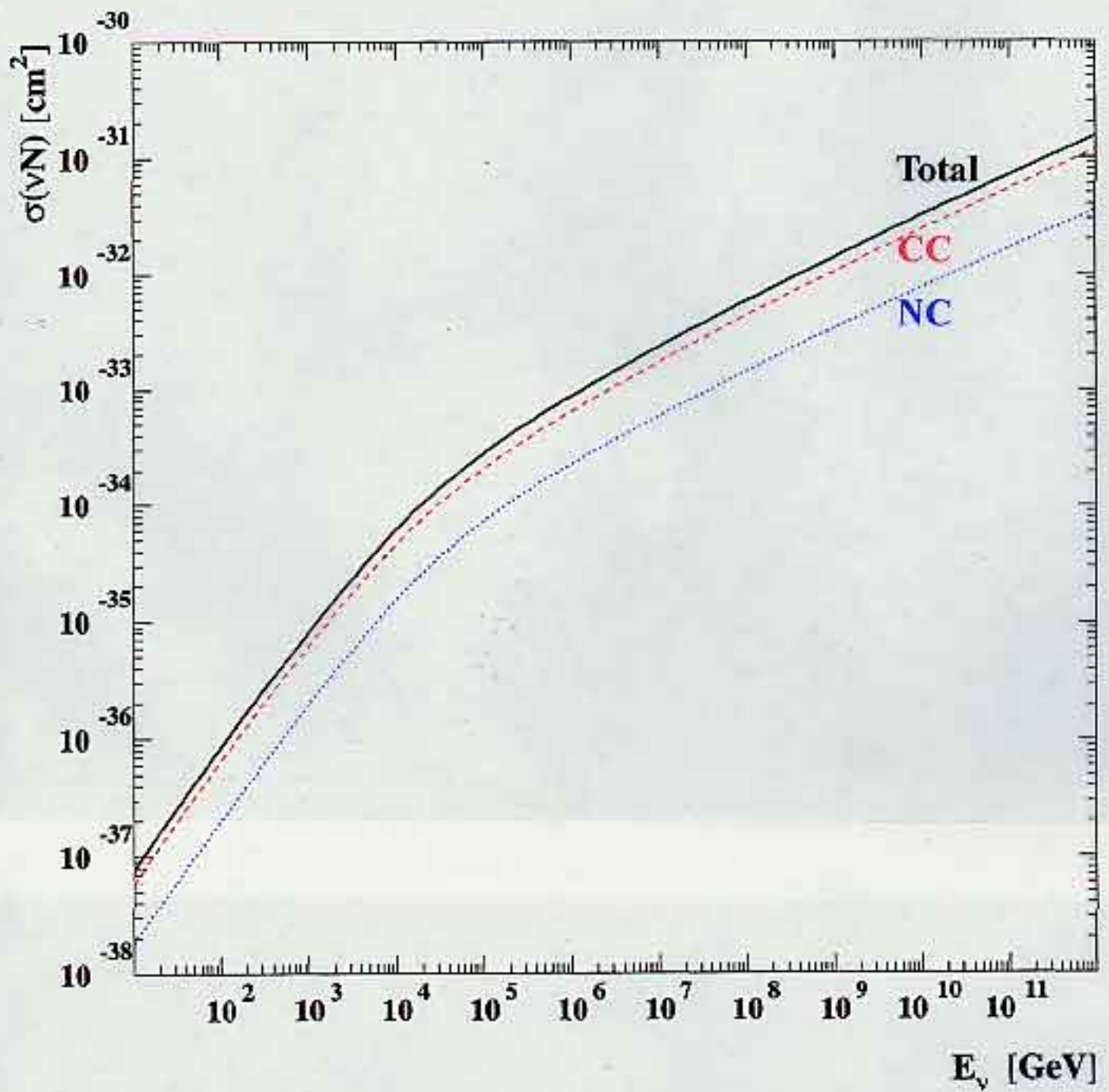


Attenuation of ν 's by the Earth







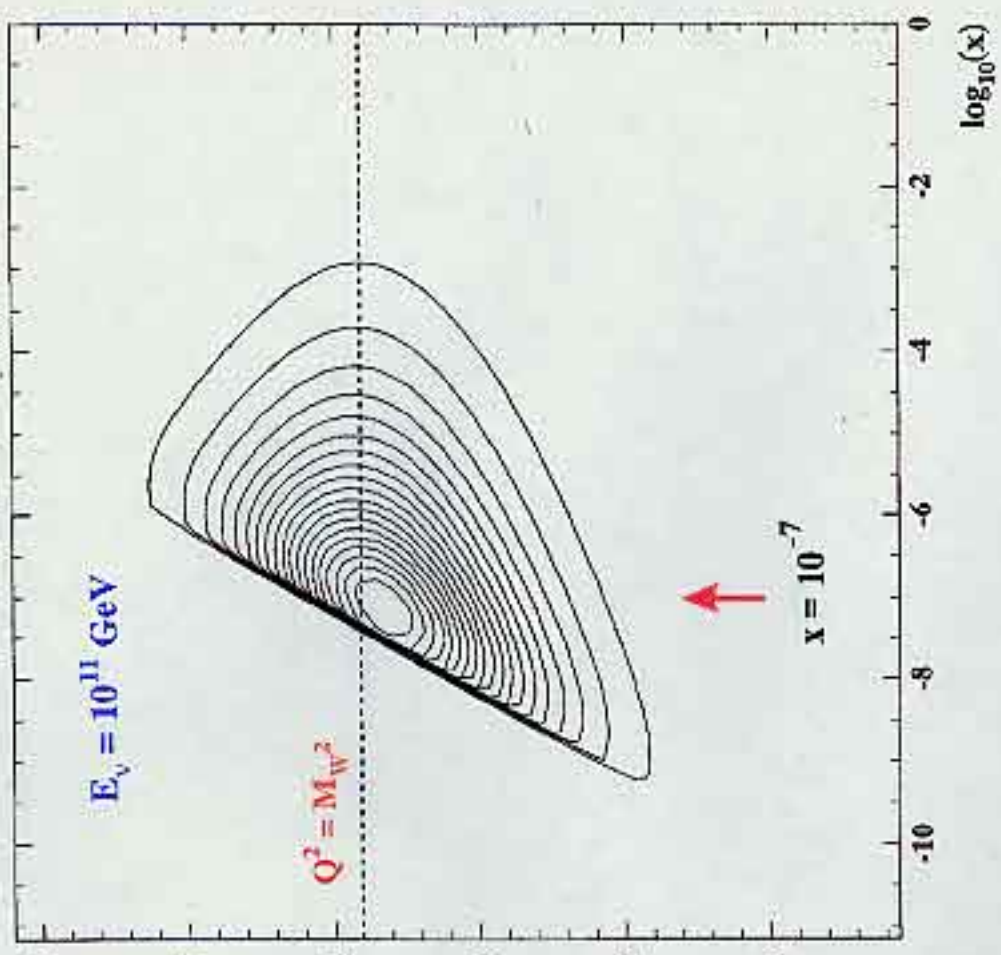
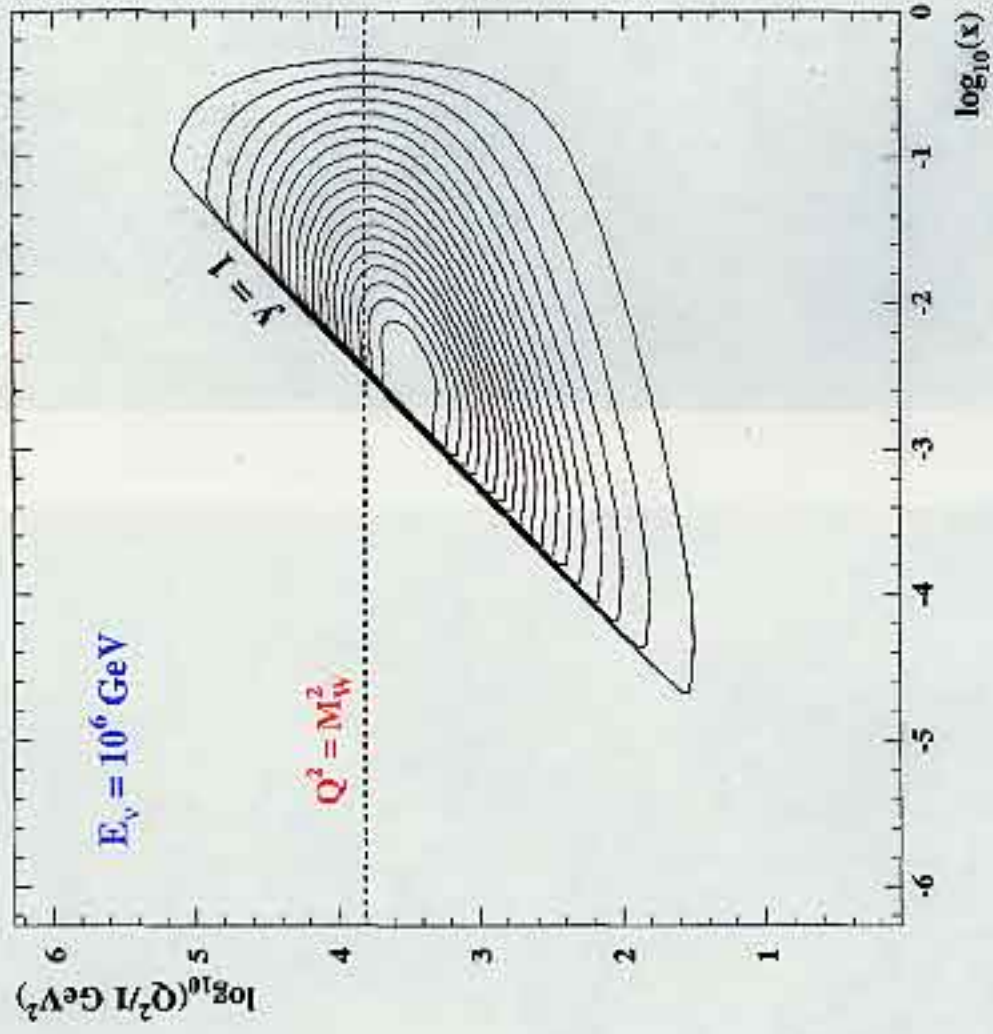


$$\frac{d\sigma(2N \rightarrow \mu X)}{dx dQ^2} = \frac{G_F^2}{\pi} \left(\frac{M_W^2}{Q^2 + M_W^2} \right)^2 \left[y(x, Q^2) + (1-y)^2 \bar{q}_1(x, Q^2) \right]$$

$$x = \frac{Q^2}{ys} \sim \frac{M_W^2}{2m_p E_\nu}$$

High E_ν : $Q^2 \sim M_W^2$

Kwiecinski, Martin, Stasto



$\sigma_{\nu N}^{\text{tot}}$ at ultrahigh E_{ν}

- V. small x
 $\ln(1/x)$ effects? OK Kwiecinski et al.
- $gg \rightarrow g$ + recombⁿ?
 saturation? small Reno et al.
- perturbative unitarity?

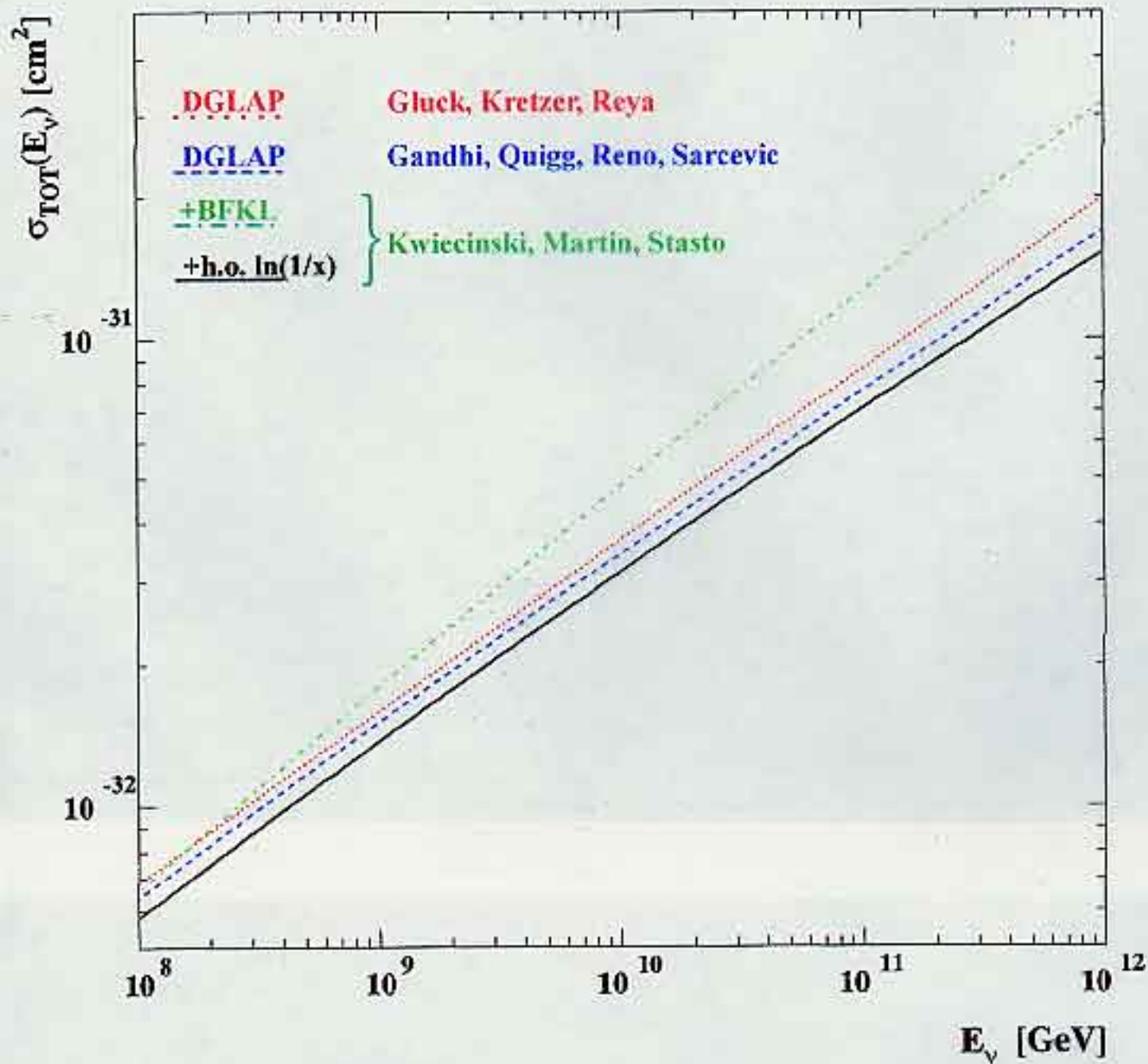
questioned by Dicus et al.

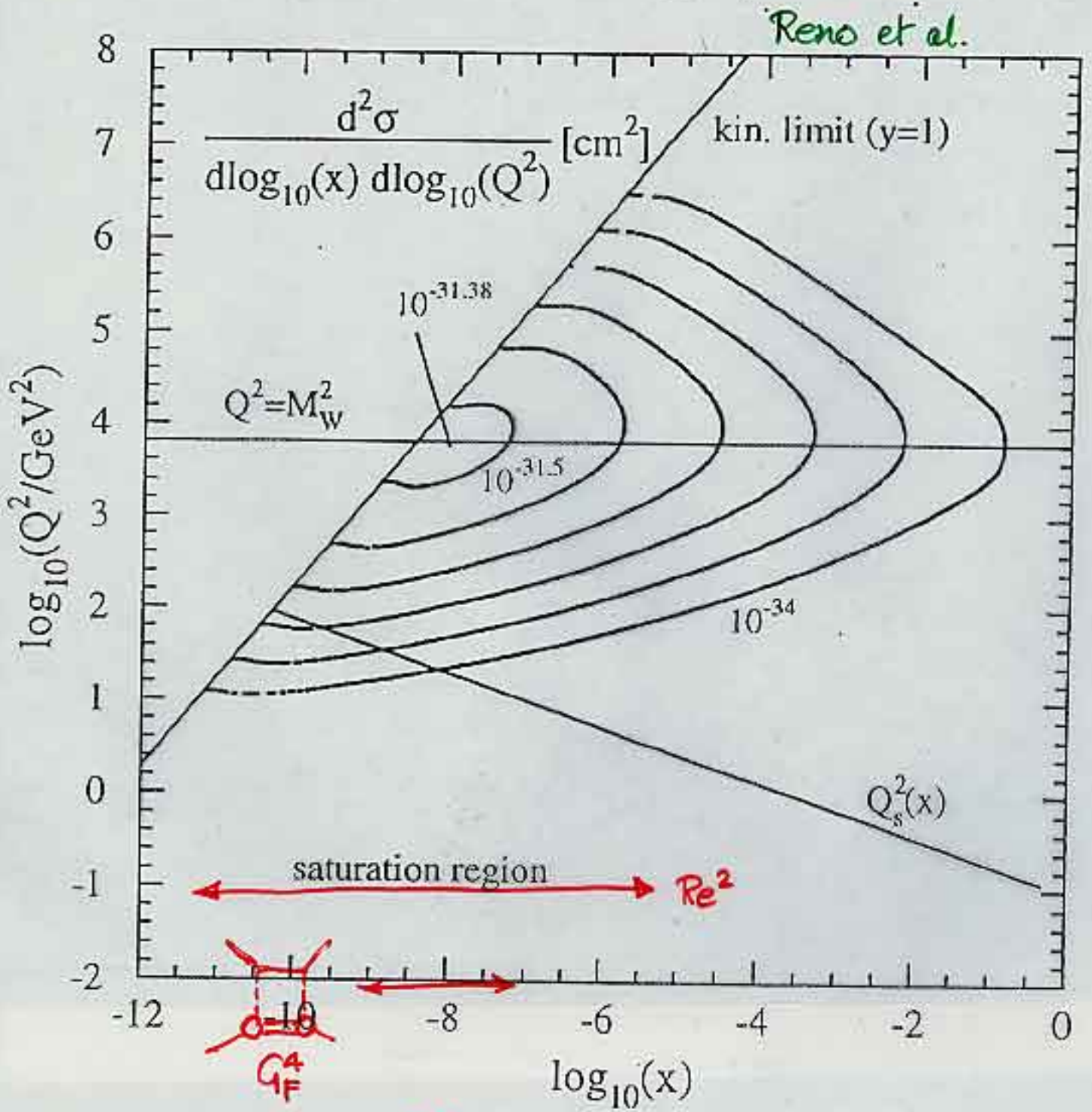
$$\frac{d\sigma_{el}(\nu N \rightarrow \nu N)}{dt} \stackrel{\text{Im}^2 + \text{Re}^2}{\geq} \frac{1}{16\pi} \sigma_{\text{tot}}^2$$



G_F^4

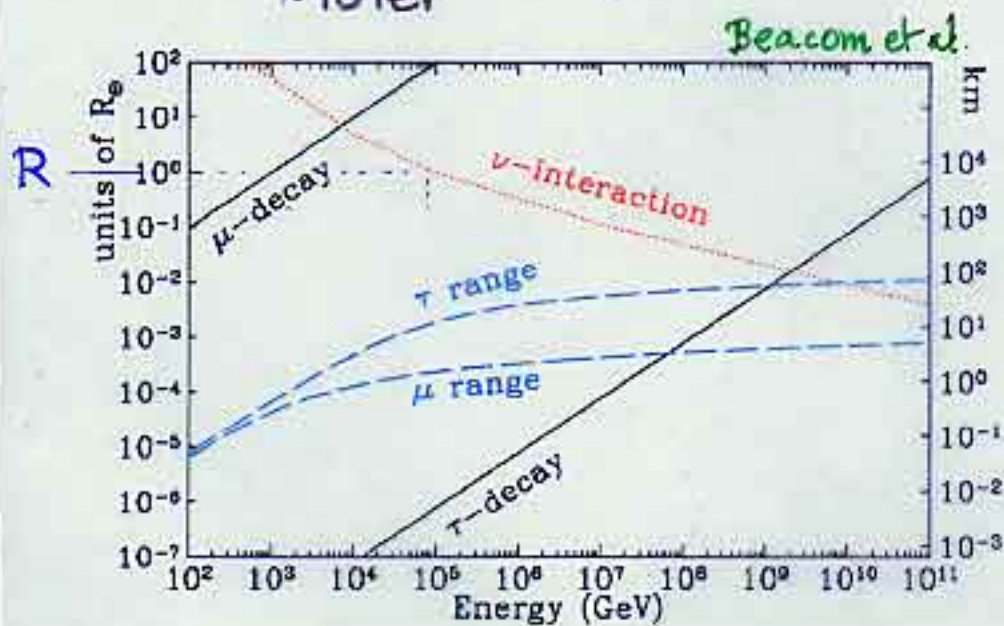
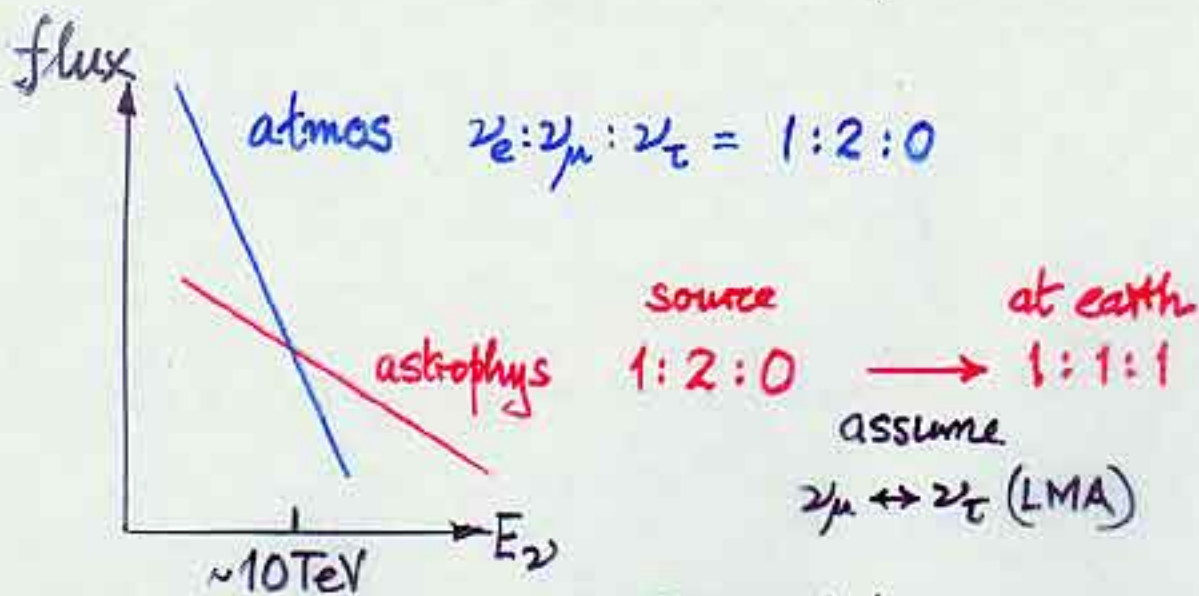
← σ_{tot} violates this inequality at $E_{\nu} \gtrsim 2 \times 10^8 \text{ GeV}$?
 Dicus et al.





OK
Reno et al.

Conclusion: no reason to doubt extrapolation of σ_{2N}^{tot} above HERA domain



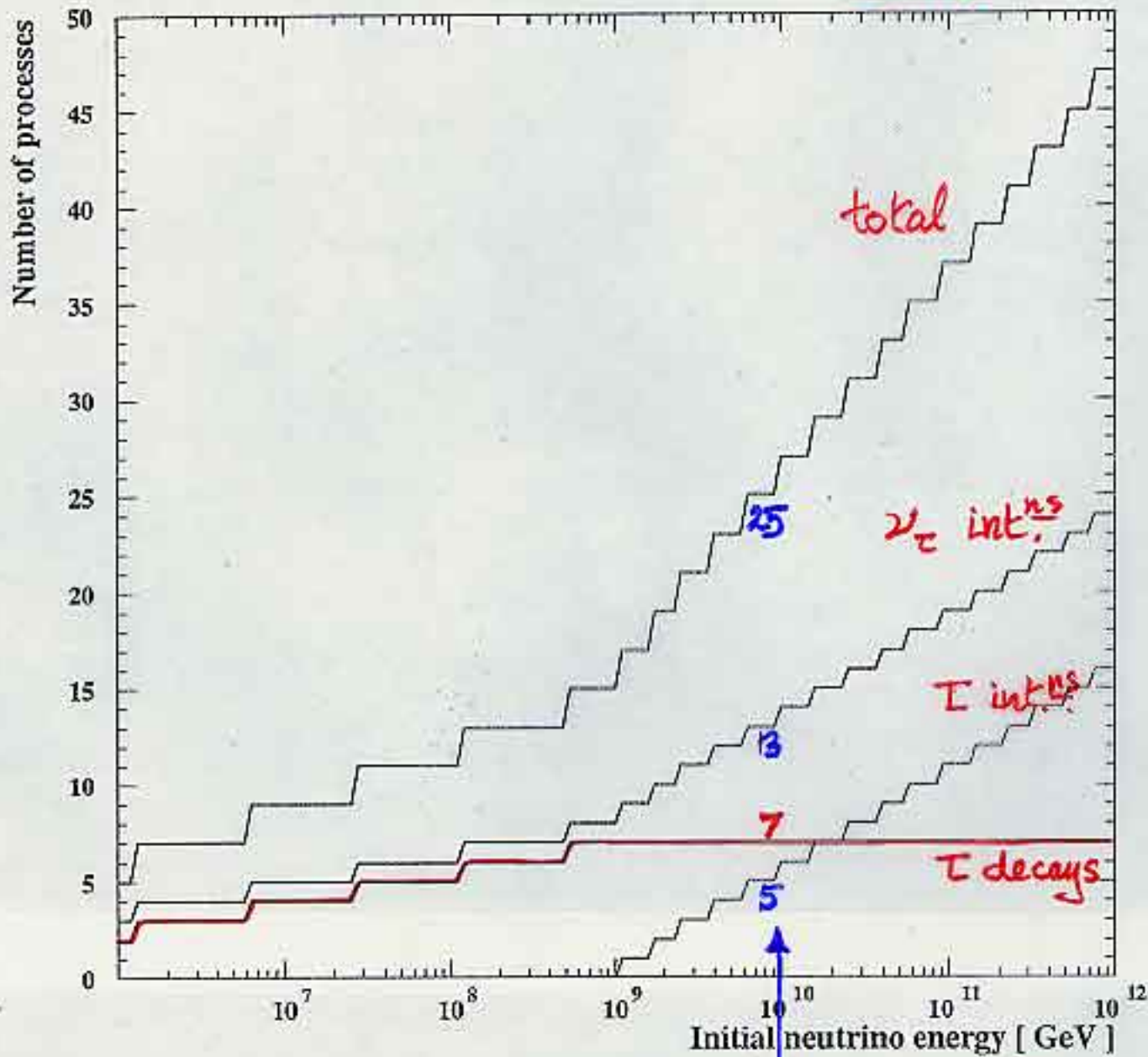
$\nu_e \rightarrow e$ quickly brought to rest } earth opaque
 $\nu_\mu \rightarrow \mu$ quickly slowed down } $E_2 \gtrsim 100\text{TeV}$

$\nu_\tau \rightarrow \tau \xrightarrow{\text{fast}} \nu_\tau \rightarrow \tau \rightarrow \nu_\tau \dots$ until earth transparent
 $E_2 \sim 10^{14}\text{eV}$ 'pile-up'
 for flat spectrum

Halzen, Saltzberg

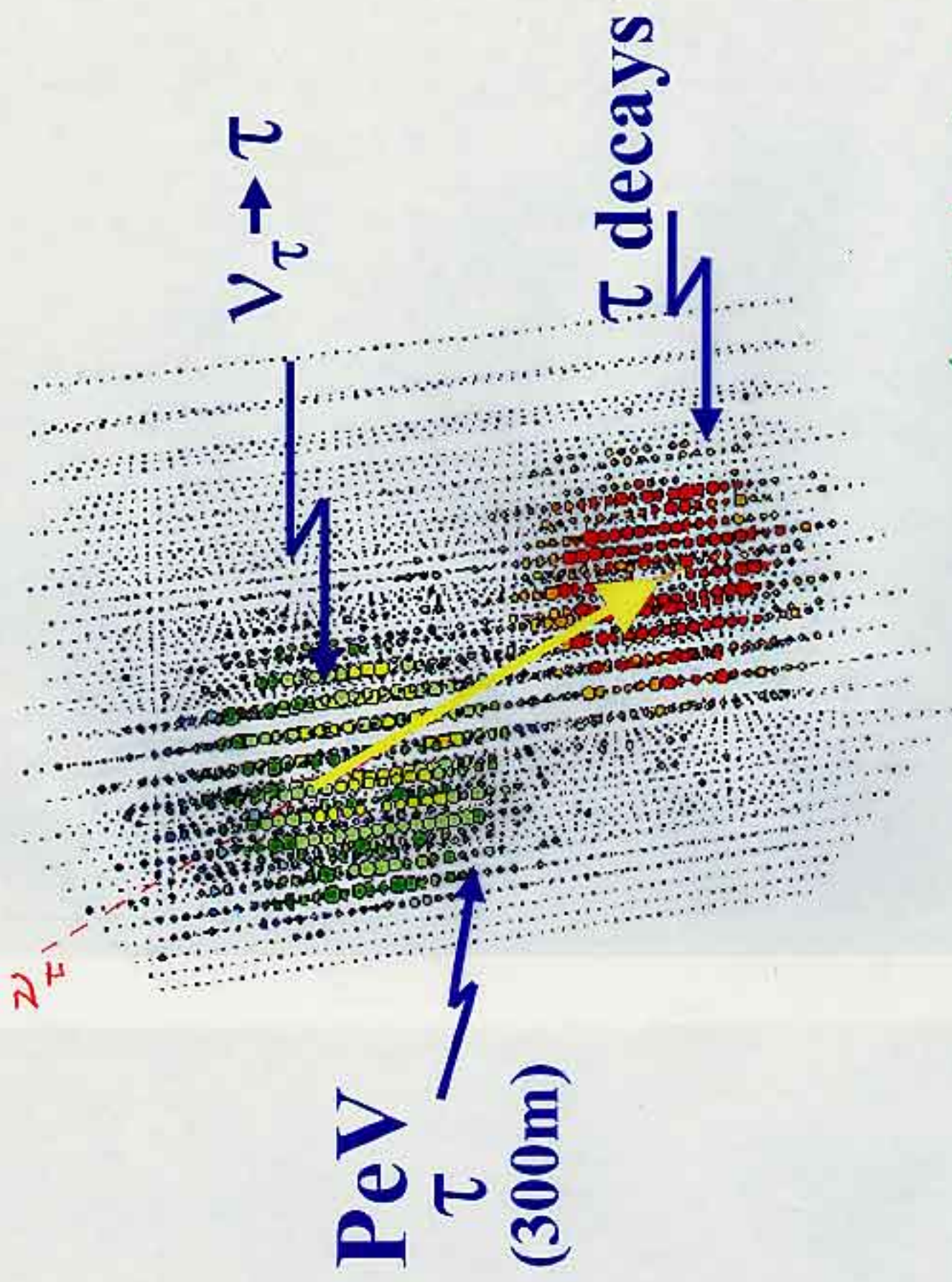
[EHE $\nu_\tau \rightarrow \tau \rightarrow \tau (E \sim 10^{17}\text{eV})$ then $\rightarrow \nu_\tau \dots$]

Stasto



$E_\nu = 10^{19} \text{ eV}$

E_ν

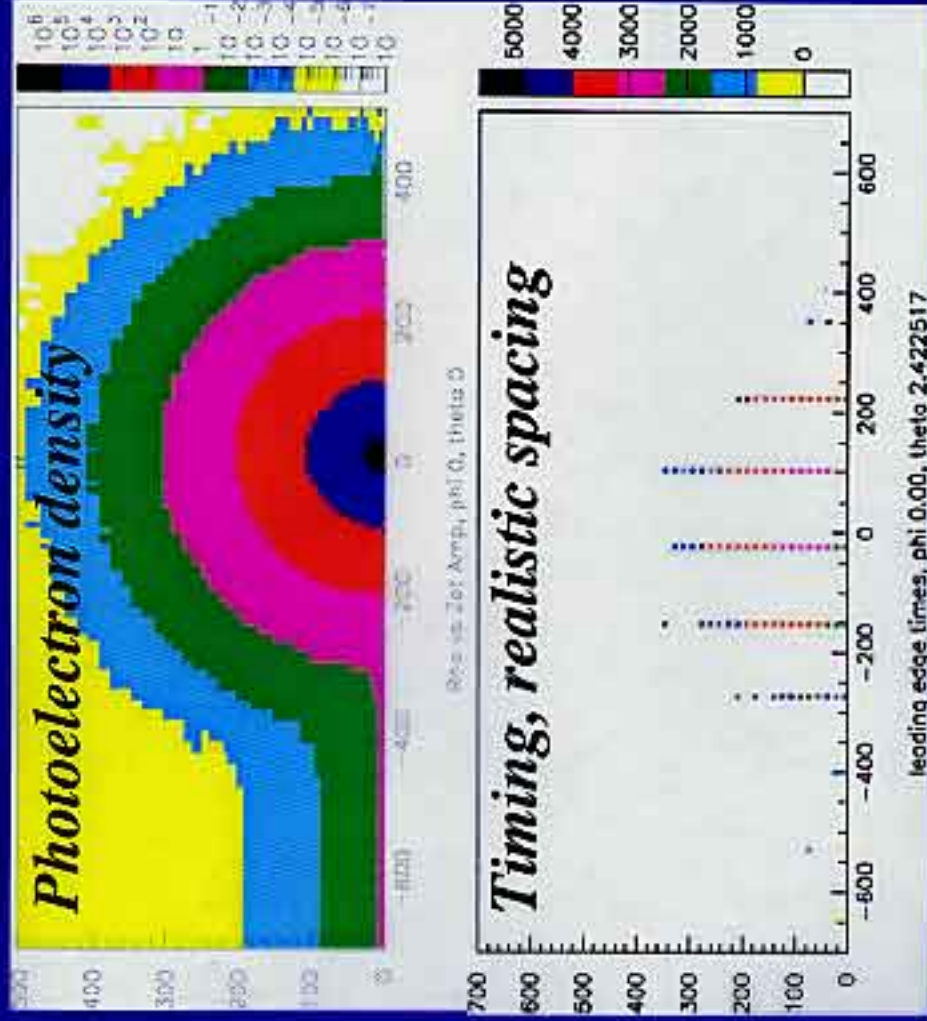


'Double-bang' events

'Single-bang Lollipop' events may be visible upto $E_2 \sim 10^{17}$ eV

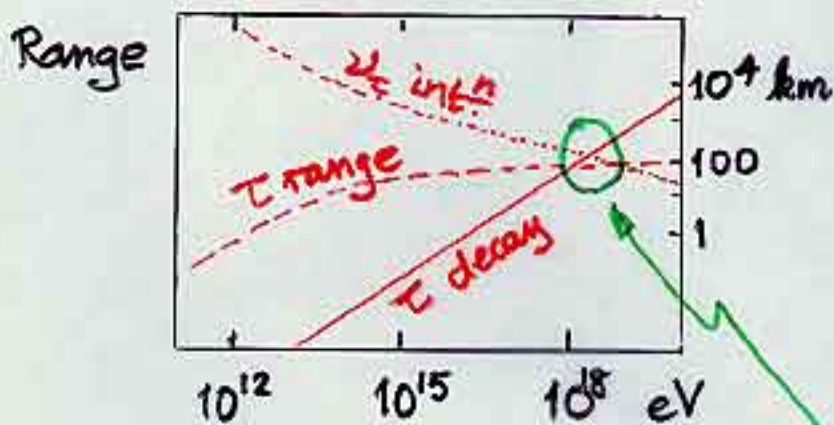
ν_τ at $E > \text{PeV}$: Partially contained

- The incoming tau radiates little light.
- The energy of the second bang can be measured with high precision.
- Clear signature
- Muon Brem would be much brighter than the tau (compare to the PeV muon event shown before)



Result:
*high effective volume;
only second bang seen in Ice3*

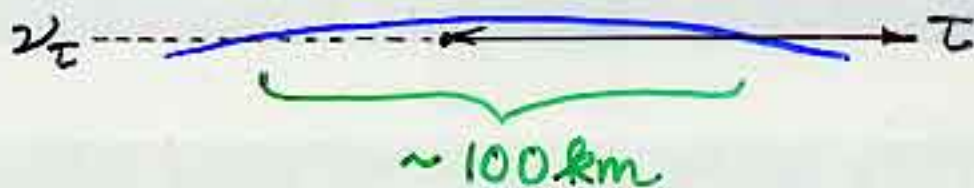
To avoid $\nu_\tau \rightarrow \tau \rightarrow \nu_\tau \rightarrow \tau \rightarrow \nu_\tau \dots$ pile-up
 at 10^{14} eV, use shorter path length thru earth



all comparable
 for $E \sim 10^{18}$ eV

~ 100 km

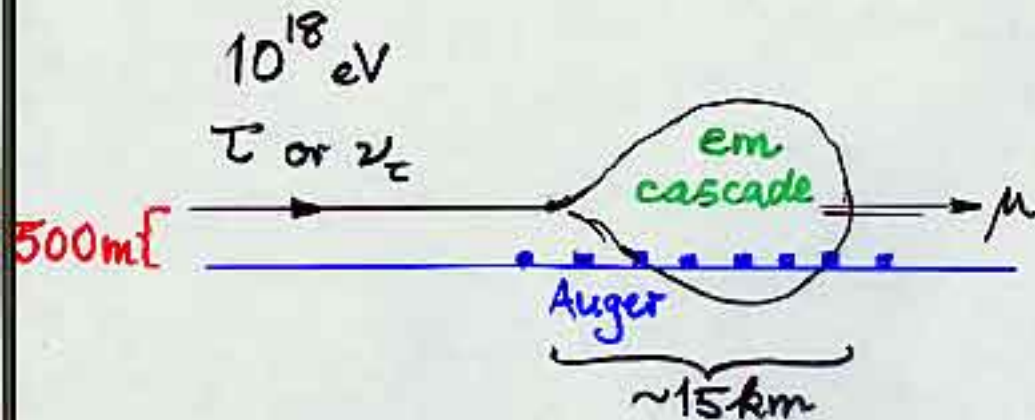
\rightarrow horizontal ν_τ 's



(Icecube OK up $\sim 10^{17}$ eV)

Horizontal ν_e 's in Auger

Bertou et al.



hadrons interact at top of atmos.
Em cascade does not reach Auger

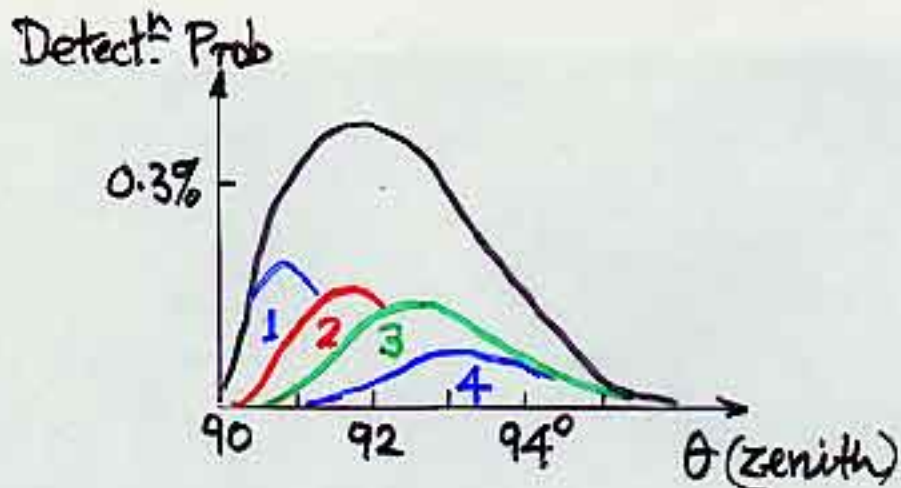


Detectable if $10^{17} \text{ eV} < E < 10^{19} \text{ eV}$

← not enough Auger triggers

→ decay too slow

few events/yr?



1, 2, 3, 4 int.^{ns} in earth.

Conclusions:

- TeV+ ν astronomy taking off
- $\sigma(\nu N)$ reliable
- ν_{τ} allow UHE to be probed.